



李政道研究所
TSUNG-DAO LEE INSTITUTE



TRIDENT
海 | 铃 | 计 | 划

The TRIDENT Deep-sea Neutrino Telescope

Donglian Xu (TDLI)

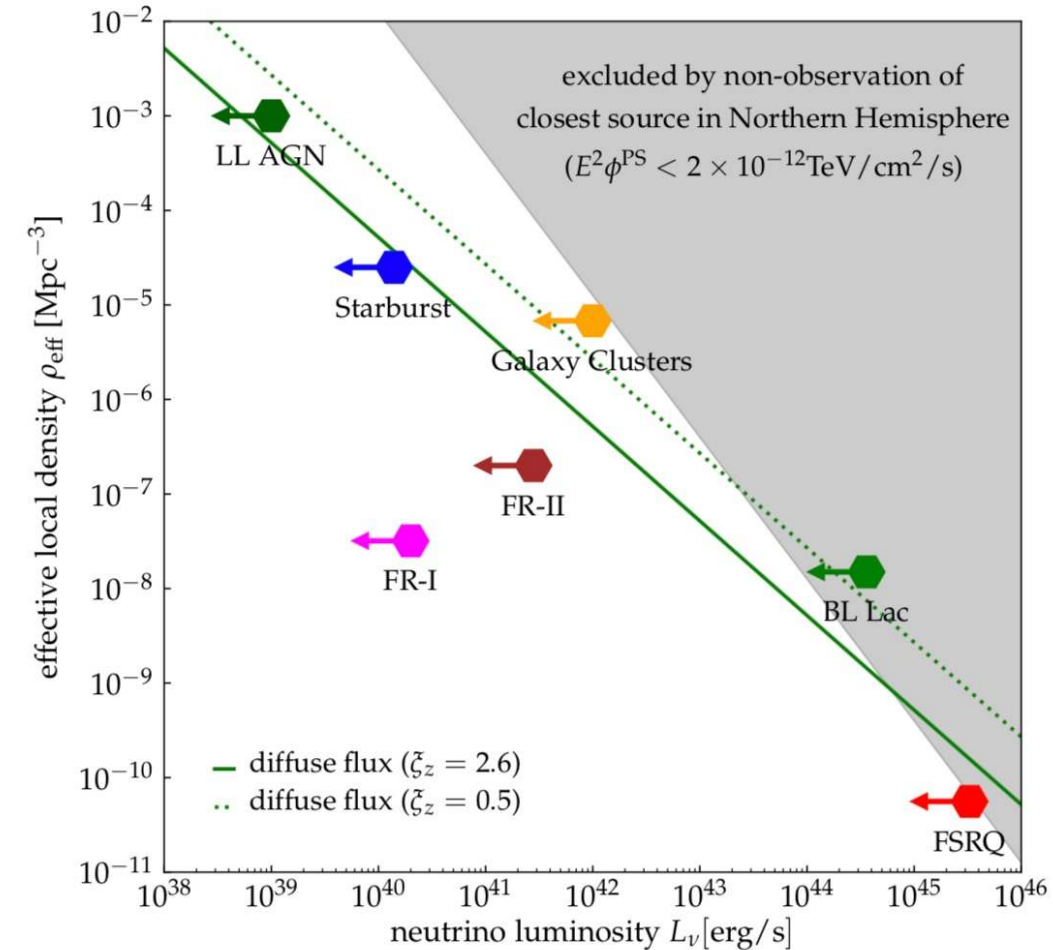
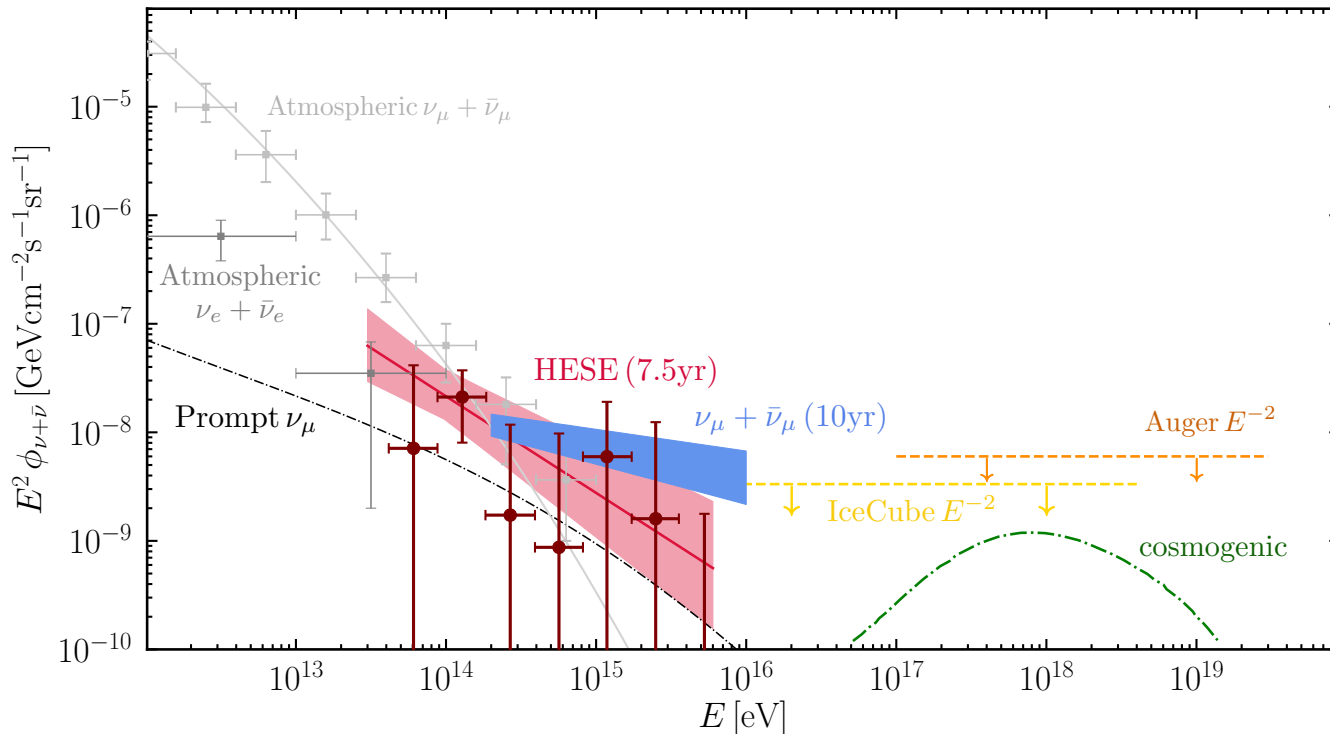
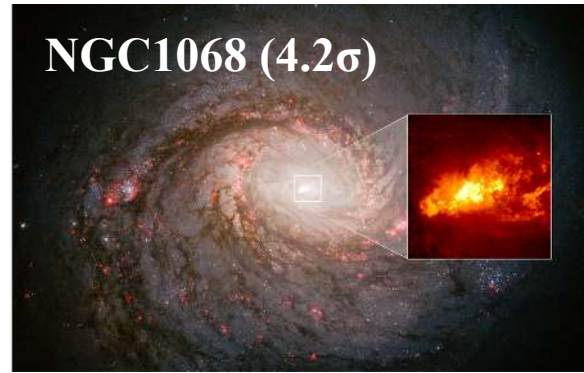
TeVPA 2024

2024. 08. 26-30

Chicago, Illinois, USA

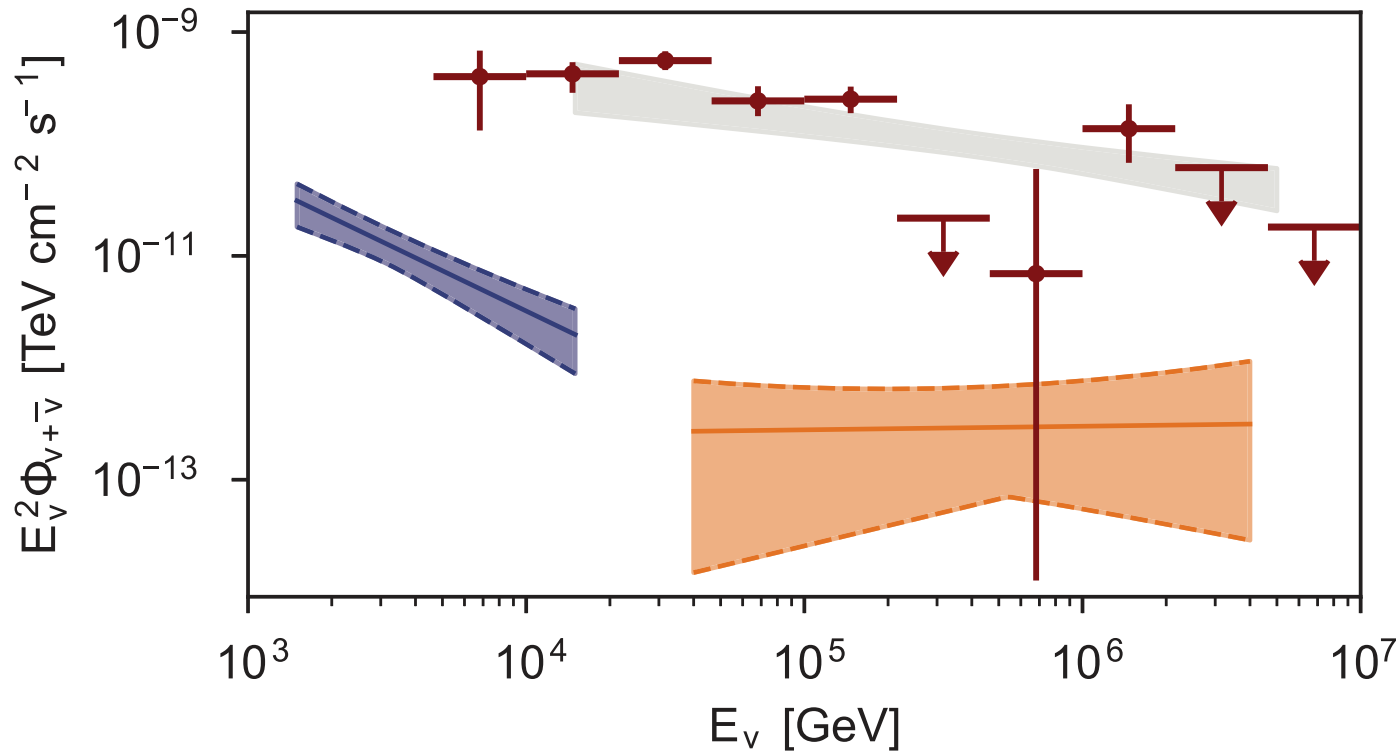
donglianxu@sjtu.edu.cn

A new era of neutrino astronomy



Halzen & Khierandish, arXiv:2202.00694

< 14% of the diffuse flux can come from the Galaxy



IceCube Collaboration, *Science* 378, 538 (2022)

- At least two distinctive categories of sources

- Diffuse flux largely unresolved



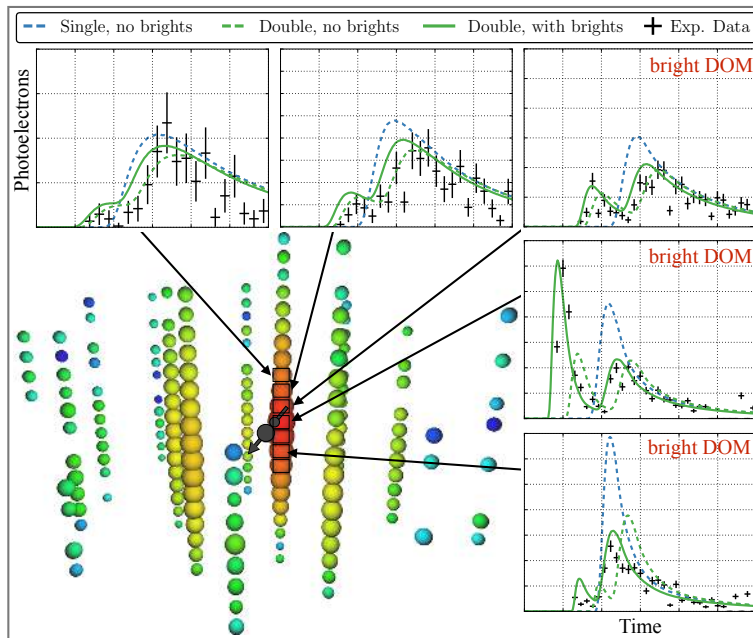
1) How to **optimize** next-gen neutrino telescopes for **low** and **high energies**?

2) Room to **improve** on **angular resolution**? Need better than 0.1° @ 100TeV to resolve the diffuse flux

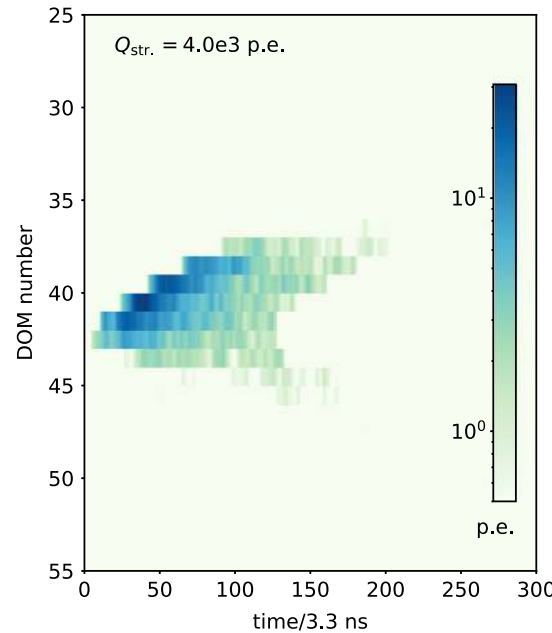
3) How to **boost flavor identification** for discovered sources?

How can we improve? →

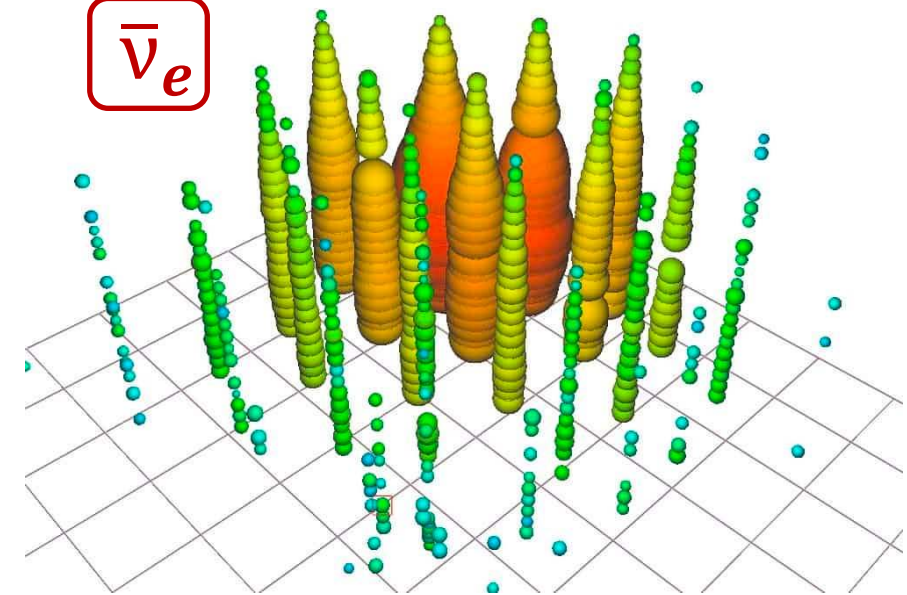
- 1) Larger detectors for more high-energy events
- 2) Pixelized DOMs with waveforms for recording kinetic info



ν_τ



$\bar{\nu}_e$



IceCube Collaboration, *Eur. Phys. J. C* 82, 1031 (2022)

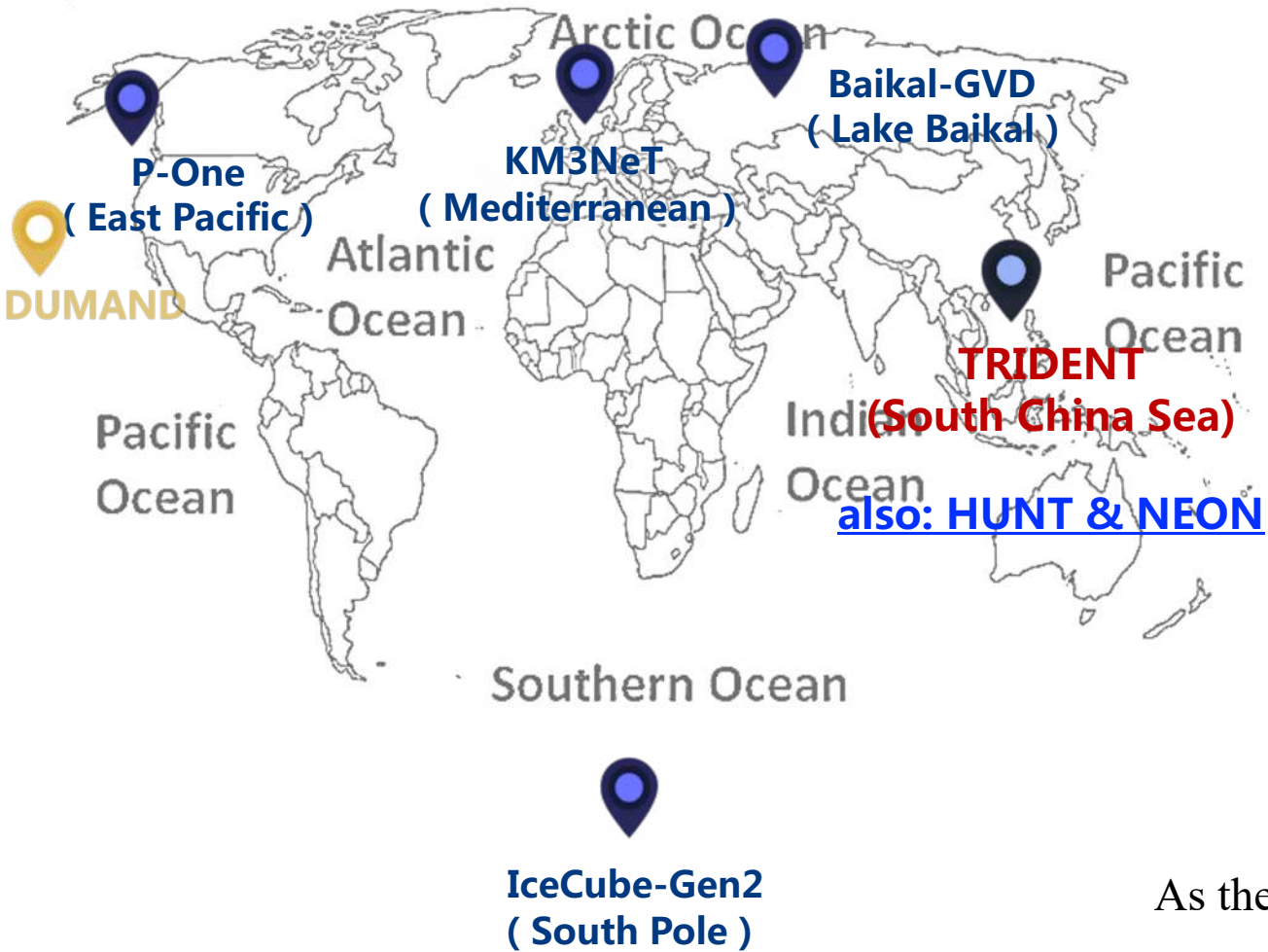
M. Meier, J. Soedingrekso, PoS (ICRC2019) 960

L. Wille, D.-L. Xu, PoS (ICRC2019) 1036

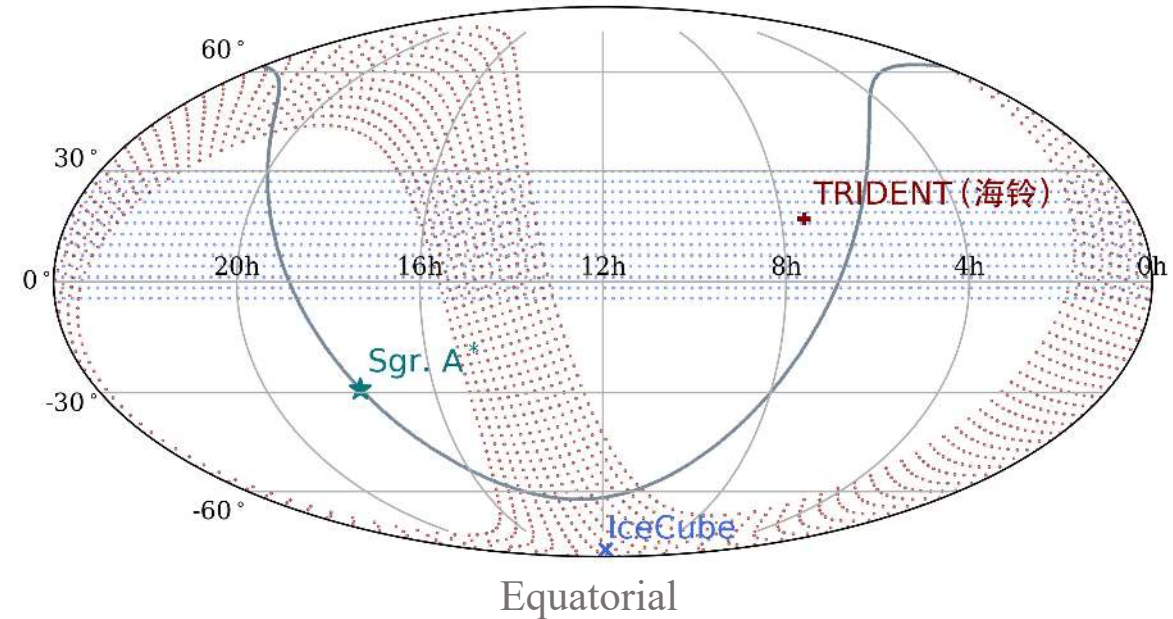
IceCube Collaboration,
Phys. Rev. Lett. 132, 151001 (2024)

IceCube Collaboration,
Nature 591, 220–224 (2021)

(Next-gen) neutrino telescopes under planning



TRopIcal DEep-sea Neutrino Telescope



As the Earth rotates, TRIDENT's best sensitivity band will sweep through the entire sky, complementing IceCube-Gen2 well

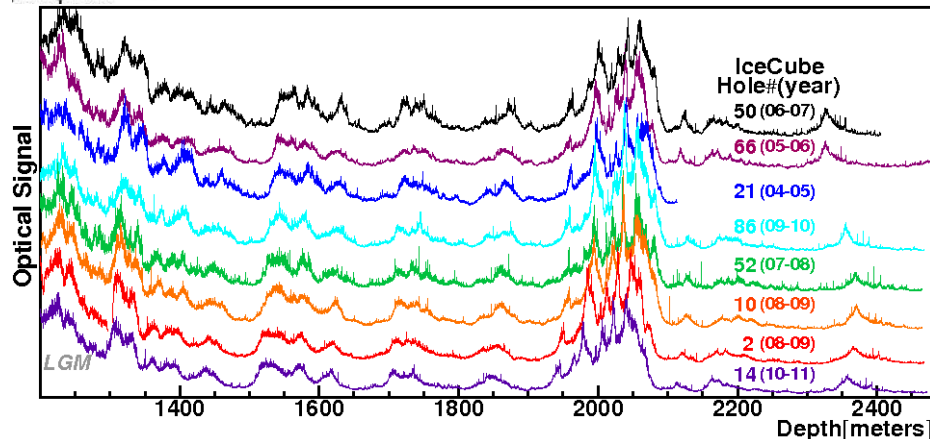
Interaction medium: Ice vs Water

Glacial ice

Most transparent medium on Earth!

Scattering length: $\sim 25\text{m}$

Absorption length: $>100\text{m}$



Lake/sea water

Lake Baikal

Water properties:

Abs. length: $22 \pm 2 \text{ m}$

Scatt. length: $L_s \sim 30\text{-}50 \text{ m}$

$L_s / (1 - \langle \cos\theta \rangle) \sim 300\text{-}500 \text{ m}$



Mediterranean Sea

UV Scattering length: $>100\text{m}$

UV Absorption length: $\sim 25\text{m}$



On average, ice is more transparent / less absorbing, while water is less scattering

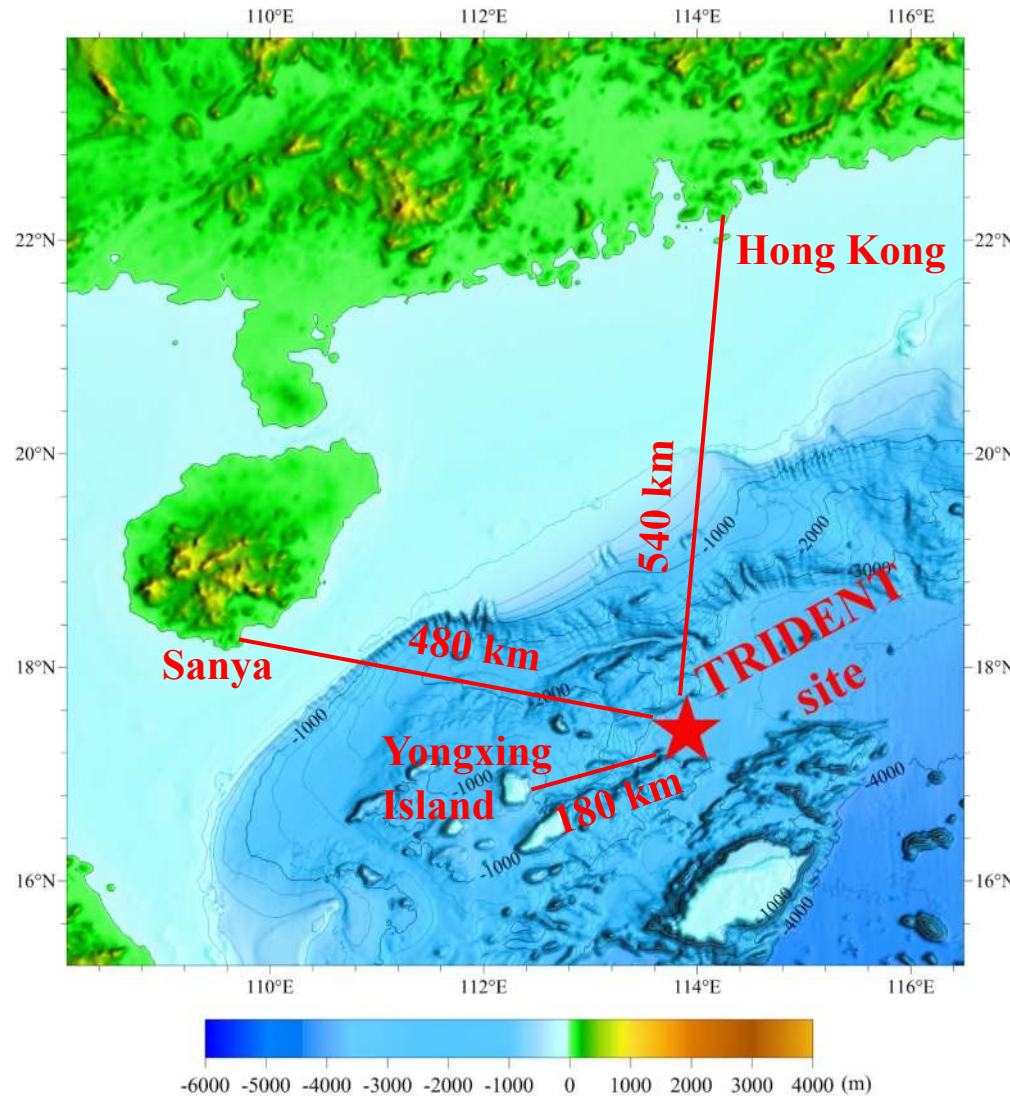


More “direct” photons in water-based telescopes \rightarrow intrinsically better pointing can be achieved

TRIDENT Explorer : T-REX



September, 2021



Pre-selected site conditions

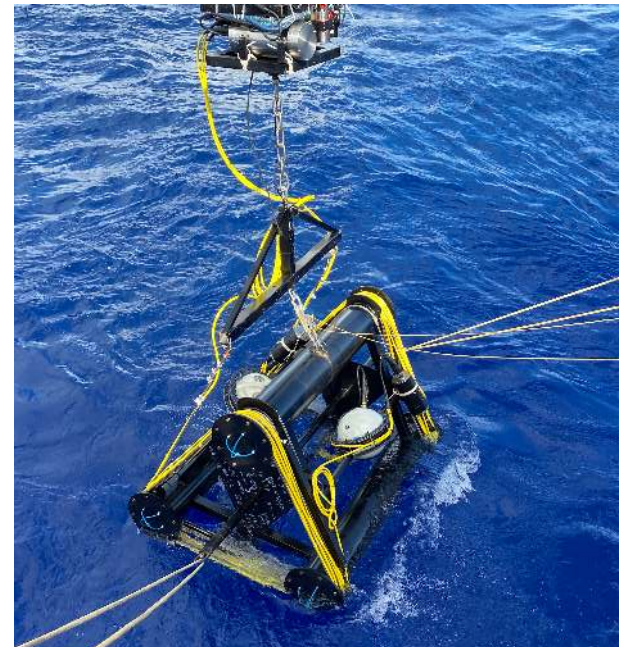
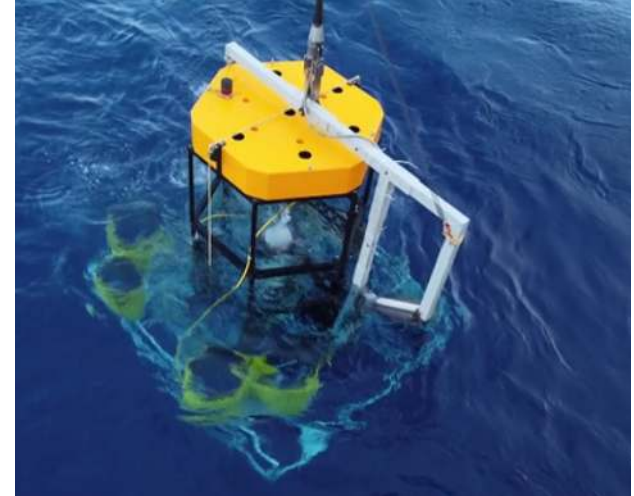
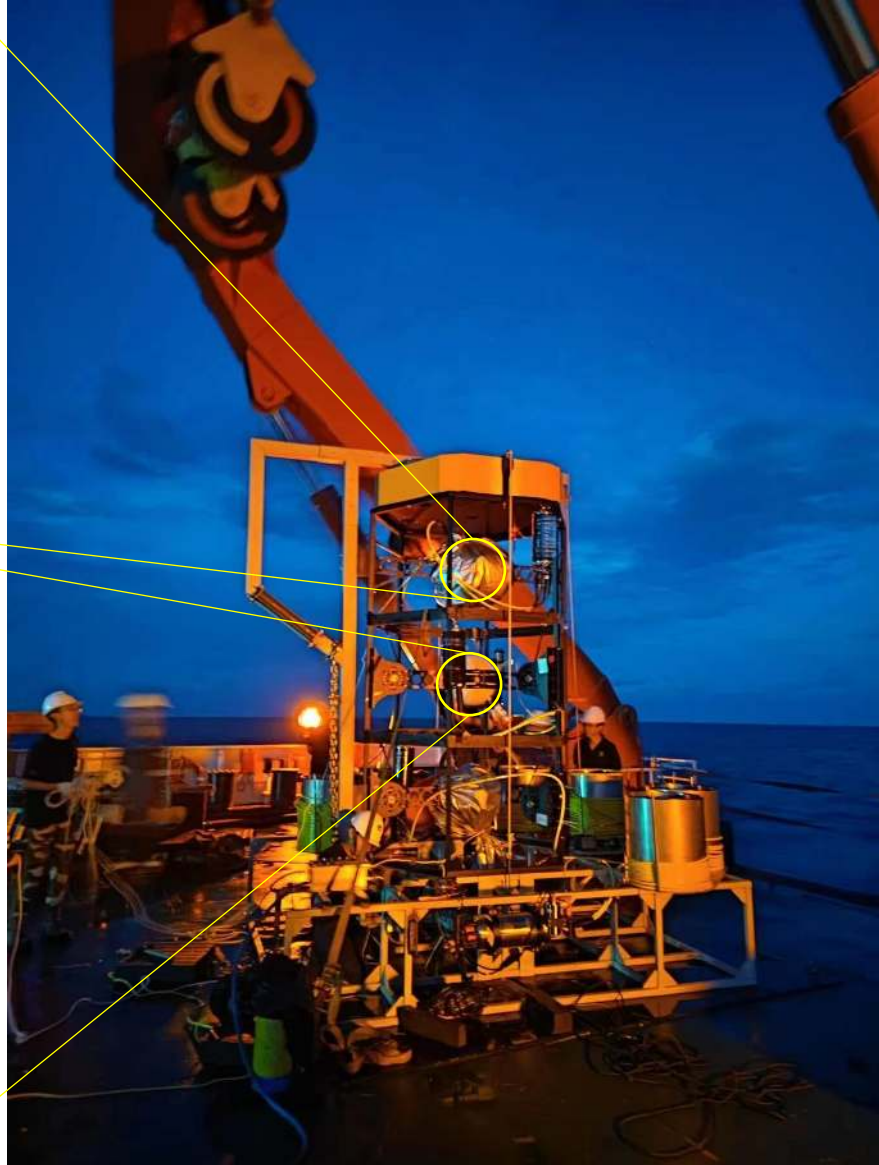
- Flat seabed
- No nearby high rises or deep trenches
- Depth >3km
- Close proximity to a shore

Measured params

- Optical properties
- Current field
- Radioactivity

<https://trident.sjtu.edu.cn/en>

TRIDENT Explorer : T-REX Apparatus



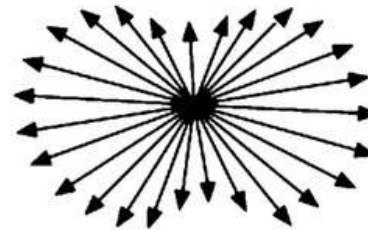
Absorption process (λ_{abs})

kill the photons, spacing design

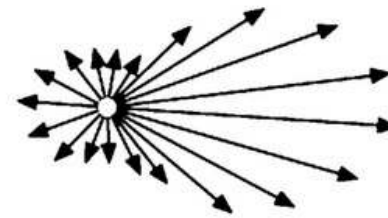
Scattering process (λ_{sca})

photon direction, angular resolution

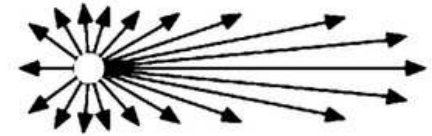
Rayleigh Scattering



Mie Scattering



Mie Scattering,
Larger Particles



→ Direction of Incident Light
← Direction of Backscatter

Rayleigh scattering (λ_{Ray}):

$$I = I_0 \frac{8\pi^4 \alpha^2}{\lambda^4 R^2} (1 + \cos^2 \theta)$$

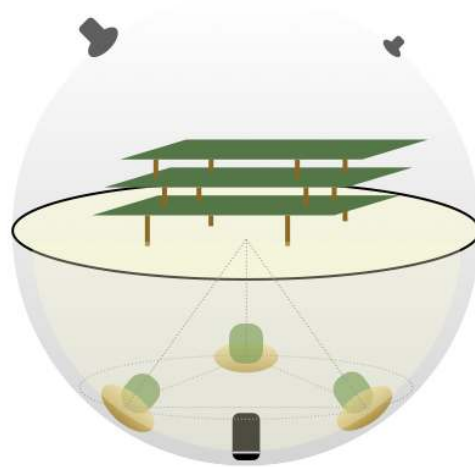
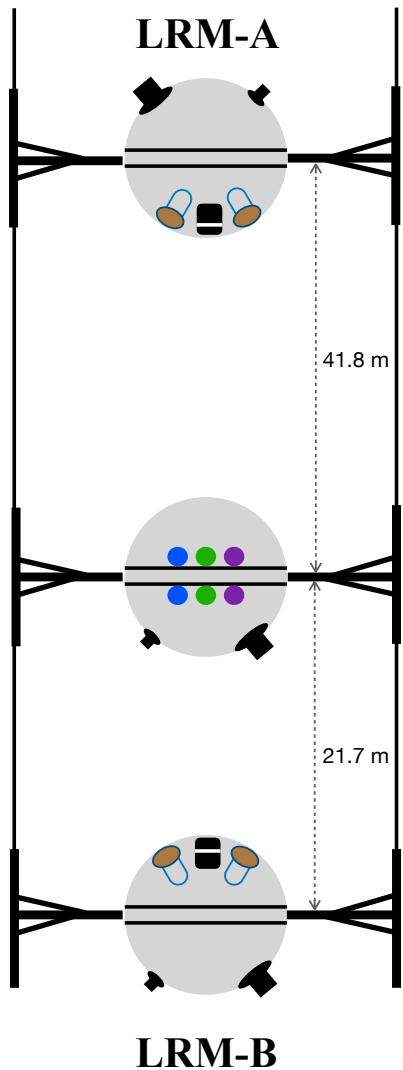
Mie scattering ($\lambda_{Mie}, \langle \cos \theta_{Mie} \rangle$):

$$\widetilde{\beta}^{HG}(g, \cos \theta) = \frac{1}{4\pi} \frac{1 - g^2}{(1 + g^2 - 2g \cos \theta)^{3/2}}$$

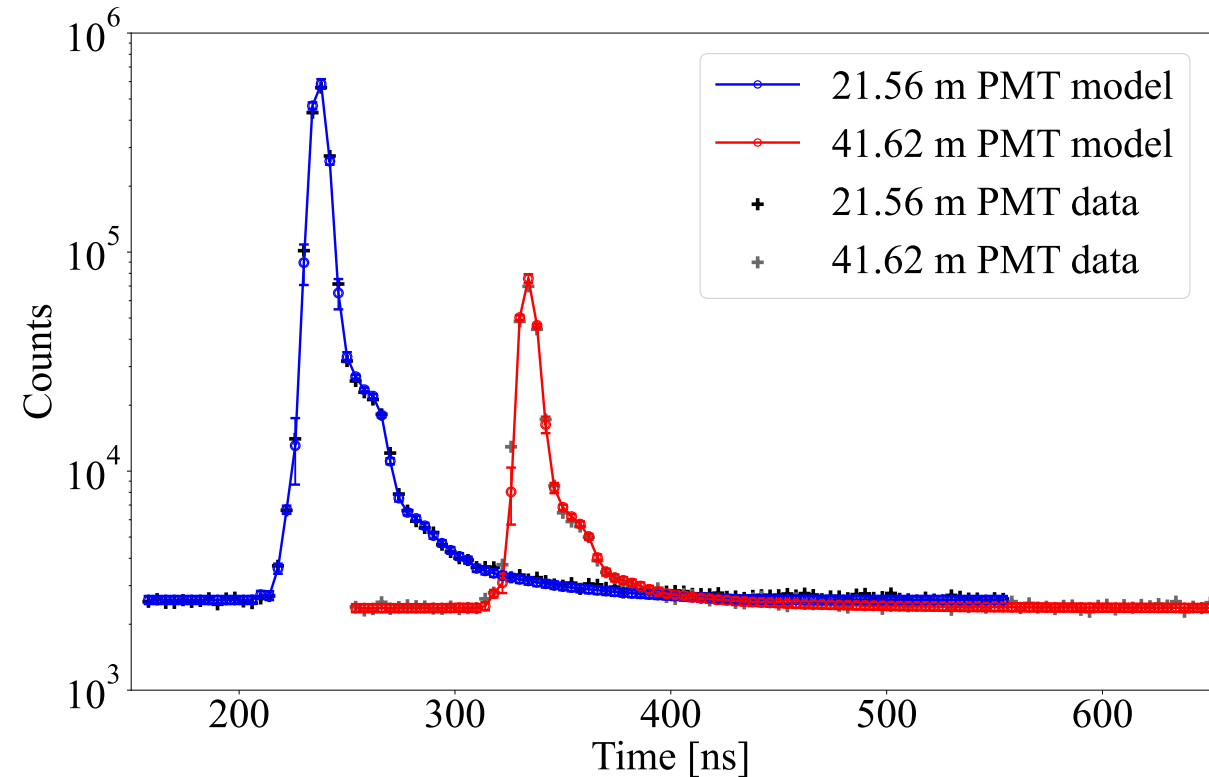
Attenuation length:

$$I(L) = I_0 \cdot e^{-\left(\frac{L}{\lambda_{abs}} + \frac{L}{\lambda_{sca}}\right)} = I_0 \cdot e^{-\frac{L}{\lambda_{att}}}$$

F. Hu *et. al.*, *Simulation study on the optical processes at deep-sea neutrino telescope sites*, NIMA 1054 (2023) 168367



Use relative measurement method to mitigate hidden systematics



Electronics: J. N. Tang *et al.*, *Journal of Instrumentation*, vol.18 T08001 (2023)

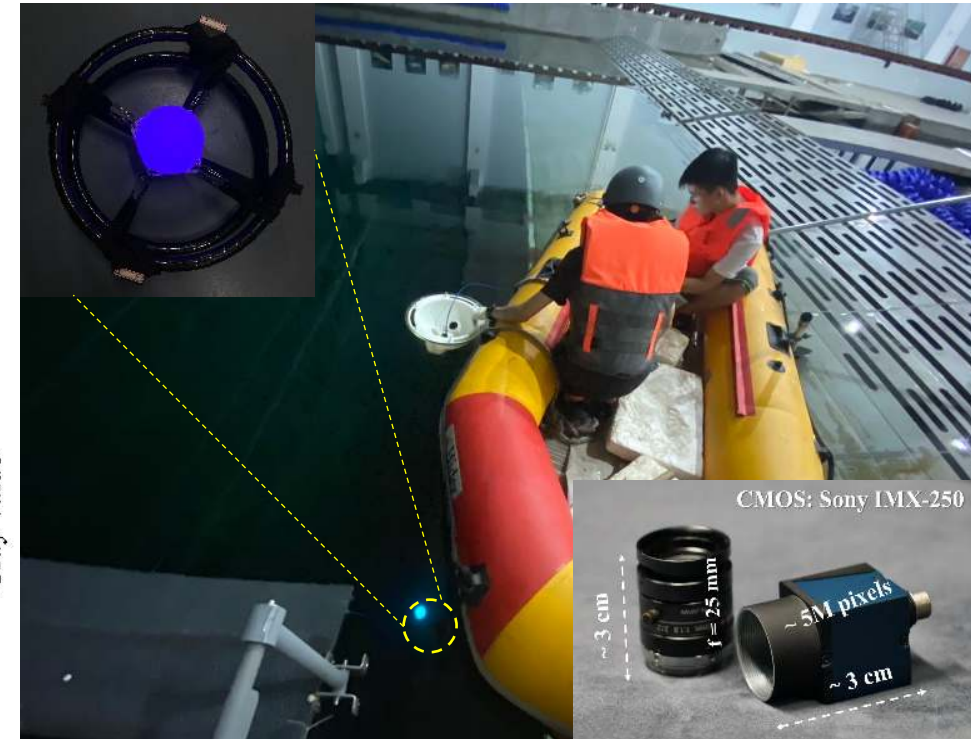
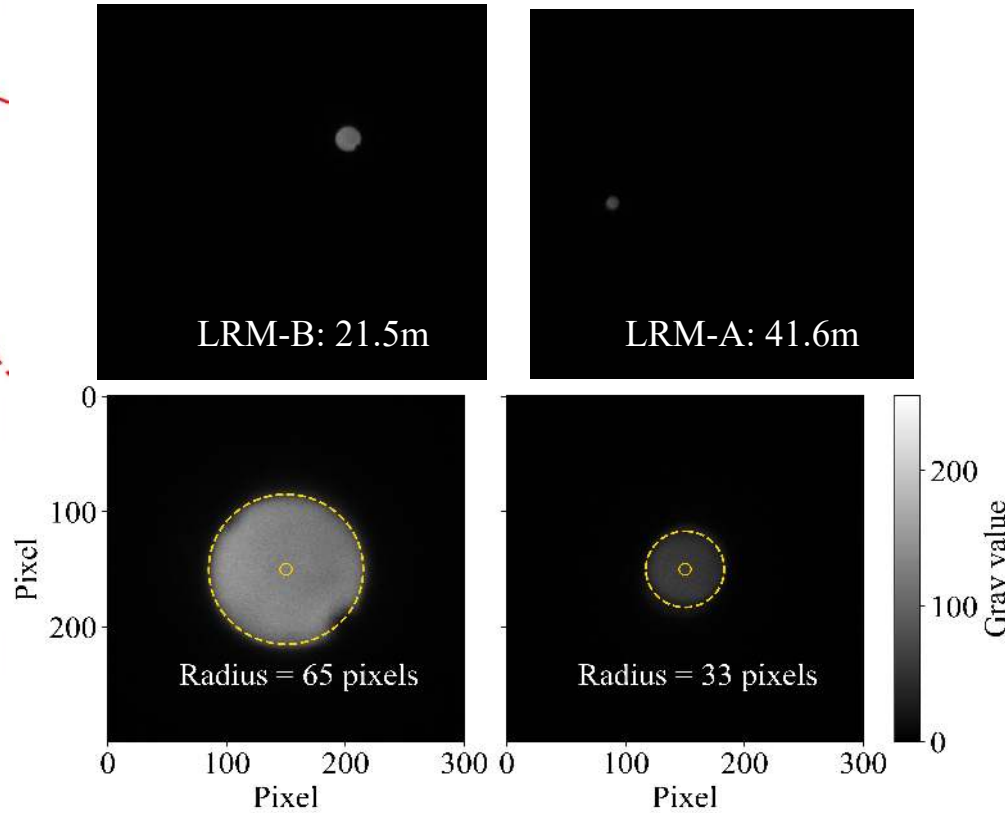
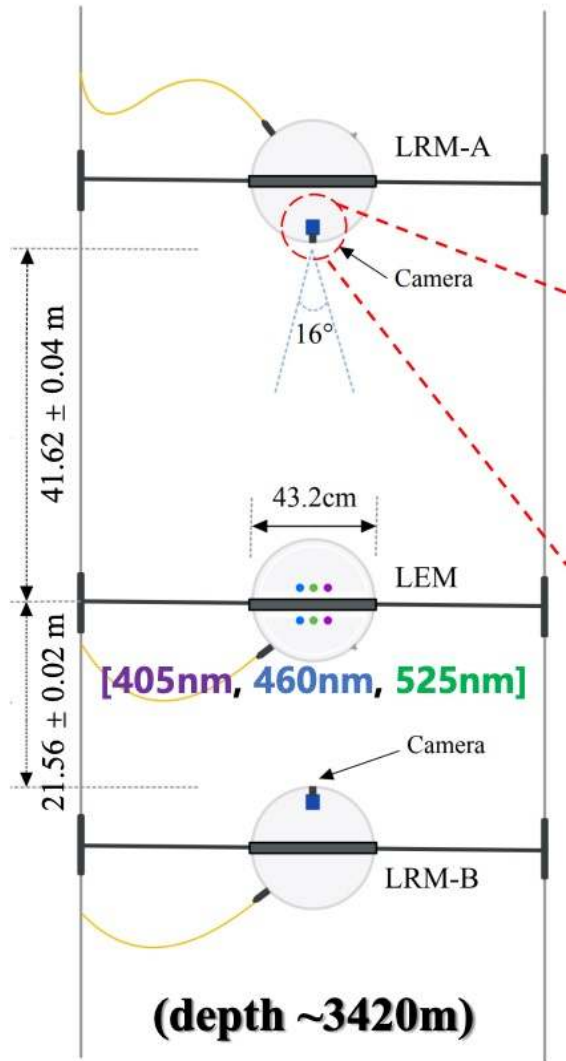
M. X. Wang *et al.*, *IEEE Transactions on Nuclear Science*, vol. 70, 2240–2247 (2023)

Light source: W. L. Li *et al.*, *The Light Source of the TRIDENT Pathfinder Experiment*, *NIMA* 1056 (2023) 168588



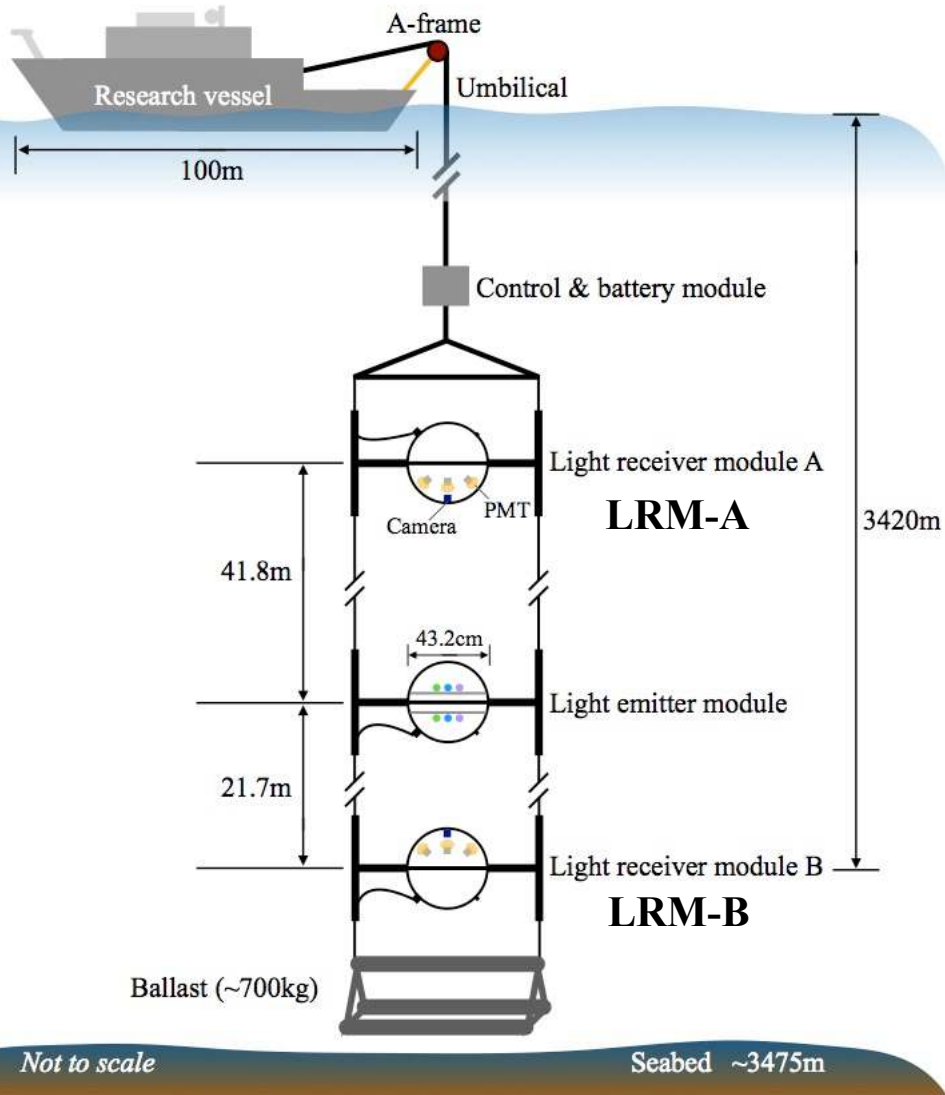
Images captured at depth of 3420m

Camera-calibrating in a ship towing tank

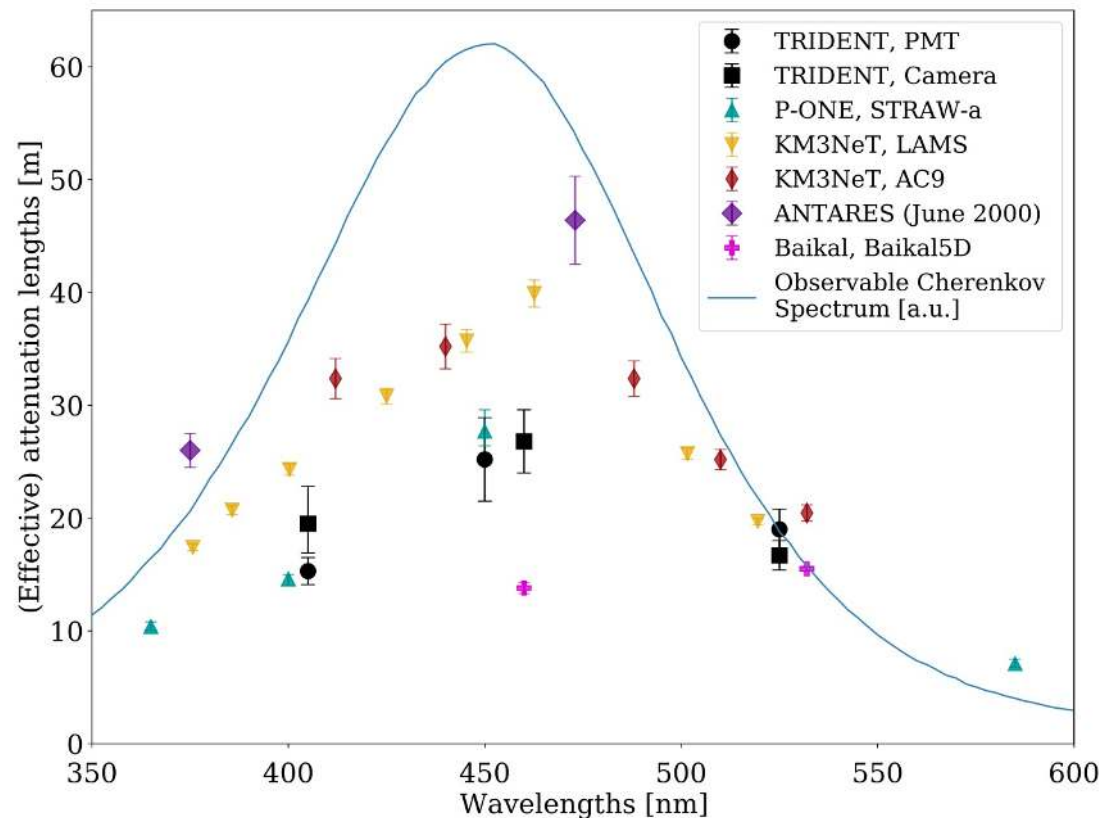


W. Tian *et al.*, *A camera system for real-time optical calibration of water-based neutrino telescopes*
arXiv 2407.19111 (submitted to NIMA)

TRIDENT Explorer : Optical Properties



- Dedicated analytical and numerical modeling
- Exp. data: ~ 1TB ↔ Simulated data: ~ 100 TB, 10M files



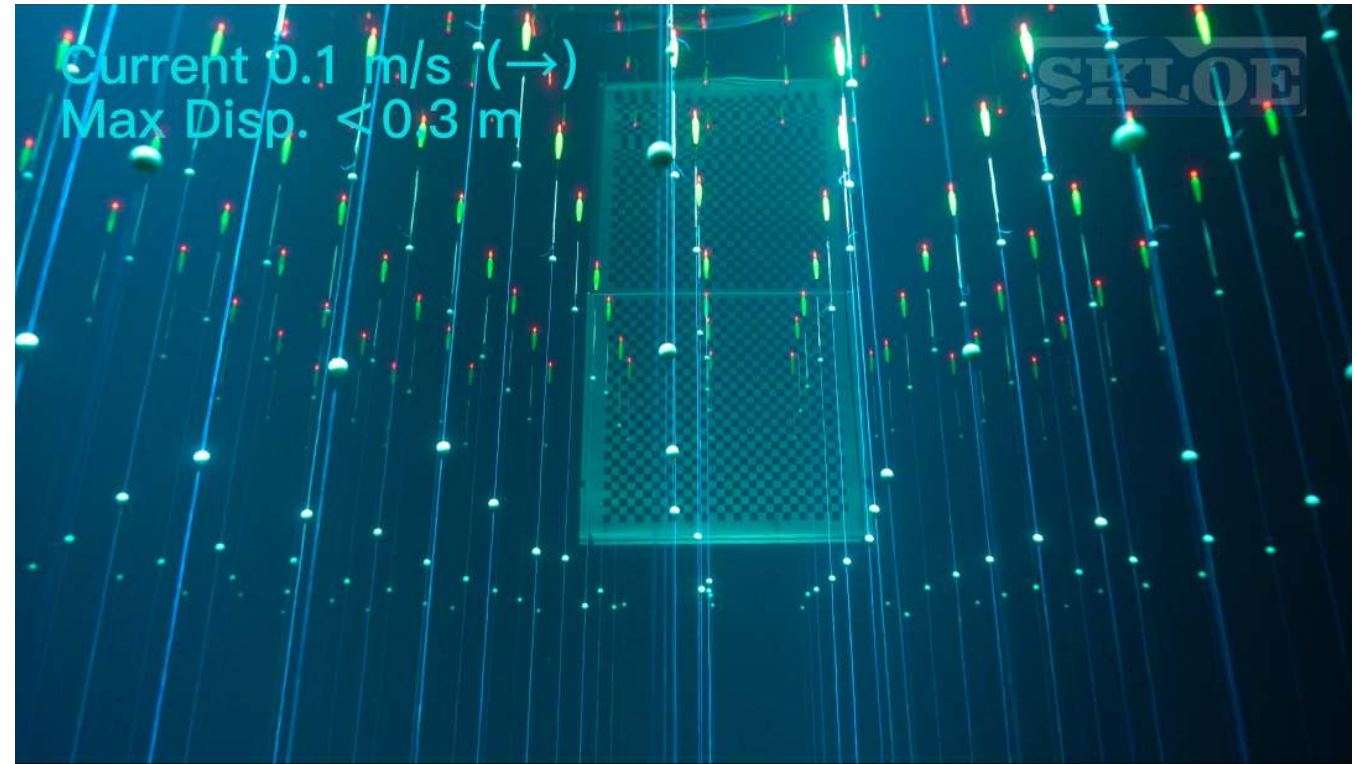
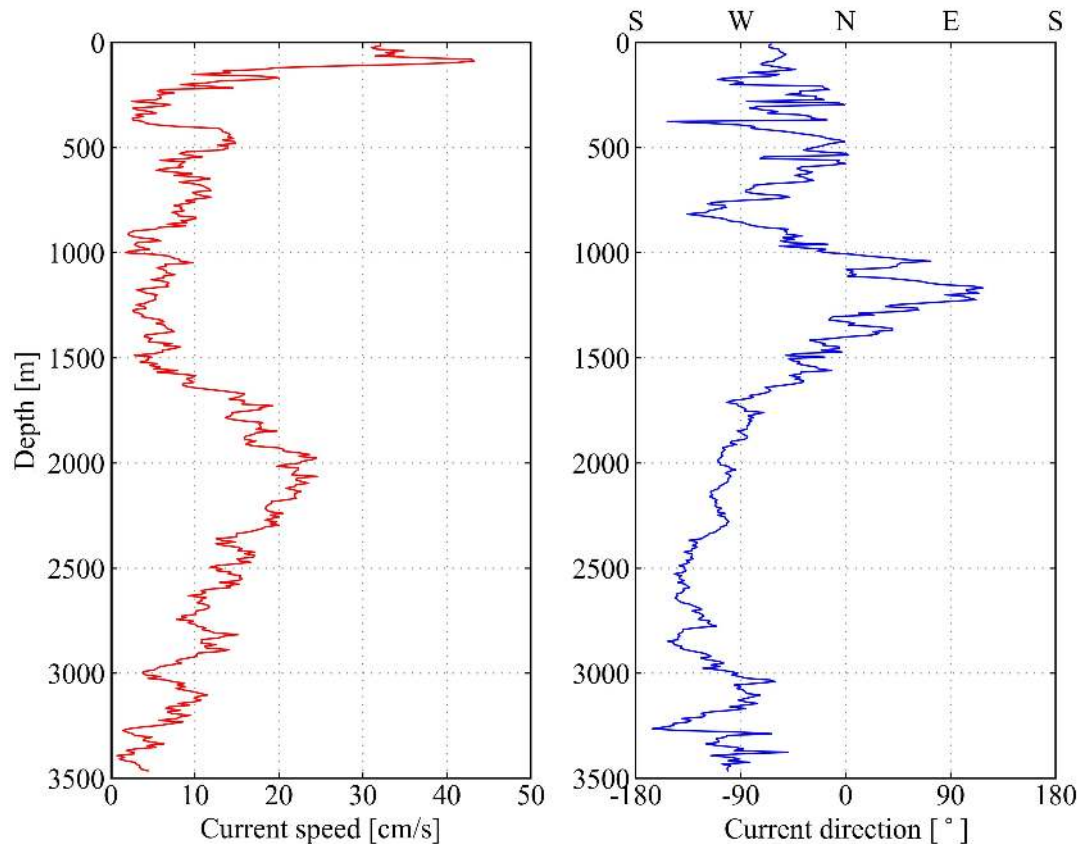
Nature Astronomy 7, 1497-1505 (2023)



Site current field measured on Sep. 6, 2021

Simulation (30-yr): ave. 6 cm/s, max < 26 cm /s

Scaled-down (1:25) experiments in a ship towing tank on SJTU campus



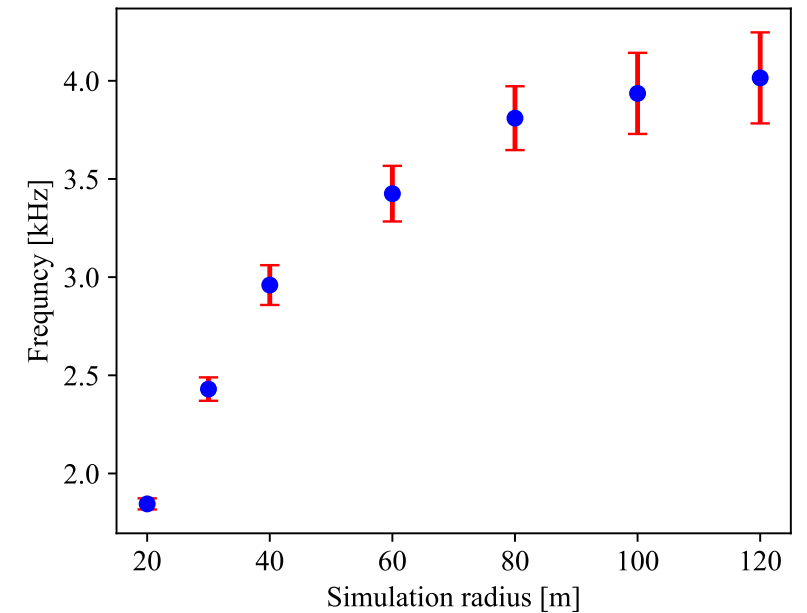
[Animation link](#)

TRIDENT Explorer : Radioactivity

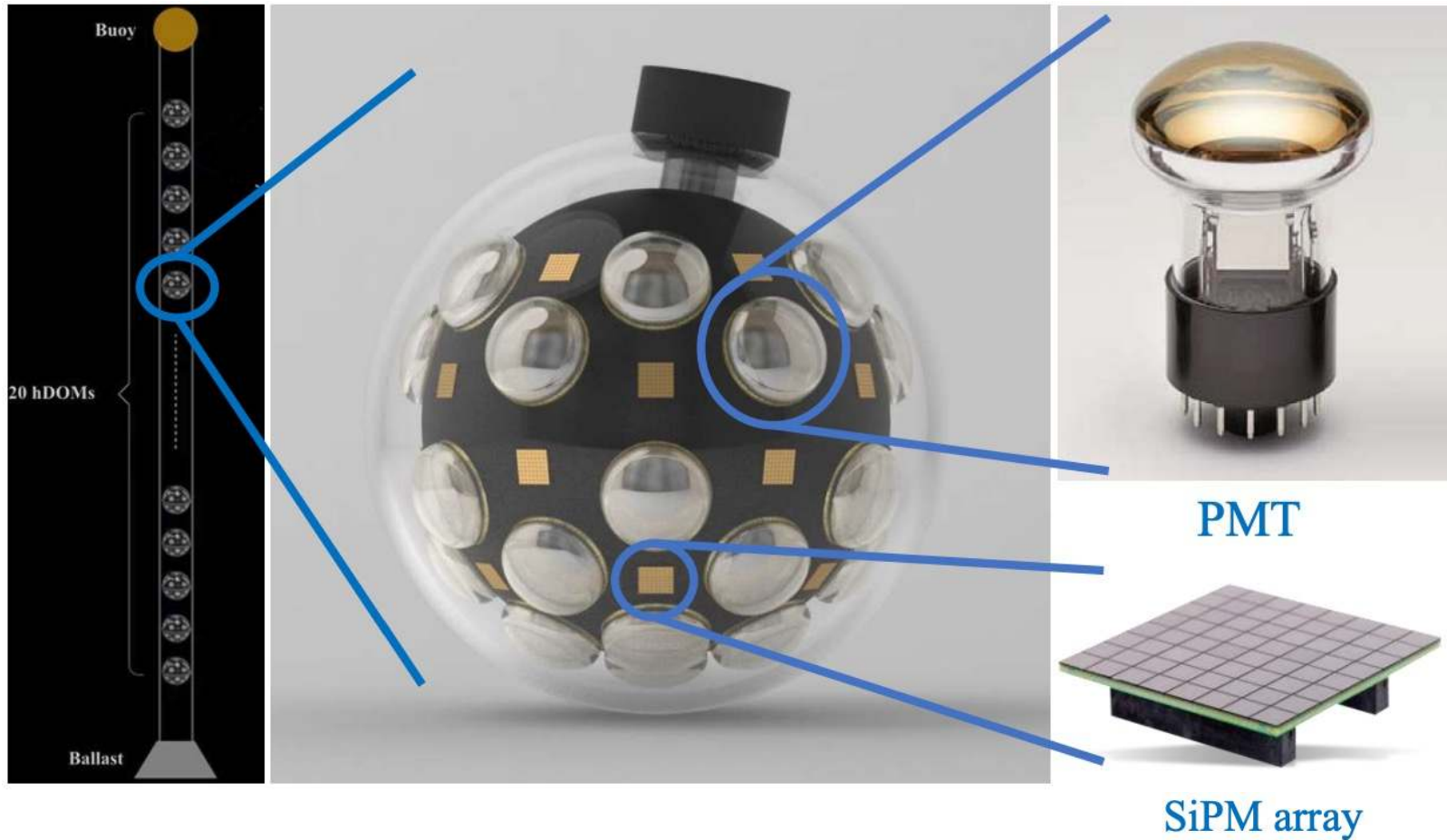


Simulated hit on each PMT caused by ^{40}K

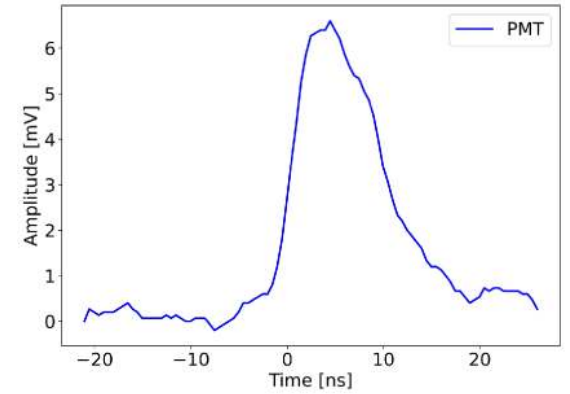
	West Pacific	Mediterranean	East Pacific
^{40}K Radioactivity [Bq/m^3]	11101 ± 119	13700 ± 200	12526 ± 752
Experiments	TRIDENT	ANTARES	P-ONE



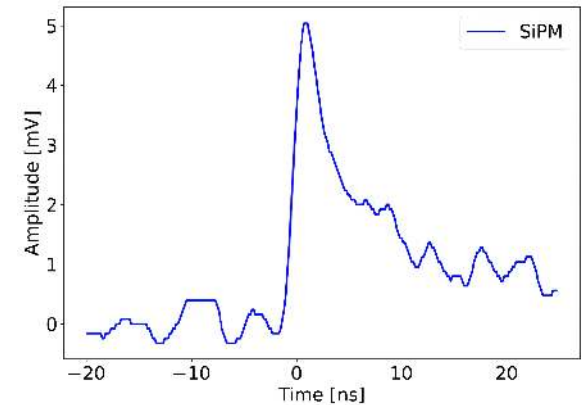
TRIDENT hybrid DOM – hDOM



TTS: < 2ns



Array time jitter: < 300ps



- Maximize photo-sensitive area & improve timing with SiPMs
- Event-by-event tau neutrino identification with PMT waveform readouts

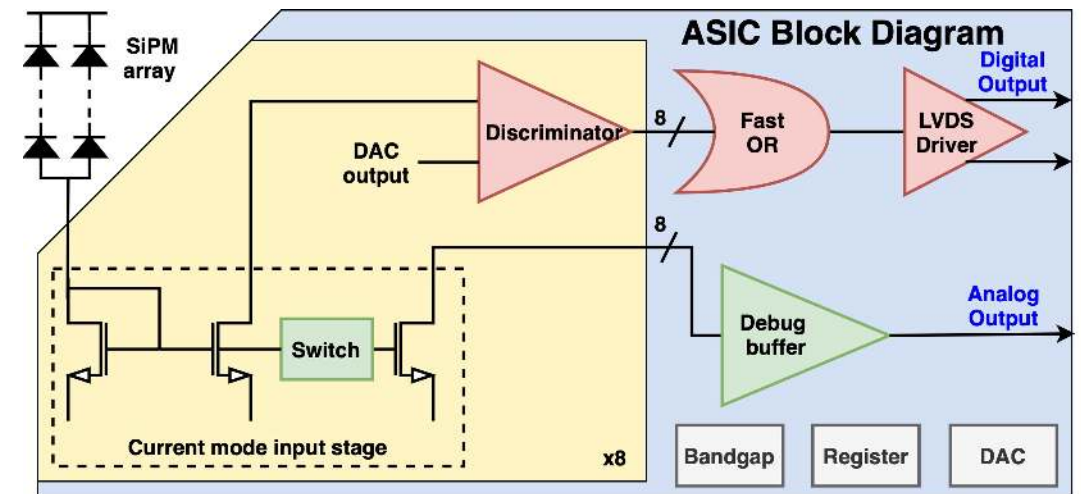
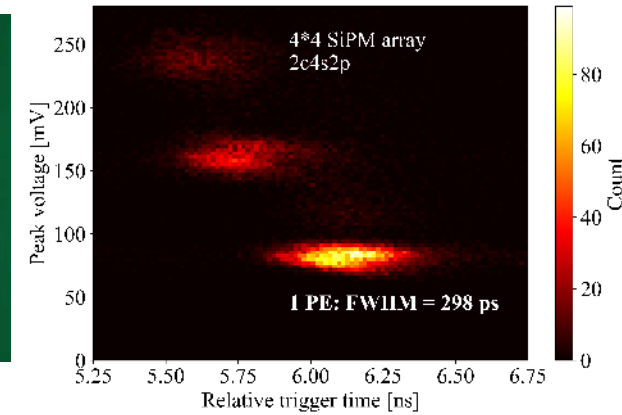
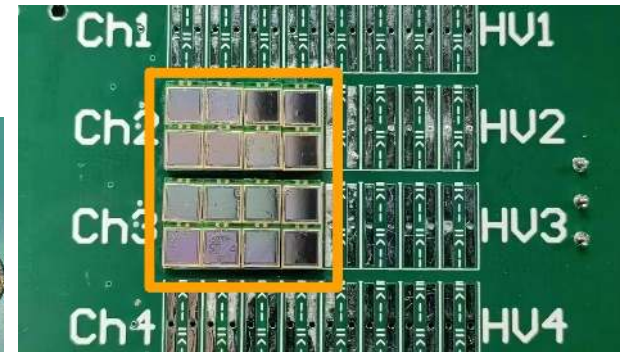


**Track better than
0.1° @ $E_\nu > 100$ TeV**

Mechanical structure of hDOM



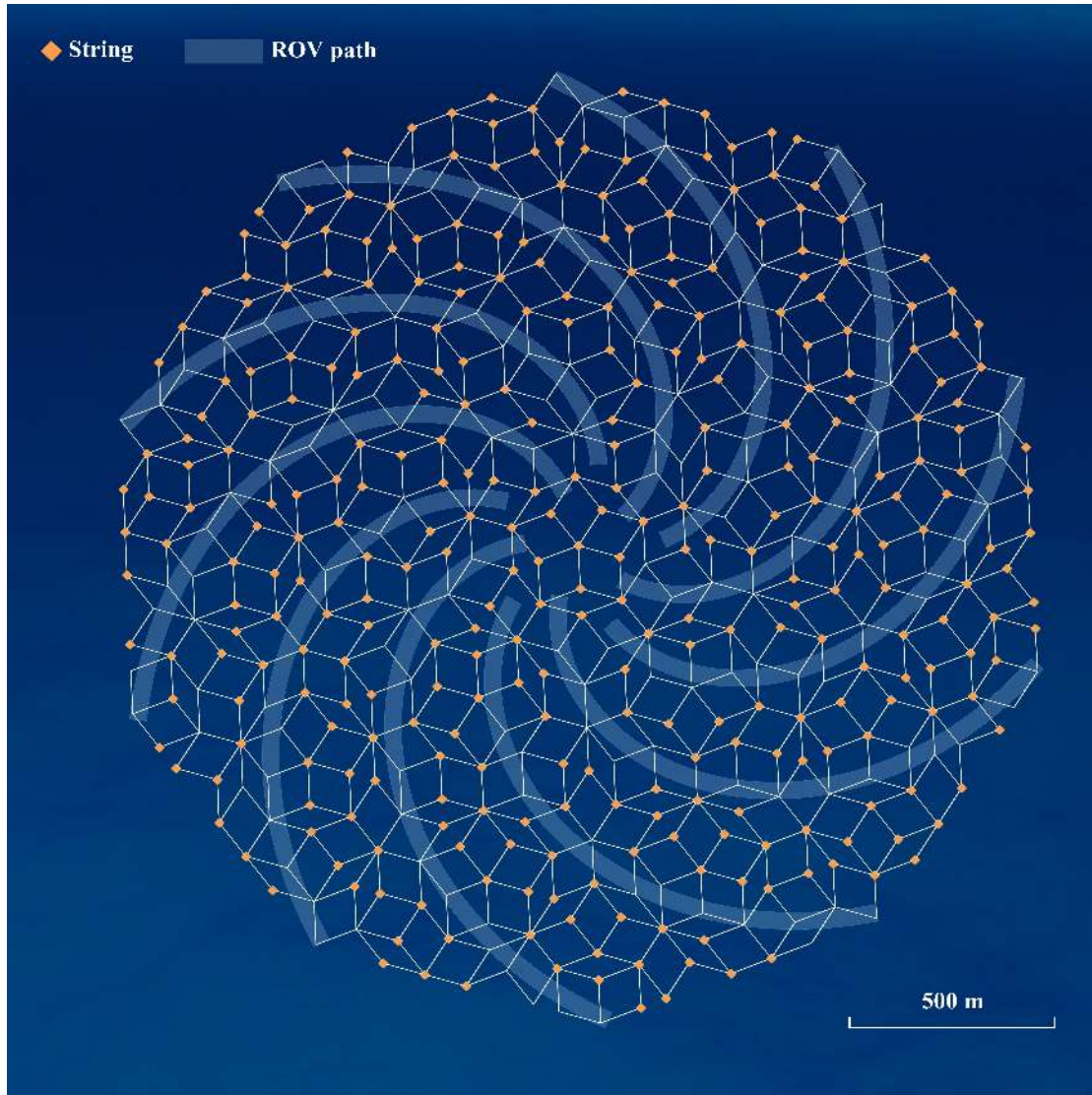
SiPM array forming & readout electronics



Development progress: PoS (ICRC2023) 1213

Conceptual design: PoS (ICRC2021) 1043

W. Zhi *et al* 2024 *JINST* 19 P06011



Penrose tiling layout

Uneven inter-string spacing **70m** and **110m**
→ expanded energy window of **sub TeV – EeV**

No translational or rotational **symmetry**
→ better rejection of “corridor” atmo. muons

1200 strings; **20** hDOMs / string

Volume: $\sim 8 \text{ km}^3$

Underwater ROV for deployment & maintenance

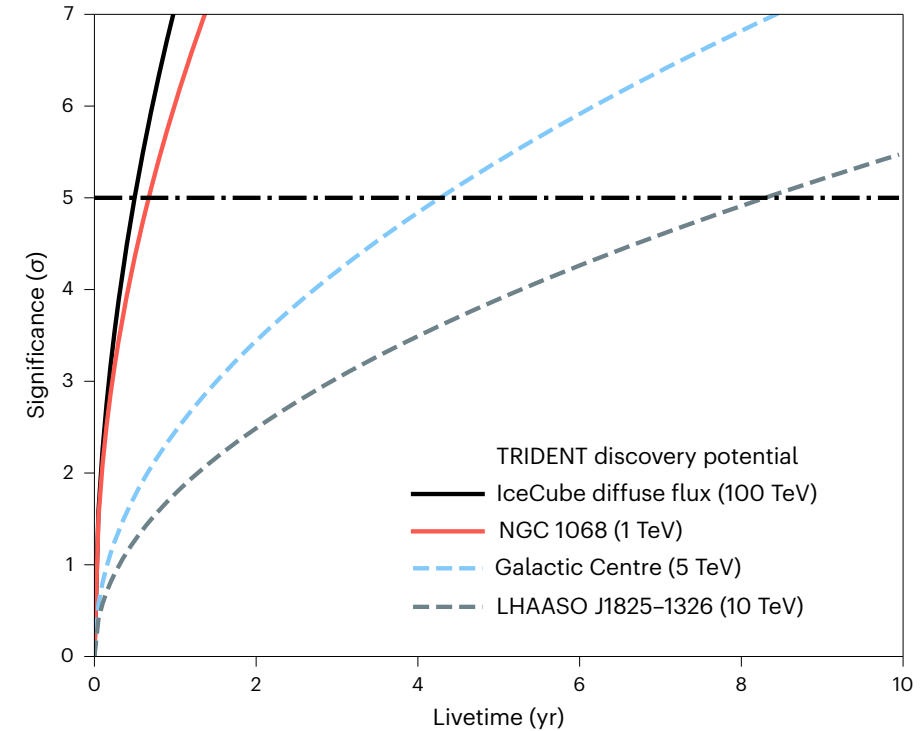
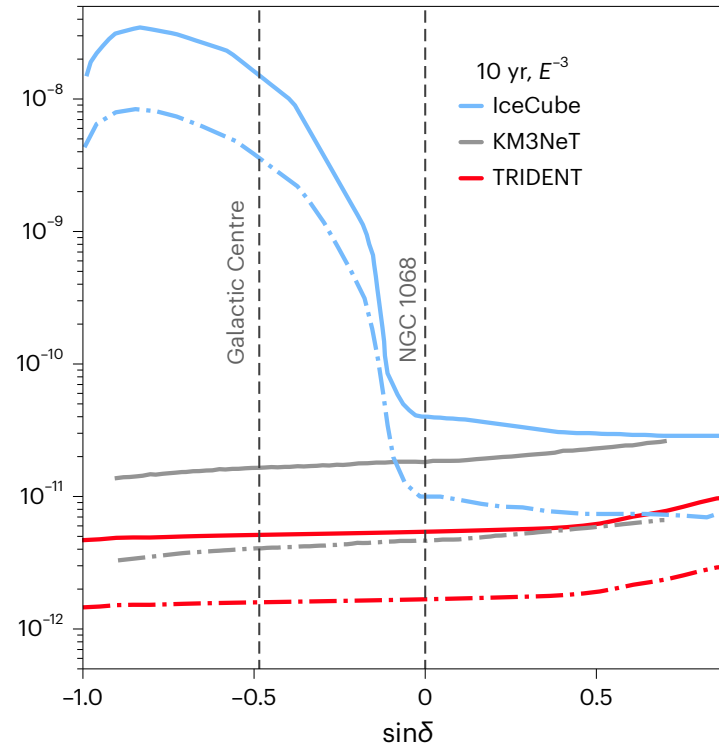
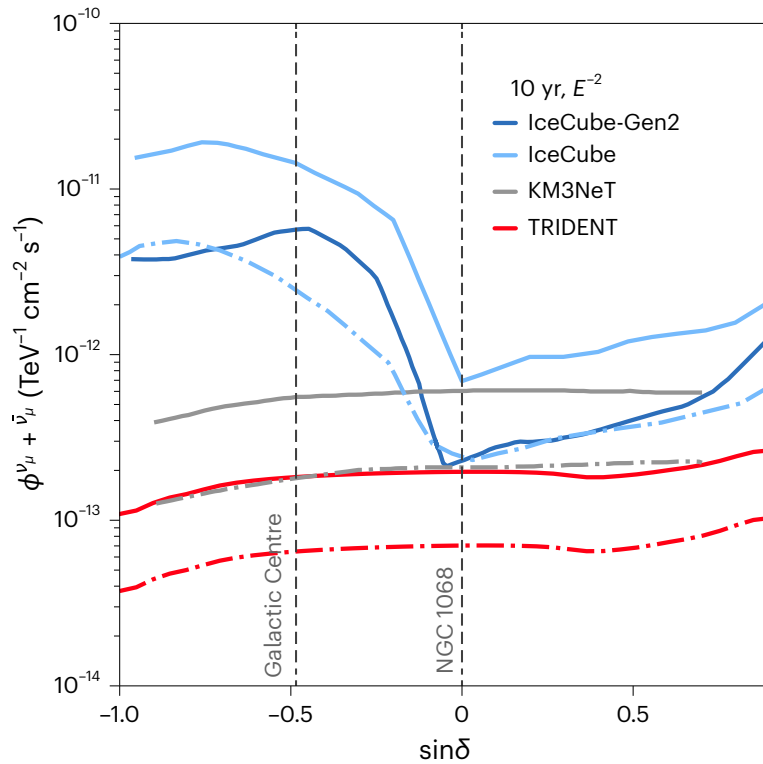
Nature Astronomy 7, 1497-1505 (2023)

Geometry comparison: PoS (ICRC2023) 1203

TRIDENT Source sensitivity & discovery potentials



Track events only



TRIDENT is expected to detect the IceCube steady source candidate **NGC1068 at 5 σ level**
within one year of operation

Nature Astronomy 7, 1497-1505 (2023)

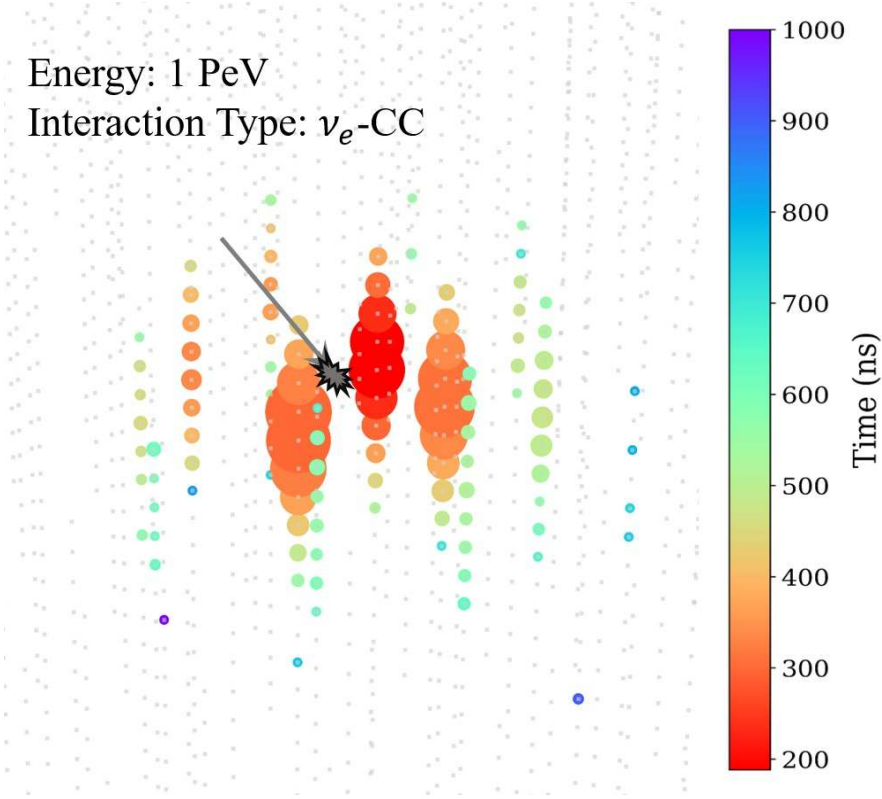
Angular resolution for cascades:

$\sim 1.8^\circ$ @ 1 PeV (likelihood)
 $\sim 1.5^\circ$ @ 100 TeV & 1 PeV (GNN)

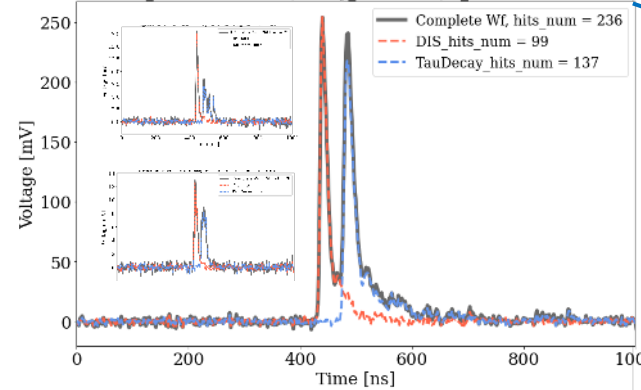
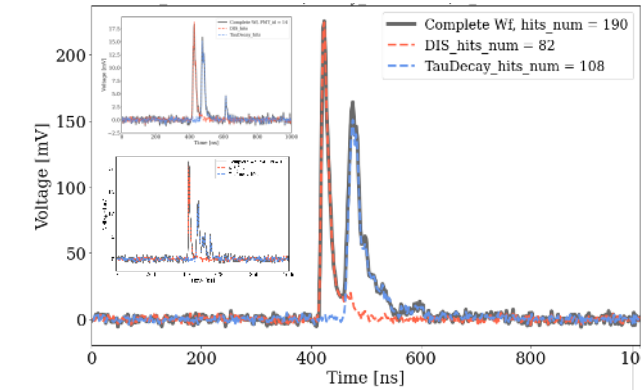
Where are the ν_e and ν_τ from NGC 1068 and TXS 0506+056 ?



Energy: 1 PeV
Interaction Type: ν_e -CC

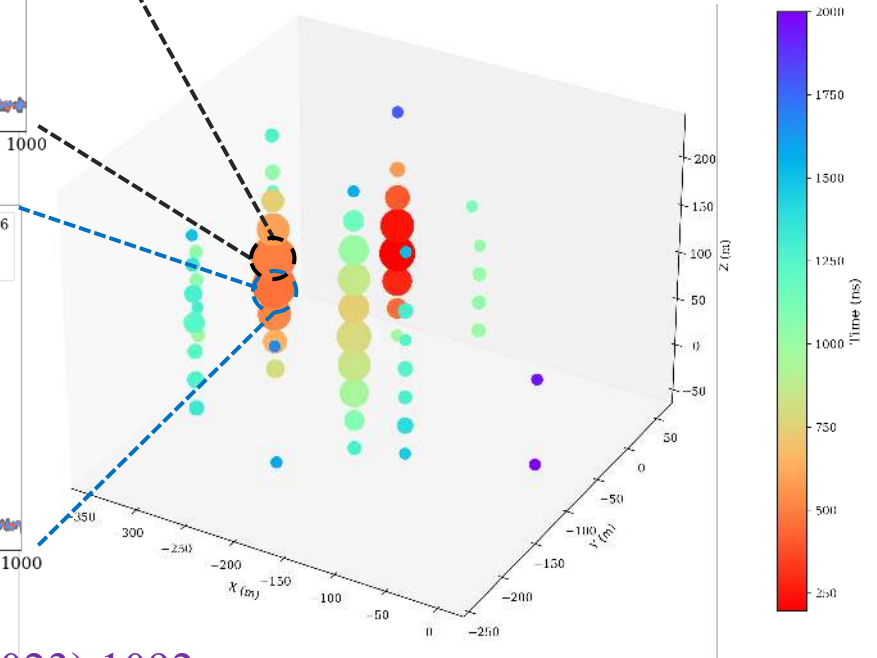


Cascade reco : PoS (ICRC2023) 1207

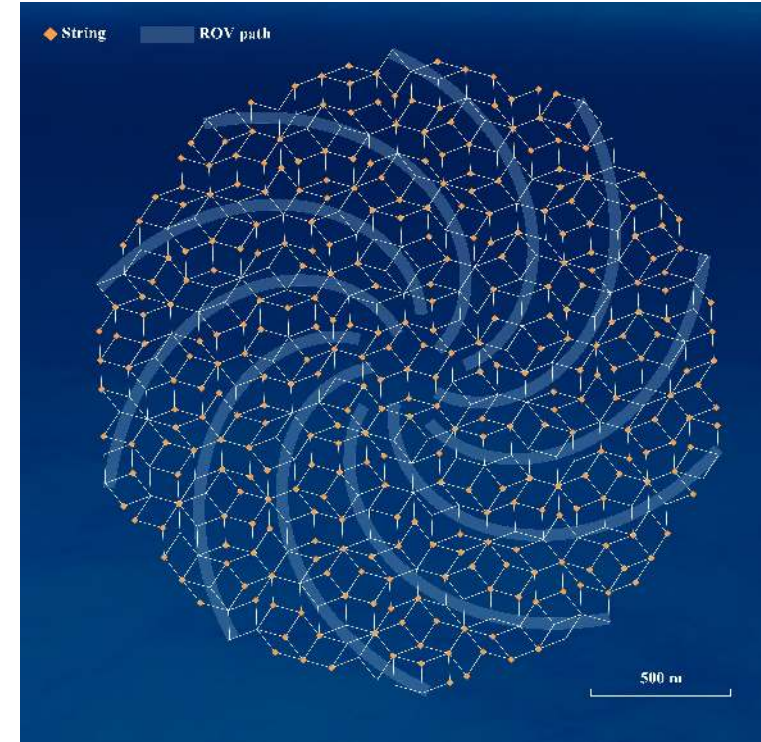
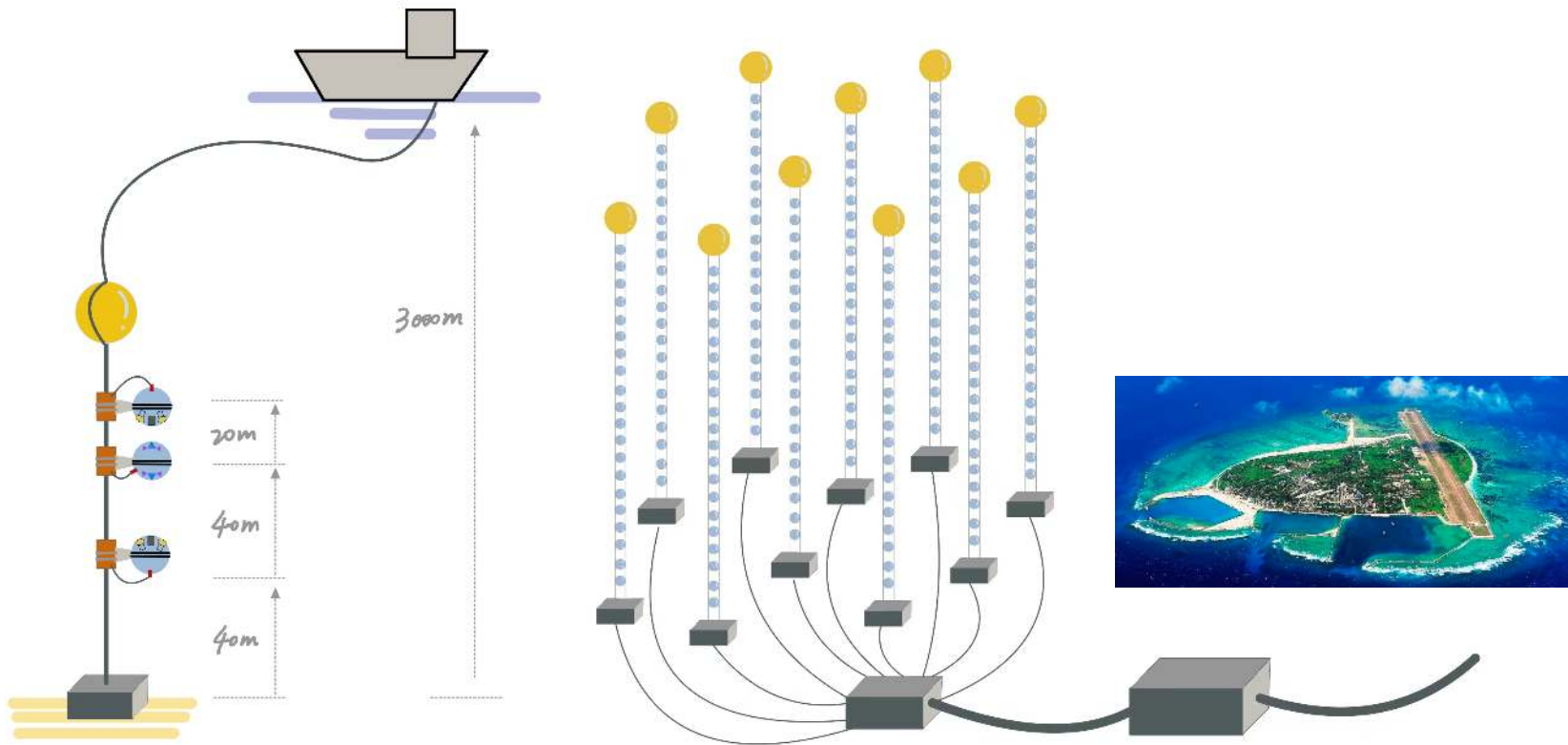


Tau neutrinos : PoS (ICRC2023) 1092

$E\nu_\tau = 334.3$ TeV



Brief timeline of TRIDENT



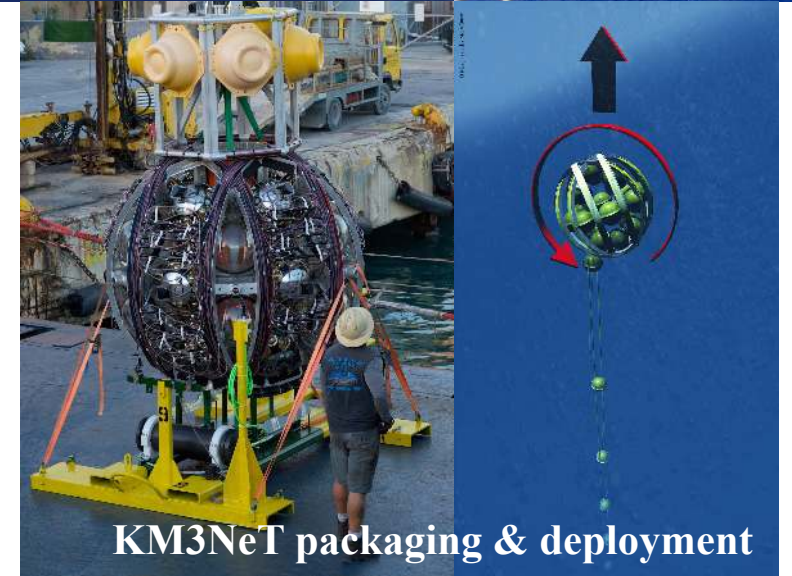
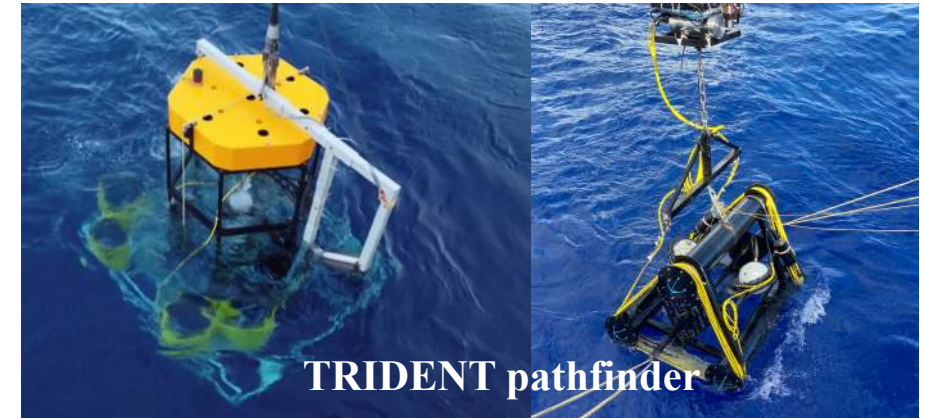
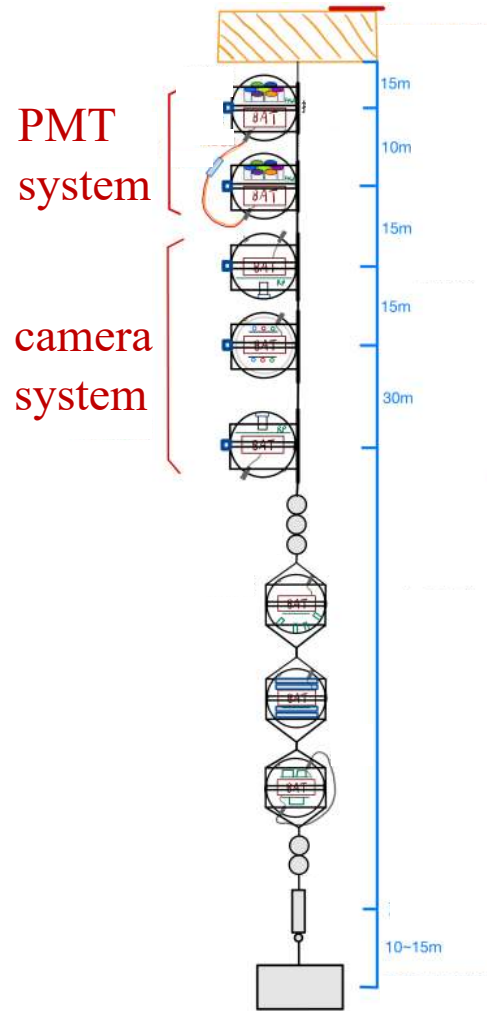
Pathfinder: 2019–2022
completed

Phase-I project: 2022–2026
in progress

Big array construction: 2026–
under planning

1) Environmental monitoring moorings 2) Prototype telescope strings 3) Deployment device & strategies

moorings



Summary

- IceCube has opened a new era for high-energy neutrino astronomy
- **More telescopes with improved detection ability are needed** to catch PLENTY of neutrinos for further scrutiny
- **New:** a viable site was found at a depth of 3.5km in South China Sea for constructing large-scale deep-sea neutrino telescopes
- A multi-cubic-kilometer next-gen deep-sea neutrino telescope (TRIDENT) has been proposed to rapidly push beyond IceCube's horizons, to exploit the discovery opportunities in the nascent field of ν astronomy

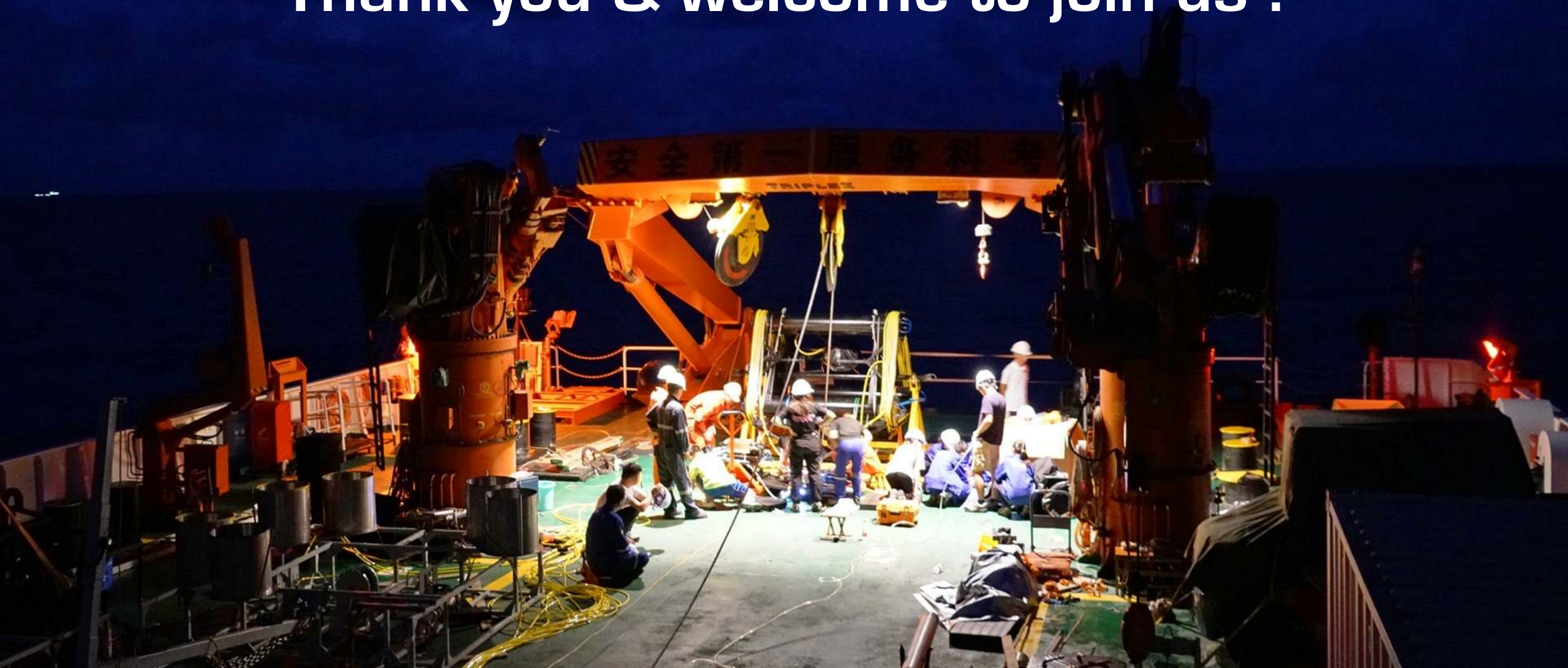


李政道研究所
TSUNG-DAO LEE INSTITUTE

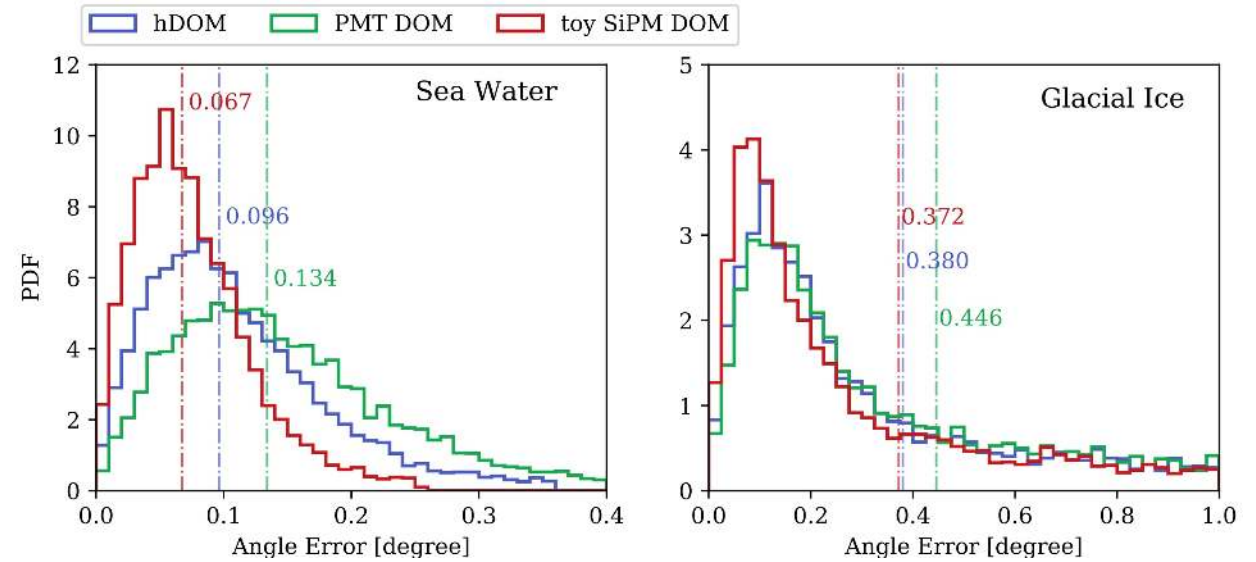
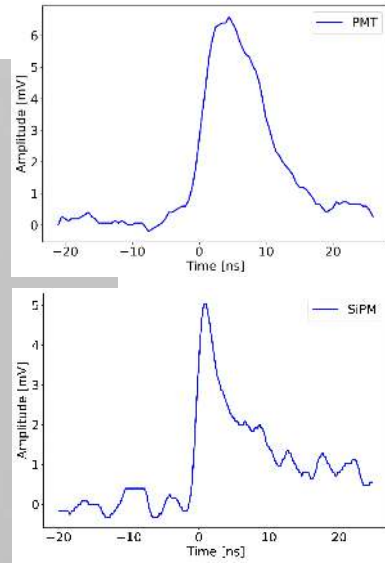
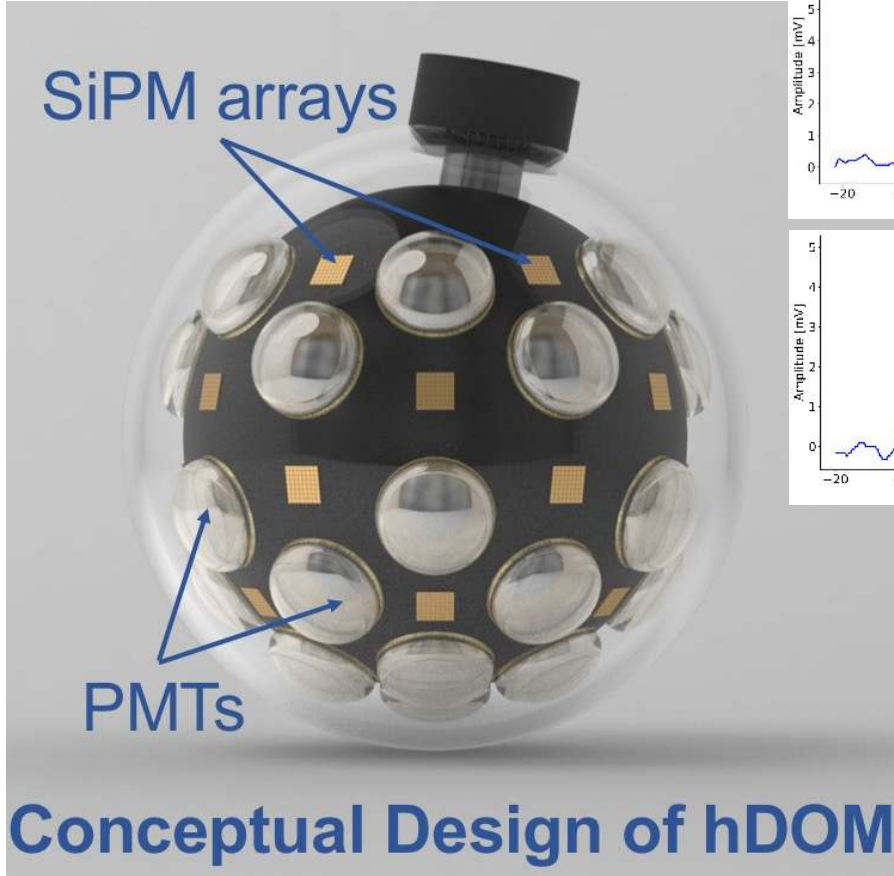
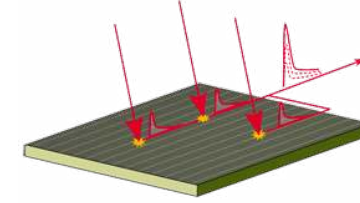
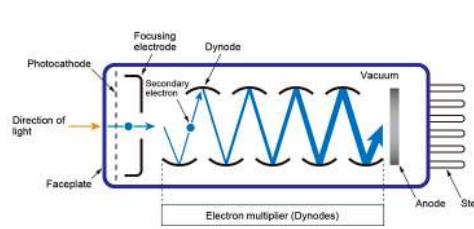


TRIDENT
海 | 铃 | 计 | 划

Thank you & welcome to join us !



TRIDENT hybrid DOM – hDOM



- Better than 0.1° @ $E_\nu > 100$ TeV
- **>40% improvement** (cf mDOM) in angular resolution, assuming PMT TTS ~ 5 ns

Updated:

PMT TTS ~ 3 ns + 10cm hDOM position smearing: 40% \rightarrow 30%

Conceptual design: PoS (ICRC2021) 1043

Development progress: PoS (ICRC2023) 1213