

# Neutrino mass bounds from DESI 2024 are relaxed by Planck PR4 and cosmological supernovae



BROWN



Chicago 2024

Itamar J. Allali

Department of Physics, Brown University



August 14, 2024

Based on

IJA, A. Notari 2406.14554

- 1 Neutrino Mass from Cosmology
- 2 Including Planck PR4 and Cosmological Supernovae Data
- 3 Bonus: Dark Radiation

- 1 Neutrino Mass from Cosmology
- 2 Including Planck PR4 and Cosmological Supernovae Data
- 3 Bonus: Dark Radiation

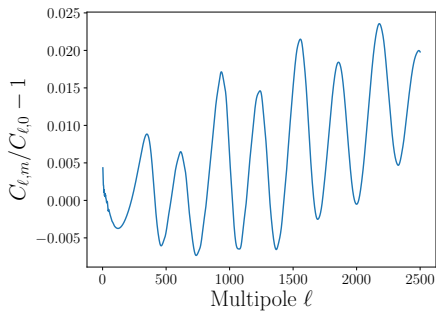
# Sensitivity of Cosmology to Neutrino Mass

Neutrinos impact cosmic expansion history; data are becoming more sensitive to the effects of massive neutrinos

# Sensitivity of Cosmology to Neutrino Mass

Neutrinos impact cosmic expansion history; data are becoming more sensitive to the effects of massive neutrinos

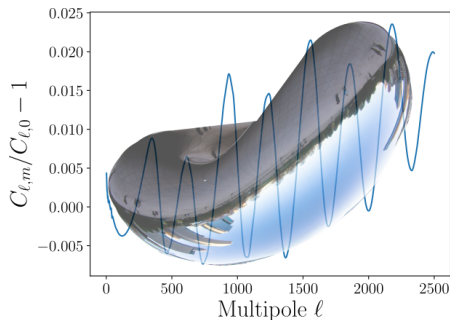
- Transition of (some) neutrinos from relativistic to non-relativistic increases matter abundance
- Enhances the early ISW effect, suppressing the first CMB peak
- In addition, shift of all peaks to lower  $\ell$



# Sensitivity of Cosmology to Neutrino Mass

Neutrinos impact cosmic expansion history; data are becoming more sensitive to the effects of massive neutrinos

- Transition of (some) neutrinos from relativistic to non-relativistic increases matter abundance
- Enhances the early ISW effect, suppressing the first CMB peak
- In addition, shift of all peaks to lower  $\ell$



# Constraints from CMB and BAO

Fitting to cosmological data can provide a strong upper bound on the sum of neutrino masses

- Assuming the  $\Lambda\text{CDM} + \sum m_\nu$  model
- Using cosmic microwave background data from Planck 2018 (Aghanim et al 2020)
- and baryon acoustic oscillations measurements from SDSS BOSS DR12 (Alam et al 2017)

$$\sum m_\nu < 0.12 \text{ eV (Planck18 + SDSS DR12)} \quad (1)$$

# Recent constraints from DESI BAO

Recent measurements of BAO from the Dark Energy Spectroscopic Instrument (DESI)

- 10x targets, 5-10x precision, redshift range  $(0,1) \rightarrow (0,2.1)$



# Recent constraints from DESI BAO

Recent measurements of BAO from the Dark Energy Spectroscopic Instrument (DESI)

- 10x targets, 5-10x precision, redshift range (0,1)→(0,2.1)

Report stringent bound on neutrino mass

$$\sum m_\nu < 0.072 \text{ eV (DESI + Planck18 + PR4/ACT lensing)} \quad (2)$$

(Adame et al 24 (DESI VI))

# Recent constraints from DESI BAO

Recent measurements of BAO from the Dark Energy Spectroscopic Instrument (DESI)

- 10x targets, 5-10x precision, redshift range (0,1)→(0,2.1)

Report stringent bound on neutrino mass

$$\sum m_\nu < 0.072 \text{ eV (DESI + Planck18 + PR4/ACT lensing)} \quad (2)$$

(Adame et al 24 (DESI VI))

Essentially claims to rule out inverted neutrino mass hierarchy  
( $\sum m_\nu > 0.1 \text{ eV}$ )

# Data Analysis

Goal of this work

Evaluate  $\sum m_\nu$  upper bounds with broad set of cosmological data

# Data Analysis

## Goal of this work

Evaluate  $\sum m_\nu$  upper bounds with broad set of cosmological data

Cosmologies are computed with Einstein-boltzmann solver

CLASS: (Blas + Lesgourgues + Tram 11)

Markov Chain Monte Carlo analysis using

MontePython (Audren et al 12, Brinckmann + Lesgourgues 18)

and Cobaya (Torrado + Lewis 20)

- 1 Neutrino Mass from Cosmology
- 2 Including Planck PR4 and Cosmological Supernovae Data
- 3 Bonus: Dark Radiation

# Planck PR4 Likelihoods

Final release (PR4) Planck data has new likelihoods, including Hillipop+Lollipop ( $P20_H$ )

(Tristram et al 24)

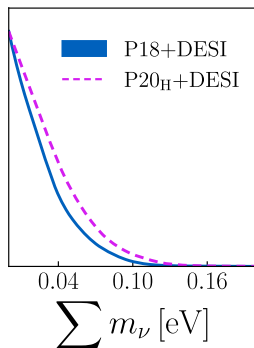
- Incorporates final Planck CMB data (Akrami et al 20)
- Alleviates significantly the  $A_L$  “anomaly”

# Planck PR4 Likelihoods

Final release (PR4) Planck data has new likelihoods, including Hillipop+Lollipop (P20<sub>H</sub>)

(Tristram et al 24)

- Incorporates final Planck CMB data (Akrami et al 20)
- Alleviates significantly the  $A_L$  “anomaly”



Comparing to Planck 2018 data, bounds are significantly relaxed

$$\Sigma m_\nu < 0.086 \text{ eV (DESI + P20}_H\text{)} \quad (3)$$

# Supernovae Datasets

Cosmological supernovae not included in original DESI analysis

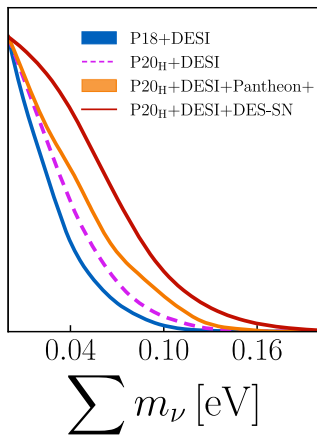
- Pantheon+ (Scolnic et al 22)
- DES-SN5YR (Abbott et al 24)



# Supernovae Datasets

Cosmological supernovae not included in original DESI analysis

- Pantheon+ (Scolnic et al 22)
- DES-SN5YR (Abbott et al 24)
- Further relaxes constraints
- Inching towards nonzero mass preference



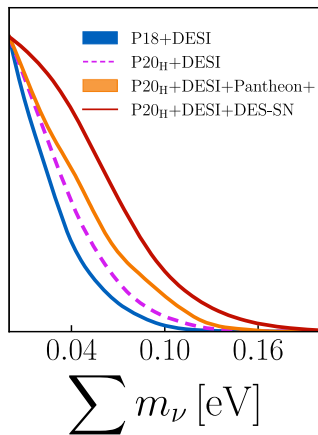
# Supernovae Datasets

Cosmological supernovae not included in original DESI analysis

- Pantheon+ (Scolnic et al 22)
- DES-SN5YR (Abbott et al 24)
- Further relaxes constraints
- Inching towards nonzero mass preference

$$\sum m_\nu < 0.099 \text{ eV} \quad (4)$$

(DESI + P20<sub>H</sub> + Pantheon+)

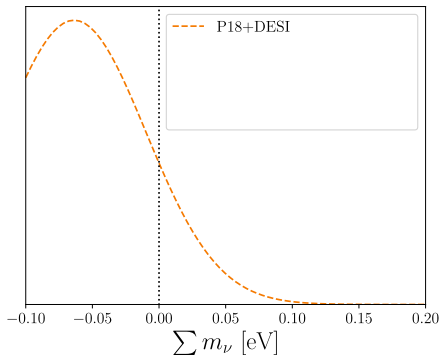


$$\sum m_\nu < 0.12 \text{ eV (DESI + P20}_H \text{ + DES-SN5YR)} \quad (5)$$

# Fitting Distribution Peaks

CMB + BAO data has shown mild preference for effectively “negative” neutrino masses

(Alam et al 21; Craig et al 24)

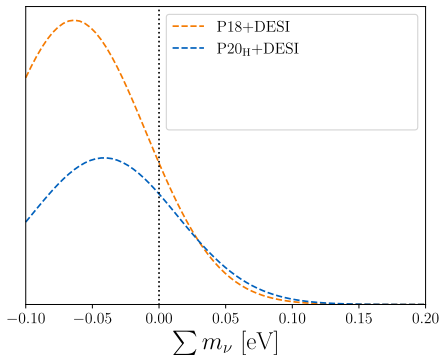


# Fitting Distribution Peaks

CMB + BAO data has shown mild preference for effectively “negative” neutrino masses

(Alam et al 21; Craig et al 24)

- Planck PR4 (P20<sub>H</sub>)  
minor shift to positive

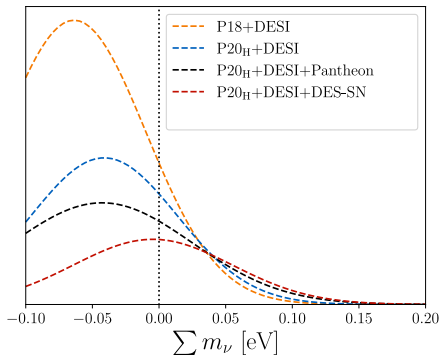


# Fitting Distribution Peaks

CMB + BAO data has shown mild preference for effectively “negative” neutrino masses

(Alam et al 21; Craig et al 24)

- Planck PR4 (P20<sub>H</sub>) minor shift to positive
- DES-SN5YR data causes large shift

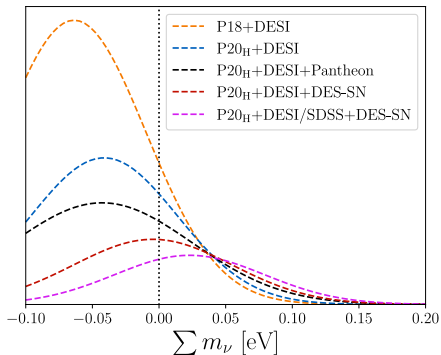


# Fitting Distribution Peaks

CMB + BAO data has shown mild preference for effectively “negative” neutrino masses

(Alam et al 21; Craig et al 24)

- Planck PR4 (P20<sub>H</sub>) minor shift to positive
- DES-SN5YR data causes large shift
- DESI/SDSS combination shift also



- 1 Neutrino Mass from Cosmology
- 2 Including Planck PR4 and Cosmological Supernovae Data
- 3 Bonus: Dark Radiation

# Beyond $\Lambda$ CDM + $\sum m_\nu$ : Fluid Dark Radiation Model

Simple extended model beyond  $\Lambda$ CDM: dark radiation (DR)

- One-parameter extension  
 $\Delta N_{\text{eff}}$
- Example: Fluid-like DR  
(self-interacting)
- In light of DESI, potential  
resolution to Hubble tension

Grey bands show direct  $H_0$  measurement  
(Riess et al 22)

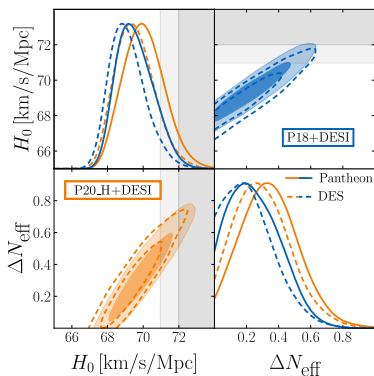


# Beyond $\Lambda$ CDM + $\sum m_\nu$ : Fluid Dark Radiation Model

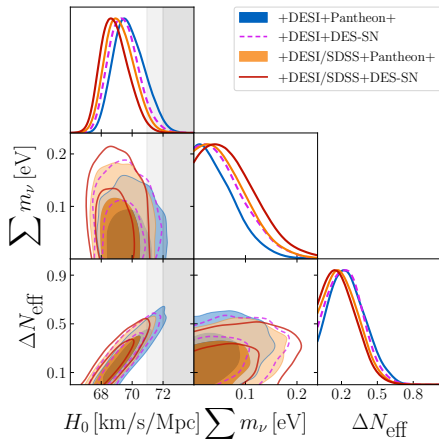
Simple extended model beyond  $\Lambda$ CDM: dark radiation (DR)

- One-parameter extension  $\Delta N_{\text{eff}}$
- Example: Fluid-like DR (self-interacting)
- In light of DESI, potential resolution to Hubble tension

Grey bands show direct  $H_0$  measurement (Riess et al 22)

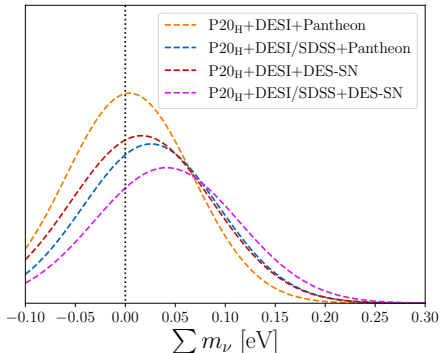


# $\sum m_\nu$ Constraints with Fluid DR



Dataset $P20_H+$	$\sum m_\nu$ [eV]
<b>DESI</b>	
+Pantheon	$< 0.13$
+DES-SN	$< 0.15$
<b>DESI/SDSS</b>	
+Pantheon	$< 0.15$
+DES-SN	$< 0.17$

# $\sum m_\nu$ Constraints with Fluid DR



Dataset P20 <sub>H</sub> +	$\sum m_\nu$ [eV]
<b>DESI</b>	
+Pantheon	< 0.13
+DES-SN	< 0.15
<b>DESI/SDSS</b>	
+Pantheon	< 0.15
+DES-SN	< 0.17

# Summary

- Claims about neutrino mass constraints are highly dependent on choice of data
- Planck PR4, supernovae datasets relax constraints
  - $\sum m_\nu < 0.072$  eV (P18+DESI)
  - $\sum m_\nu < 0.12$  eV (P20<sub>H</sub>+DESI+DES-SN5YR)
- Not yet able to rule out inverted mass hierarchy
- Showing signs of future detection with more data, especially in  $\Lambda$ CDM+ $\sum m_\nu$ +DR