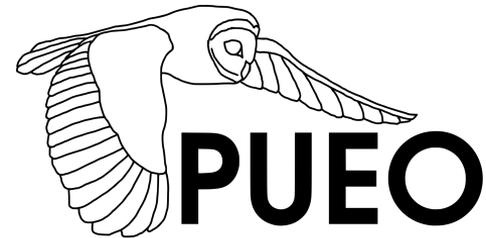
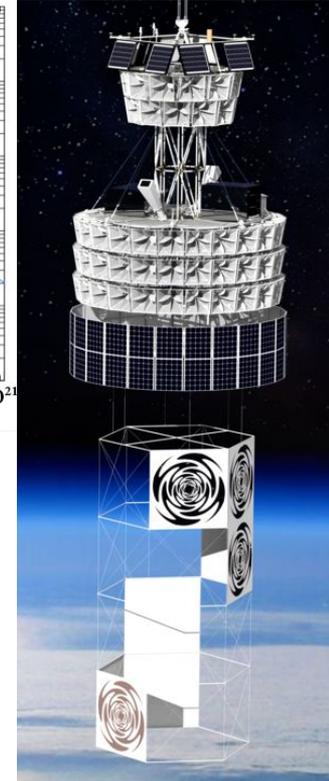
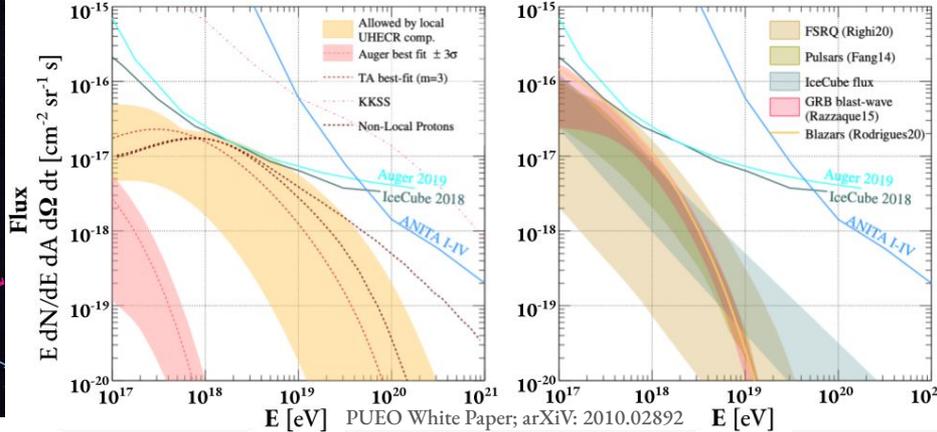
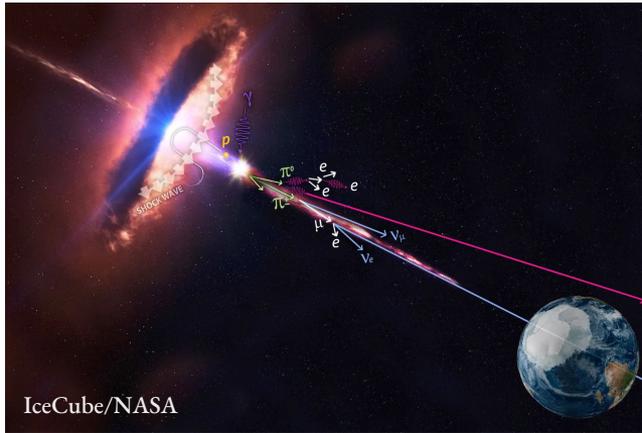


# Design and Hardware Overview of PUEO

Rachel Scrandis for the PUEO Collaboration



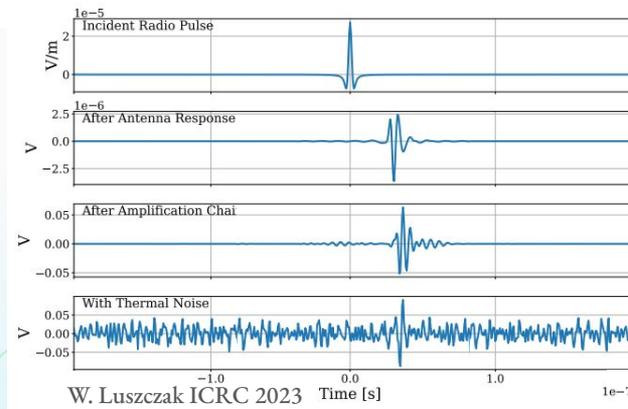
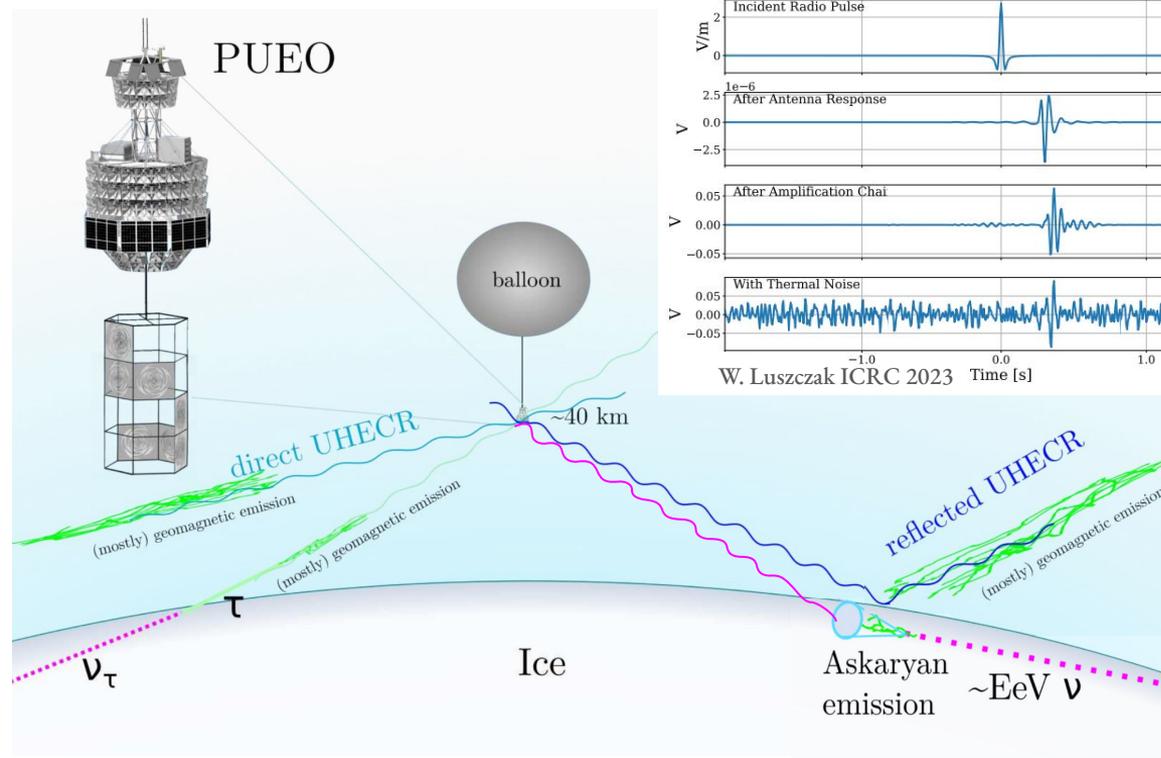
# PUEO's Science Goals



PUEO's goal is to use radio to measure (or set the best limits on) the UHE neutrino flux in order to understand the most extreme environments of our universe

Based off of ANITA heritage, it will leverage the Askaryan emission mechanism to measure these neutrinos

# PUEO's Science Signal



ANITA-IV Recorded Data ( $O(10^7)$ events )	
Thermal Noise	~99% of data
Anthropogenic	~0.9% of data
Cosmic Rays	~30 events
Neutrino Candidates	~a few events

A. Ludwig Thesis

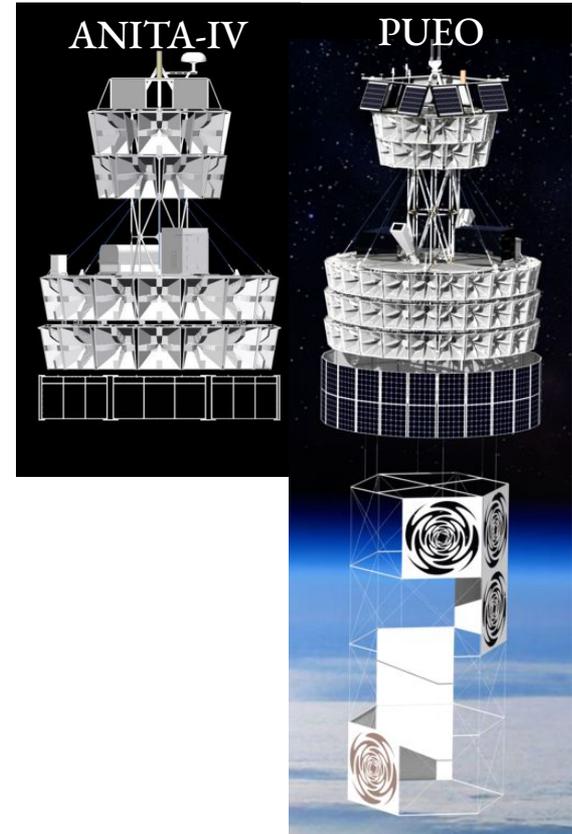
Over 4 successful flights, ANITA saw 100s of UHECRs, no excess in neutrino-like events, & a few upward-going air shower candidates

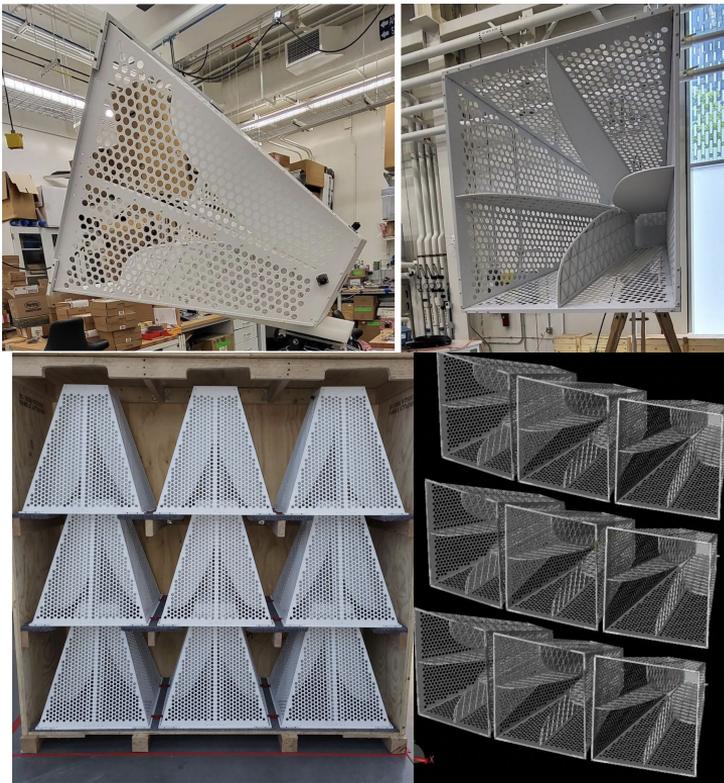
PUEO will improve on background rejection & trigger threshold to increase discovery potential for UHE neutrinos

PUEO will be searching for impulsive radio signals in the 300-1200 MHz range and will be capable of reconstructing neutrino direction, energy, and potentially flavor

PUEO will feature:

- 96 quad-ridge horn antennas (192 RF channels)
- 8 sinuous low frequency antennas (16 RF channels)
- Phased array trigger
- Telemetry capabilities
- Navigation suite
- On board power system & housekeeping



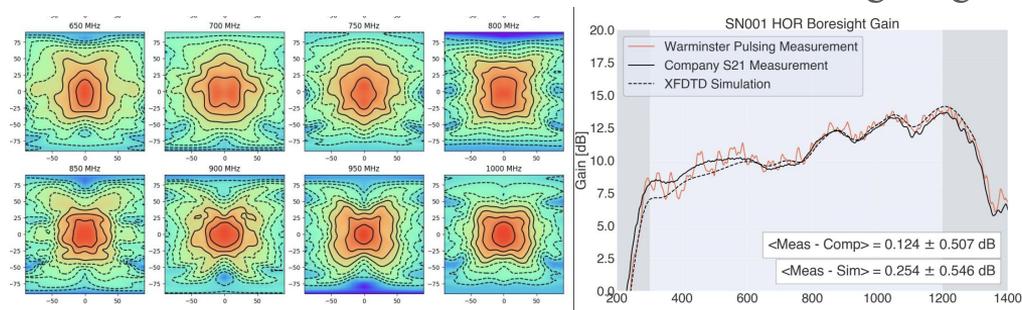


R. Scrandis

Custom antennas feature (in 300-1200 MHz band):

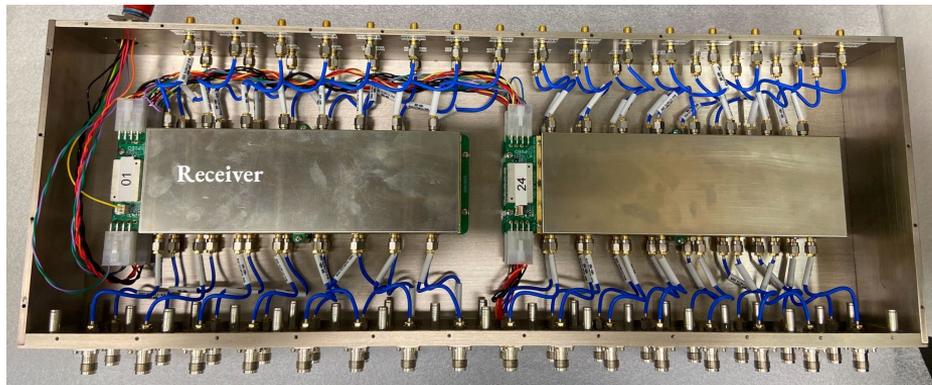
- Dual polarization
- $< -10\text{dB}$  in return loss ( $S_{11}$ )
- $< -30\text{dB}$  in port isolation ( $S_{21}$ )
- $\sim 10\text{dBi}$  boresight gain (ANITA-IV gain:  $\sim 8\text{dBi}$ )

Characterization measurements and sims ongoing

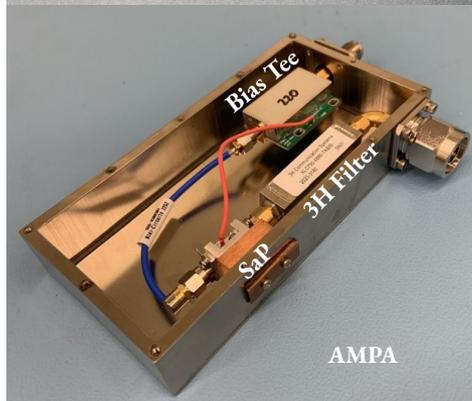


Design & Hardware of PUEO

Work by Zack Martin, Scott Mackey, & Natalie Orrantia



Full Flight Chain Noise Temp



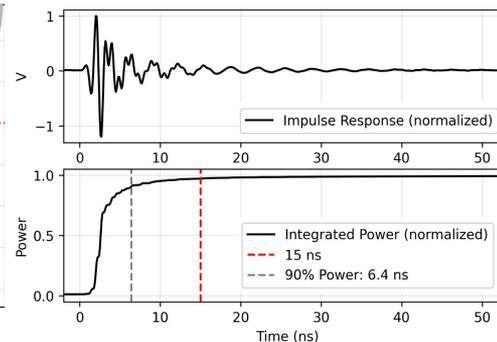
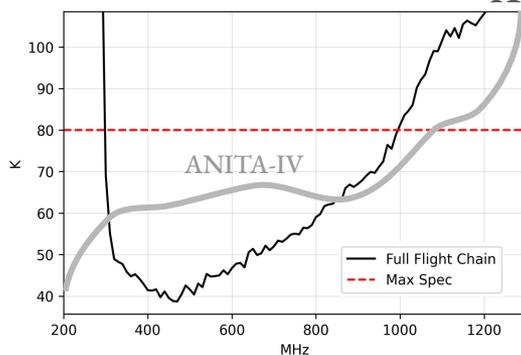
AMPA

Signal conditioning happens in 2 stages:

1. Antenna Mounted Pre-Amplifier (AMPA) gives  $\sim 40\text{dB}$
2. Octal Receiver (2nd stage amp) gives  $\sim 30\text{dB}$

With cable losses, chain presents  $68\text{dB}$  of gain & 23% lower noise temp (inherit chain noise) than

ANITA-IV Full Flight Chain

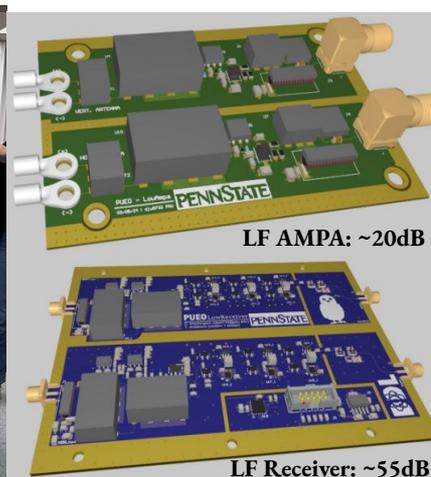
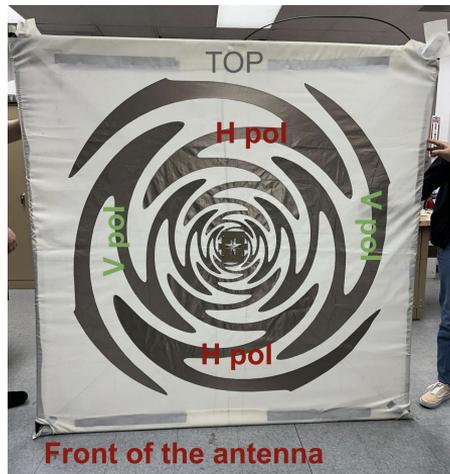
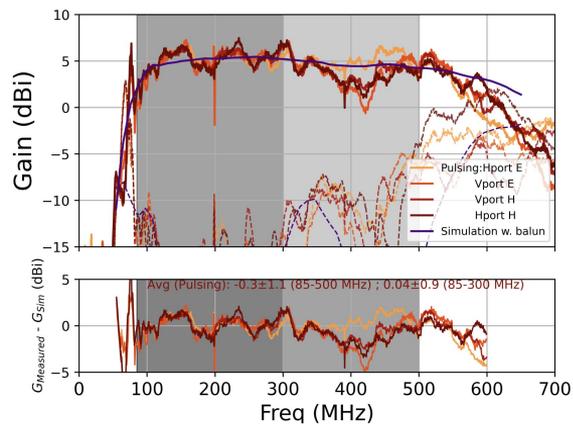


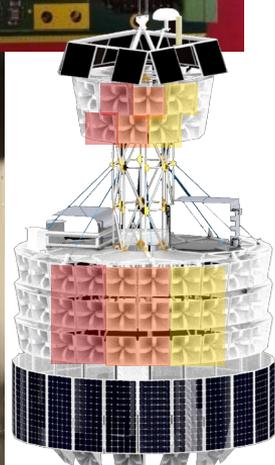
Work by me, Eric Oberla,  
Peter Gorham, Claire  
Kaplan, and Tori Ankel

LF Instrument will target lower frequency air showers (CRs and upgoing Taus). Antennas are sinuous design, sensitive to 50-300MHz.

- Conductive fabric (Nylon ripstop coated with Ni/CuAg) on polyester
- ~5dBi boresight gain and 30° beam width
- Will be stowed in gondola during launch, and deployed at float

Boresight Gain (front), angle=0°





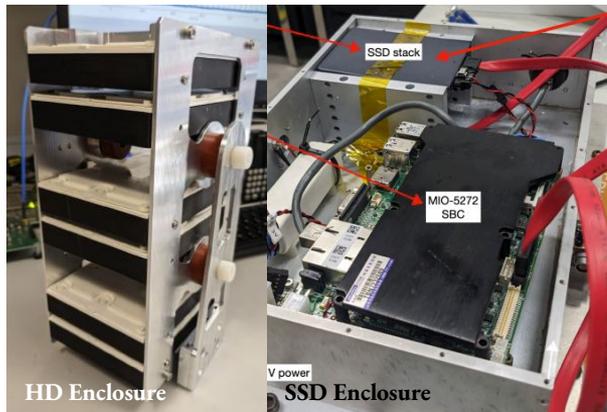
## Sampling Unit for Radio Frequencies (SURF)

- Uses RFSocS (x8 ADC @ 3GHz) to digitize 8 RF channels (8 antennas, one pol = 2 phi sectors)
- Conditions signal with FIR low pass filter (boost SNR), 2 biquad notch filters (satellite noise rejection), and 12-to-5 bit conversion (offset & gain control)
- On board FPGA handles beamforming & triggering:
  - L1 Trigger: 100kHz, beam forming with 2 phi sectors (1 SURF)
  - L2 Trigger: 100Hz, beam forming with 4 phi sectors (pulls from neighboring SURF)

## Triggering Unit for Radio Frequencies (TURF)

- L3 Trigger: under design (CW elimination, polarization check, etc...)
- Communicates with SFC to record triggered data

# Data Storage & Downlinking

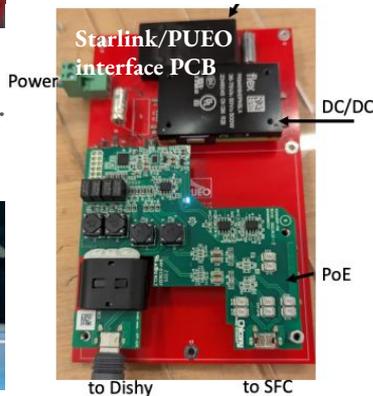
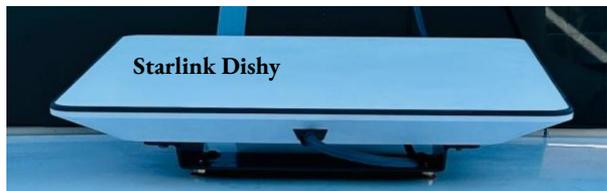


Triply redundant data storage onboard

- 2 HD enclosures (128TB of storage each)
- 1 SSD enclosure (80TB of storage)

Quick connect mounting/demounting for fast takeaway

Starlink used recently on balloon flights → exploring use for PUEO (after mitigating RFI & thermal concerns)



Data downlinking rates dependant on what we decide to use:

- LOS EVTM (up to 12.8 Mbps or so)
- Iridium (255 bytes/every 15M)
- TDRS (6 kbps)
- Starlink (1-40 Mbps downlink)

# Location & Orientation Tracking



## ABX2:

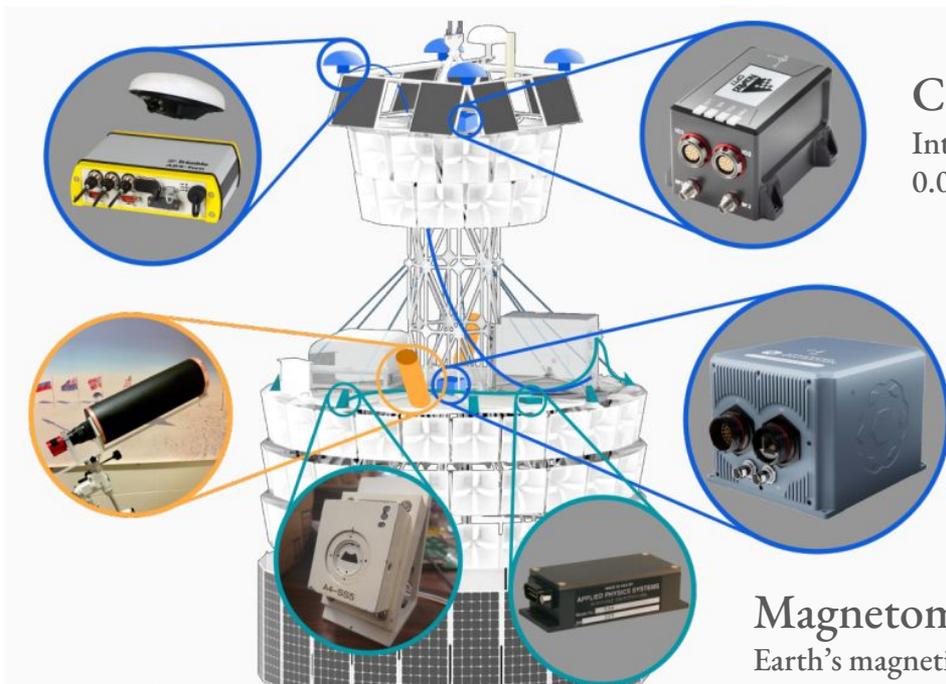
Differential GNSS  
0.02° heading, 0.075° pitch/roll

## Star Tracker:

Mag <6 stars  
0.001° heading, 0.001° pitch/roll

## Sun Sensor:

Sun  
0.17° heading, ~0.3 ° pitch/roll



## CPT7 IMU:

Int gyroscope and accelerometer  
0.03° heading, 0.01° pitch/roll

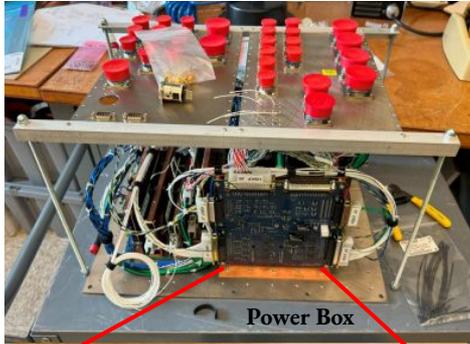
## Boreas IMU:

Int gyroscope and accelerometer  
0.005° heading, 0.006° pitch/roll

## Magnetometer:

Earth's magnetic field  
~0.1 ° heading, ~0.1 ° - 0.5 ° pitch/roll

# Power & Housekeeping

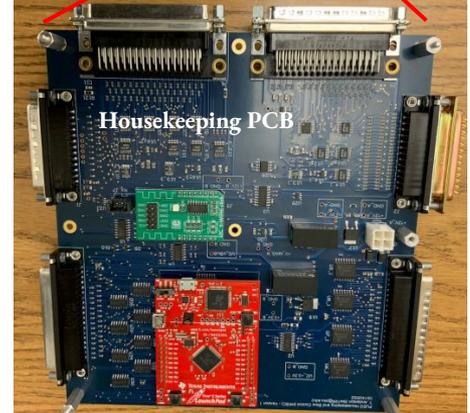


Power comes from PV array → designed to generate 1800W

- Power box supplies 24V, 12V, 5V, and 4V across payload

Housekeeping board also lives in power box

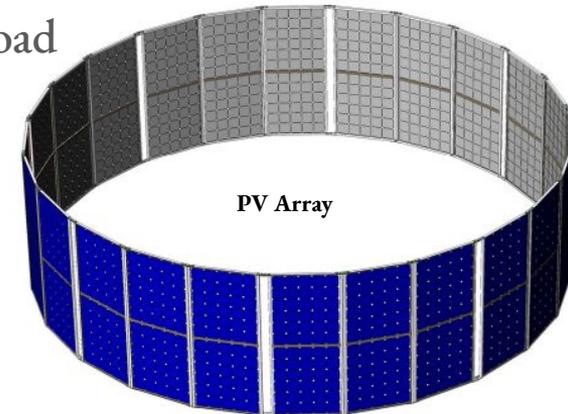
- Monitors power consumption across payload
- Monitors temperatures across payload
- Toggles on/off RF chain



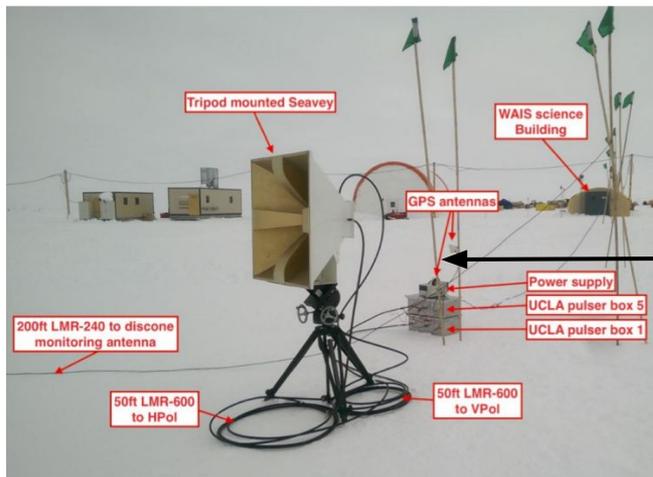
temp sensor

Work by Brian Rauch, Keith McBride, and Zack Martin

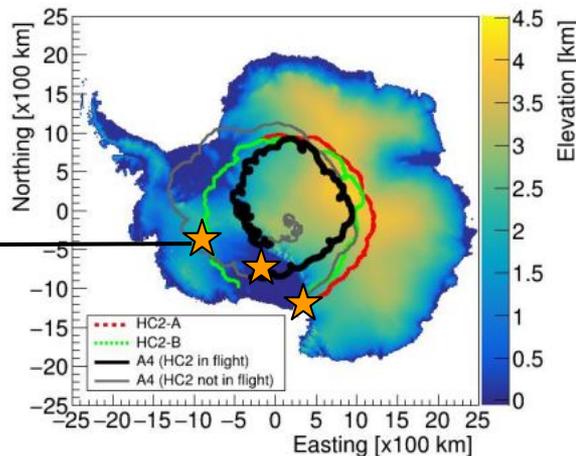
Design & Hardware of PUEO



## Ground Calibration

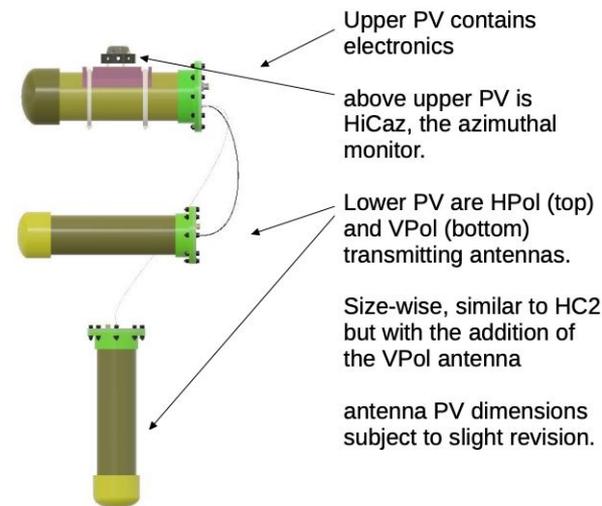


Field stations along PUEO path that pulse as payload flies overhead



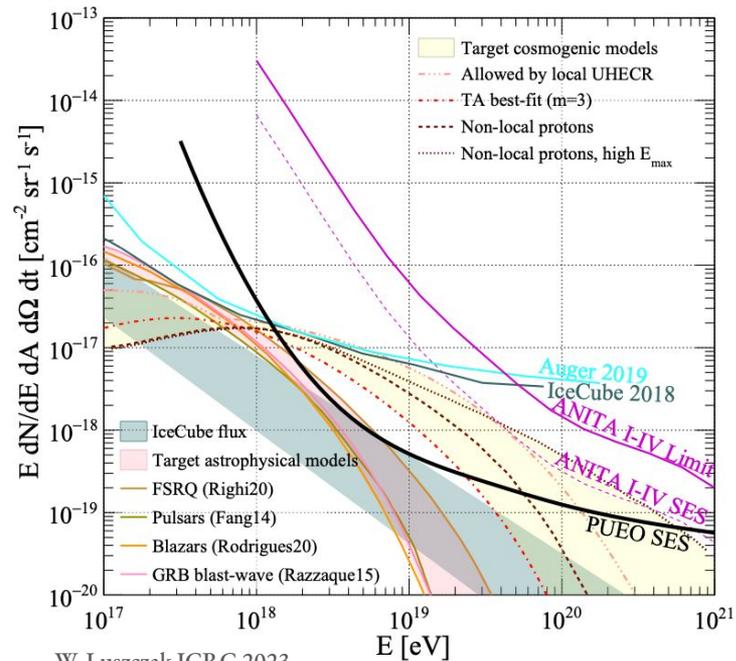
Work by Steven Prohira, Steph Wissel, Julien Alfaro, Shoukat Ali, Dave Besson, Kenny Couberly, Rob Young, & Andrew Zeolla

## High-altitude Calibration



Small hand-launched balloons that follow main payload

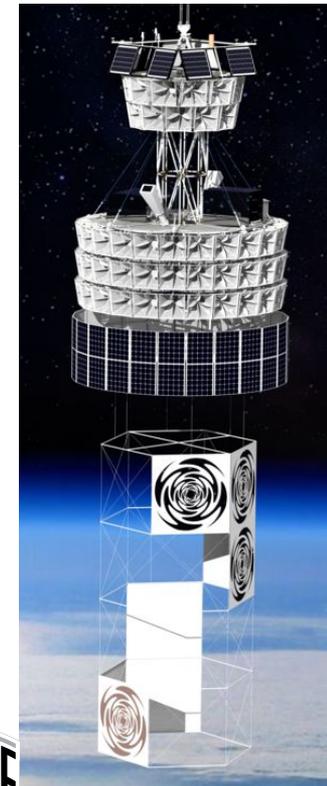
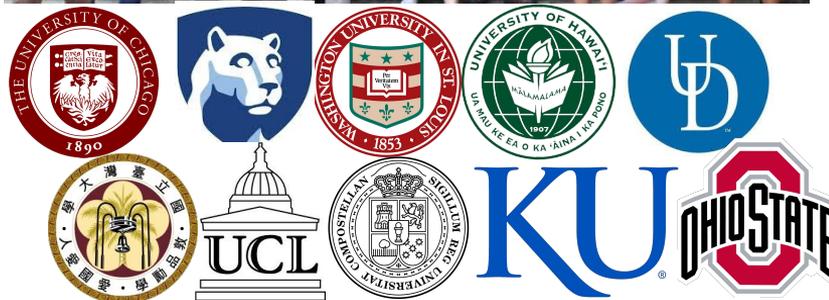
# Outlook



W. Luszczyk ICRC 2023

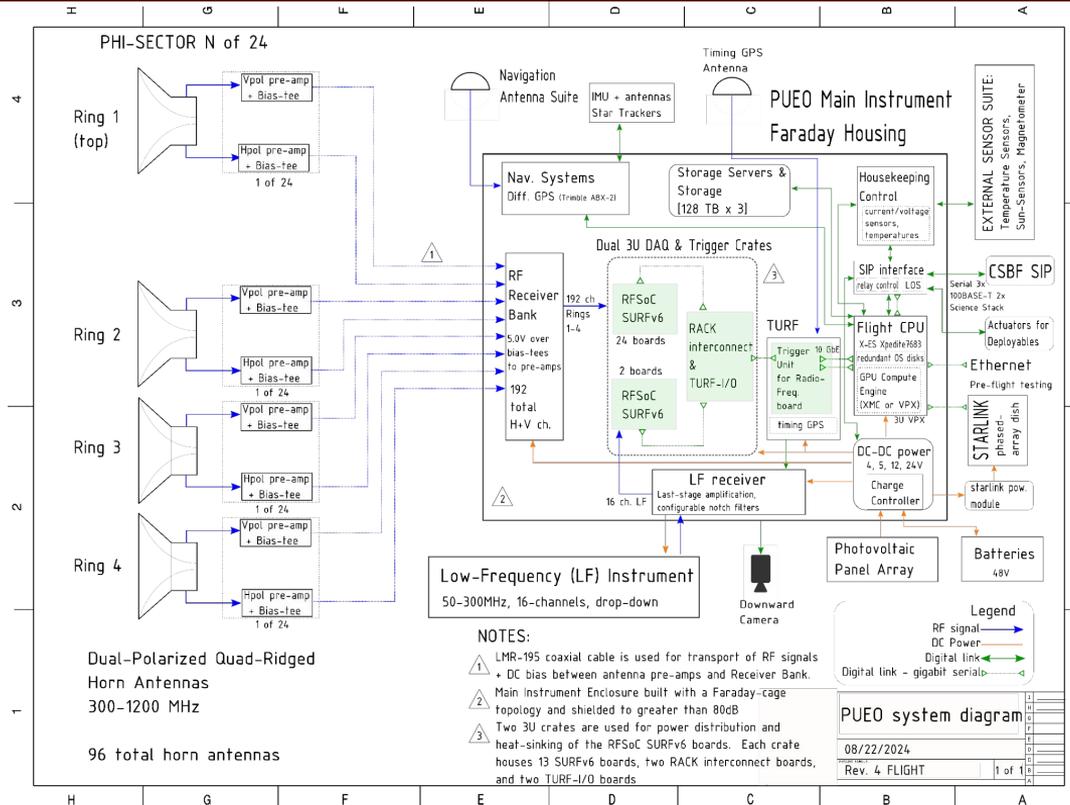
We gratefully acknowledge NASA Awards #80NSSC20K0775 and #80NSSC20K0925.

R. Scrandis

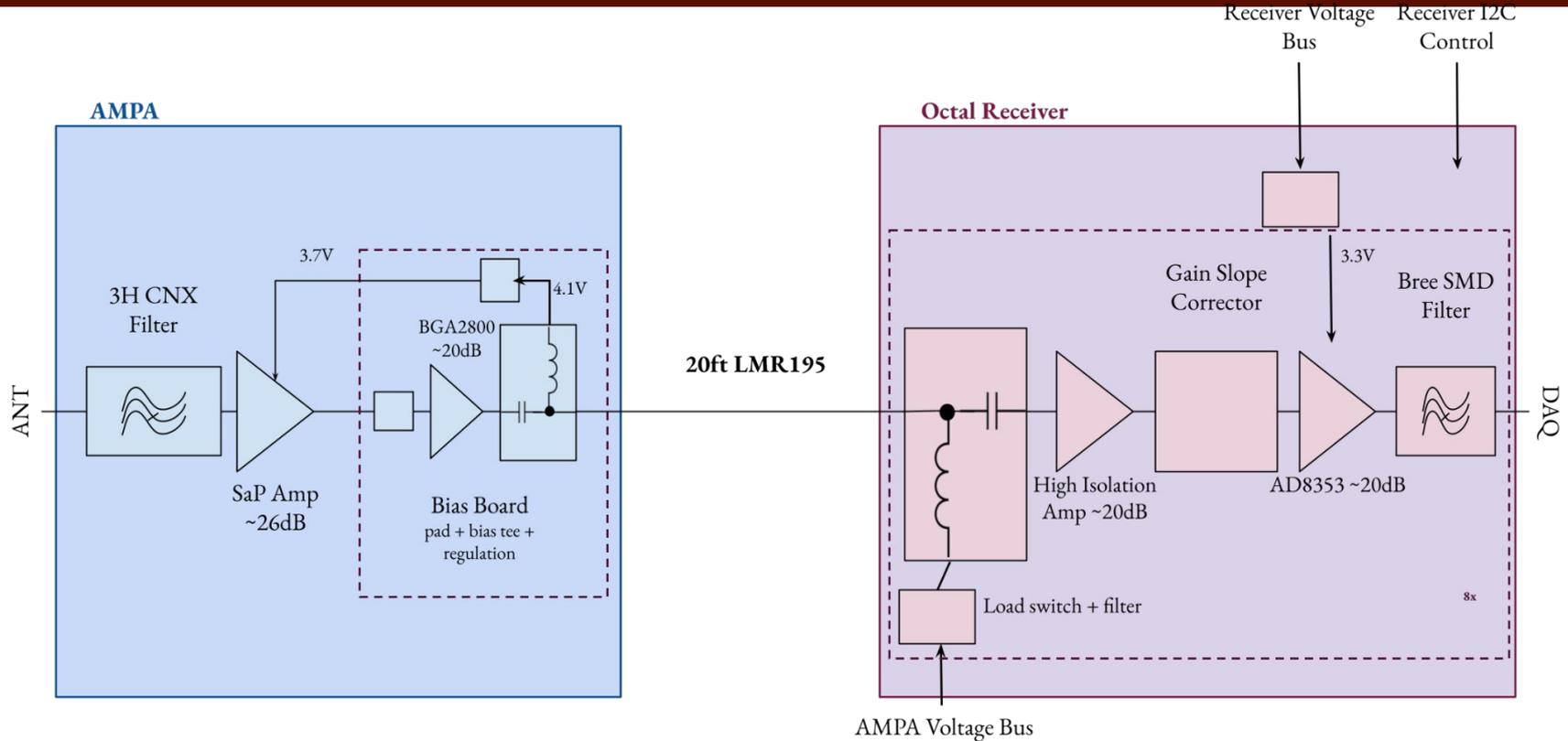


# Extra Slides

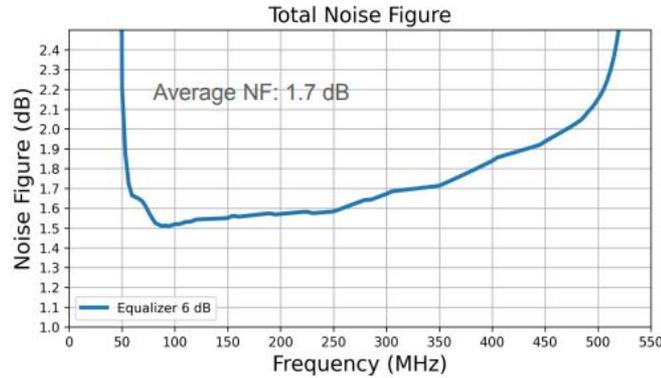
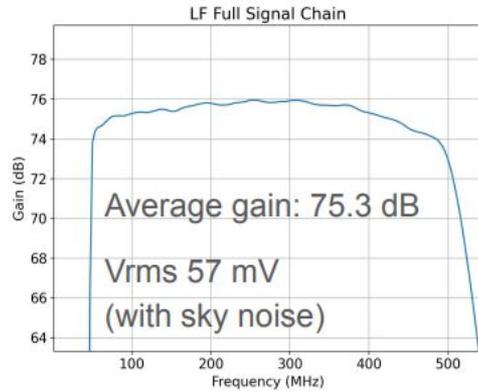
# Full System Diagram



# RF Chain Block Diagram

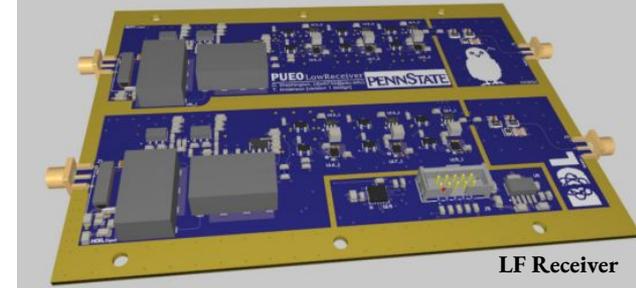
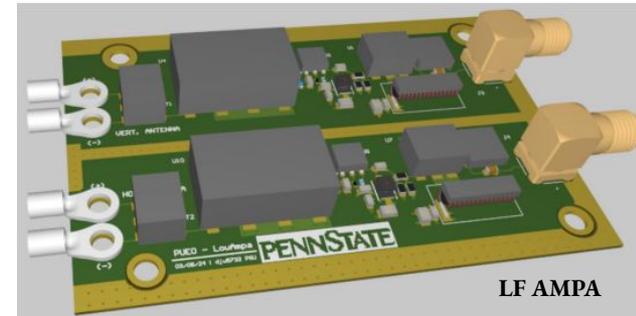


# LF RF Chain



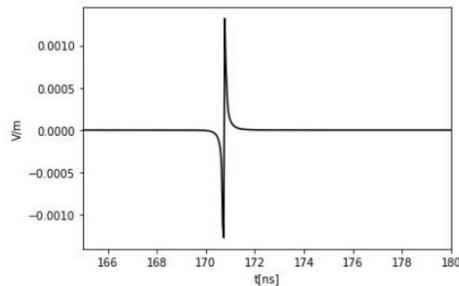
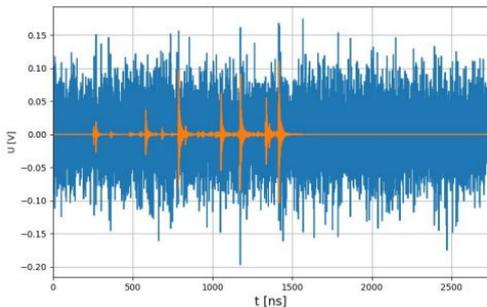
LF signal conditioning happens in two stages

- LF AMPA mounted directly behind antenna gives ~20dB
- LF Receiver gives ~55dB

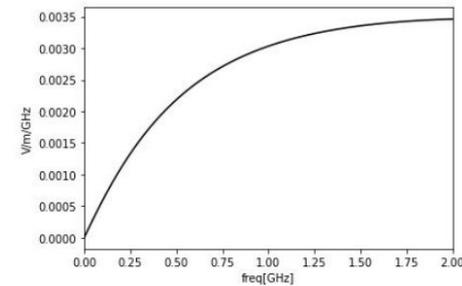


Work by Daniel Washington

# Flavor Reconstruction

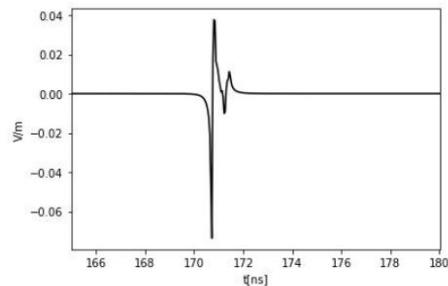
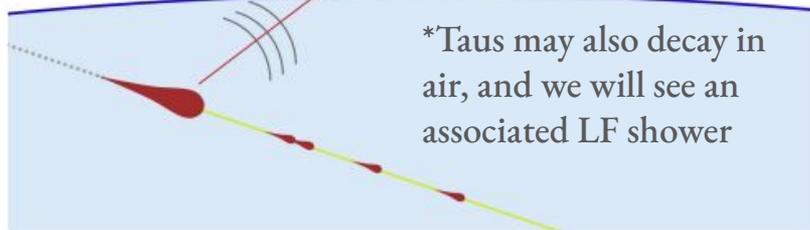


Hadronic shower

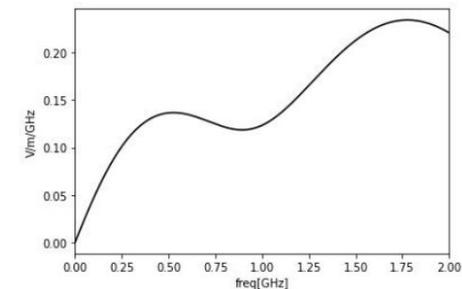


C. Welling ARENA 2024

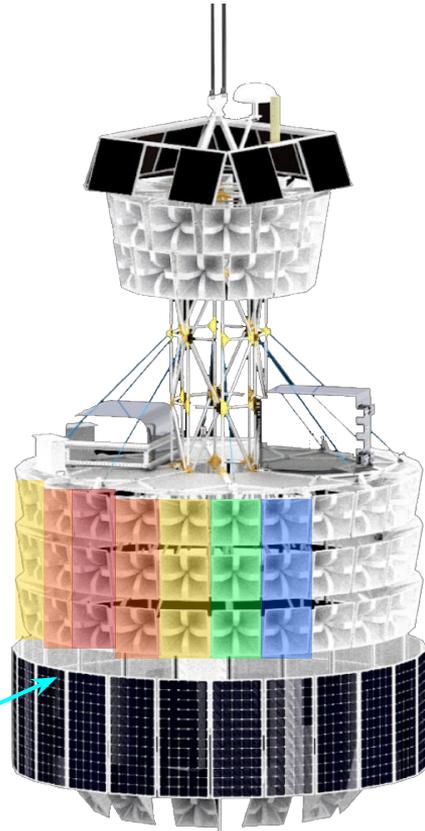
Work done by Christoph Welling,  
Austin Cummings, and me



Hadronic + EM shower

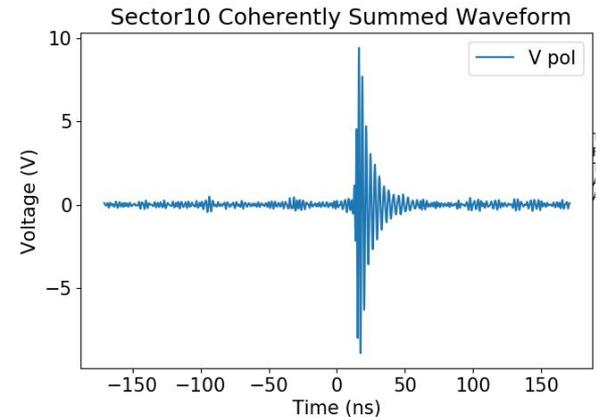
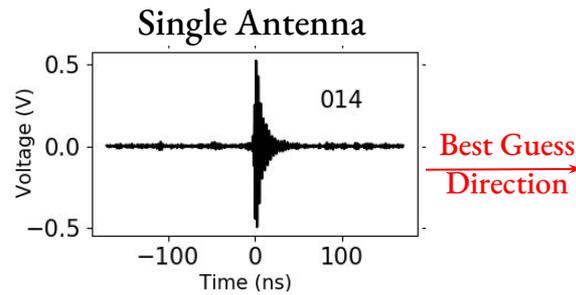


# Phased Array Triggering



On board, preset arrival directions will be checked by applying relevant time delays and summing each channel in a sector

- If any of the beams (arrival directions) passes some voltage threshold, event will pass the trigger
- This beam forming + coherent summing (summing each channel) will help reject thermal noise triggers (SNR goes as  $\sqrt{N(\text{ant})}$ )





# DAQ Hardware Specs



SURF: Trenz TE0835 with a Xilinx ZU47DR RFSoc

- FPGA Package: E1156 (35 x 35 mm)
- Engine: Quad-core Arm Cortex-A53 MPCore up to 1.3GHz
- 8 RFADCs with 4.096 GSPS

TURFIO: Xilinx Artix-7 FPGA

TURF: Xilinx Zynq Ultrascale+ MPSoC

# Nav Suite Block Diagram

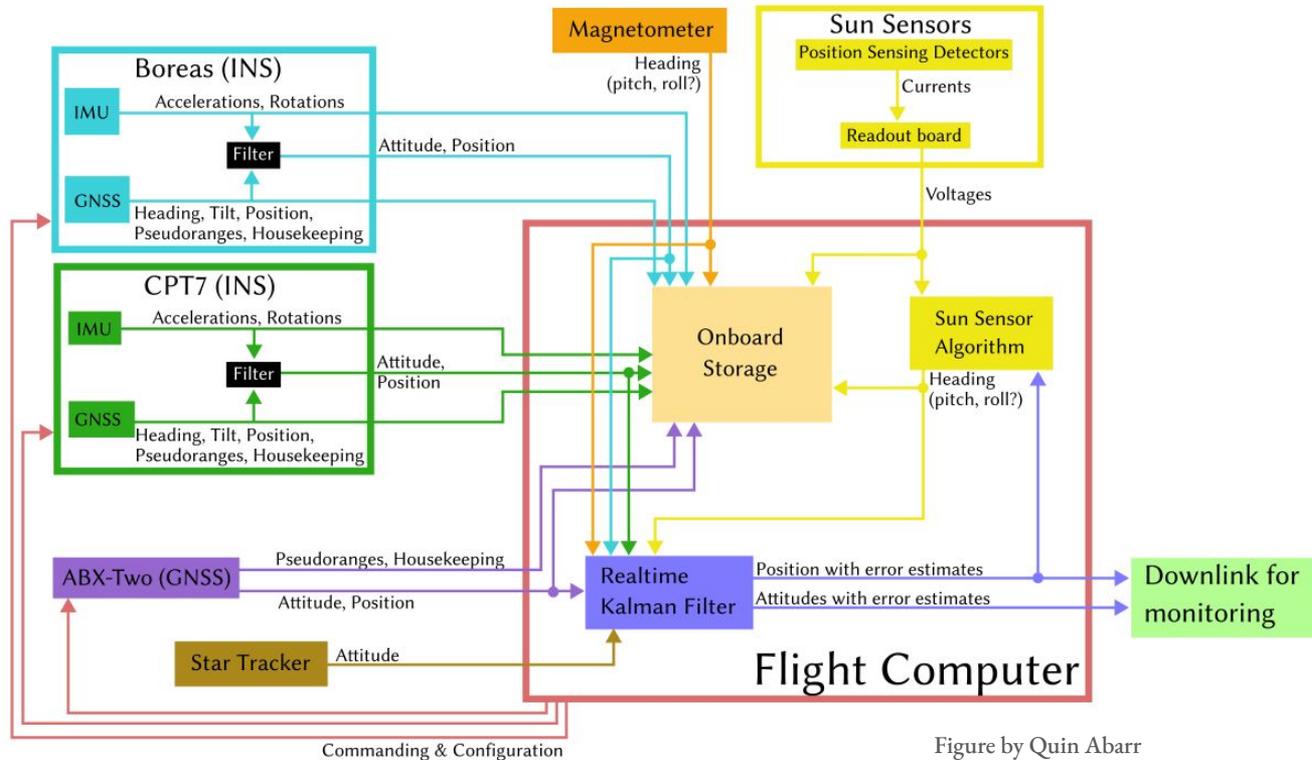


Figure by Quin Abarr

# Power Block Diagram

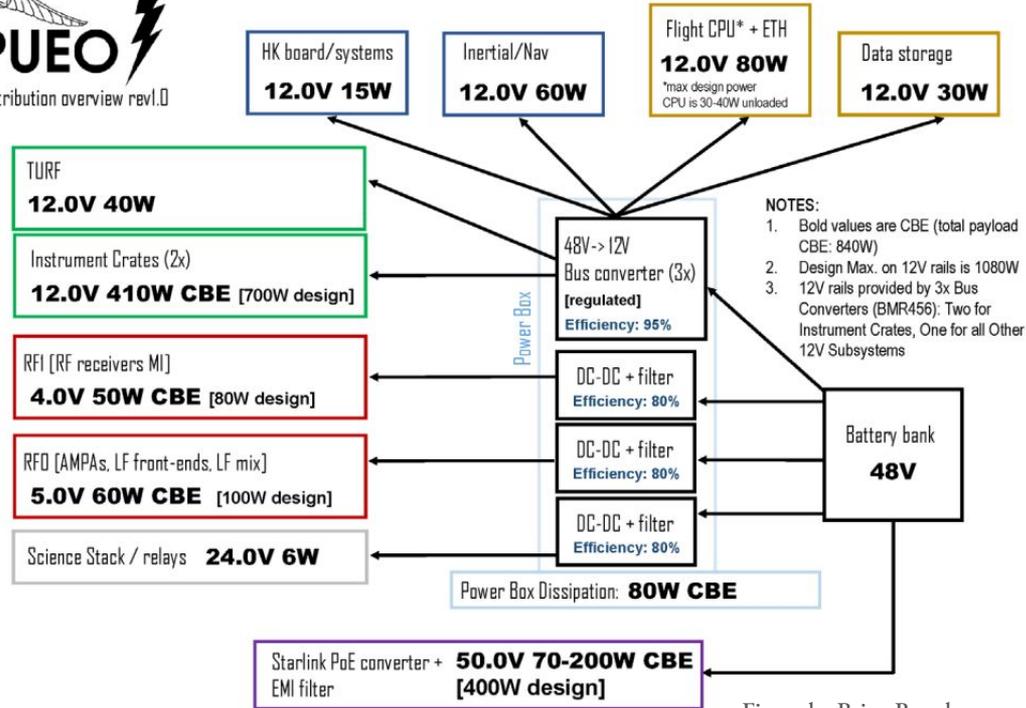


Figure by Brian Rauch