Space-Based Gamma-ray Observations for the Next Decade

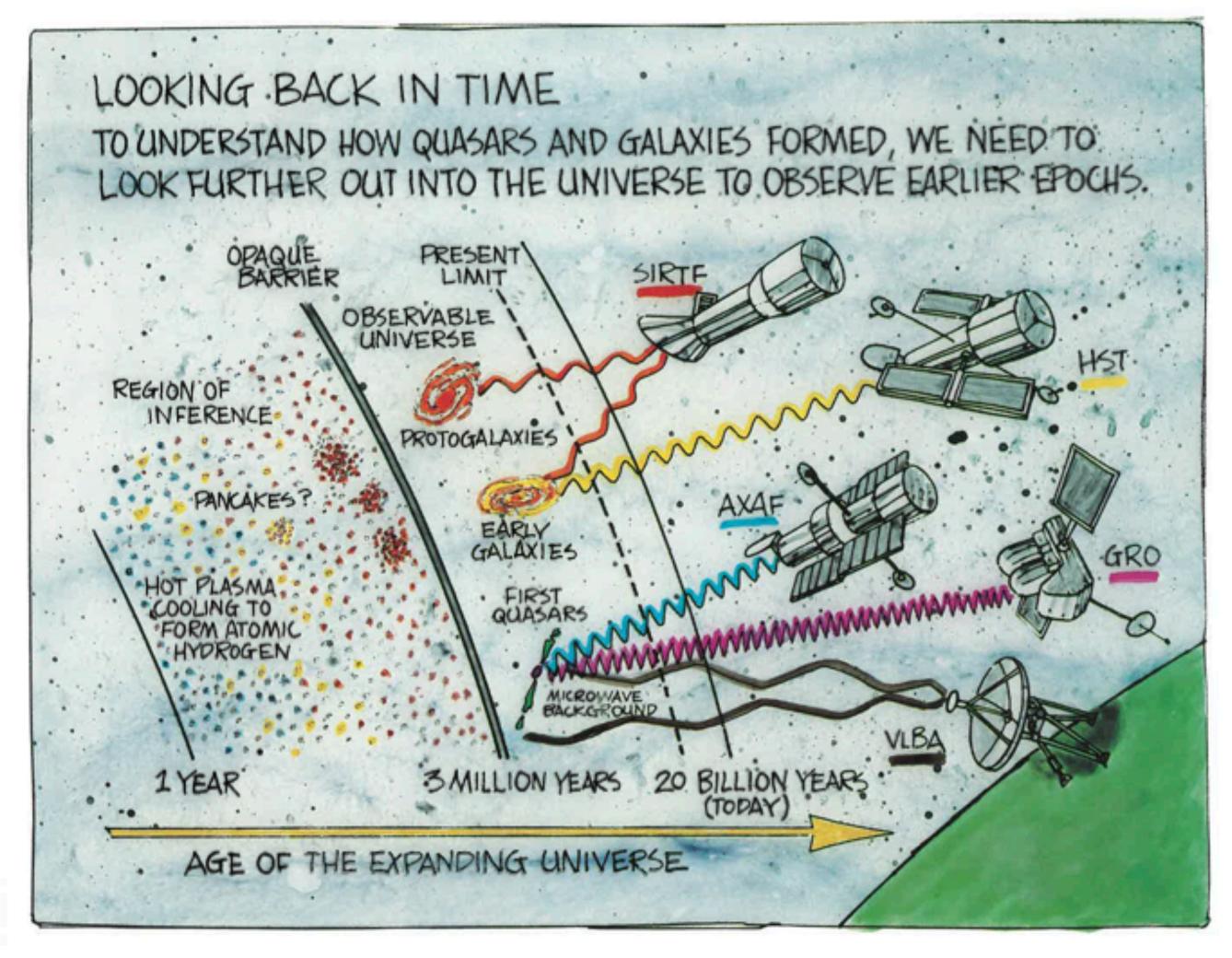
Where we've been and where we're going

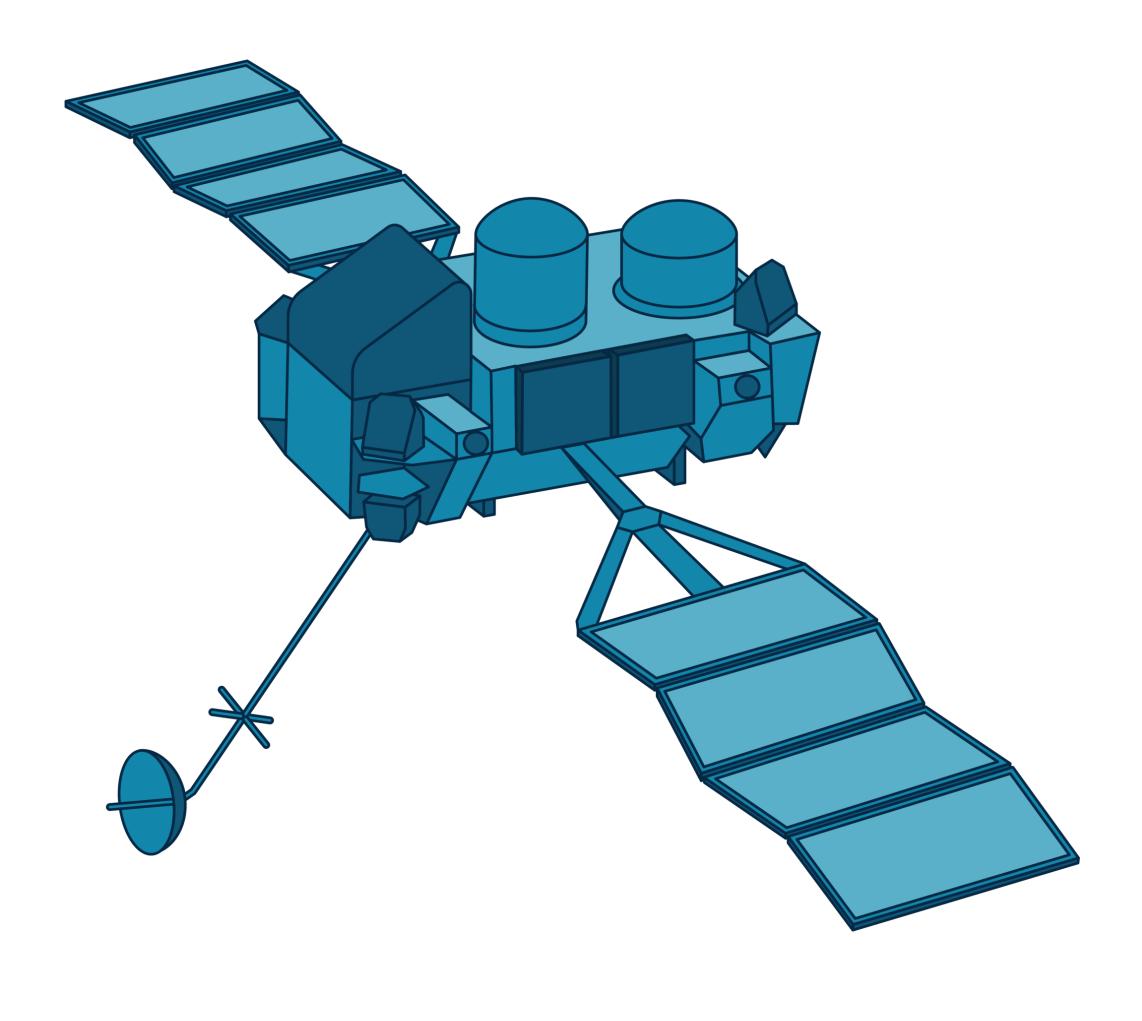
R. Caputo, NASA GSFC

TeV Particle Astrophysics Chicago, IL August 26, 2024

Where have we been?*

Great Observatories Program

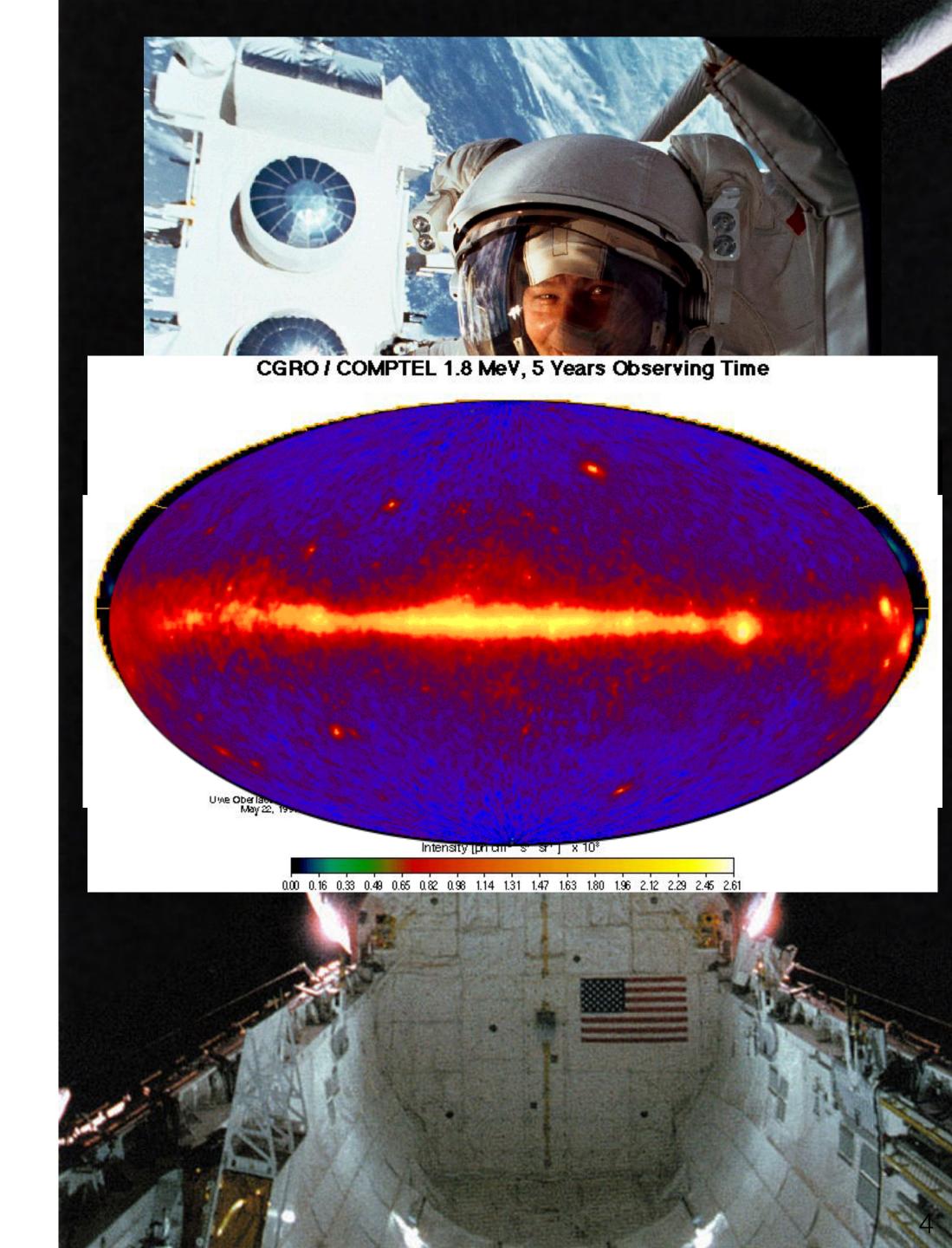




Launching from 1990 to 2003

Compton Gamma-ray Observatory

- One of the original Four Great Observatories. Launched 1991 and de-orbited in 2000
- Four Instruments:
 - The Burst Alert and Transient Source Experiment (BATSE) an all sky monitor 20 keV to 1 MeV
 - The Oriented Scintillation Spectrometer Experiment (OSSE) for the 0.05 to 10 MeV range
 - The Compton Telescope (CompTel) in the 0.8 to 30 MeV range capable of imaging 1 steradian.
 - The Energetic Gamma-Ray Experiment Telescope (EGRET) in the 30 MeV to 10 GeV range.

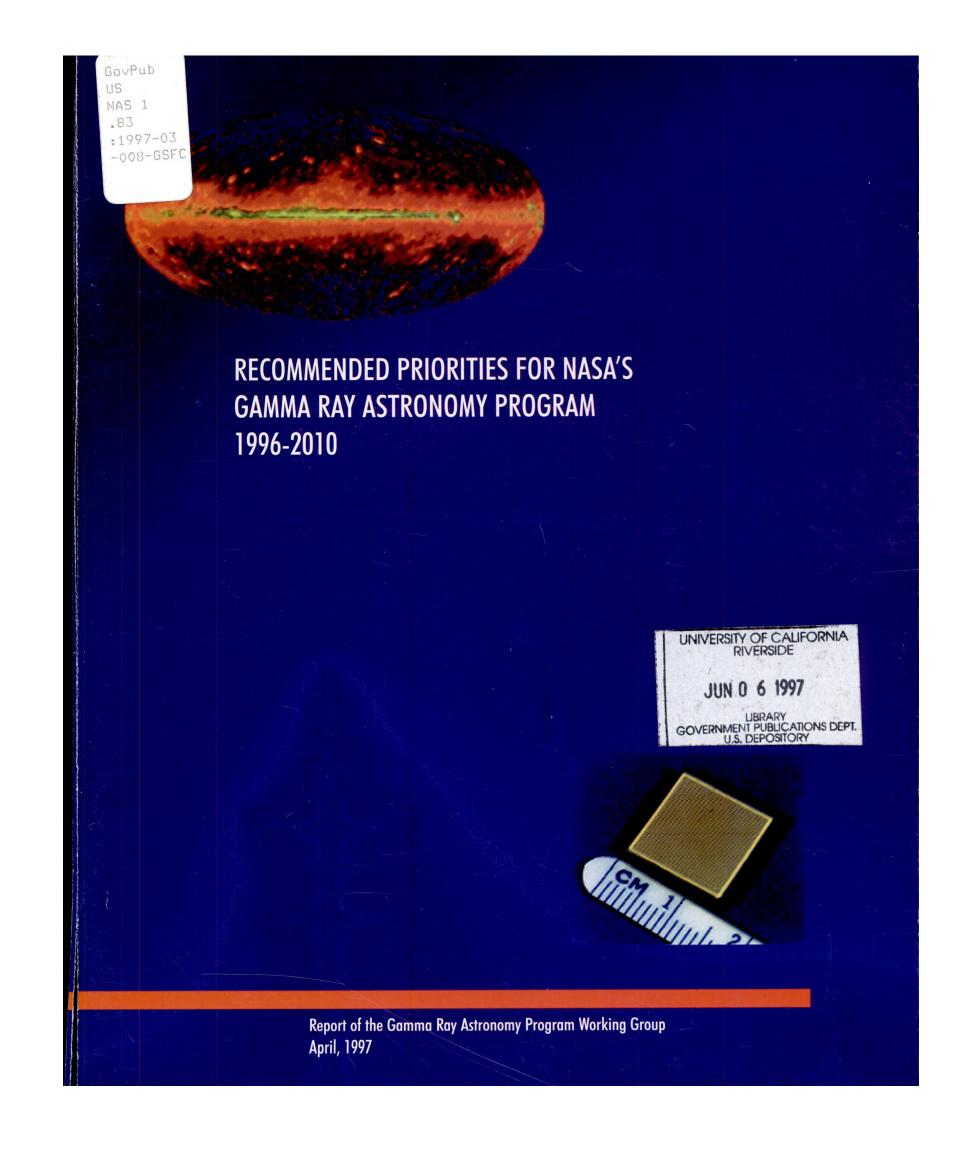


Inspired by the success of CGRO:



Bob Hartman and Dave Thompson working on **EGRET** in 1993

"The mandate of the working group is to recommend a road map to the future for use as an input to the next NASA strategic plan..."





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GAMMA-RAY ASTRONOMY PROGRAM WORKING GROUP MEMBERS:

Elena Aprile (Columbia)

Alan Bunner (NASA) [Ex-Officio (NASA Headquarters)]

Neil Gehrels (GSFC) [Co-Chair]

Jonathan Grindlay (Harvard)

Gerald Fishman (MSFC)

W. Neil Johnson (NRL)

Kevin Hurley (UCB/SSL)

Steve Kahn (Columbia)

Richard Lingenfelter (UCSD)

Peter Michelson (Stanford)

Thomas Prince (Caltech) [Co-Chair]

Roger Romani (Stanford)

James Ryan (UNH)

Bonnard Teegarden (GSFC)

David Thompson (GSFC)

Trevor Weekes (Harvard/Smithsonian)

Stanford Woosley (UCSC)



Intermediate Missions

The HIGHEST PRIORITY recommendation is:

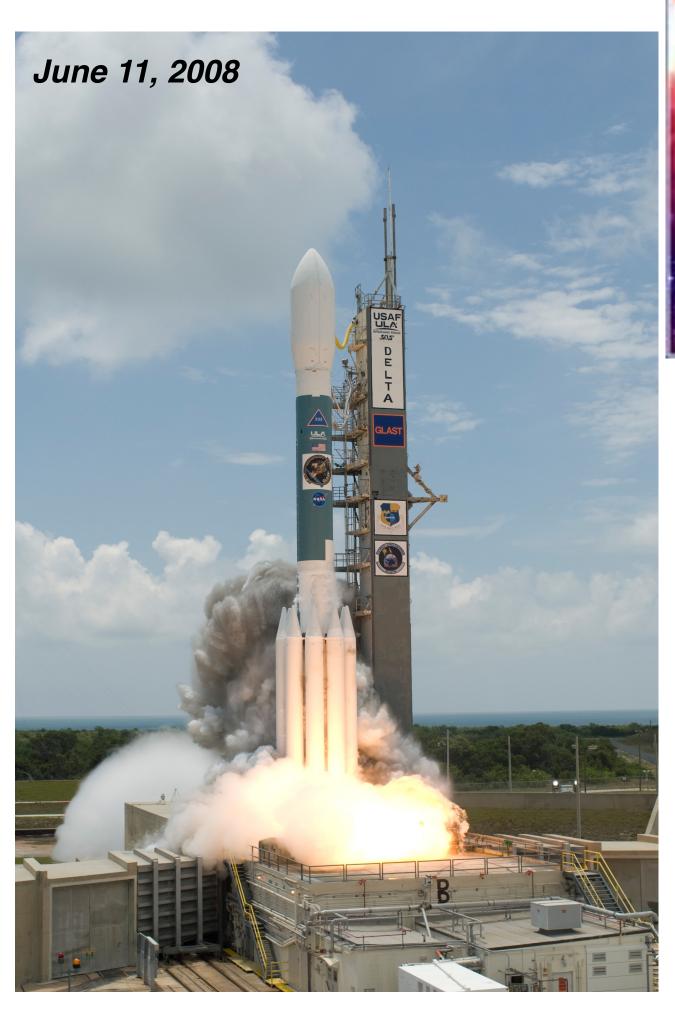
A next generation 10 MeV to 100 GeV gamma-ray mission such as GLAST.

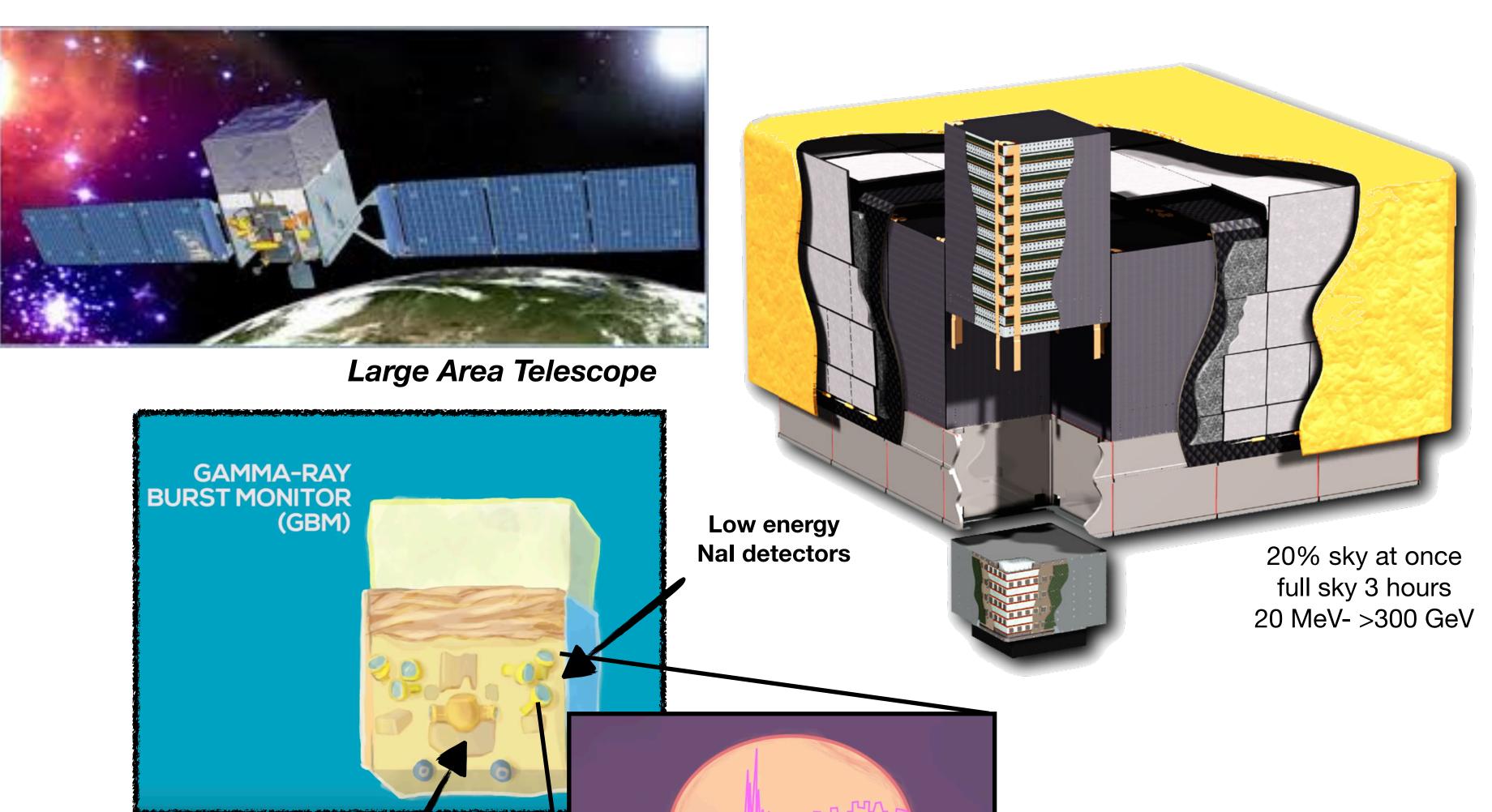
1 to 2 orders of mag improvement in sensitivity compared to EGRET.

Fermi Gamma-ray Space Telescope

High energy

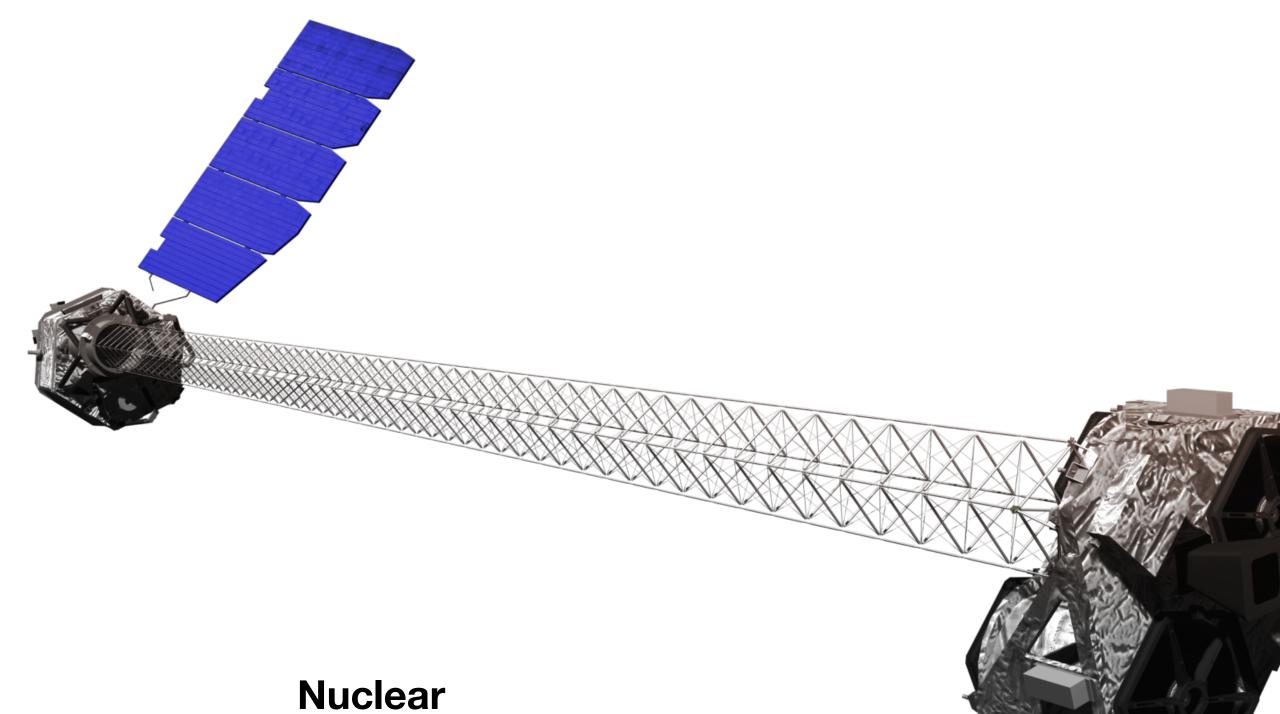
BGO detectors





full unocculted sky continuous 8 keV - 40 MeV

Intermediate Missions

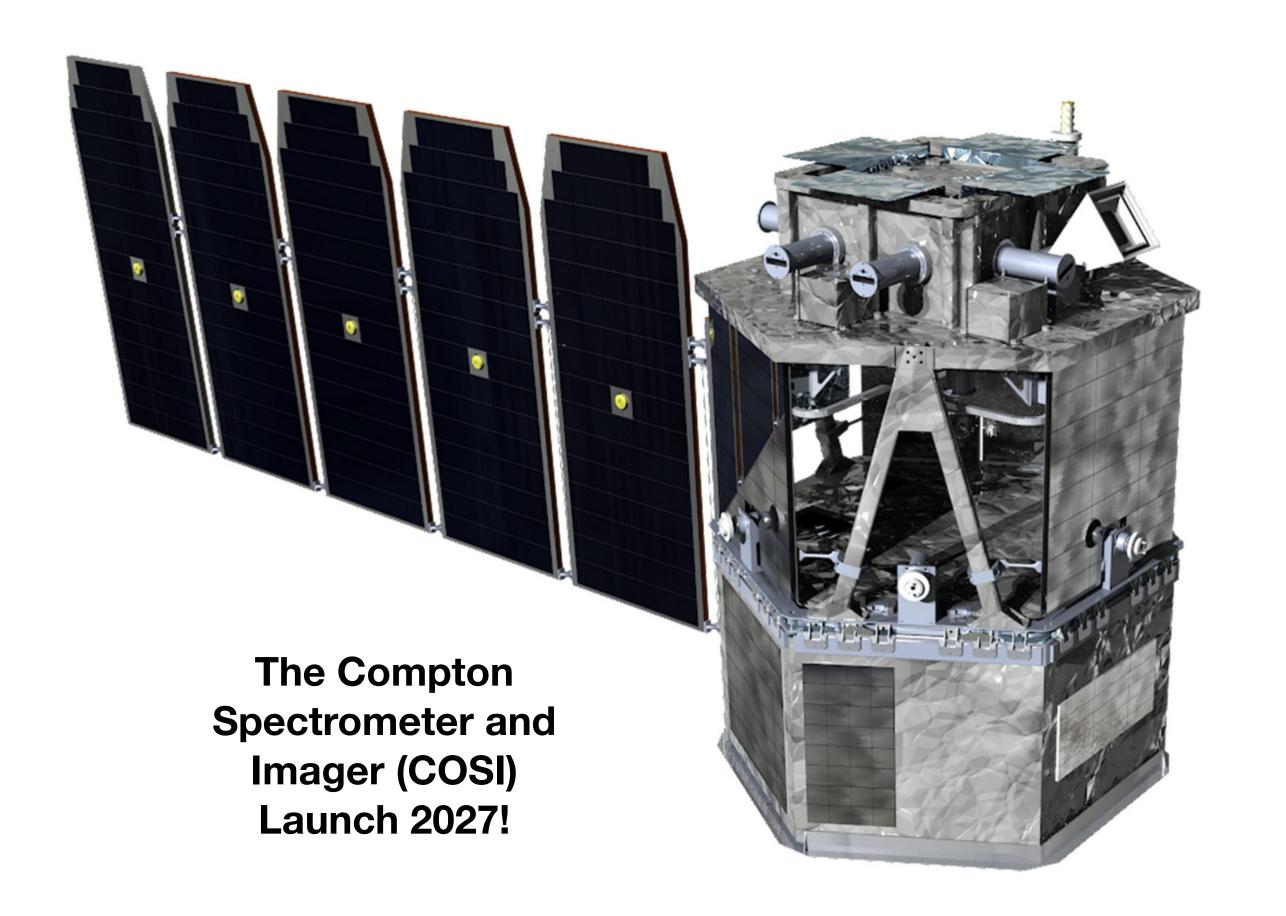


Another very-high priority:

A Focusing Hard X-ray Telescope.

Nuclear
Spectroscopic
Telescope Array
(NuSTAR)
launched 2012

Intermediate Missions



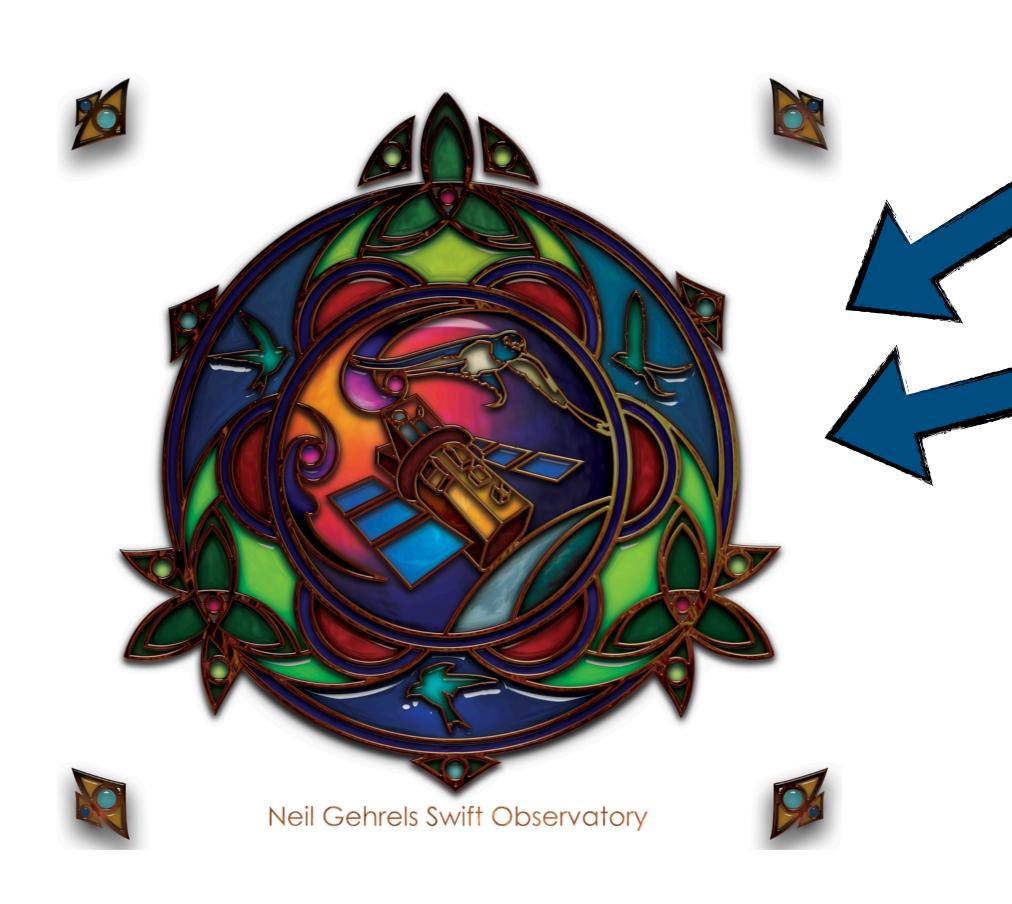
The second very-high priority:

A next-generation nuclear line and MeV continuum mission. A major step forward compared to INTEGRAL in both sensitivity and energy range.

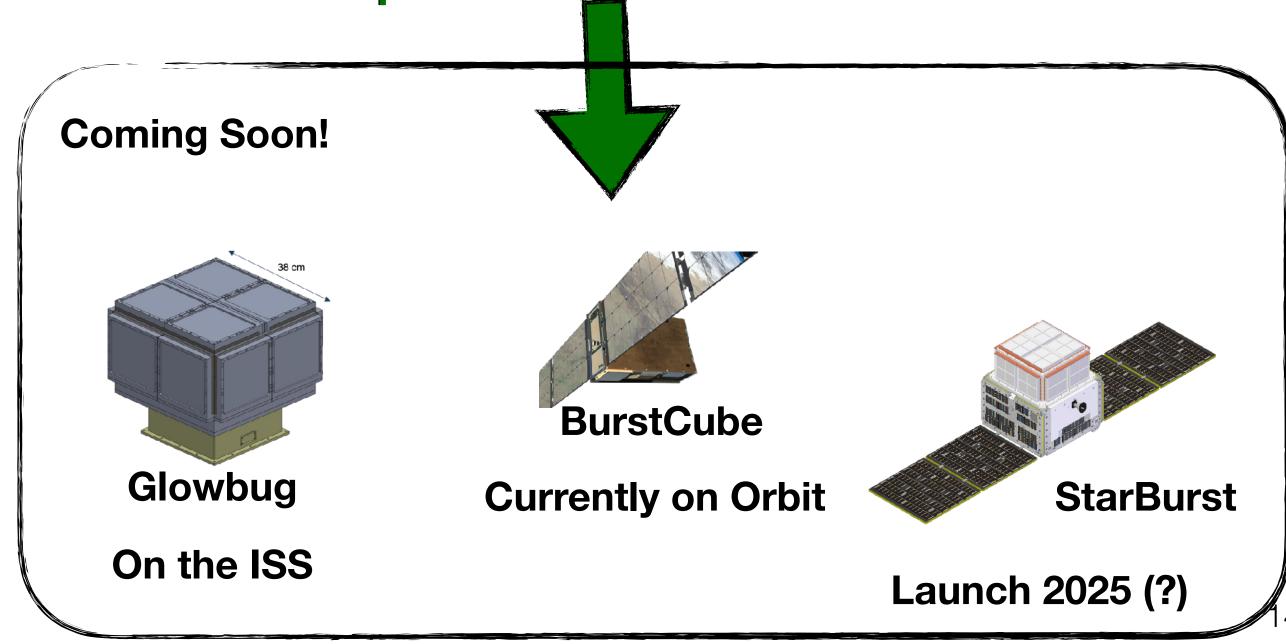
More info: https://science.nasa.gov/mission/cosi/

Participate in the COSI 2nd data challenge: https://github.com/cositools/cosi-data-challenge-2

MIDEX and SMEX Missions

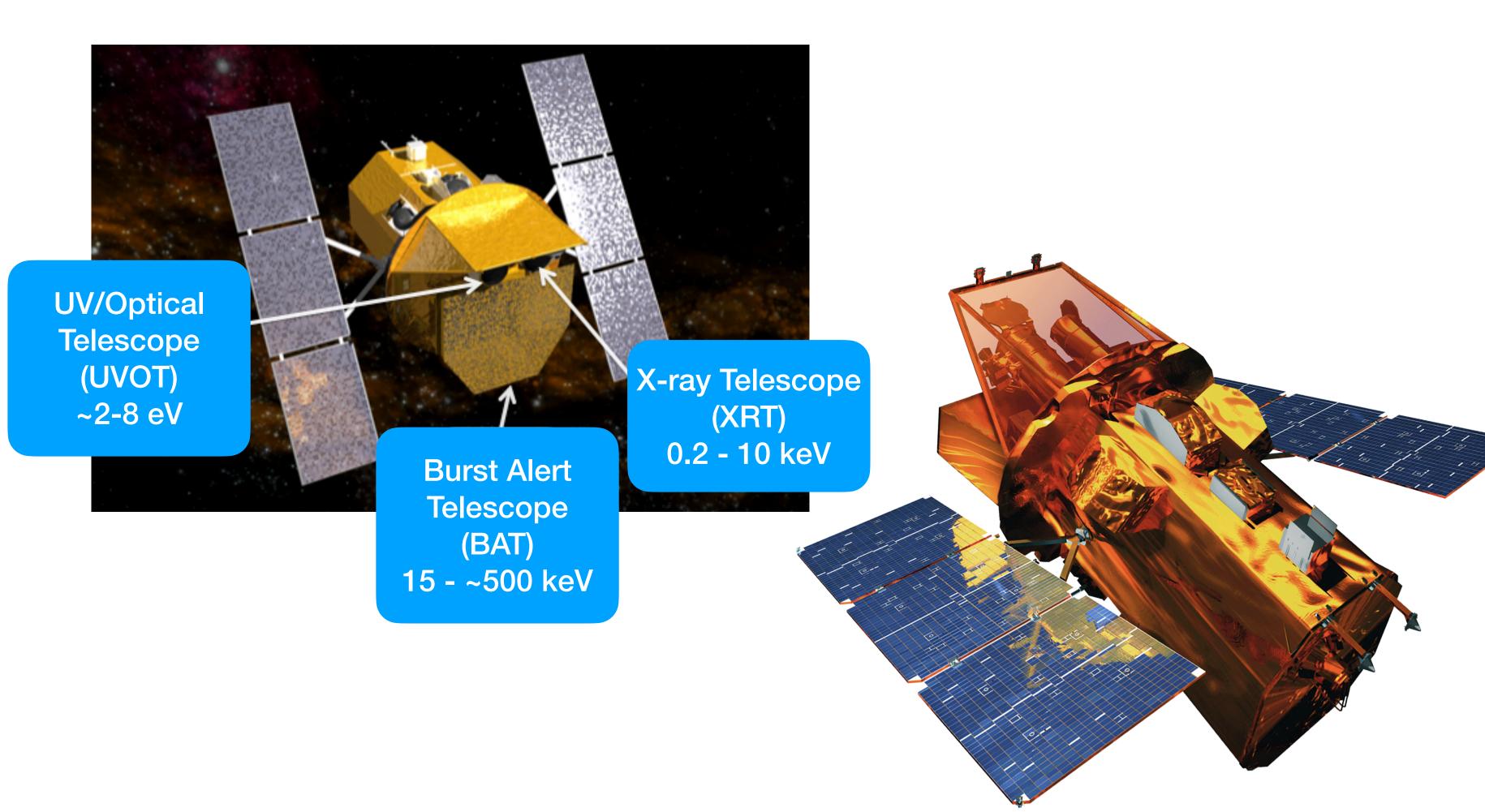


A gamma-ray burst localization mission. Such a mission would address the origin of gamma-ray bursts. Missions with coding apertures or an array of small telescopes would fill this need.



Neil Gehrels Swift Observatory





KEY QUESTIONS IN GAMMA-RAY ASTRONOMY FROM 1997

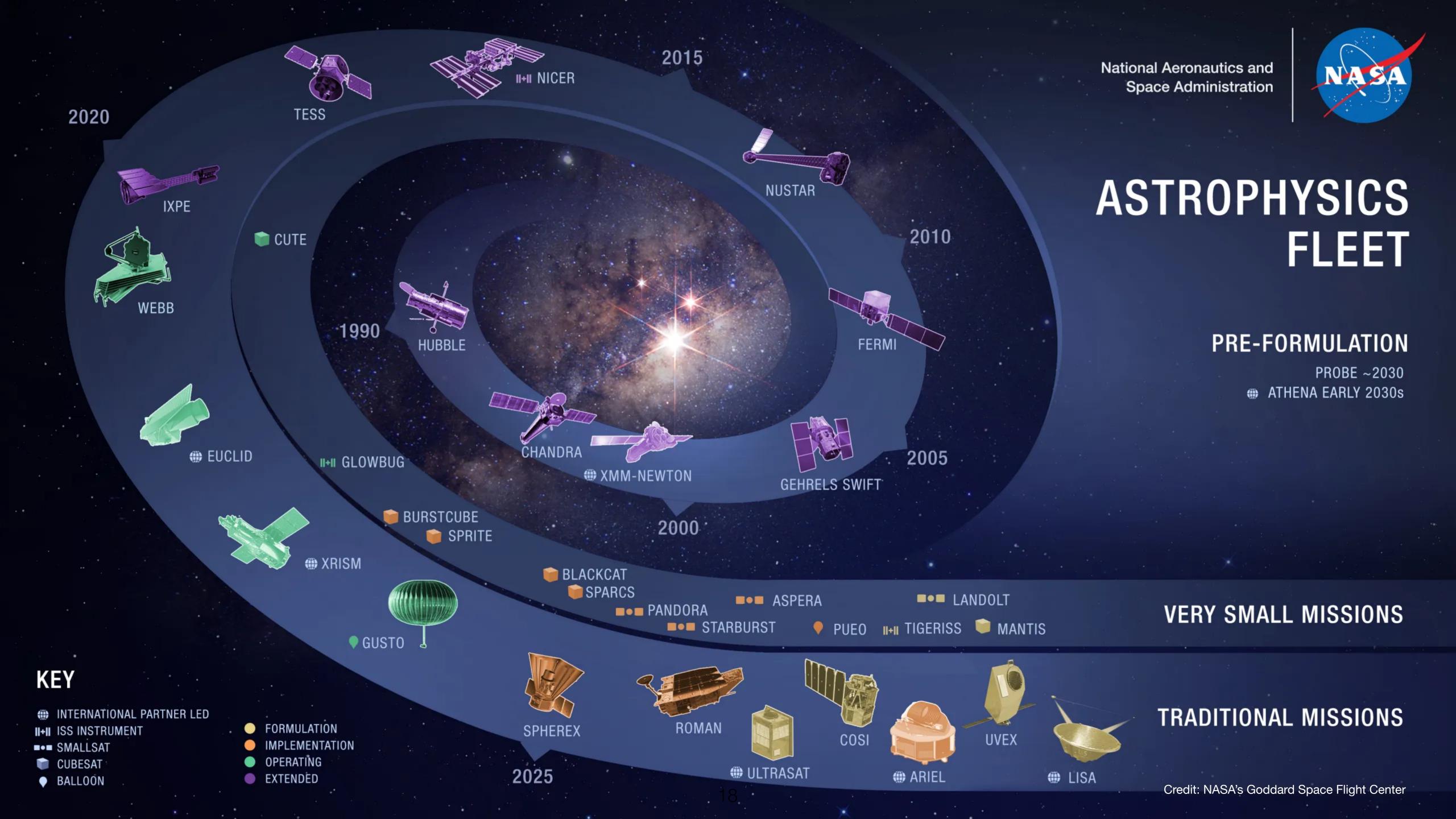
- What is the origin and nature of gamma-ray bursts?
- What are the physical conditions and processes near accreting black holes and neutron stars?
- How does matter behave in extreme conditions like those in neutron stars, supernova expulsions and active galactic nuclei?
- How do astrophysical accretion processes work and what are their instabilities, periodicities and modes?
- What is the nature of the jets emanating from galactic black holes and AGN and how are the particles accelerated?
- What is the origin of the diffuse gamma-ray background?
- What is the nature of the unidentified high energy gamma-ray sources?
- What are the sites of nucleosynthesis?
- How do supernovae work? What are the progenitors and explosion mechanisms? What has bene the rate in the last several hundred years?
- What and where are the sites of cosmic ray acceleration?

Why did they recommend these missions?

- They developed a series of Key Science
 Questions that pointed to the need for this
 diverse set of missions.
 - Lesson: Lead with the Science
 - Lesson: Don't shy away from the big problems
 - Lesson: Make strong/bold recommendations
- Many of these questions are still open but we have made significant progress.

'97 Report Checklist

- ✓ Intermediate Missions: Fermi, NuSTAR and now COSI
- ✓ MIDEX and SMEX: Swift and NICER (EXIST in the report)
- √ Technology: a robust technology development program (SiPMs, new scintillators, upgraded silicon detectors, etc)
- ✓ Balloons (+ CubeSats!): long duration balloons enabled COSI, LEAP, etc.
- ✓ Data Analysis & Theory: mainly supported through GI programs
- √ TeV Astronomy: VERITAS, HESS, HAWC, and MAGIC.



Takeaways

- The Gamma-ray sky has been observed by large observatories since 1991 via CGRO and then followed with INTEGRAL, Swift, Fermi, and AGILE
- Gamma-ray observations have enabled huge discoveries over the past ~2 decades and most recently as we have entered the era of multi messenger astrophysics
 - Progress on all fronts of the checklist!
- The next generation of discoveries in astrophysics need all-sky gammaray observatories complement CTA, and GW and neutrino observatories

Where are we going?

"The test of a first-rate intelligence is the ability to hold two opposed ideas in mind at the same time and still retain the ability to function." - F. Scott Fitzgerald



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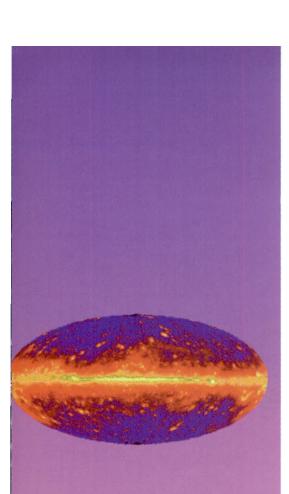
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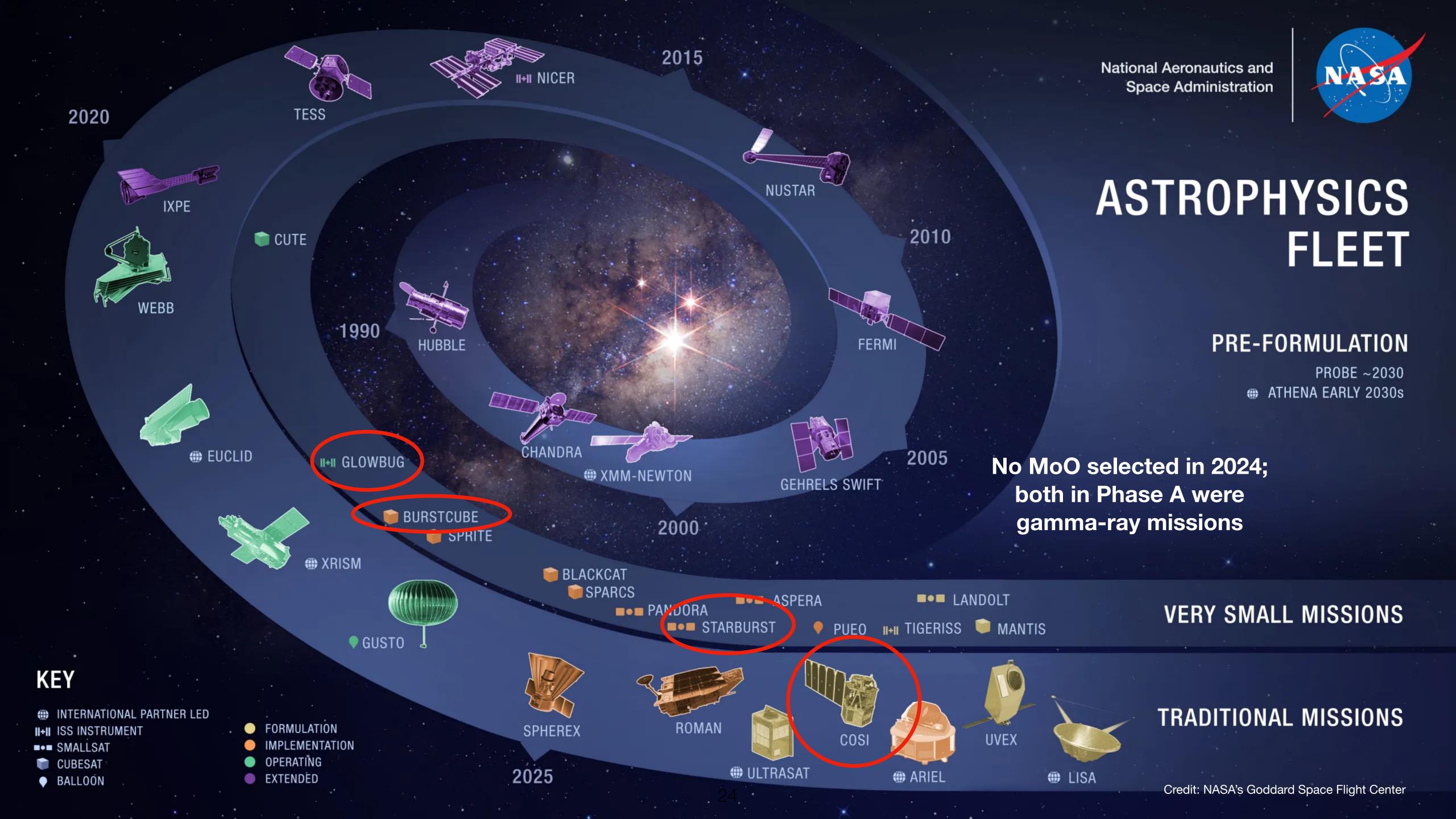
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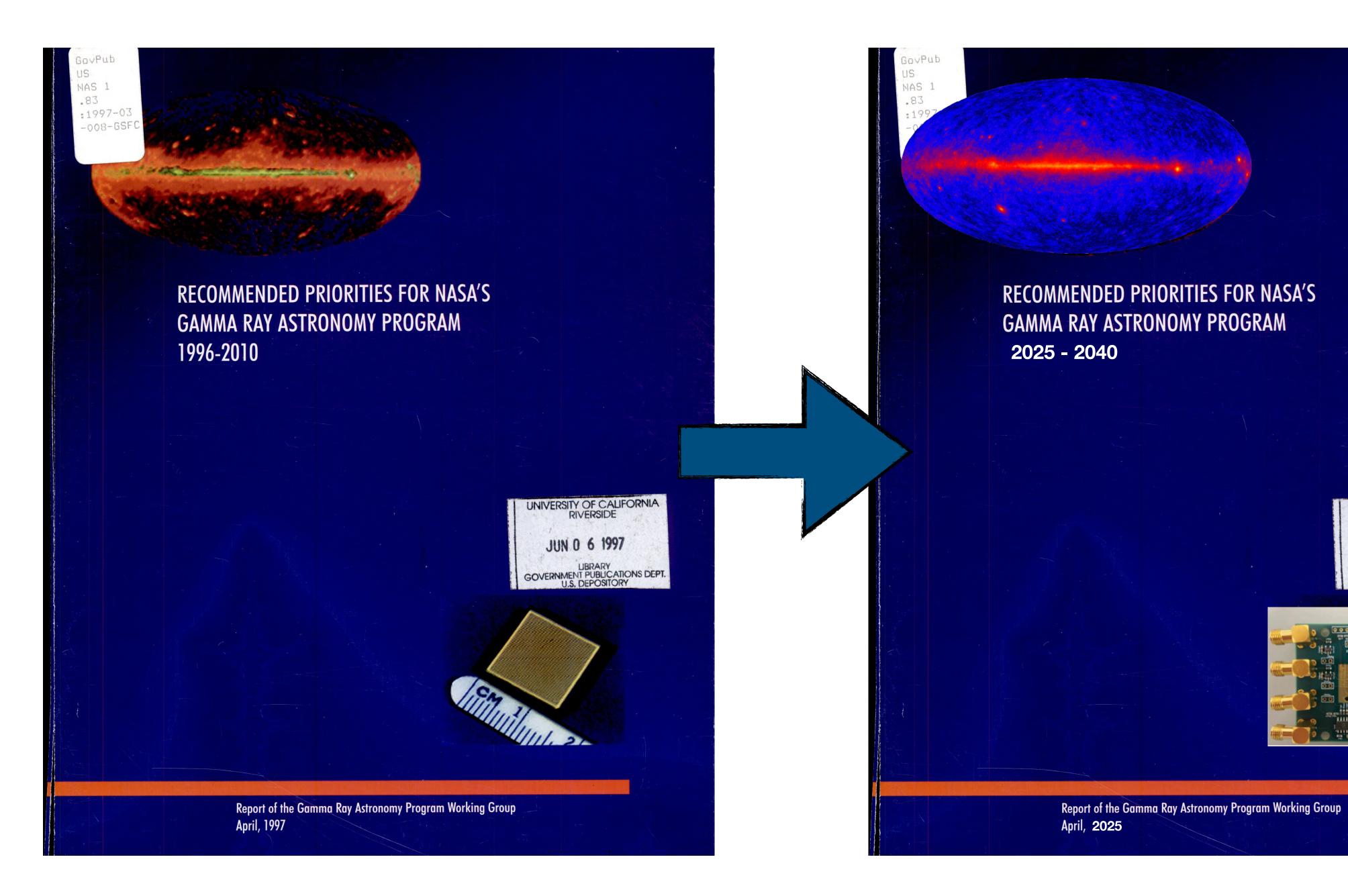
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How can we replicate this success?



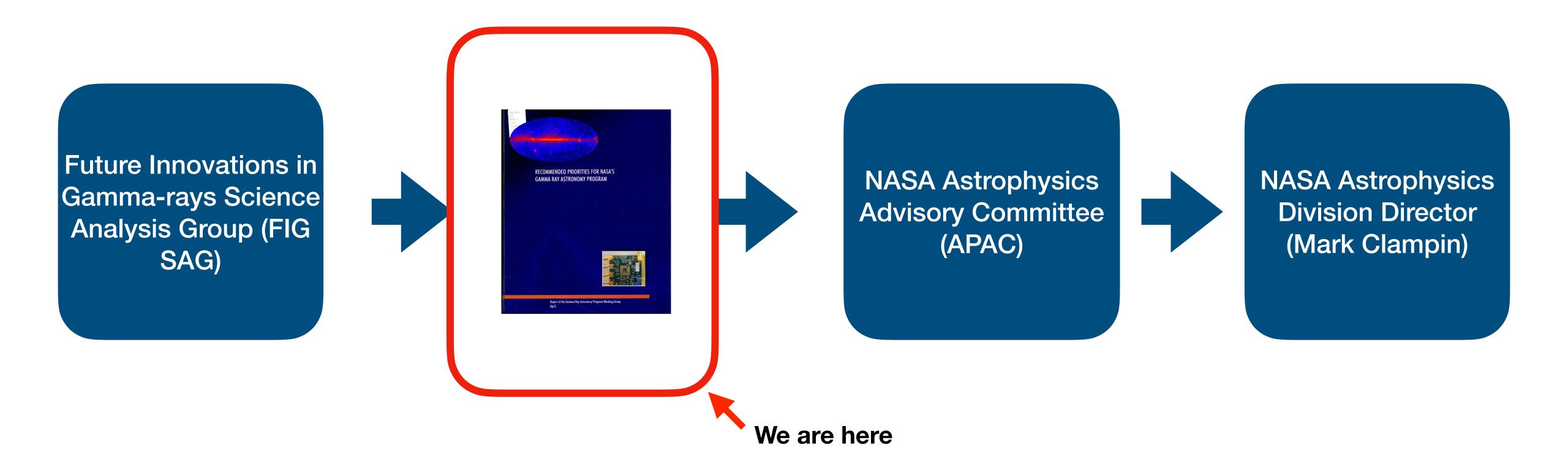
UNIVERSITY OF CALIFORNIA RIVERSIDE

JUN 0 6 2024

GOVERNMENT PUBLICATIONS DEPT. U.S. DEPOSITORY

Help develop the Roadmap

 Need the world wide gamma-ray, high-energy and multimessenger communities to contribute/provide input



Future Innovations in Gamma-ray Science Analysis Group (FIG SAG)

Astrophysical gamma rays span ten orders of magnitude in energy and capture key physics from a broad range of astrophysical phenomena. This SAG will explore gamma-ray science priorities, necessary capabilities, new technologies, and theory/modeling needs drawing on the 2020 Decadal to inspire work toward 2040.

To get involved and stay informed, please enter your contact information here: https://forms.gle/VBijBgapMRwJm9dU6



Lead Chairs: Chris Fryer & Michelle Hui

Co-chairs: Paolo Coppi, Milena Crnogorčević, Tiffany Lewis, Marcos Santander, and Zorawar Wadiasingh

Gamma-SIG: https://pcos.gsfc.nasa.gov/sigs/grsig.php
FIG SAG: https://pcos.gsfc.nasa.gov/sags/figsag.php

The FIG SAG report for 2025

- Of course, the '97 roadmap led to new questions as well. Three key ones that should be included:
 - The report recommended an MeV all-sky mission but that did not materialize <— we can emphasize that this is still missing from the portfolio
 - Multimessenger Astronomy is (of course) not mentioned.
 - This report directly led to the advent of MMA (Fermi and Swift)
 - Inclusion, Diversity, and Equity are not mentioned.

Fermi/Swift capabilities are an Astro2020 Decadal priority

Sustaining Programs (Space)

Time-Domain Program (highest priority)

 A program of competed missions and of opportunity to realize of capabiliti ph

Not the d

eve

New Physics Decadal Priority!

Probe Lin

Comp to brid strategi Ju gaps in science ___ capabilities— this decade Farand wav IR and an X-ray complement to Athena

• \$1.5 billion/mission, cadence of approx. one/decade

Astro2020: New Messengers, broader set of stems, and placing stments and those nd S.6). In space, senger program of robe missions to be

ment the European Space paoilities not possible at the Explorer scale. compared to a large strategic mission, and a cost cap of

oals. For the

mission per decade is realistic. The selection of a probe mission in replace the need for a future large, strategic mission. For ground-based projects, the priority sustaining activity is a significant augmentation and expansion of mid-scale programs, including the addition of strategic calls to support key survey priorities. The survey also

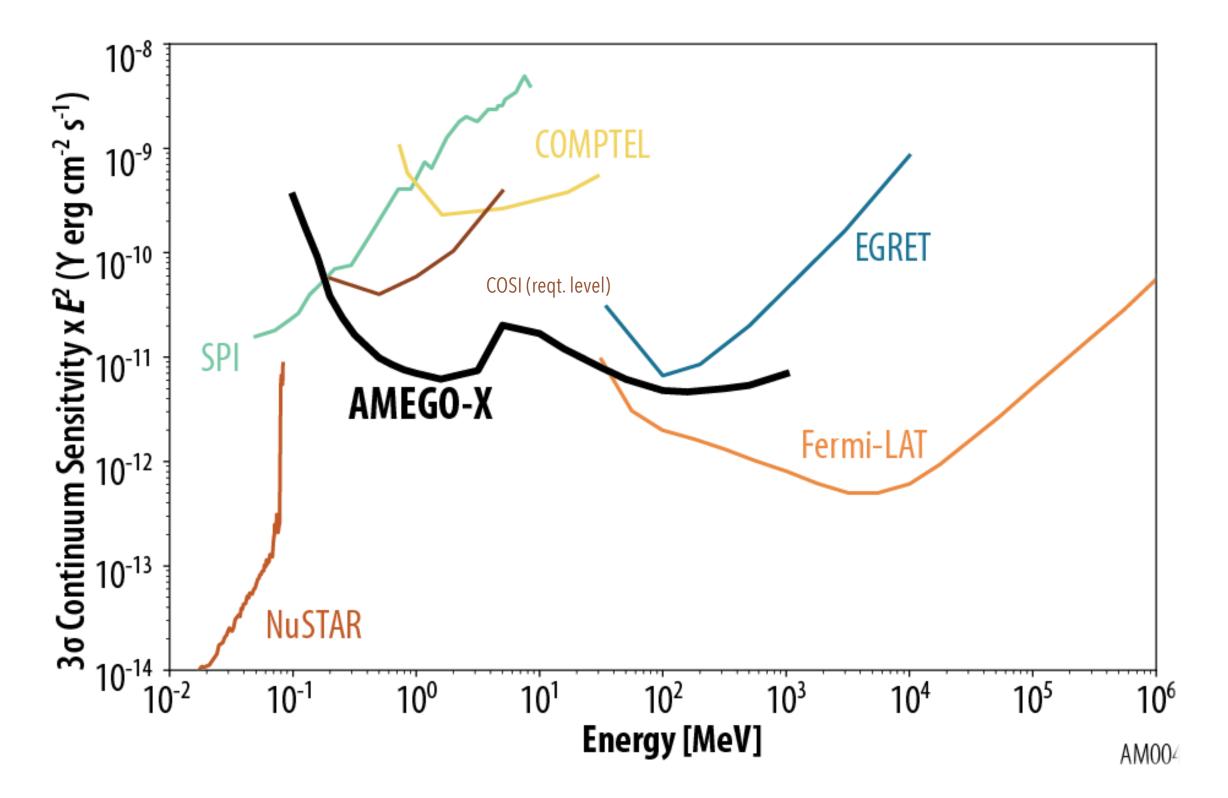
strongly endorses investments in technology development for advanced gravitational wave interferometers, both to upgrade NSF's Laser Interferometer Gravitational-Wave Observatory (LIGO),

and to prepare for the next large facility.5

30 M. Capulo | INASA GSFC, 16VFA 2024

A Telescope for the MeV Gamma-ray Regime

All-sky Medium Energy Gamma-ray Observatory eXplorer: AMEGO-X



AMEGO-X: arXiv:2208.04990 AstroPix: 2302.00101

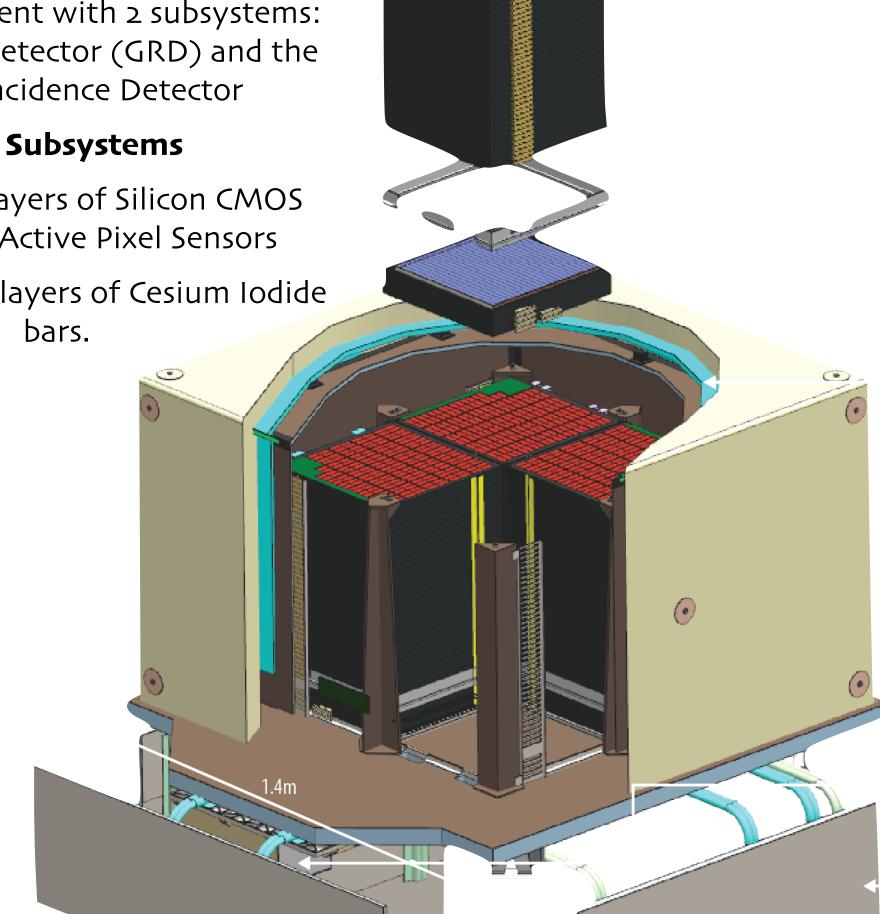
AMEGO-X

Single instrument with 2 subsystems: Gamma-Ray Detector (GRD) and the Anti-Coincidence Detector

GRD Subsystems

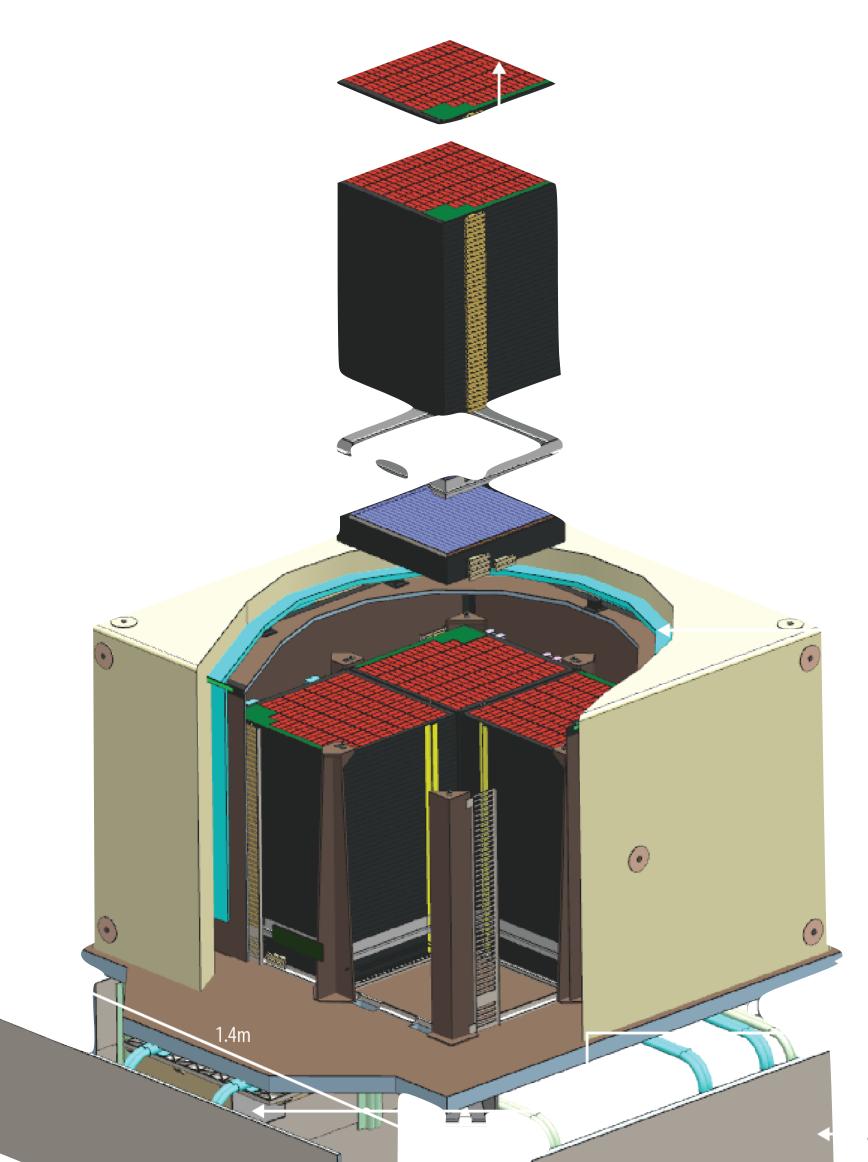
Tracker: 40 layers of Silicon CMOS monolithic Active Pixel Sensors

Calorimeter: 4 layers of Cesium Iodide



AMEGO-X: Status and Plans

- Resubmit in the next MIDEX round (~2027)
 - New folks welcome! (Email me)
 - Important for the science team to keep publishing on the need for MeV instrumentation (we're happy to share sensitivity, effective area, energy/angular resolution etc...)
- Participate in gamma-ray roadmap activities



Friends observing space





- Support Fermi/Swift! Senior review preparation NOW
 - Fermi Symposium in DC area (September 9-13)
 - FIGSAG meeting on Sept 13 (afternoon)
 - Swift@20 in Rome Spring 2025
- Join the Gamma-ray Science Interest Group (GammaSIG)
 - https://pcos.gsfc.nasa.gov/sigs/grsig.php
- Advocate with your own funding agency to support FIG SAG (Gamma-ray Roadmap) and a new Gamma-ray Observatory(s).
- Thanks to the TeVPA organizers!