

TeVPA 2024
University of Chicago, 29 August 2024

NANOGrAV AND GRAVITATIONAL WAVES FROM THE EARLY UNIVERSE

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University of Texas at Austin

on behalf of the NANOGrav collaboration



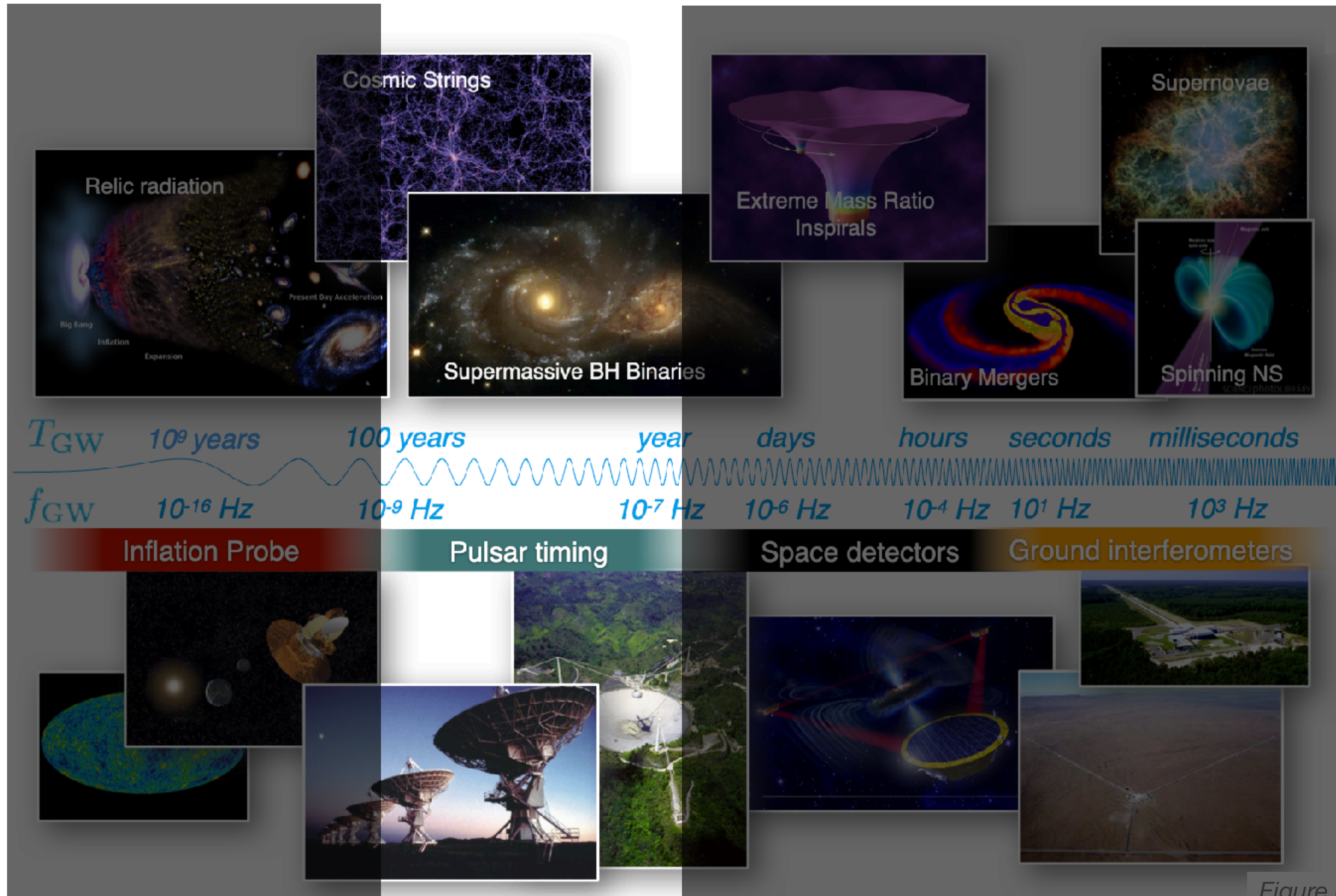


Figure credit: K. Chatziioannou

Worldwide Pulsar Timing Array (PTA) Experiments



Figure credit: T. Cromartie

Multiple pulsar timing array experiments reported evidence for background of nHz gravitational waves

NANOGrav

- ✦ [SGWB search](#) (2306.16213)
- ✦ [Observation & Timing](#) (2306.16217)
- ✦ [Detector & Noise](#) (2306.16218)
- ✦ [New physics](#) (2306.16219)
- ✦ [SMBHB](#) (2306.16220)
- ✦ [Anisotropy](#) (2306.16221)
- ✦ [Continuous waves](#) (2306.16222)
- ✦ [Pipeline](#) (2306.16223)

EPTA / InPTA

- ✦ [SGWB search](#) (2306.16214)
- ✦ [Data & Timing](#) (2306.16224)
- ✦ [Noise](#) (2306.16225)
- ✦ [Continuous waves](#) (2306.16226)
- ✦ [Signal sources](#) (2306.16227)
- ✦ [ULDM](#) (2306.16228)

PPTA

- ✦ [SGWB search](#) (2306.16215)
- ✦ [Noise](#) (2306.16229)
- ✦ [Data](#) (2306.16230)

CPTA

- ✦ [SGWB search](#) (2306.16216)



Animation by NSF

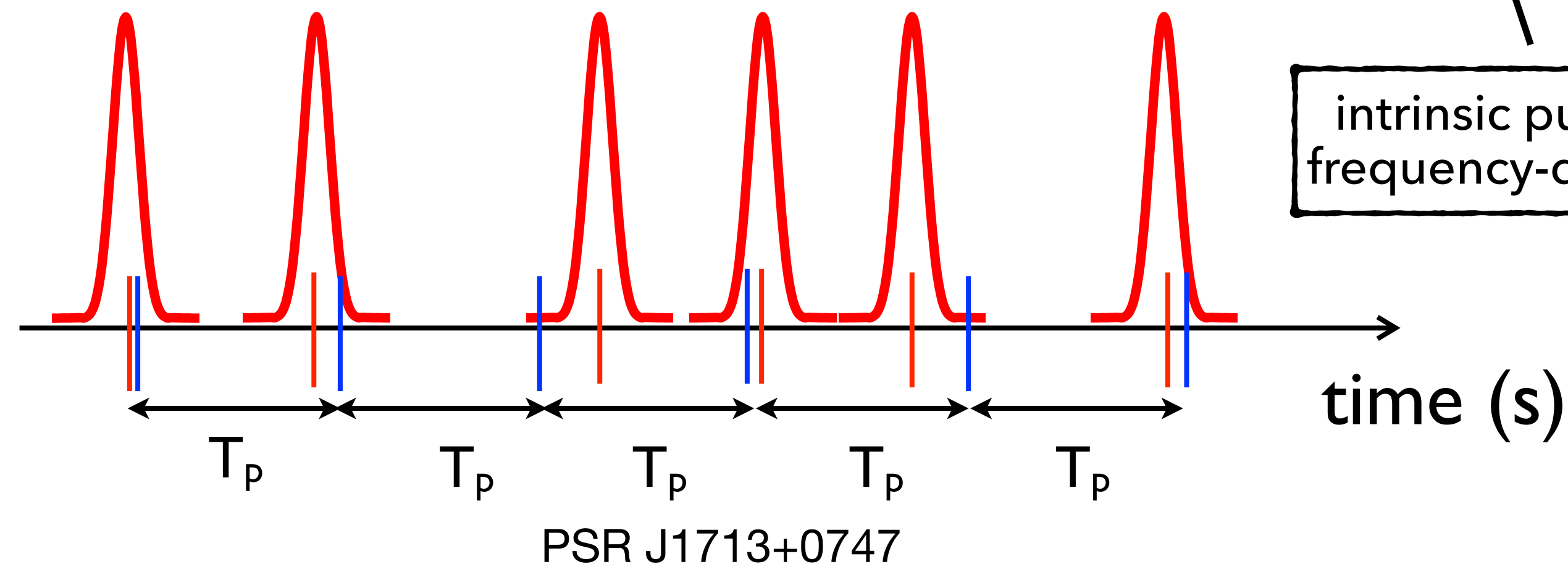
Pulsar Timing Measurements

pulsar spin period (and derivative), position in sky, proper motion, distance from Earth, etc.

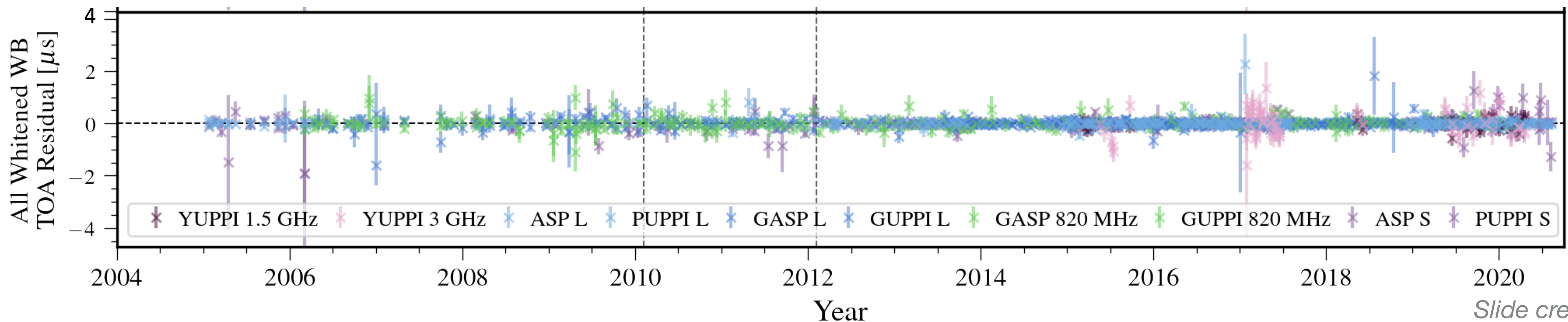
timing residual = **observed arrival time** – **predicted arrival time**
 = unmodeled deterministic processes + noise + GWs

red noise

intrinsic pulsar and measurement noise, frequency-dependent dispersion measure



PSR J1713+0747



Slide credit: J. Romano

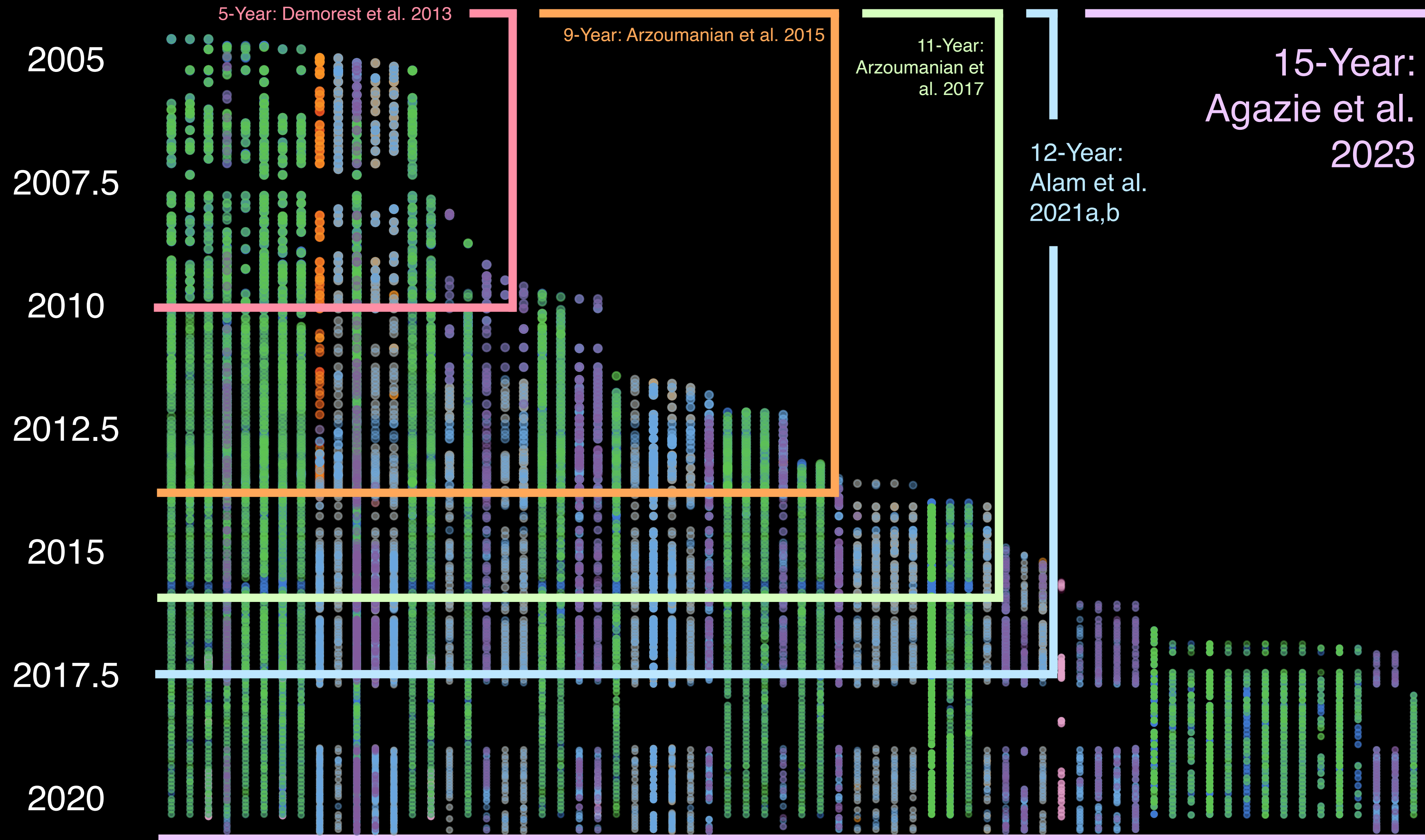
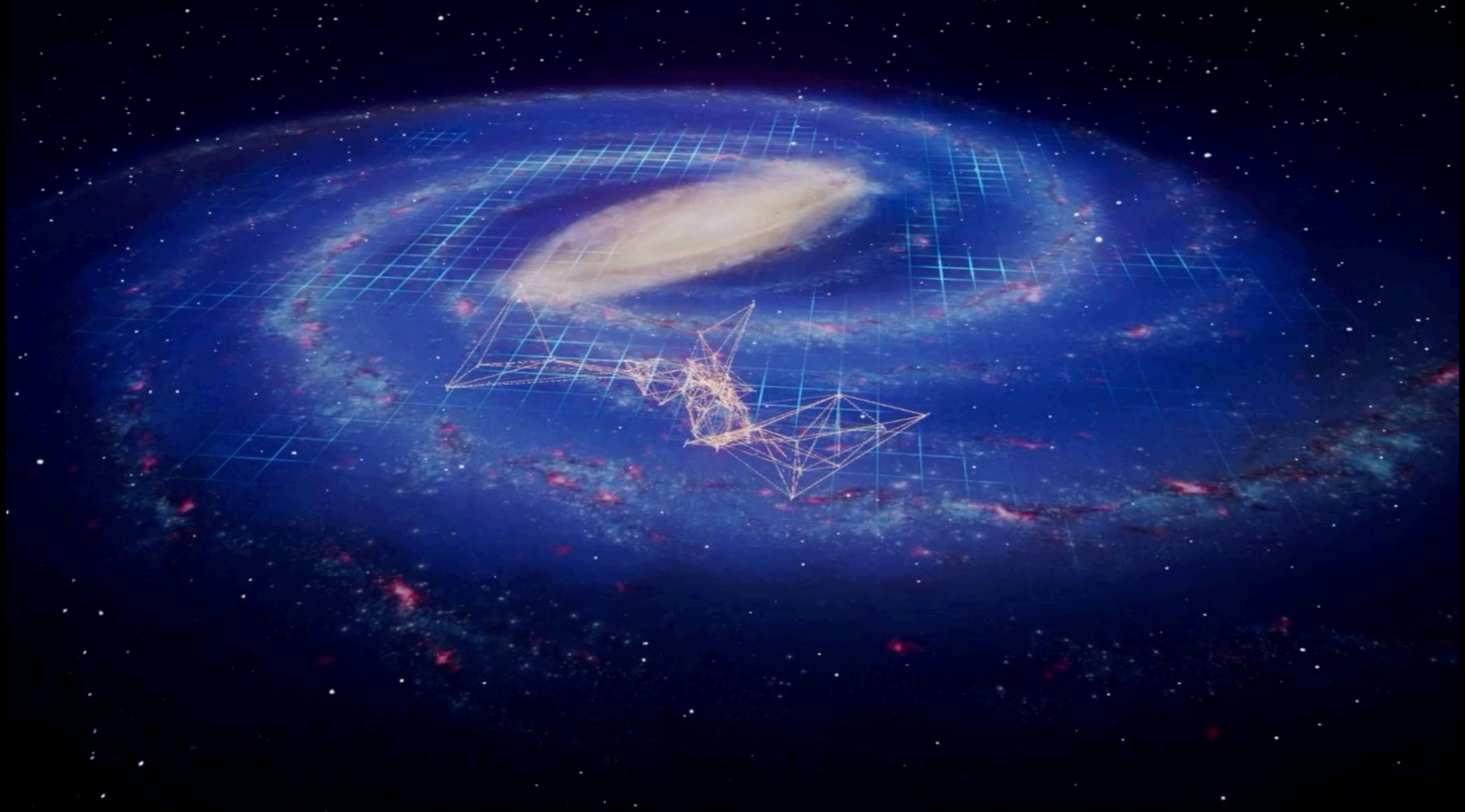
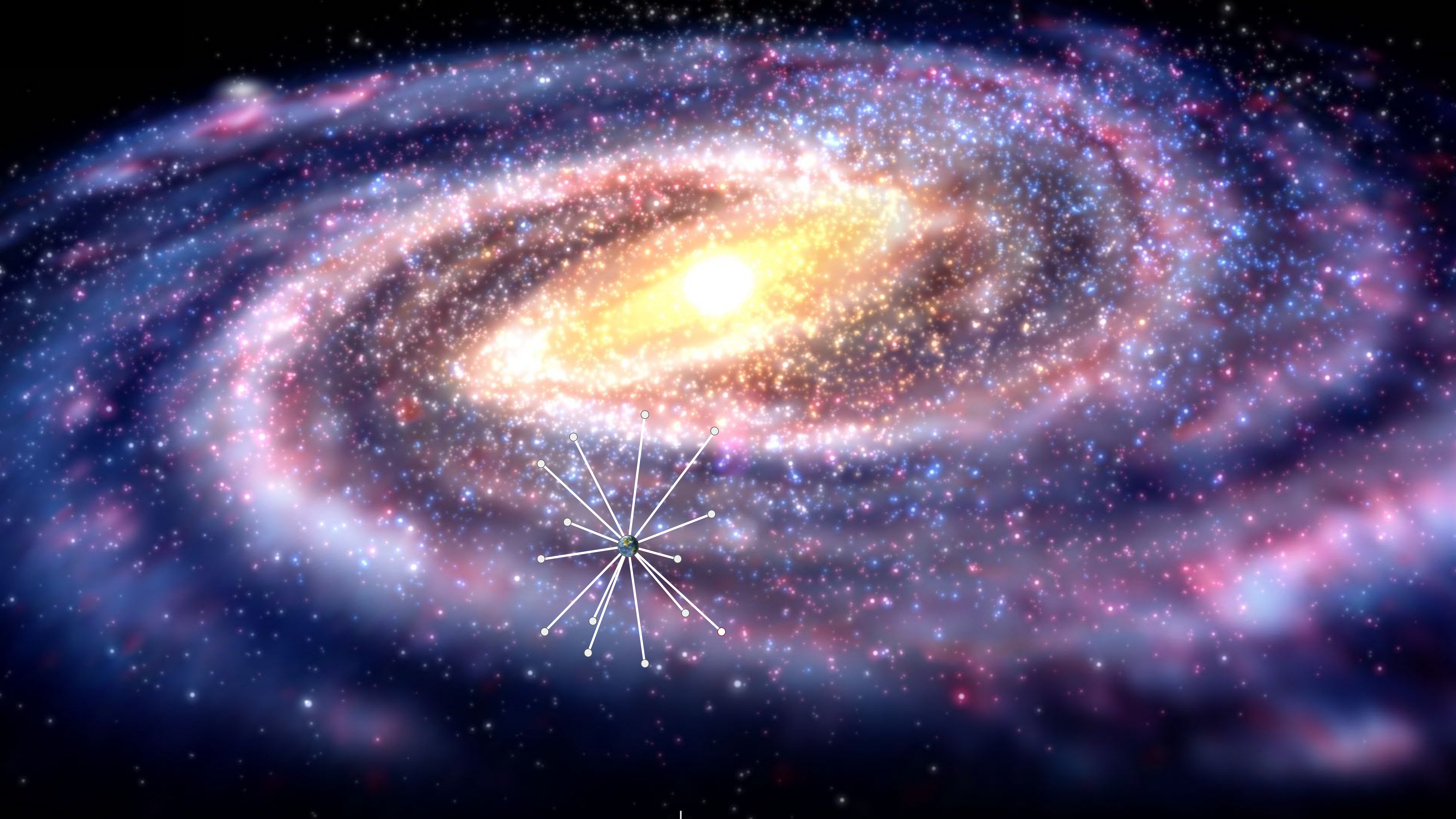
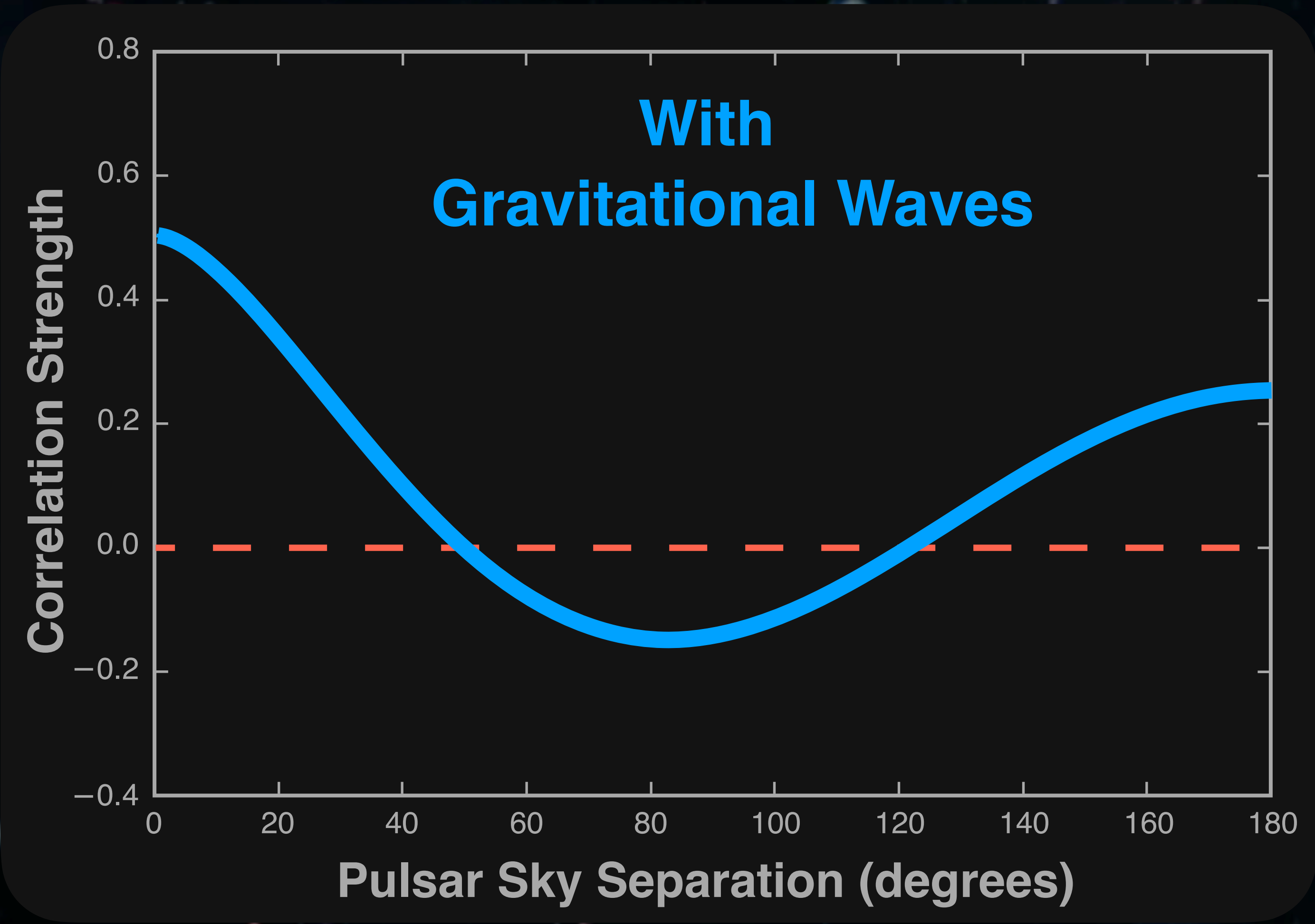
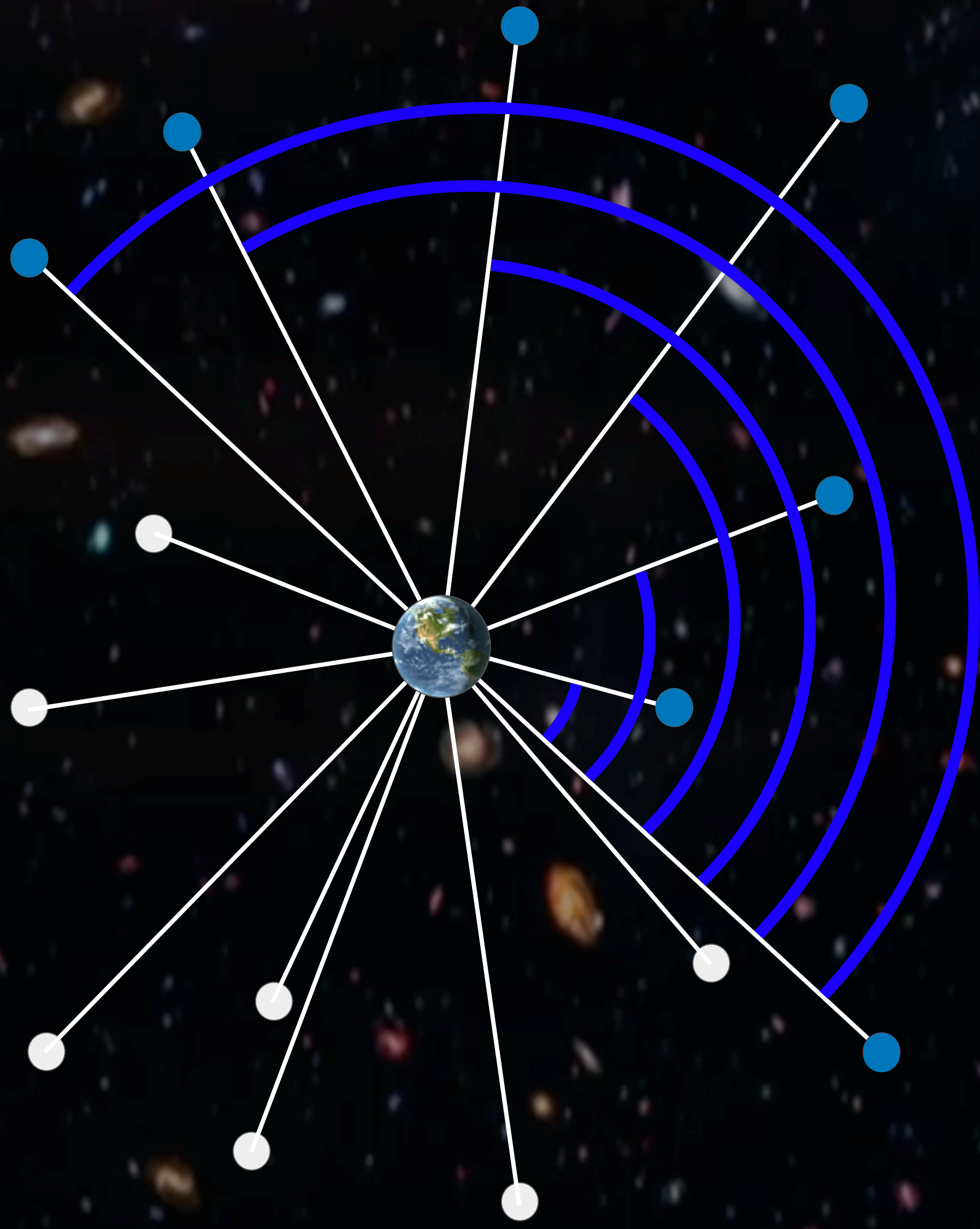
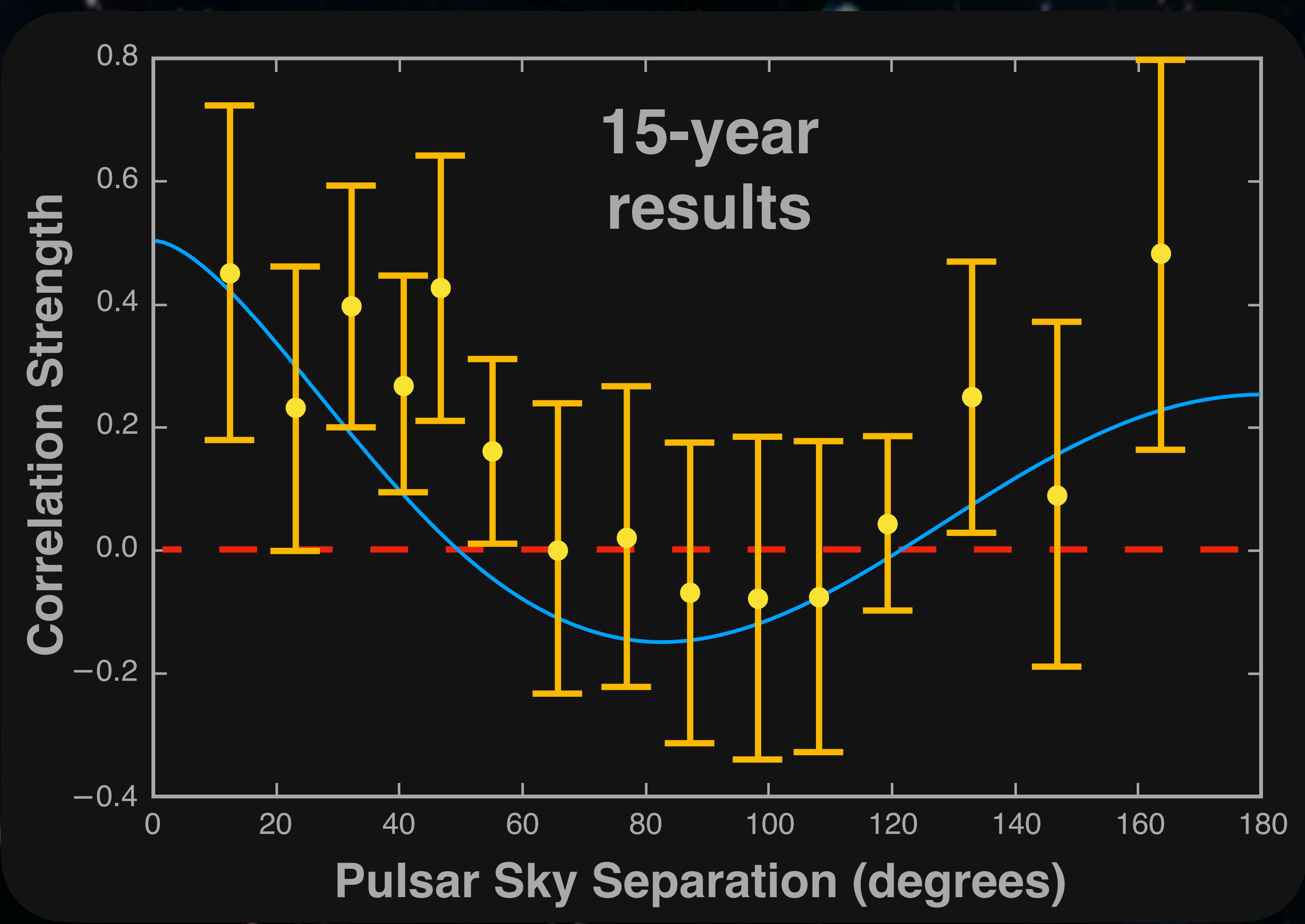
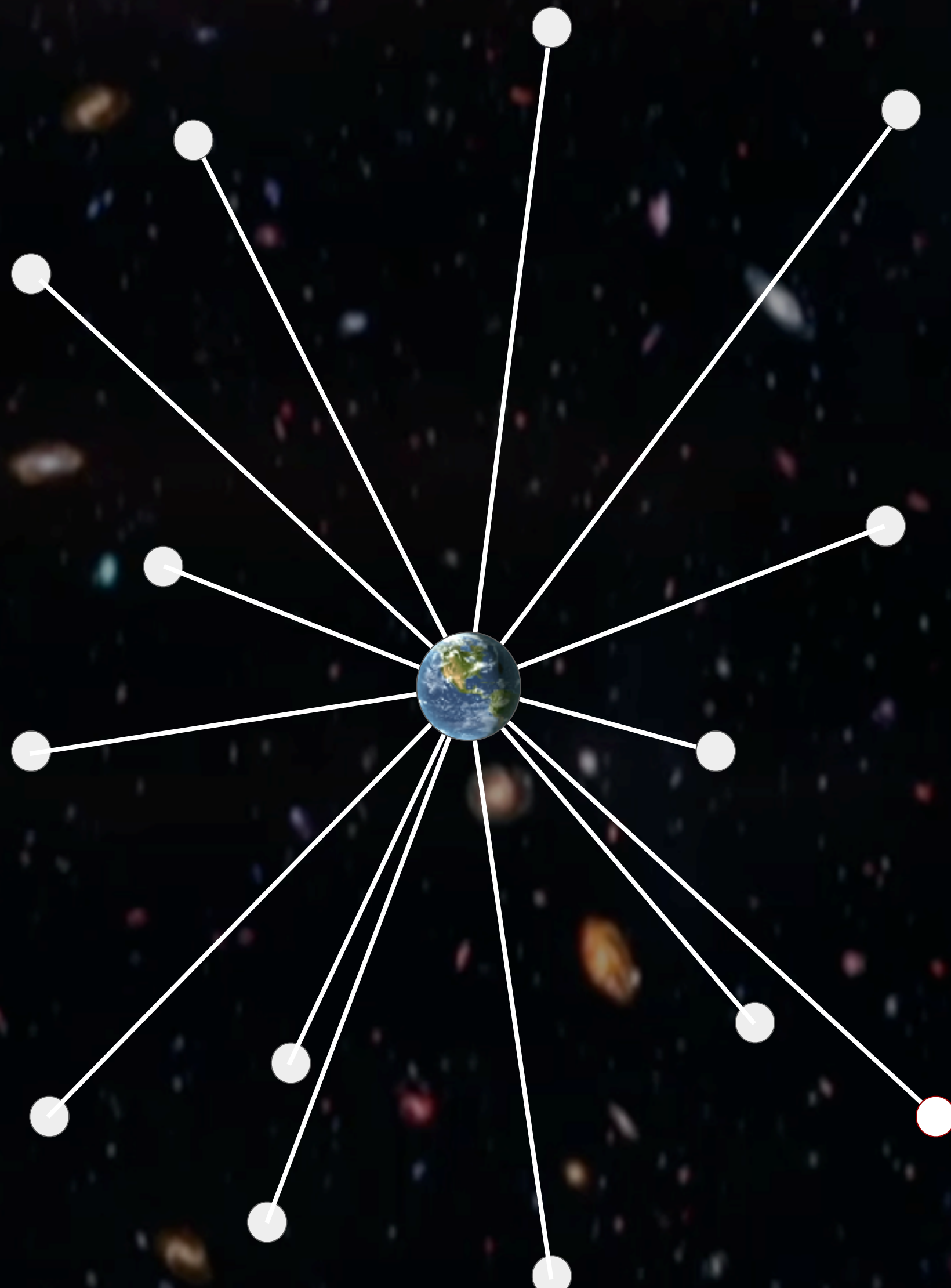


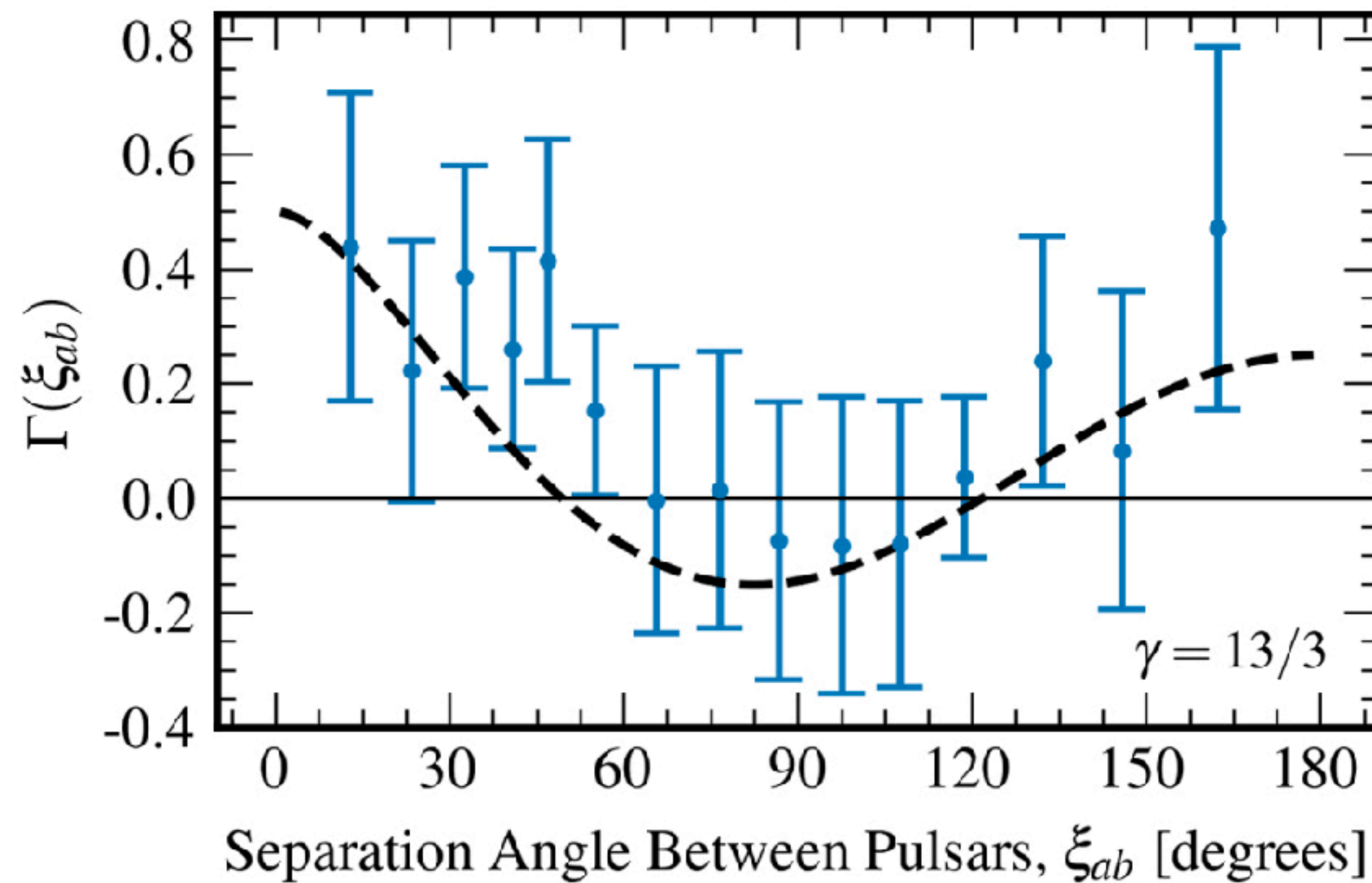
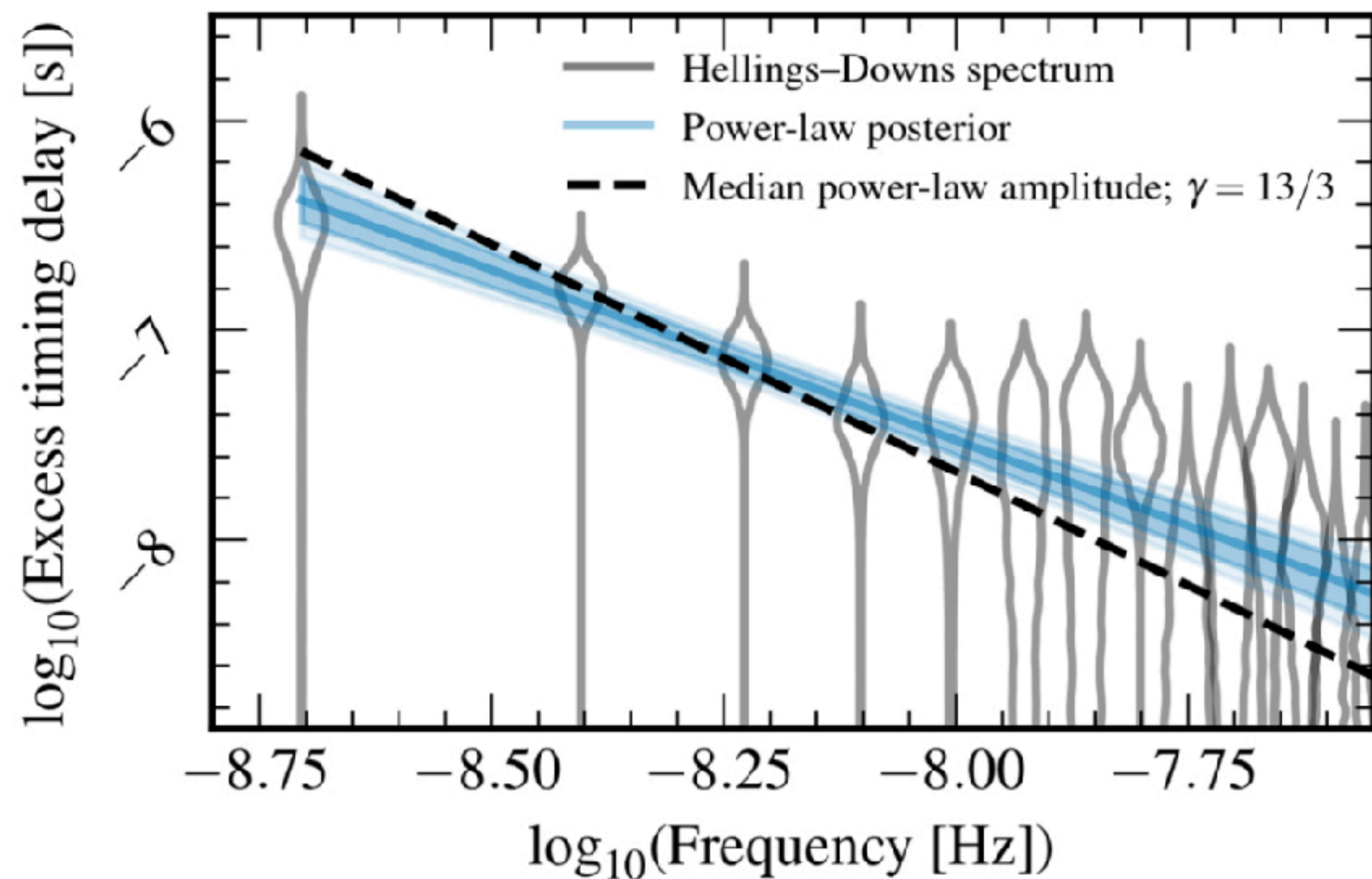
Figure credit: T. Cromartie

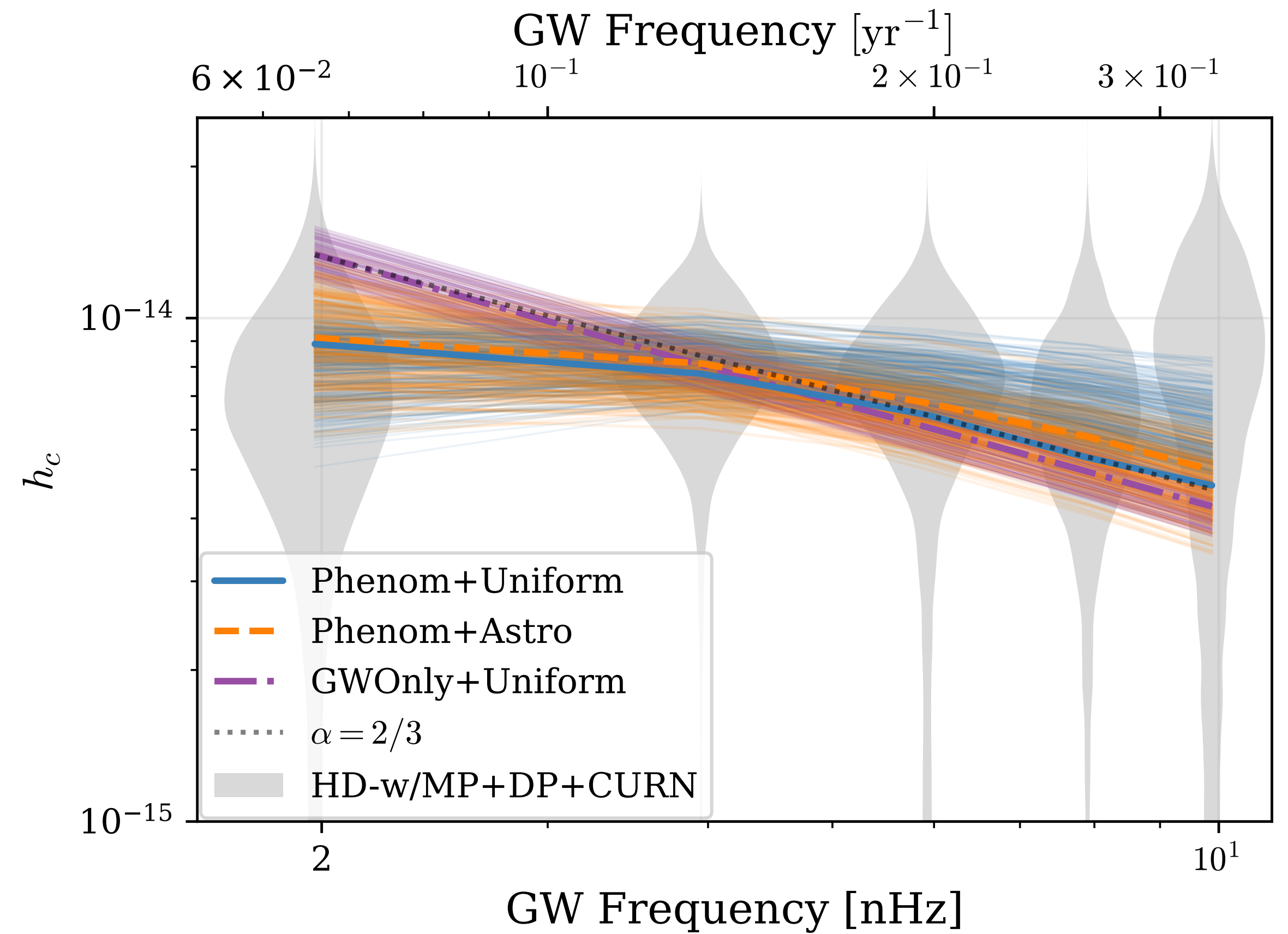
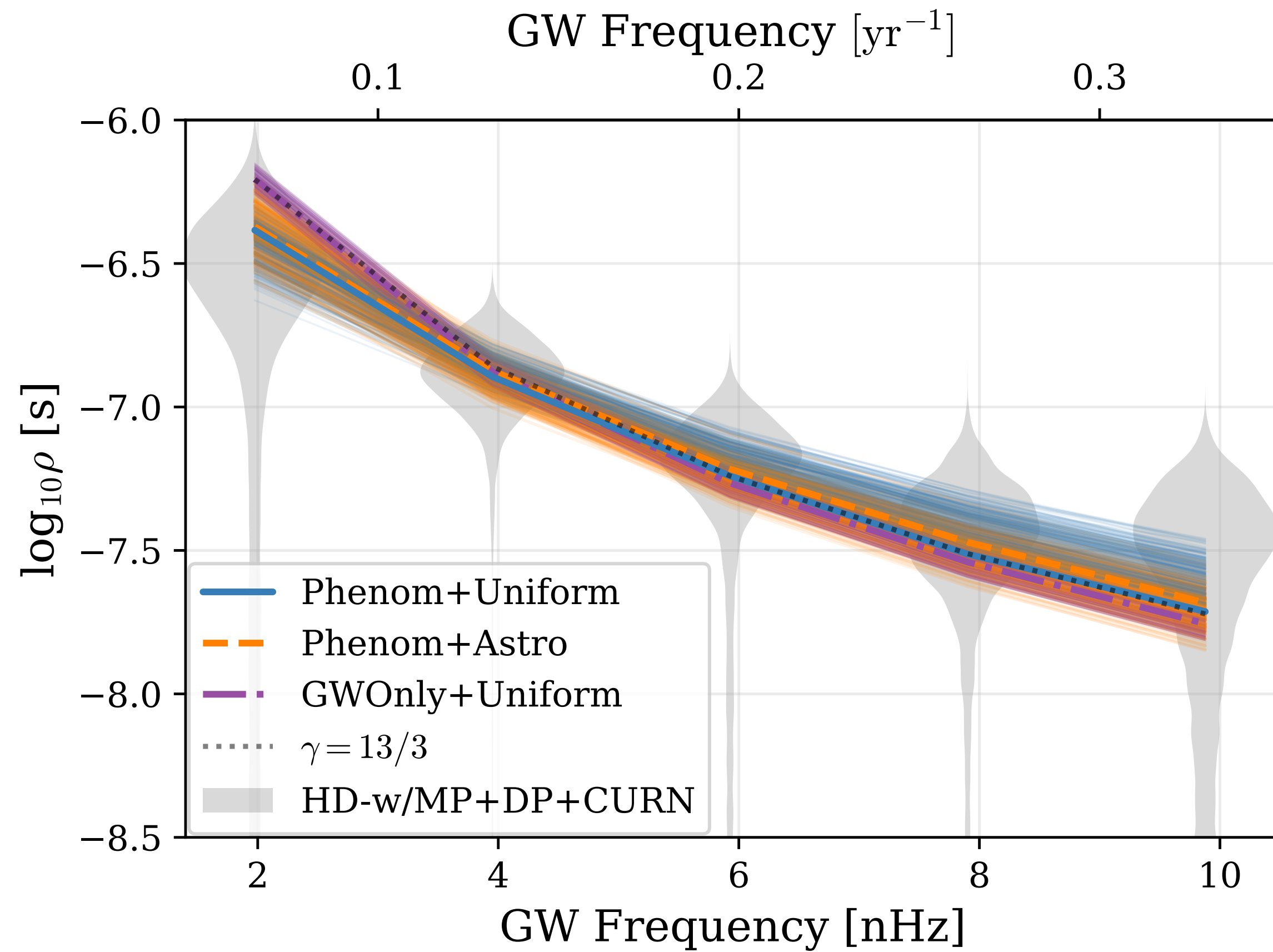












History of the Universe

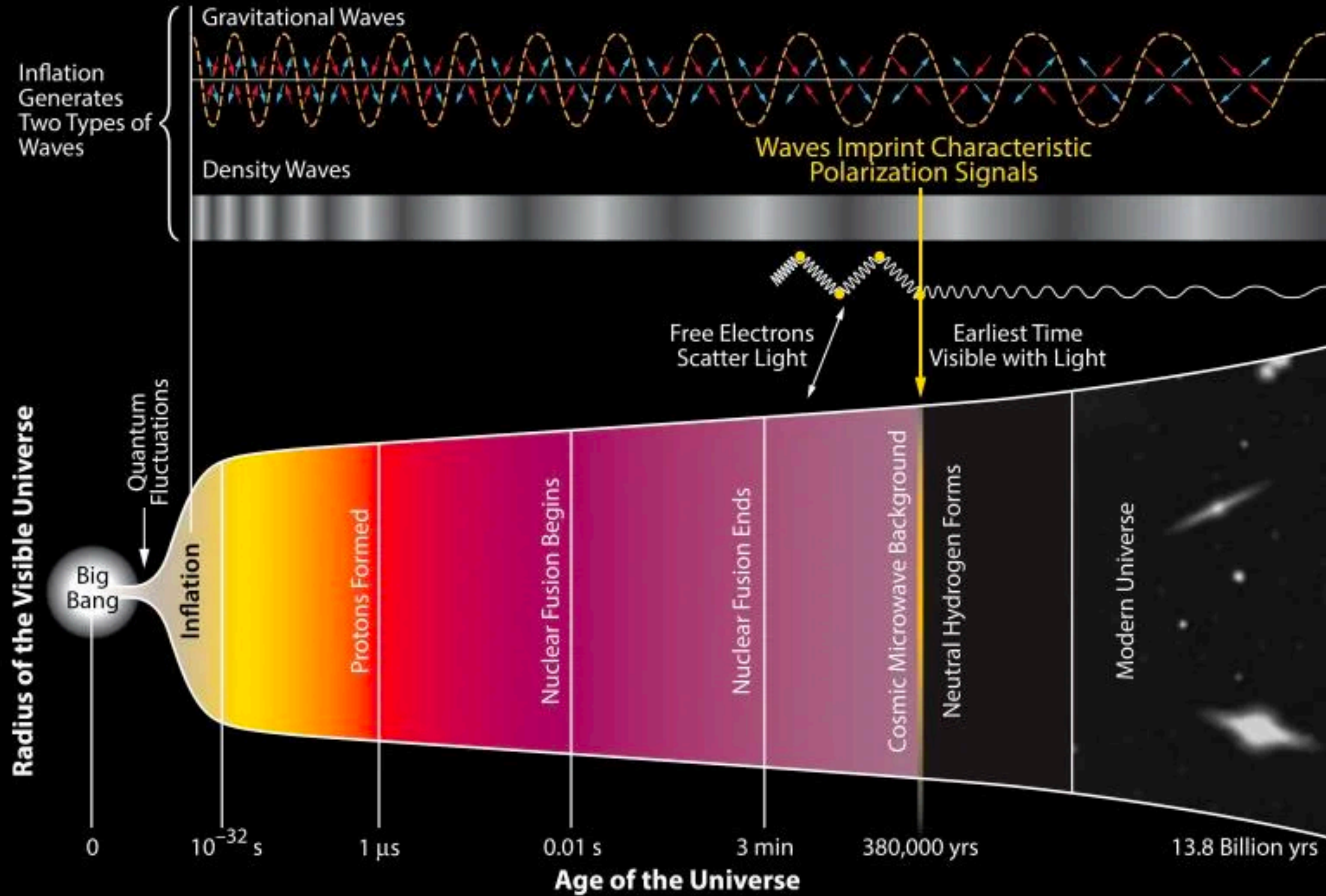
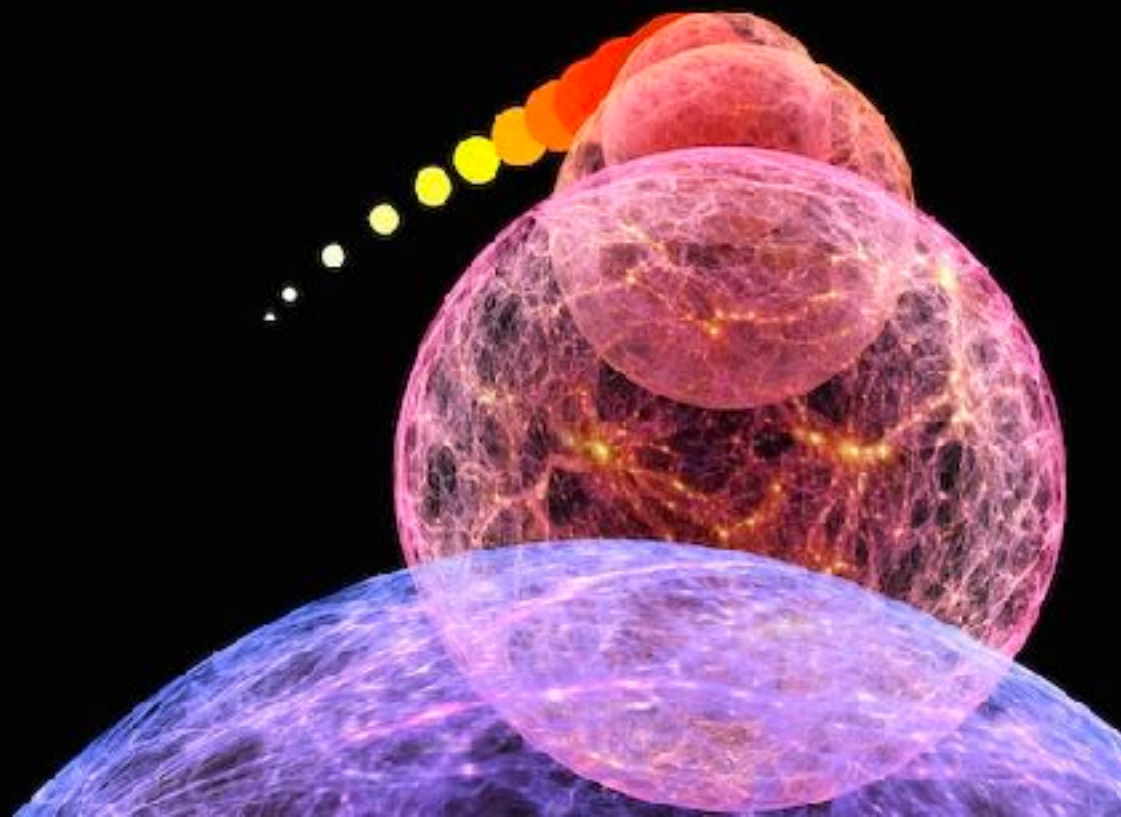


Figure credit: BICEP2

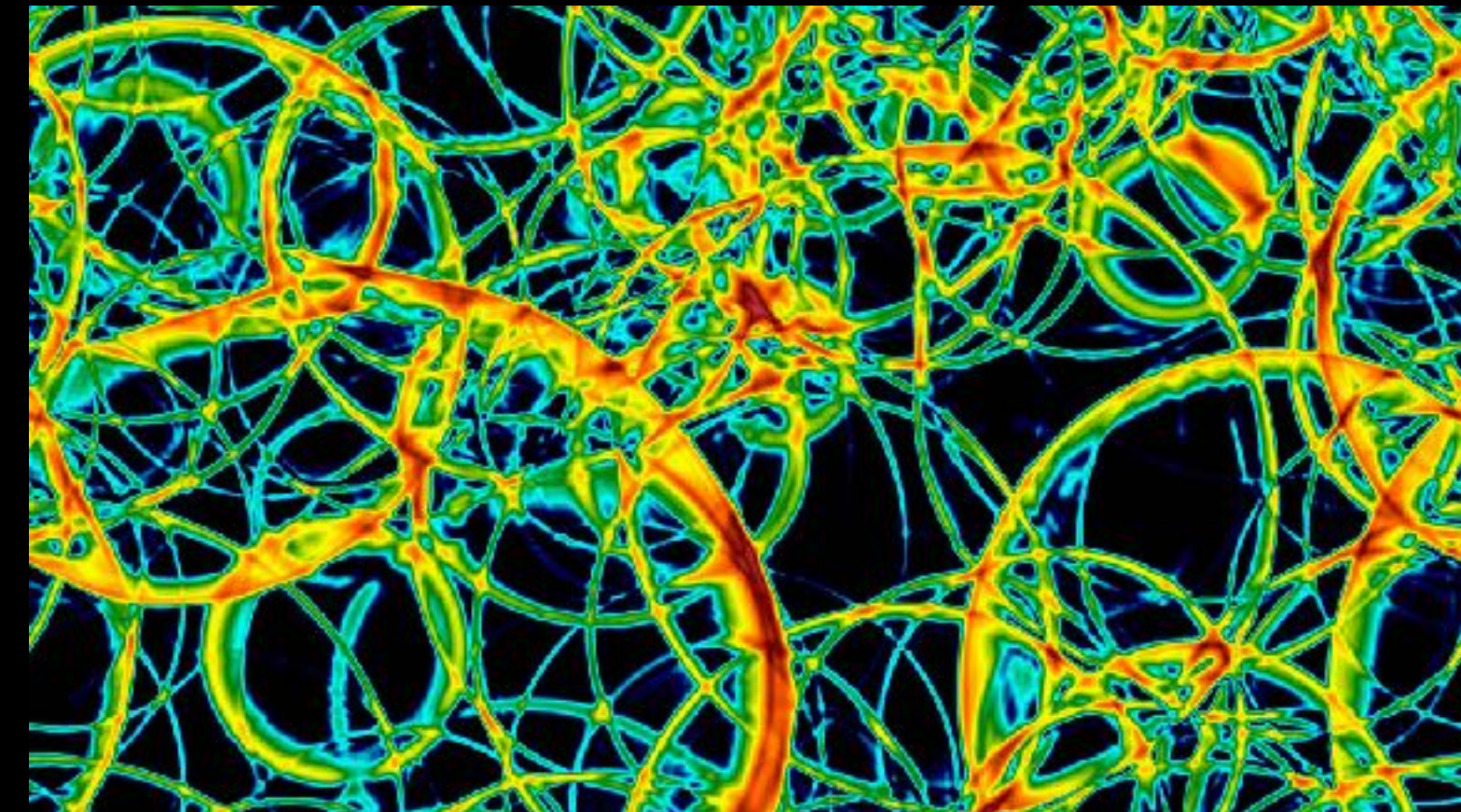
Inflation

- ◆ Non-minimal blue-tilted models



Phase transitions

- ◆ QCD transition in BSM, dark sector



Topological defects

- ◆ Cosmic strings, domain walls



Enhanced scalar perturbations

- ◆ Primordial black hole production

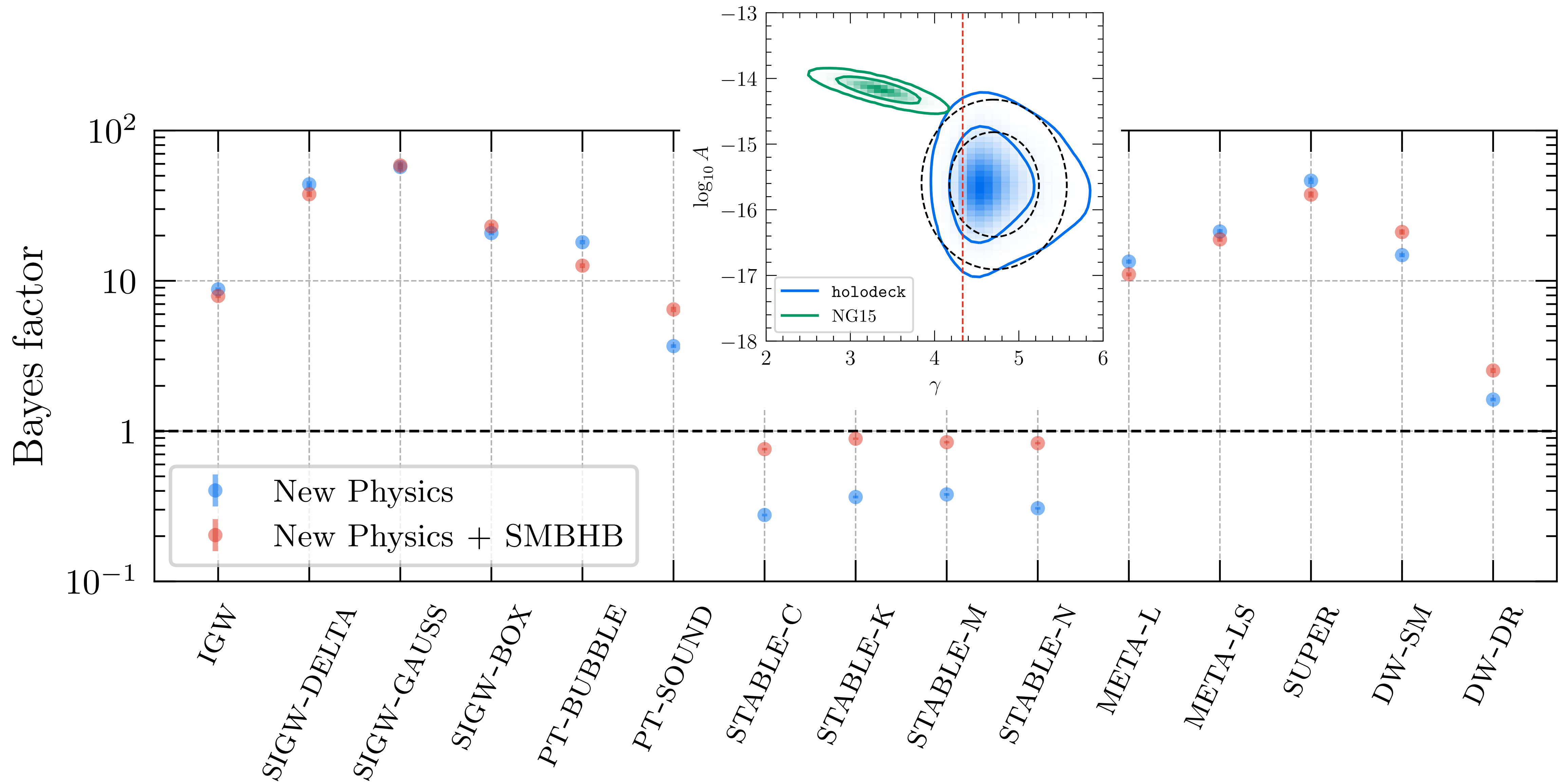


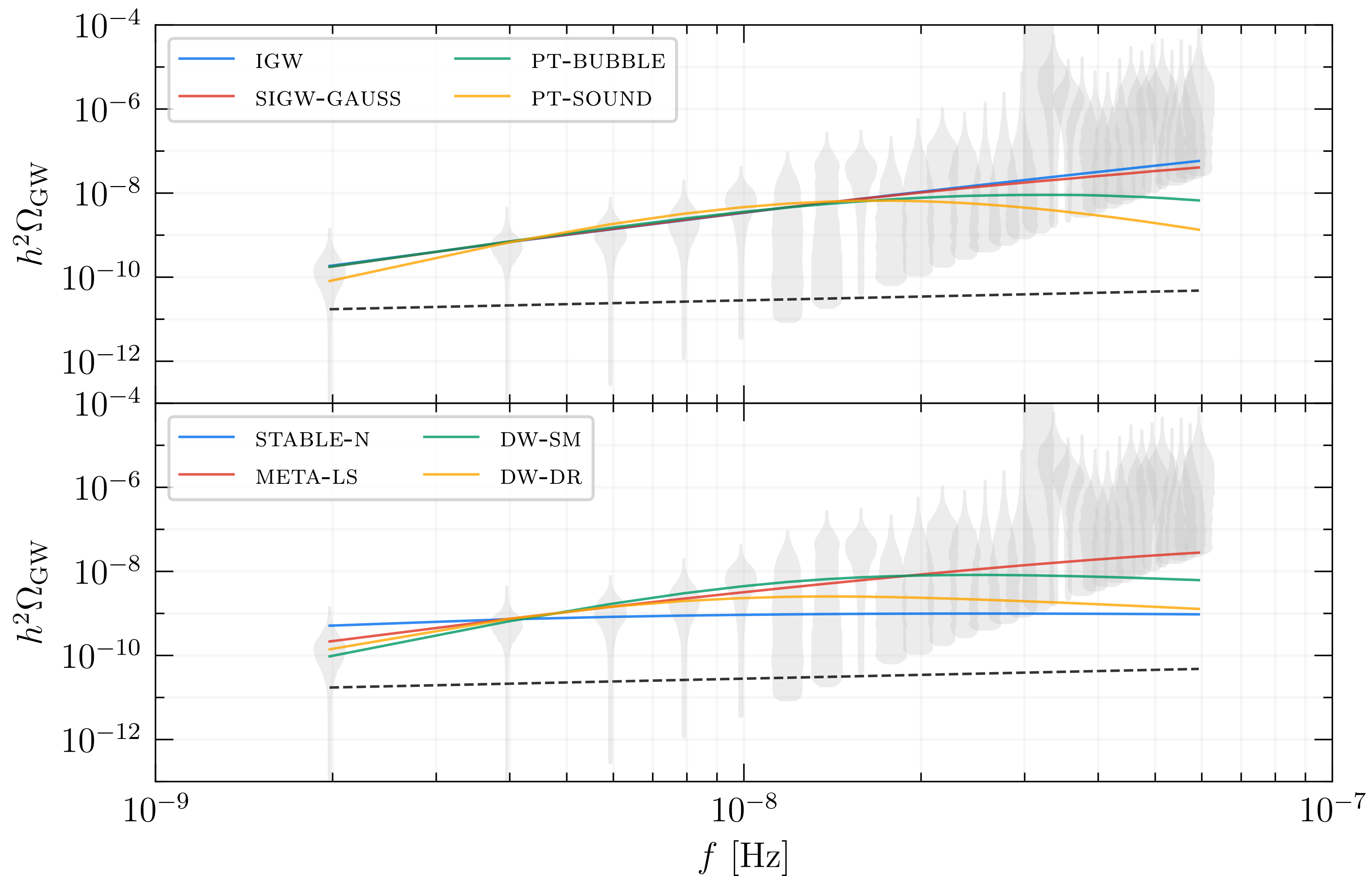
The NANOGrav 15-year Data Set: Search for Signals from New Physics

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 KATHRYN M. ZUREK⁷⁸

- ◆ IGW: Inflationary Gravitational Waves
tensor-to-scalar ratio, tensor spectral index, reheating temperature
- ◆ SIGW: Scalar-Induced Gravitational Waves
scalar amplitude, frequency shape parameters (delta, gauss, box)
- ◆ PT: Phase Transitions (sound-wave analysis & bubble-collisions only)
transition temperature and strength, bubble separation, low/high-frequency slope, spectral-shape width
- ◆ STABLE: Stable Cosmic Strings (cusps, kinks, monochromatic, numerical)
string tension
- ◆ META: Metastable Cosmic Strings (loops only, loops and segments)
string tension, decay parameter
- ◆ SUPER: Cosmic Superstrings
string tension, intercommutation probability
- ◆ DW: Domain Walls
transition temperature, energy fraction, high-frequency slope, spectral-shape width

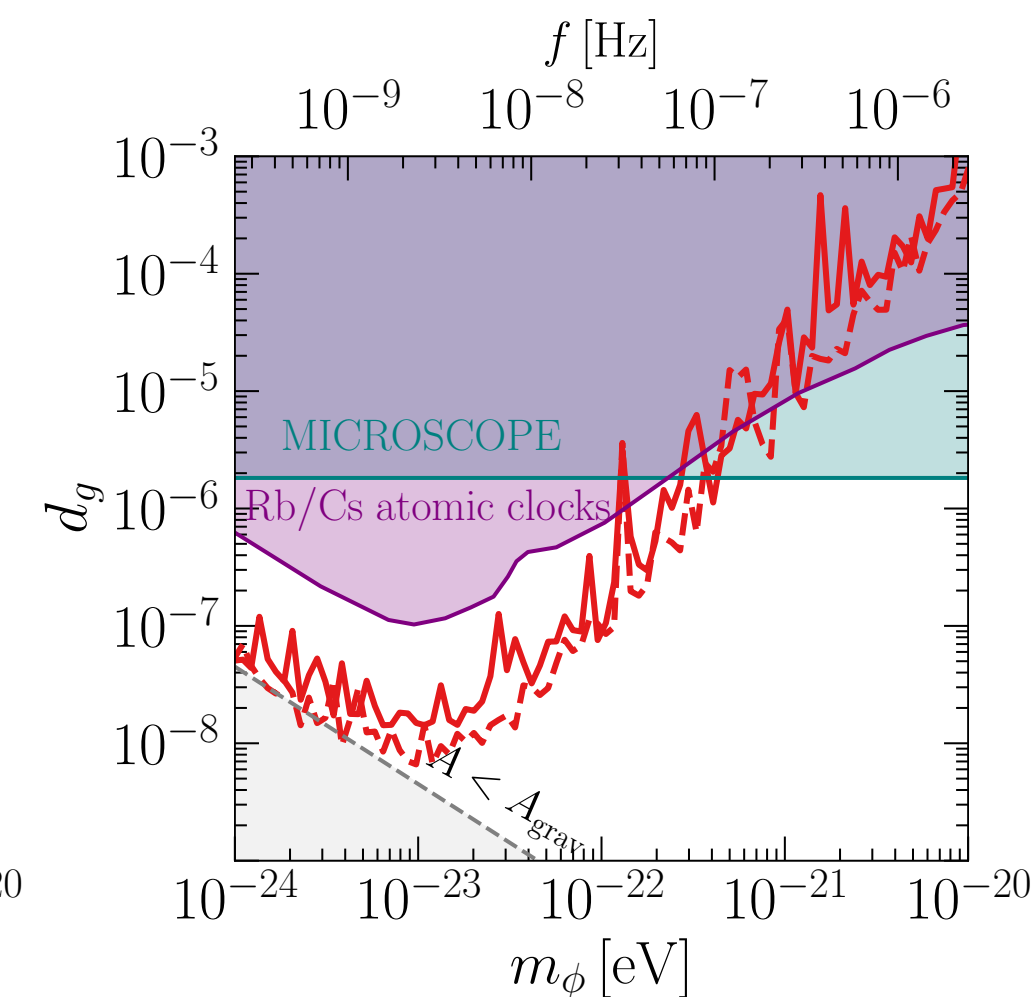
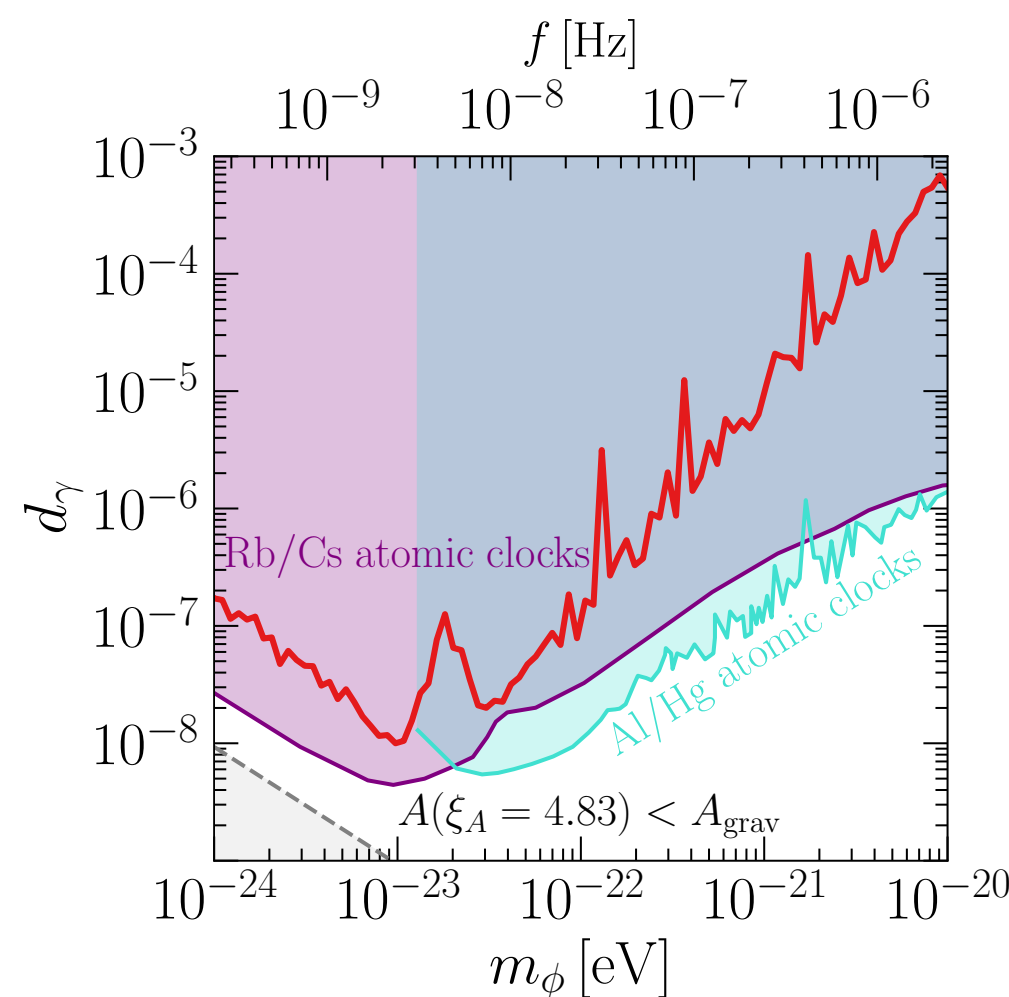
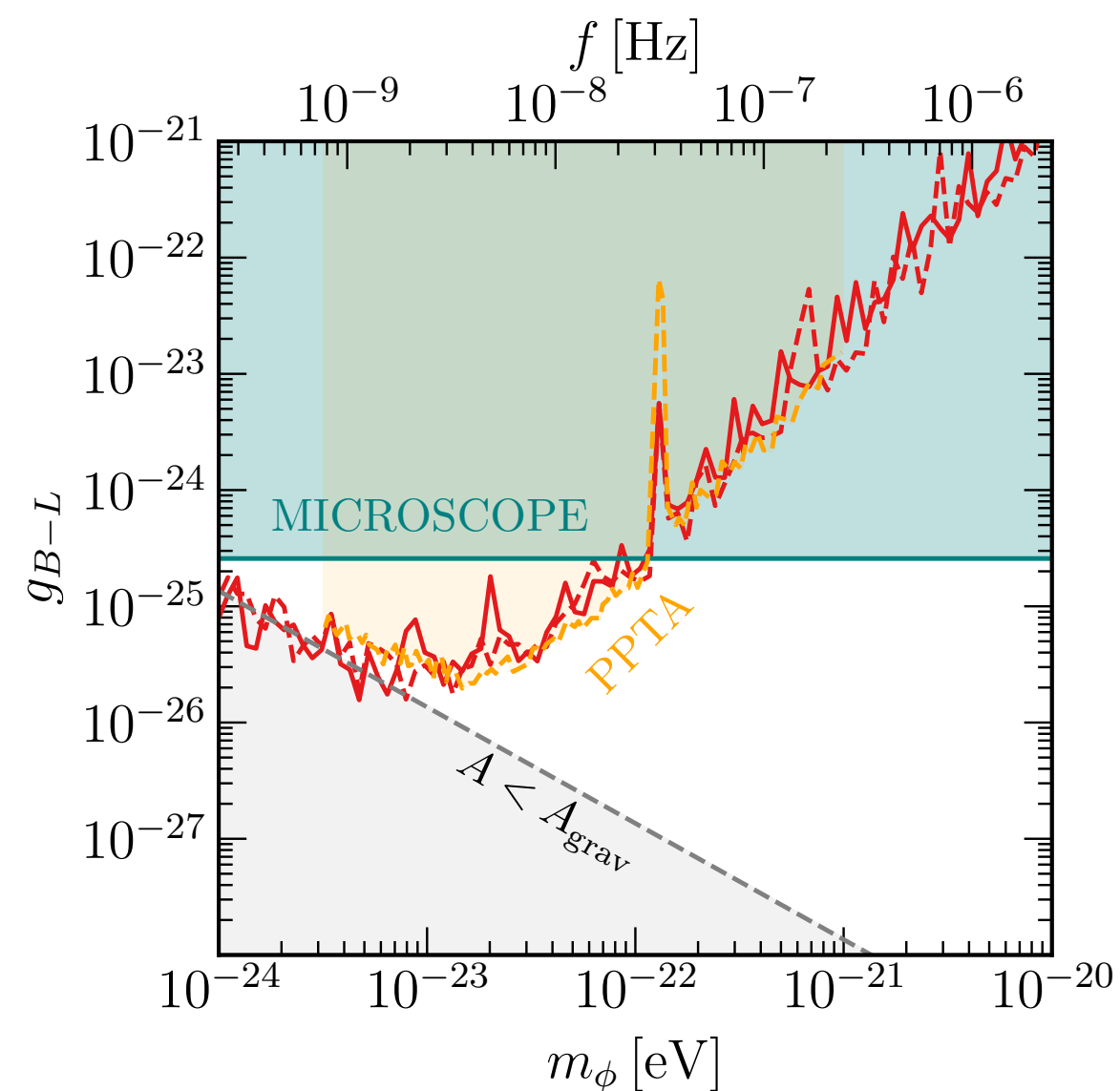
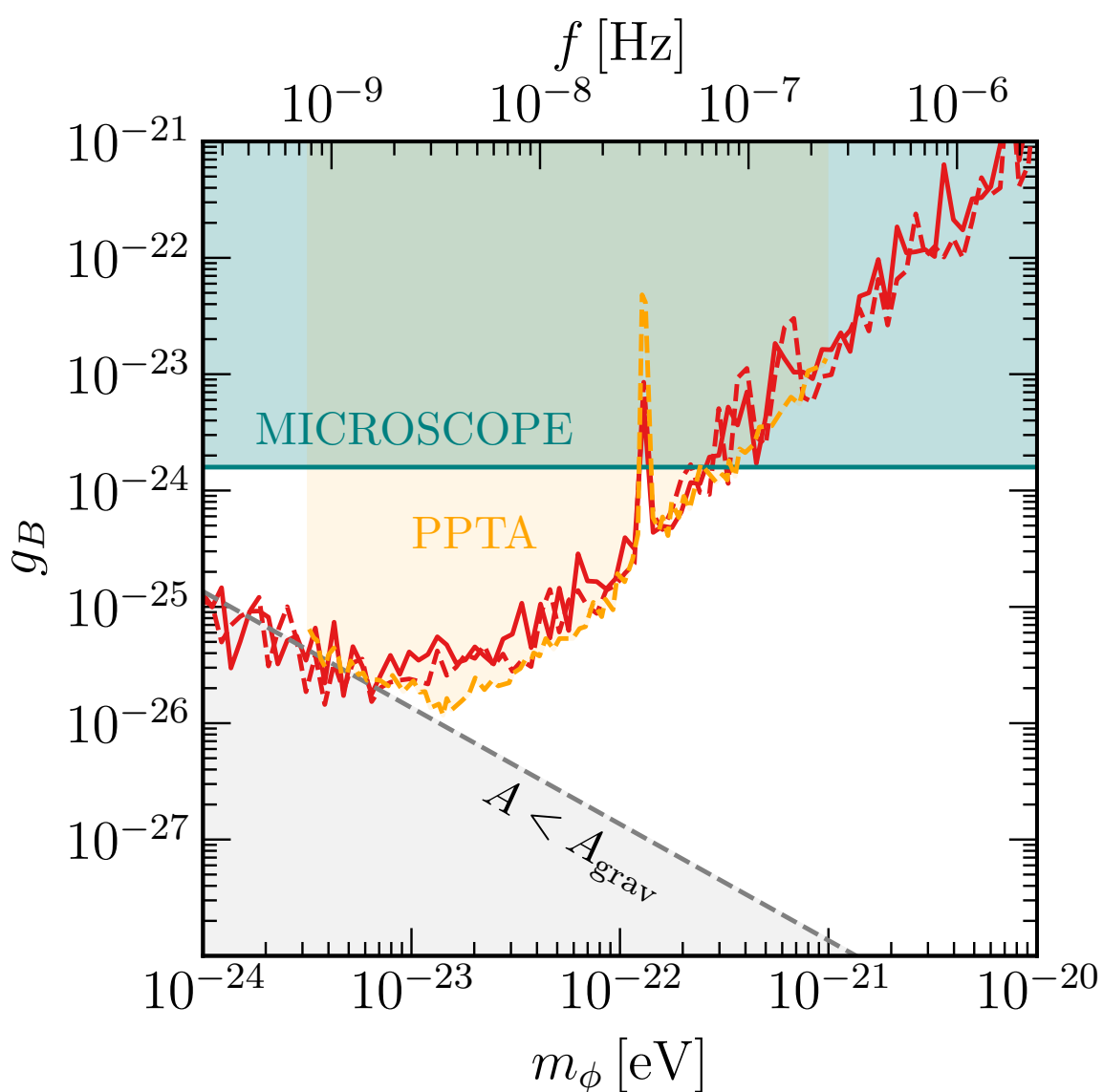
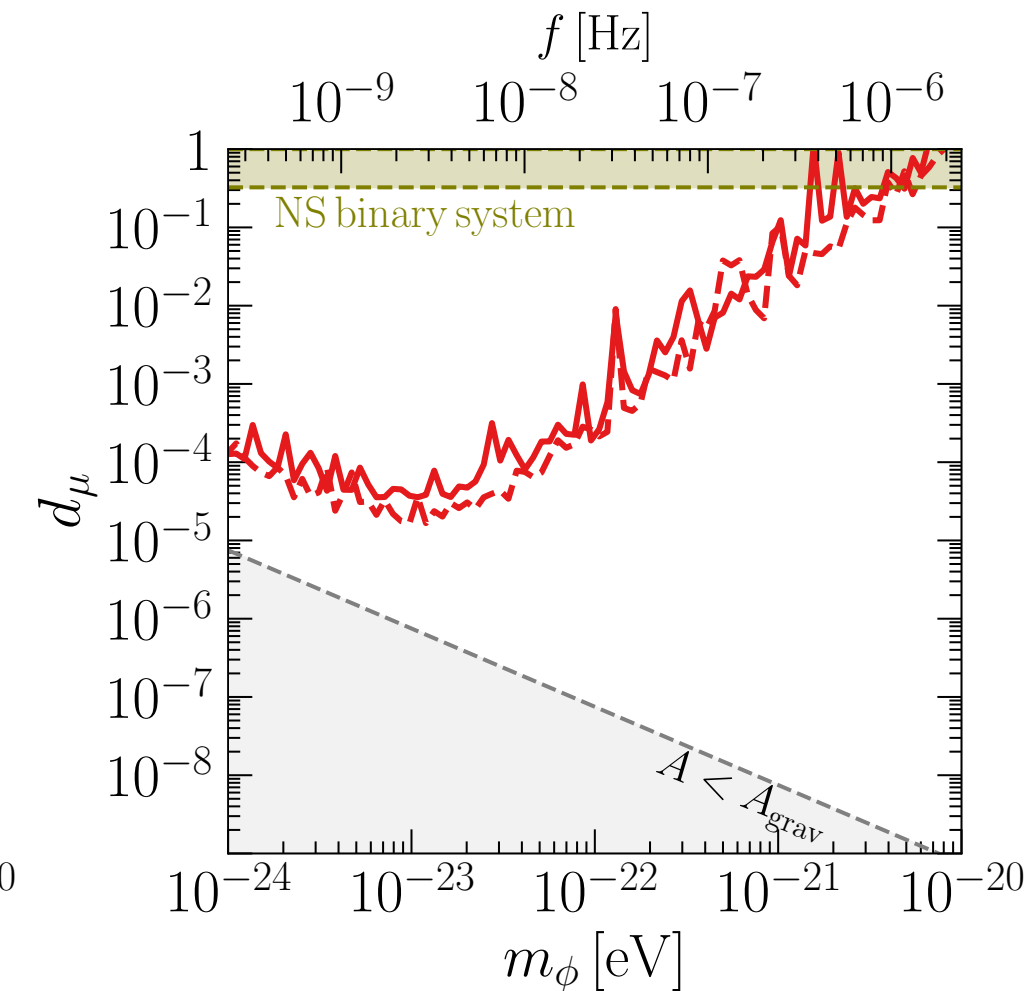
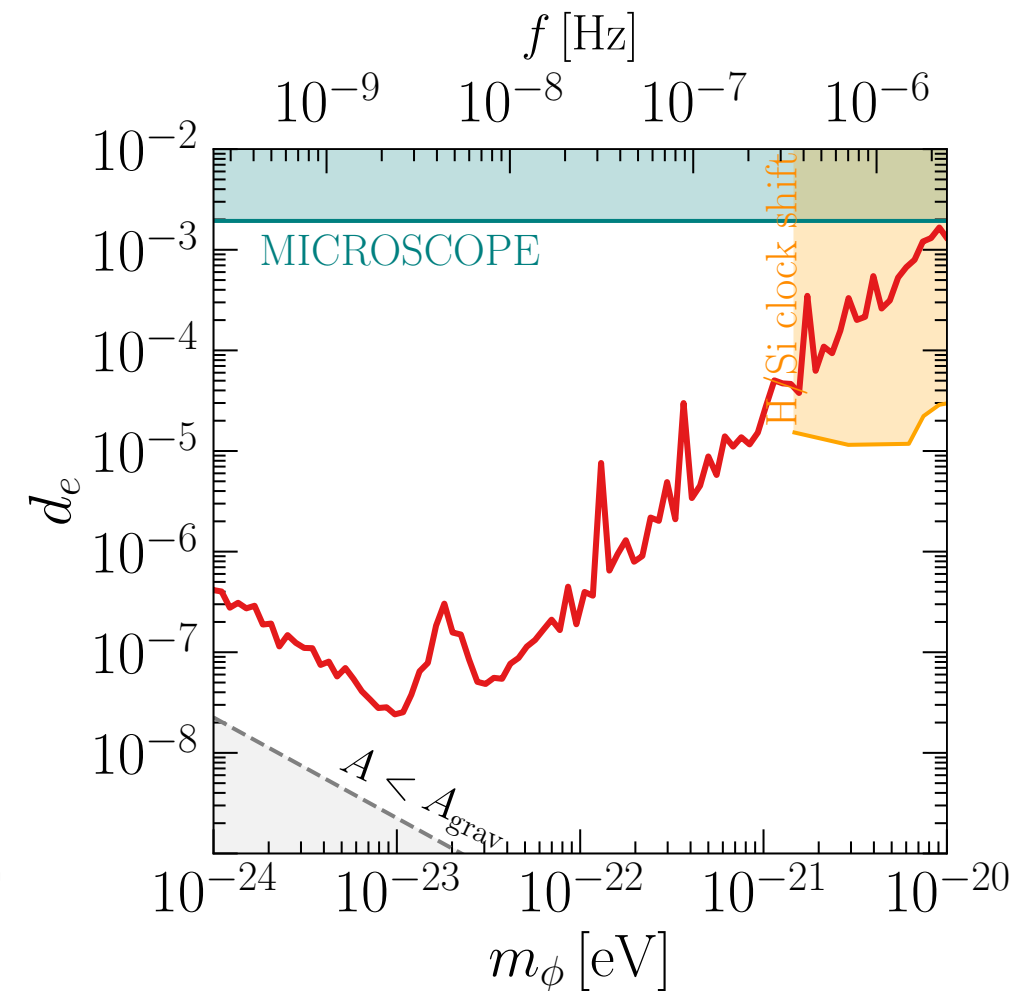
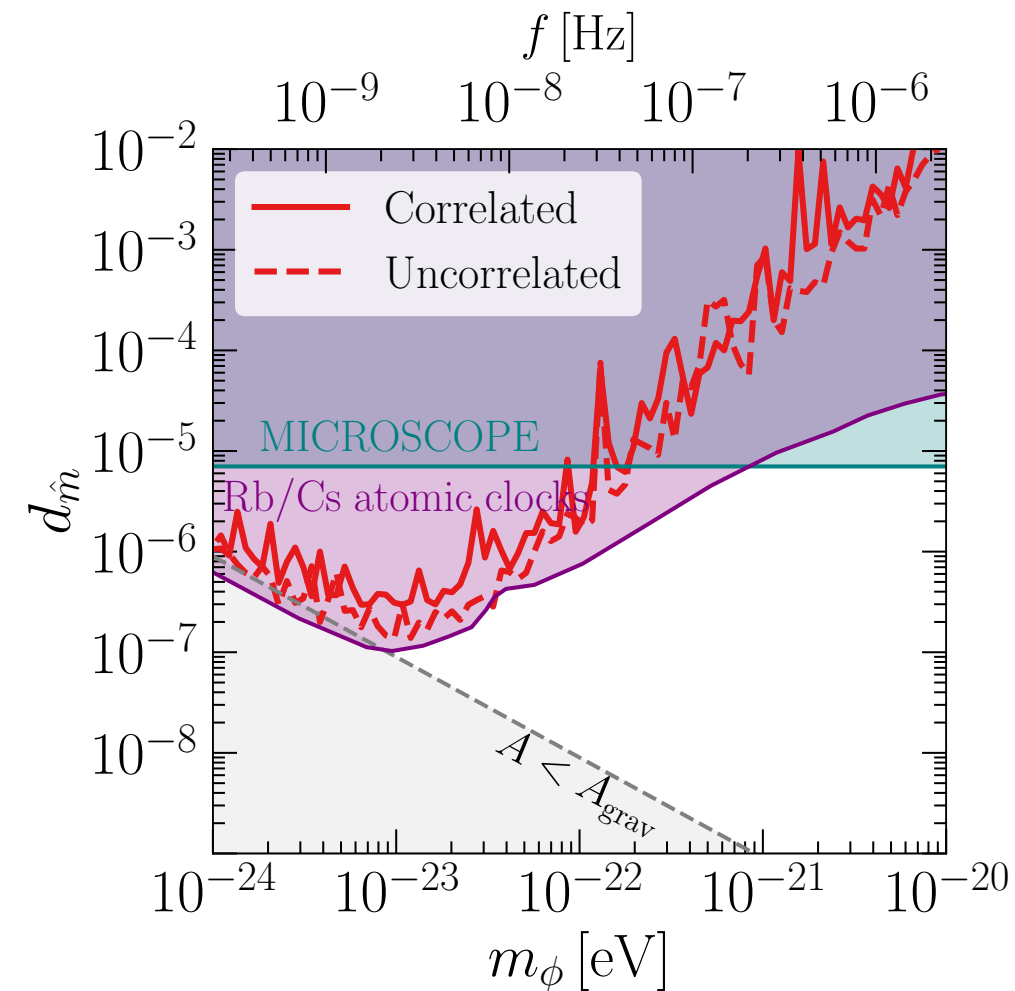
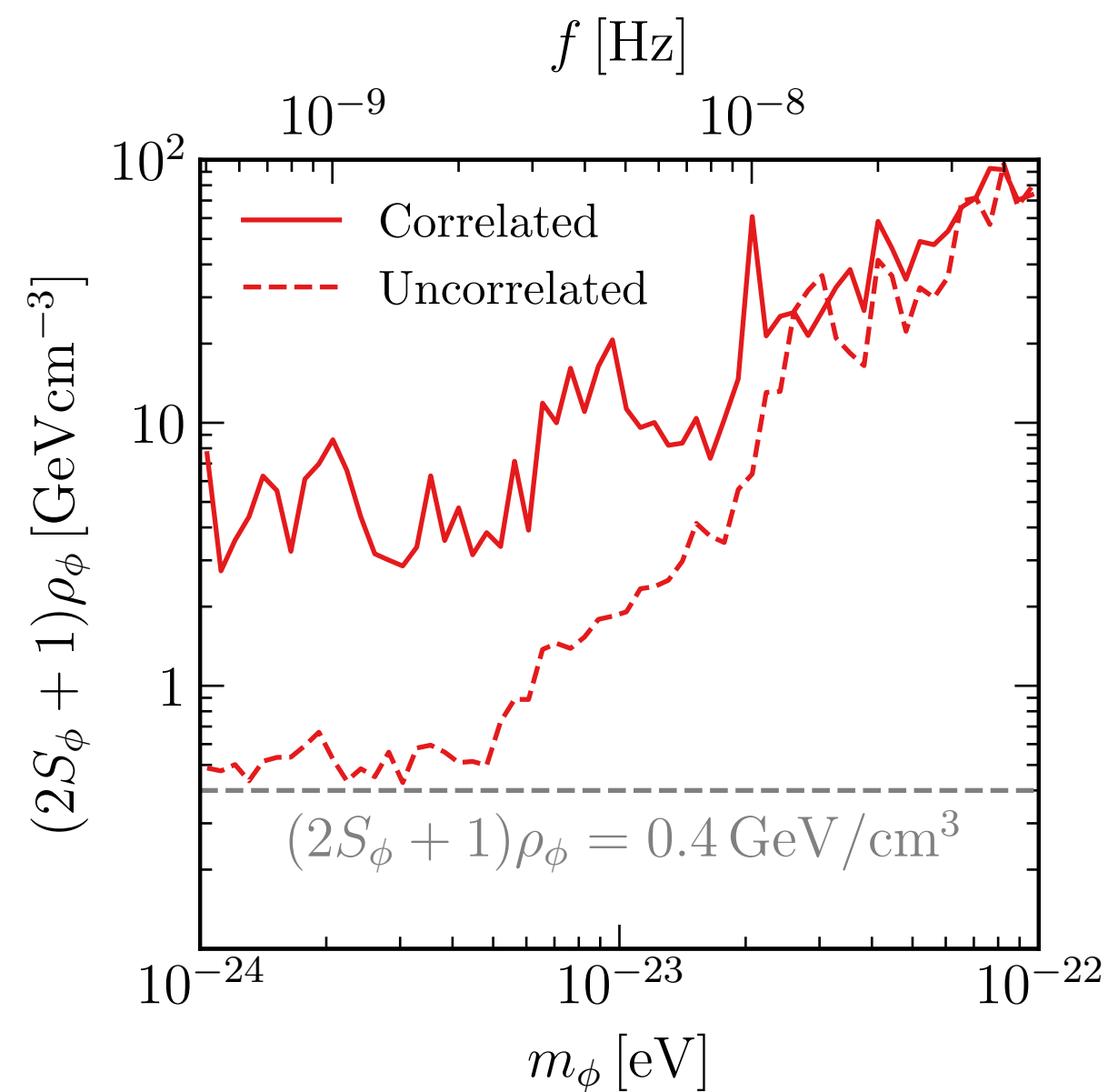
amplitude
spectral features
both



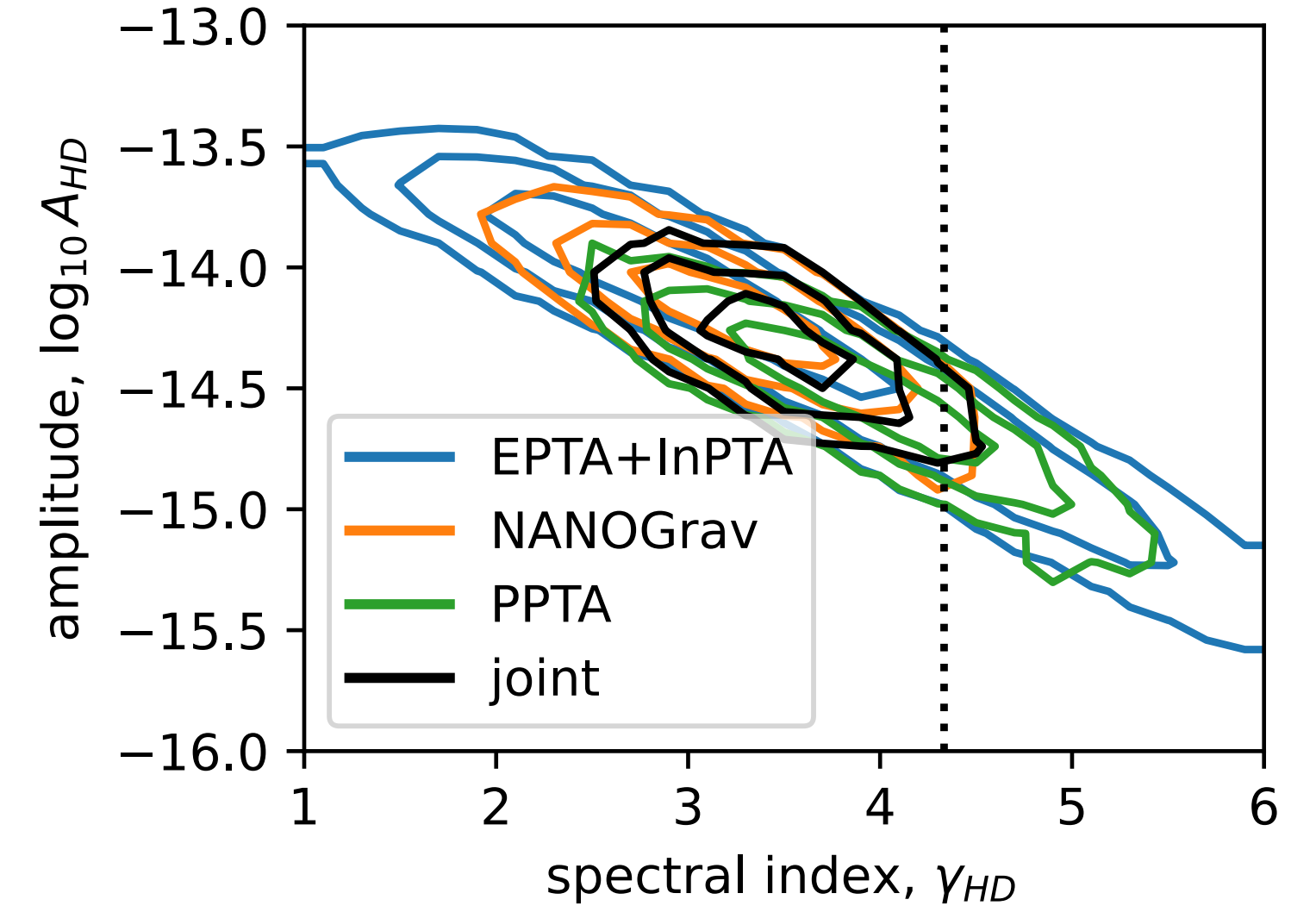
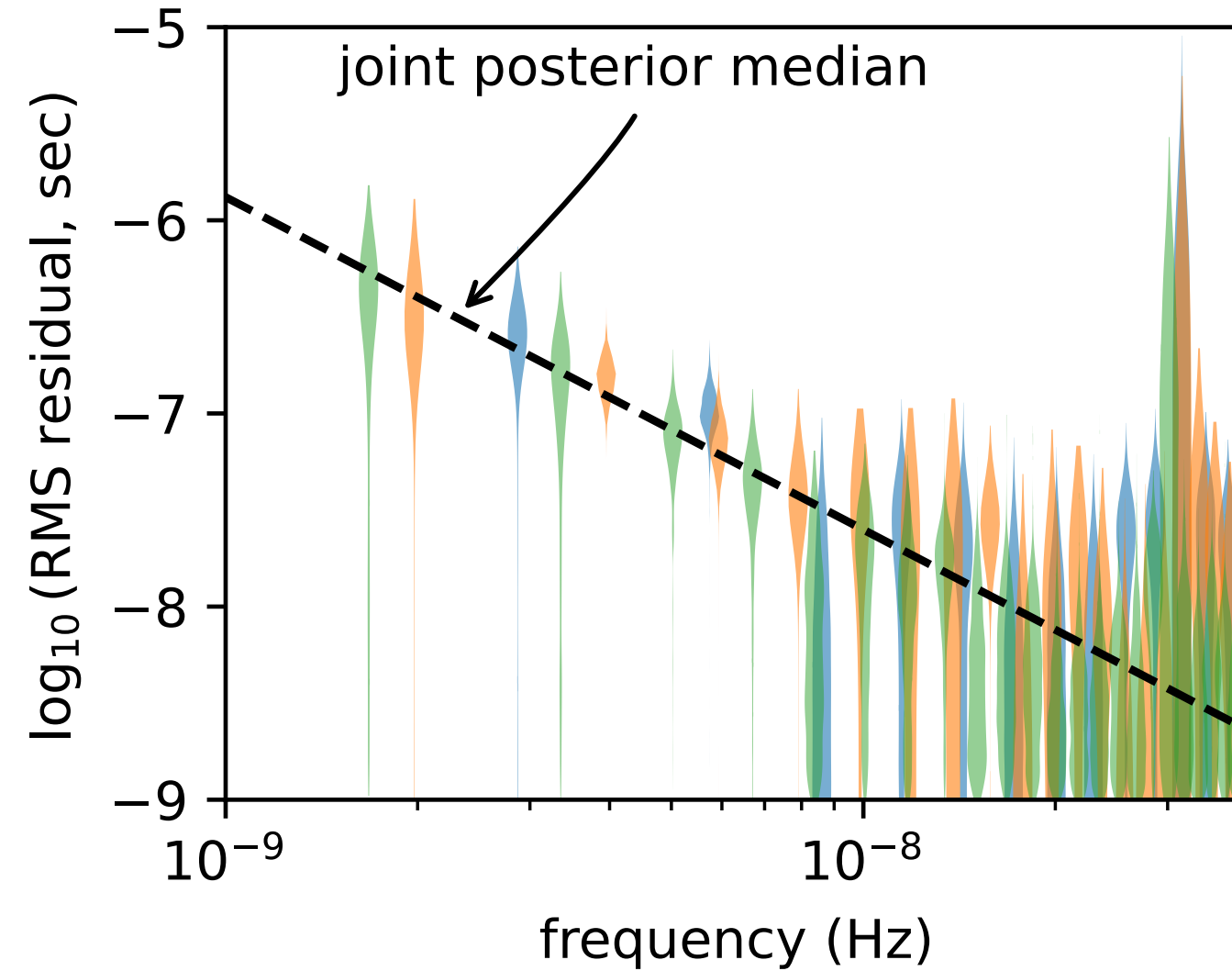
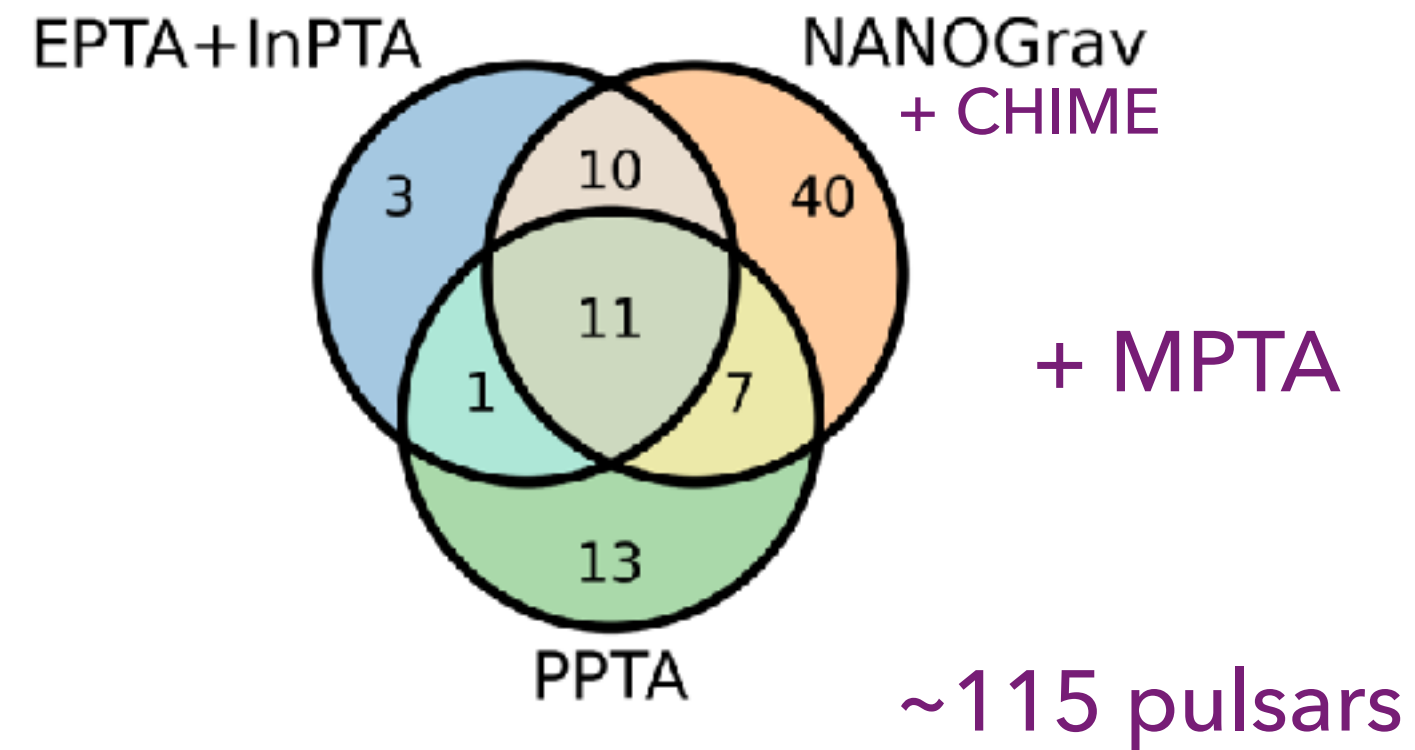


- ◆ Ultralight dark matter with gravitational coupling only
 - ◆ Substructure exhibits pressure oscillations
- ◆ Ultralight dark matter coupled to Standard Model
 - ◆ Doppler signal – vector ULDM accelerates pulsar
 - ◆ Pulsar spin fluctuations – scalar ULDM causes particle mass fluctuations
 - ◆ Reference clock shifts – scalar ULDM alters reference atomic clocks

$$\mathcal{L} \supset \frac{\phi}{\Lambda} \left[\frac{d_\gamma}{4e^2} F_{\mu\nu} F^{\mu\nu} + \frac{d_g \beta_3}{2g_3} G_{\mu\nu}^A G_A^{\mu\nu} - \sum_{f=e,\mu} d_f m_f \bar{f}f - \sum_{q=u,d} (d_q + \gamma_q d_g) m_q \bar{q}q \right]$$

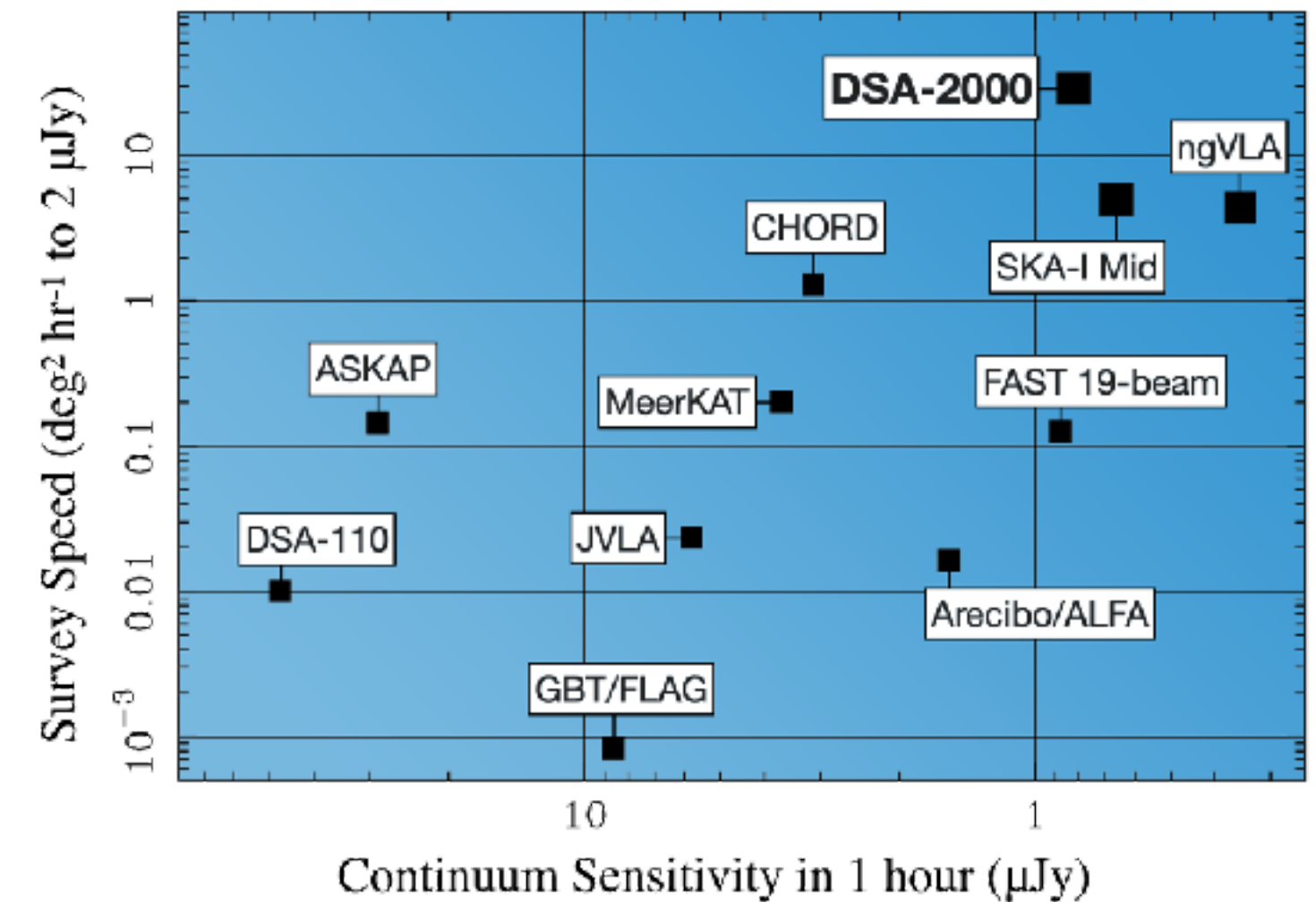


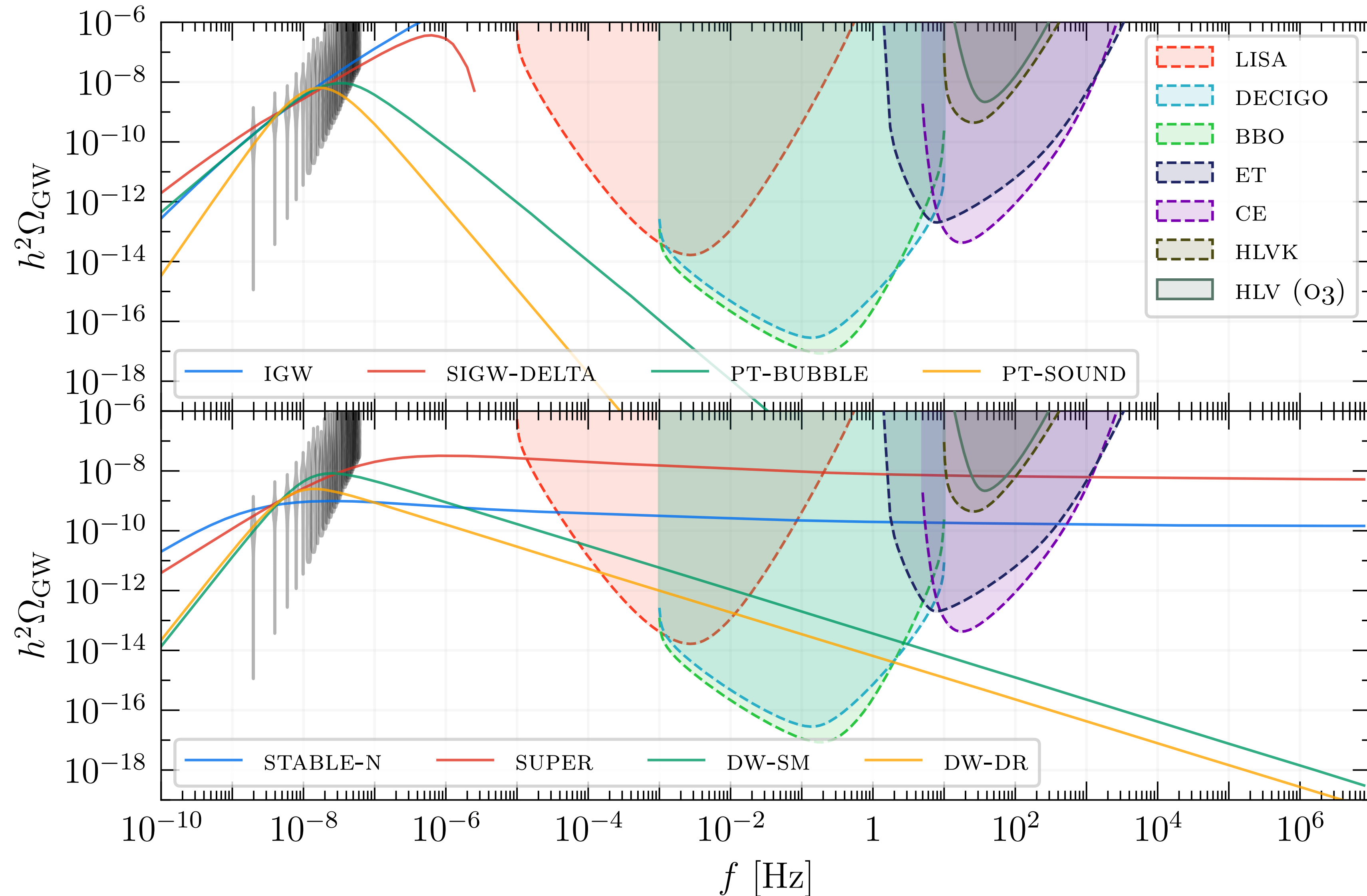
♦ IPTA is assembling DR3



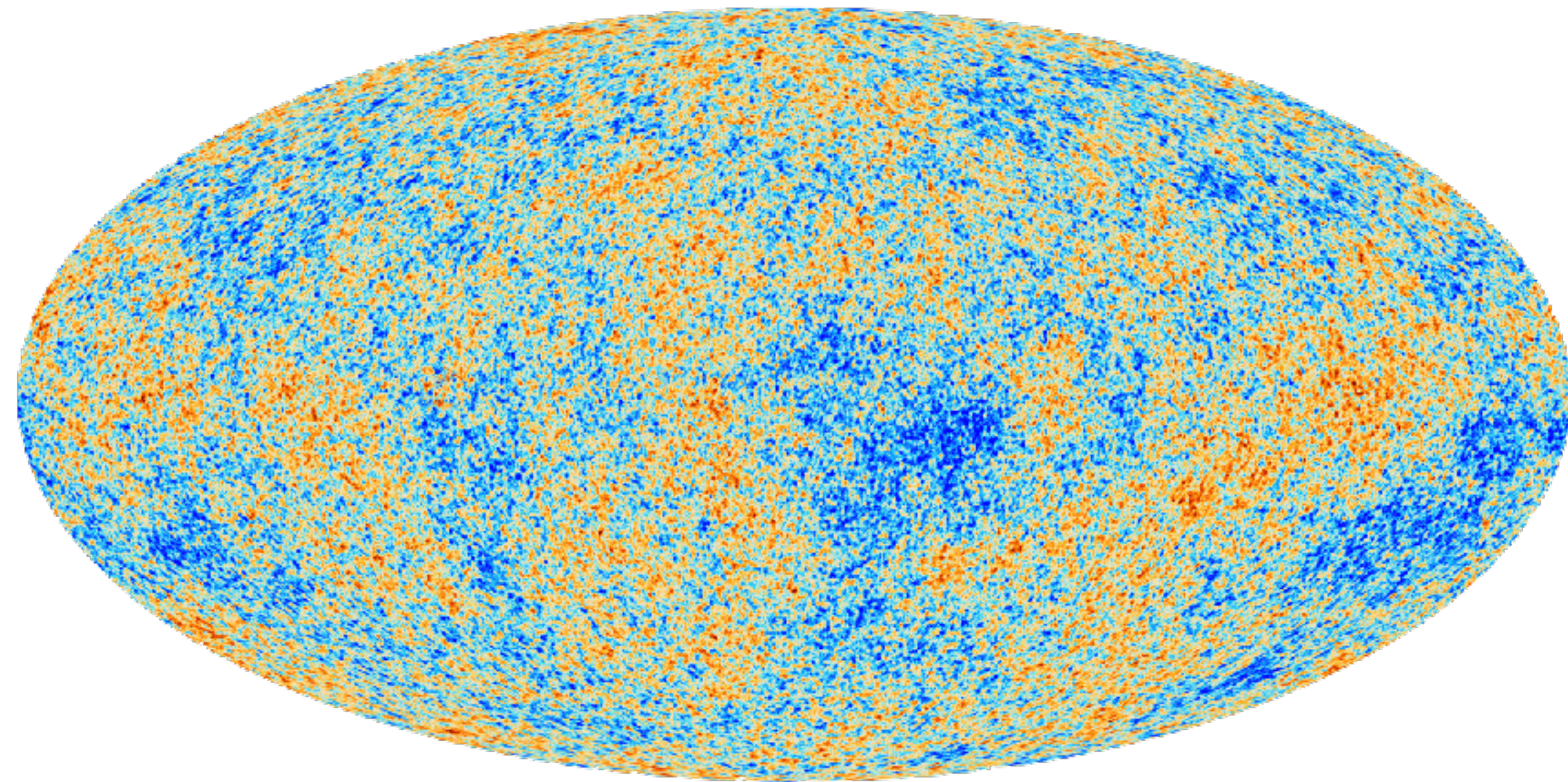
IPTA (2309.00693)

♦ Possible future of US pulsar timing: Deep Synoptic Array (DSA-2000)



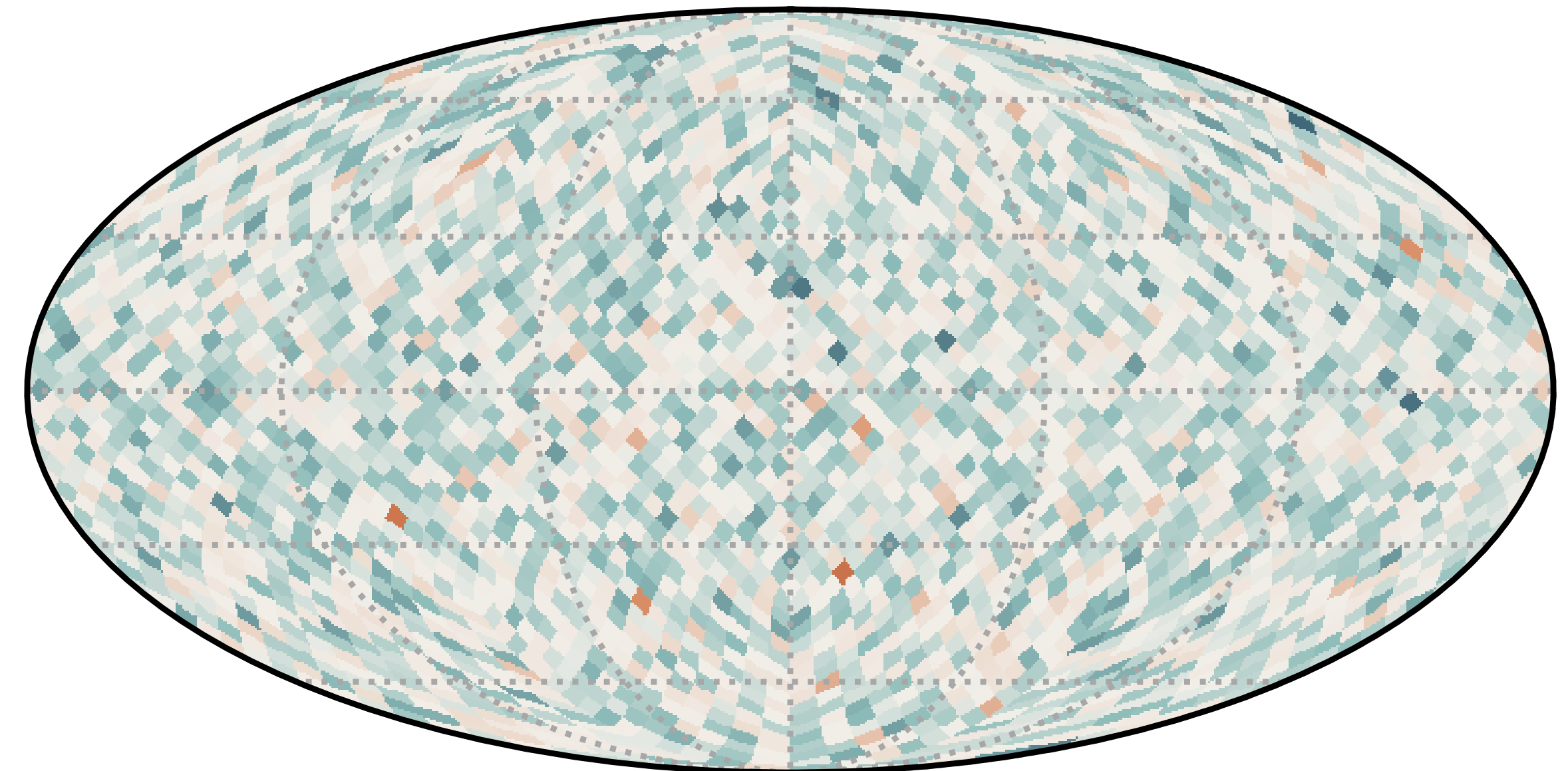


Cosmic Microwave Background



(observed)

Gravitational Wave Background



(simulated)

We have a new window into astrophysics and early Universe cosmology!

More data is incoming and more work needs to be done to extract possible primordial signals.



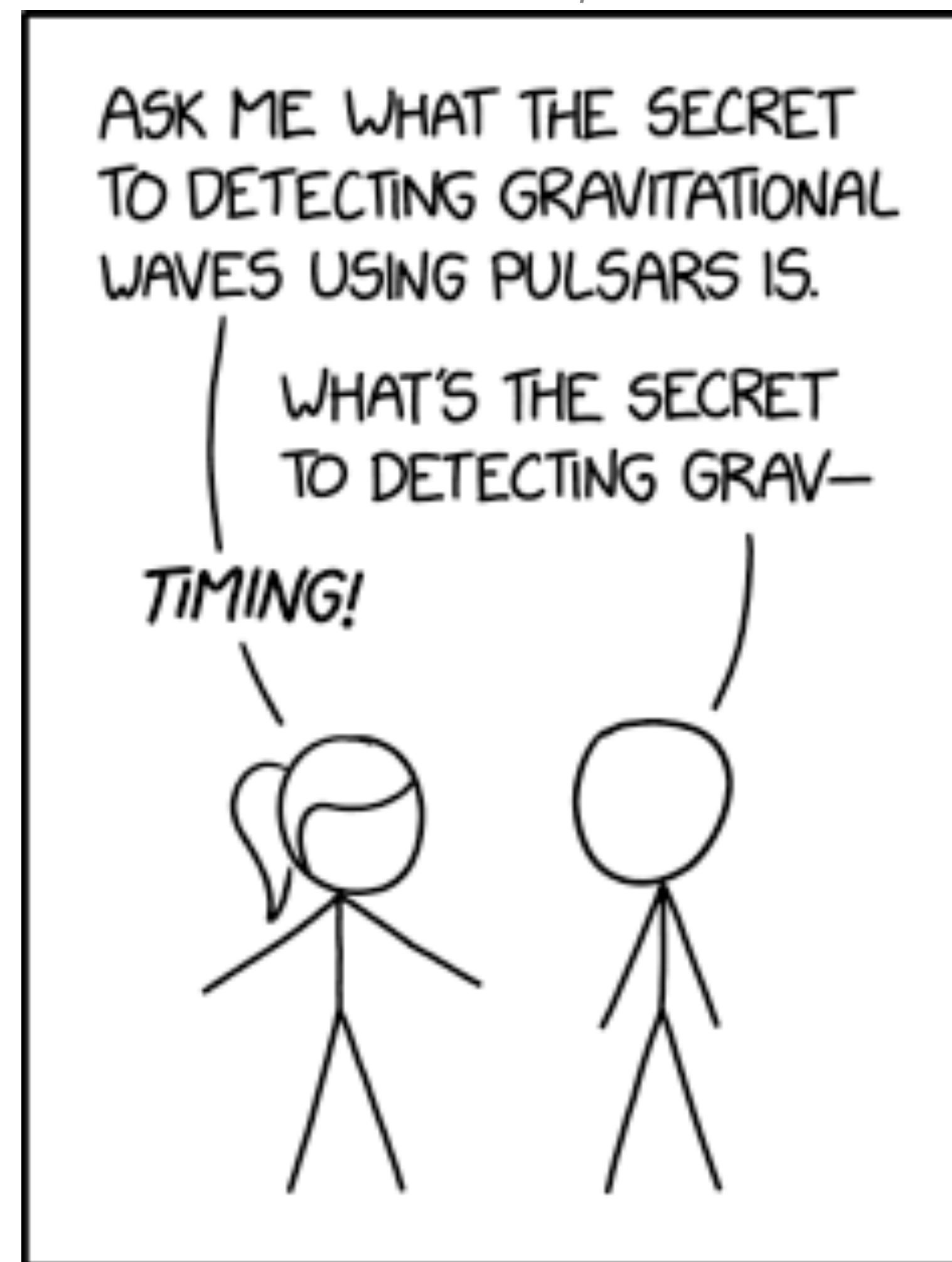
Natural Sciences and Engineering
Research Council of Canada

Canadian Institute for
Advanced Research



Gordon and Betty Moore Foundation

<https://xkcd.com/2358/>



The most important attributes of a vector in 3-space are {Location, Location, Location}