Probing the Dark Sector with Accelerator Experiments

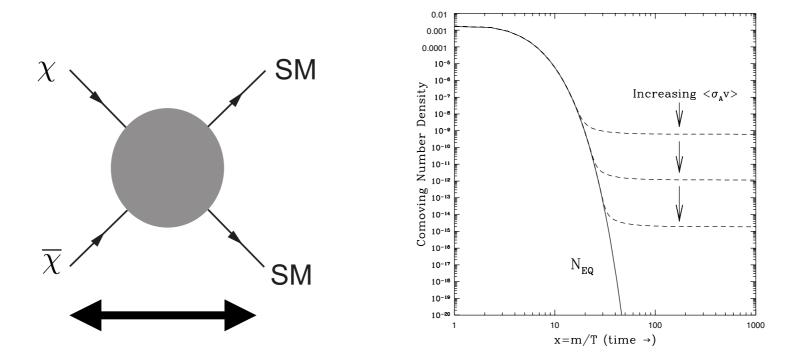
Brian Batell University of Pittsburgh



TeVPA 2024 August 26-30, 2024

Motivation: dark matter as a thermal relic

- Basic idea: dark matter produced from reactions in the plasma during the Big Bang
- Requires non-gravitational dark matter interactions with the Standard Model



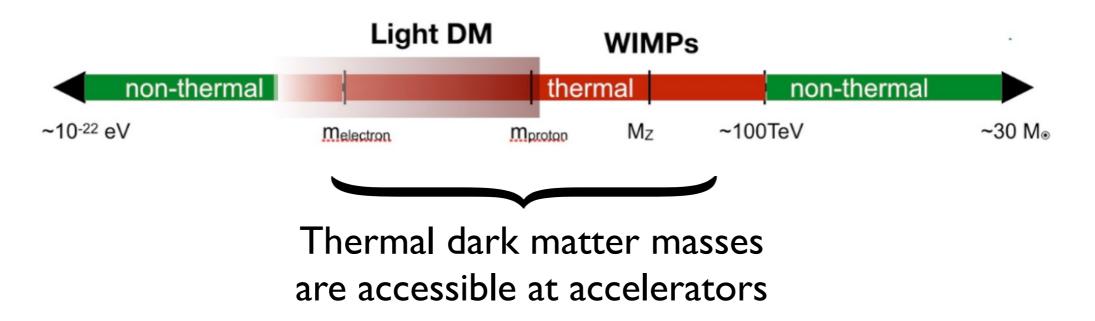
• Viable DM mass range between MeV - 10 TeV in simplest scenarios

	Light DM		WIMPs					
non-thermal			therma	al		non-thermal		
~10 ⁻²² eV	Melectron	mpr	oton	Mz	~100	TeV	~30	D M⊚

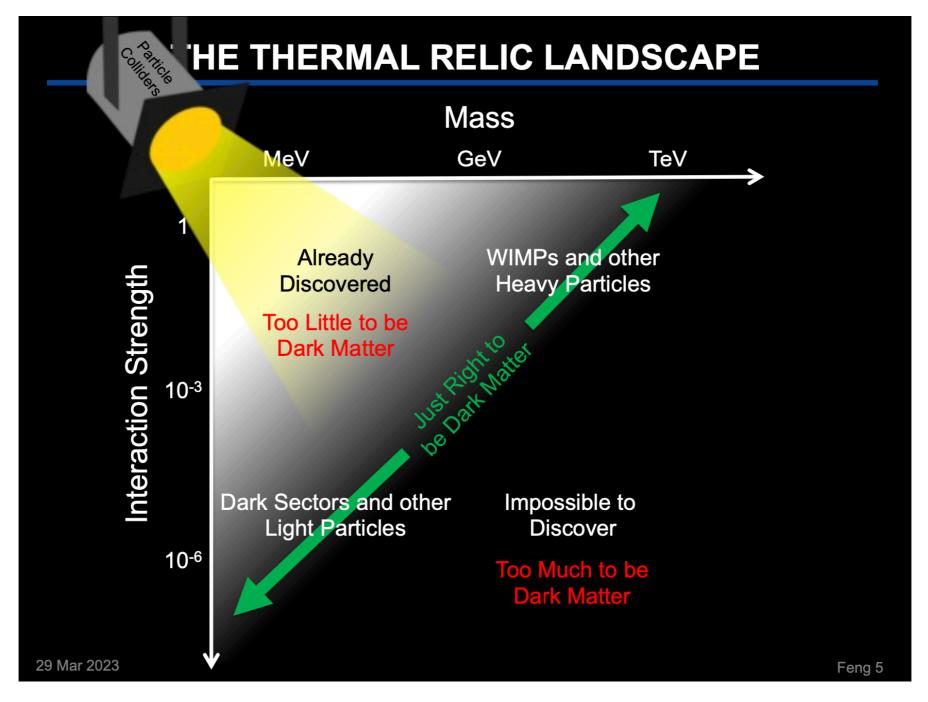
See next talk by J. Ruderman for variations of thermally produced dark matter

Thermal dark matter at accelerator experiments

- Basic idea: collide Standard Model particles produce dark matter
- Same interaction governs DM annihilation and laboratory DM production



- Upper bound: $m_{\rm DM} \sim {\rm TeV}$
 - Kinematic limit, maximum collider center-of-mass energy
- Lower bound: $m_{\rm DM} \sim {\rm MeV}$
 - Lighter DM typically strongly constrained by astrophysics and cosmology (stellar energy loss, modifications to nucleosynthesis, ...)



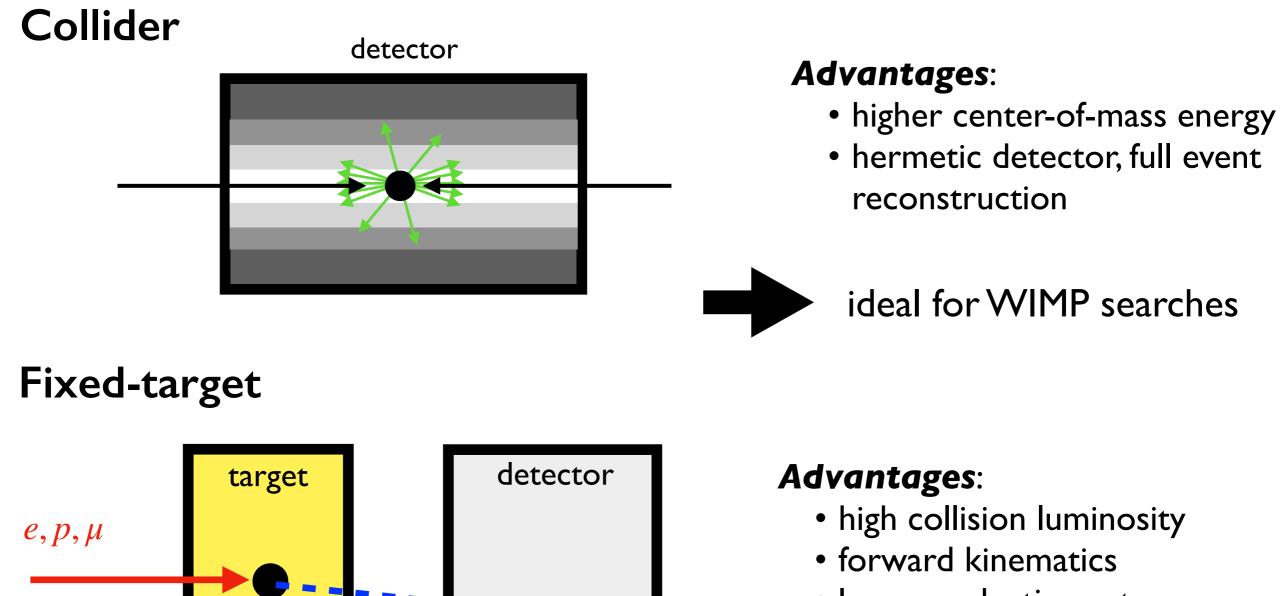
[[]Figure from J. L. Feng]

Non-gravitational coupling strengths required for thermal dark matter are accessible at accelerator experiments

Why accelerator experiments?

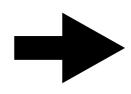
- Produce and study dark matter in a controlled laboratory setting
- Measure dark matter properties (mass, spin, couplings,...)
- Unlike direct and indirect detection, accelerator searches are not dependent on astrophysical assumptions
- Probe a variety of dark matter interaction channels
- Produce additional dark states associated with dark matter and probe the structure of the dark sector
- Dream scenario: reconstruct the dark matter thermal history!

Accelerator experiments: collider vs. fixed target



dark particle

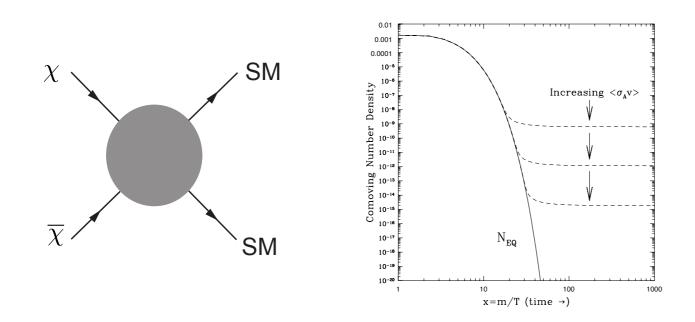
- large production rates
- clean detector environment



ideal for light sub-GeV dark matter searches

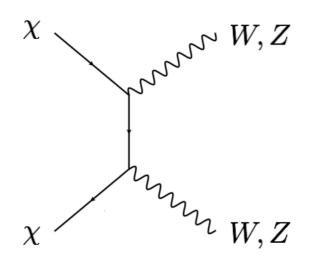
WIMPs

• "WIMP miracle" suggests dark matter mass scales in vicinity of weak scale

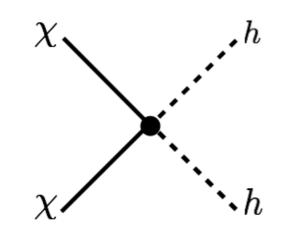


$$\langle \sigma v \rangle \sim \frac{\pi \alpha_W^2}{m_\chi^2} \sim 1 \,\mathrm{pb} \times \left(\frac{\alpha_W}{(1/30)}\right)^2 \left(\frac{\mathrm{TeV}}{m_\chi}\right)^2$$

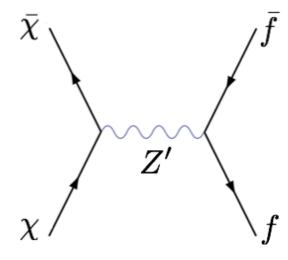
• Numerous model realizations



Electroweak DM



Higgs portal



BSM mediator (Z', sfermion, etc.)

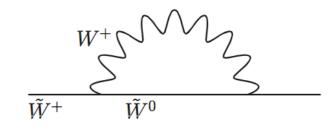
Electroweak dark matter

- Neutral component of $SU(2)_L \times U(1)_Y$ multiplet
 - Higgsino, Wino in MSSM
 - Minimal Dark Matter [Cirelli, Fornengo, Strumia]
- (co-) Annihilation to weak gauge bosons; observed relic abundance achieved for
 - ~ I TeV Higgsino and ~ 3 TeV Wino
- Indirect probes (gamma rays) are placing relevant constraints on electroweak dark matter
 [Cohen, Lisanti, Pierce, Slatyer; Fan, Reece; +more recent studies]
 See talks by T. Slatyer, M. Baumgart, W. L. Xu
- SI nuclear scattering suppressed (loop level + accidental cancellation)

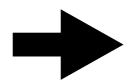
 $\sigma_n \sim 10^{-47} - 10^{-49} \ {\rm cm}^2 \quad \begin{array}{l} \mbox{[Hisano, Ishiwata, Nagata]} \\ \mbox{[Hill, Solon]} \end{array}$

• Radiative mass splitting of electroweak multiplet:

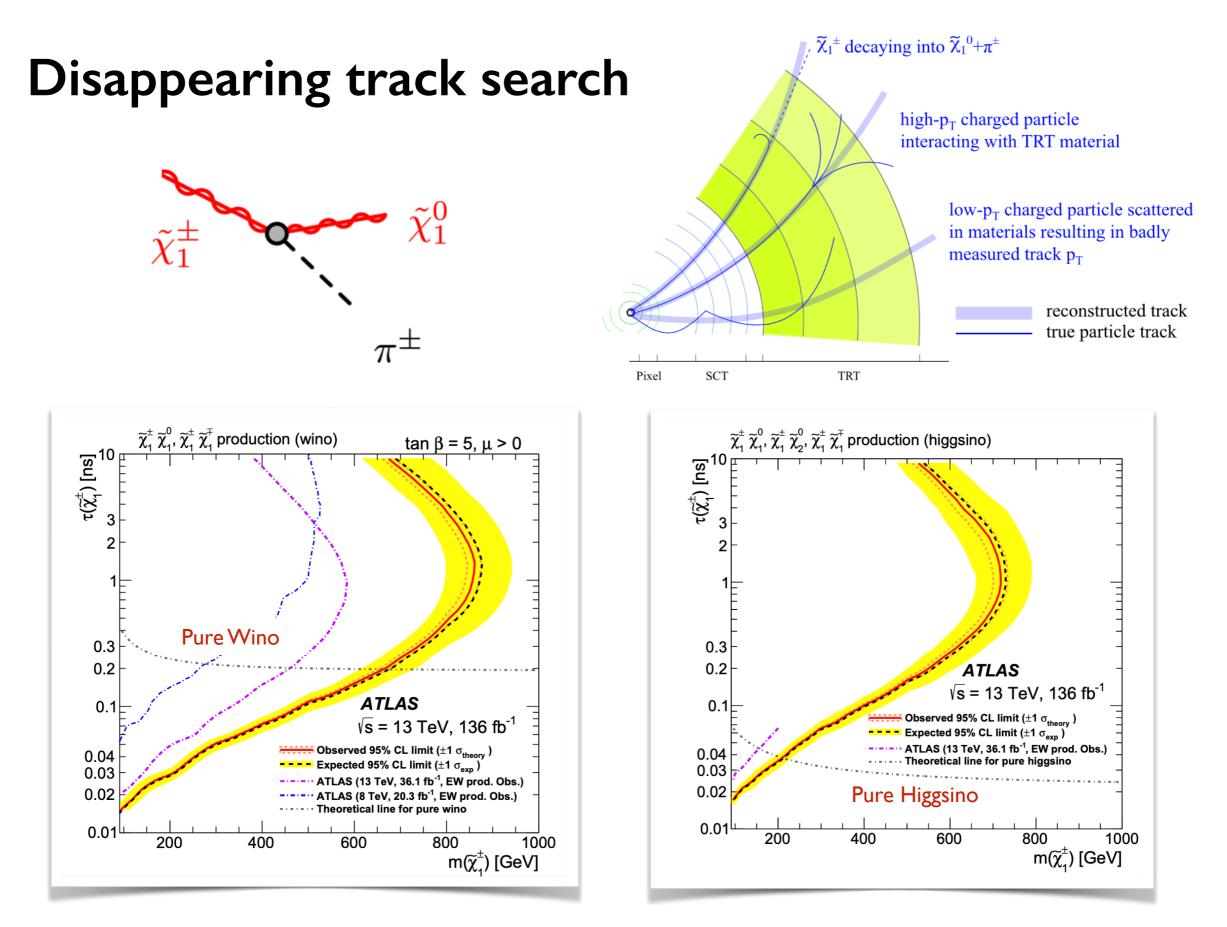
$$M_{\pm} - M_0 \approx \begin{cases} 300 \,\mathrm{MeV} & (\mathrm{Higgsino}) \\ 160 \,\mathrm{MeV} & (\mathrm{Wino}) \end{cases}$$



[Thomas, Wells; Ibe, Matsumoto, Sato]

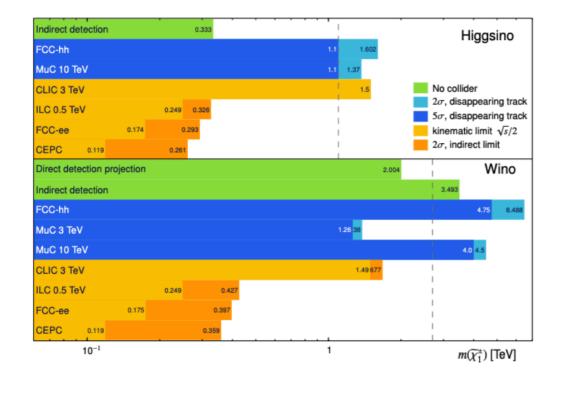


renders charginos long-lived, visible decay products soft

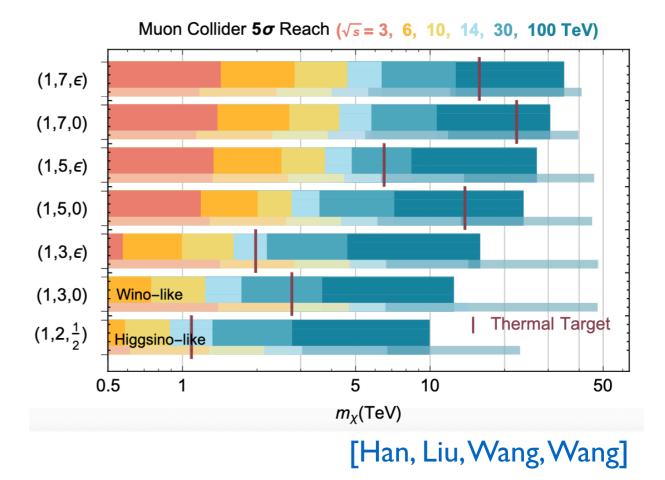


LHC has pushed beyond LEP limits, but thermal target out of reach

Electroweak DM at future colliders

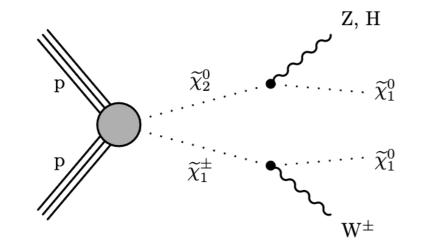


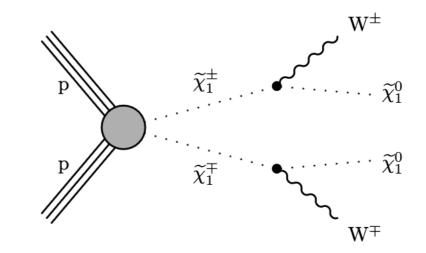
[Capdevilla, Meloni, Simoniello, Zurita]

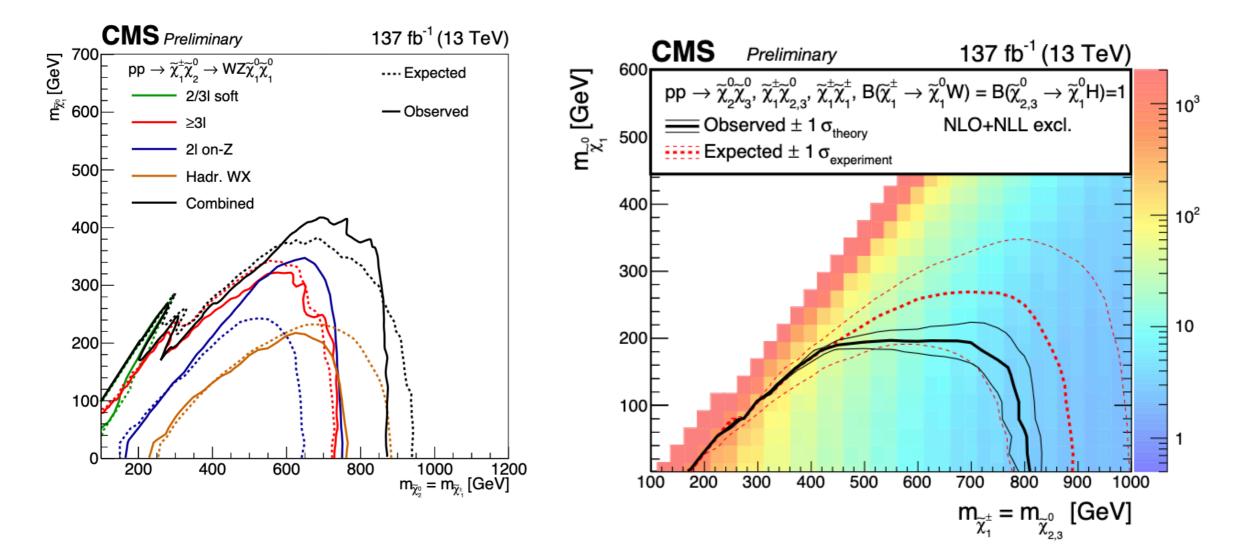


Both FCC-hh and a high-energy muon collider have the potential to test the Higgsino and Wino thermal targets

Electroweakino searches







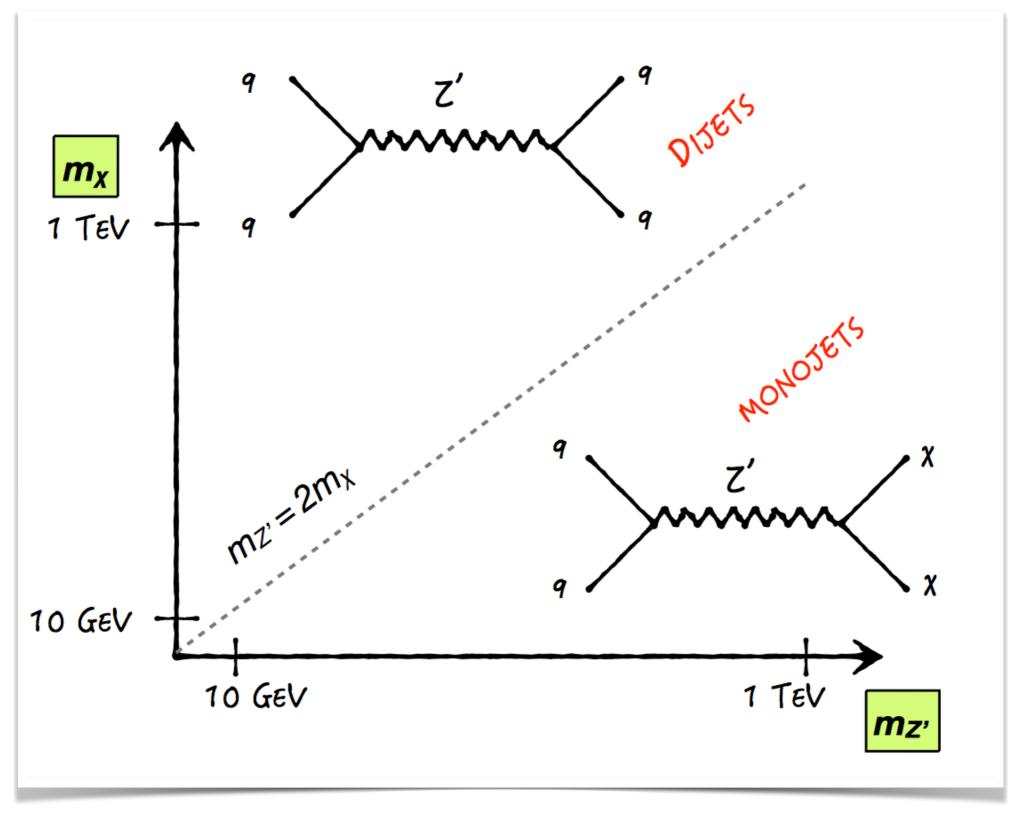


- Bottom up approach (as opposed to SUSY DM)
- EFT approach led to Mono-jet and a host of other Mono-X signatures

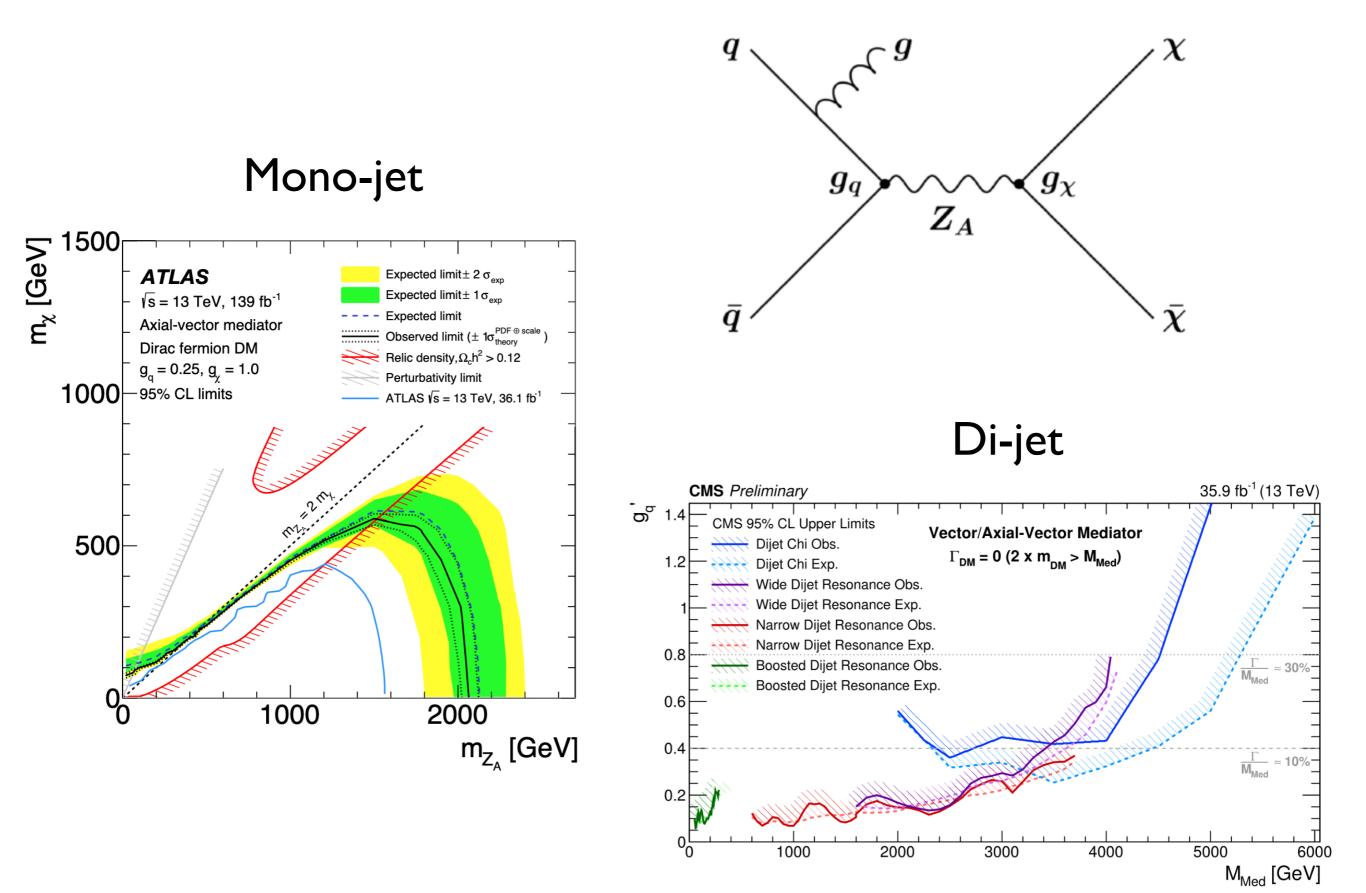
[Beltran, Hooper, Kolb, Krusberg, Tait; Goodman, Ibe, Rajaraman, Shepherd, Tait, Yu; Bai, Fox, Harnik; + many other studies]

- Concerns regarding validity of EFT, unitarity violation, etc. prompted exploration of simplified models
 - Different kinematics, adjustments to search strategies, and new signals/searches for mediators
- Explorations of UV completions of simplified models additional model-dependent phenomenology (but often different from SUSY DM)

Monojet vs. Dijet



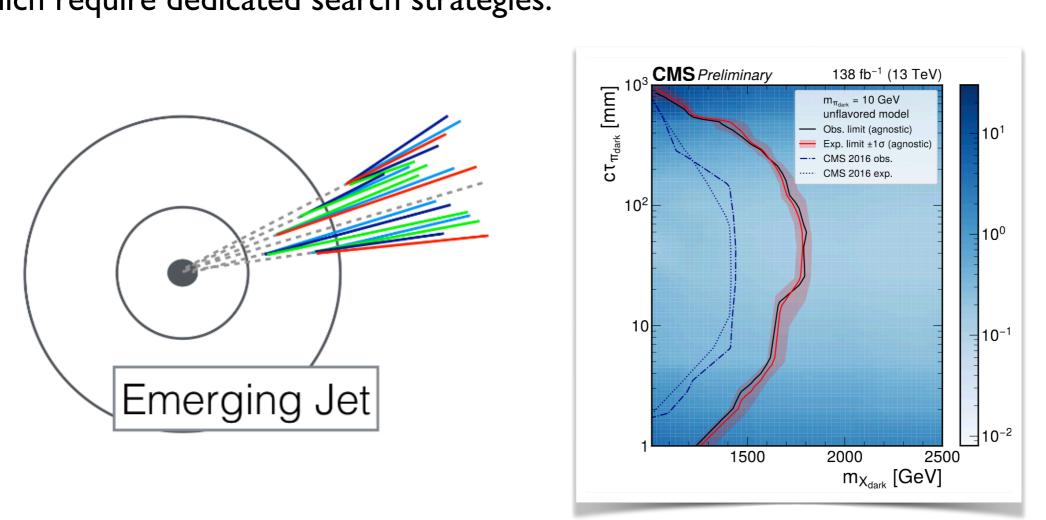
[Figure from N. Saoulidou]



Dark Showers, Emerging Jets, Semi-visible Jets, ...

[Schwaller, Stolarski, Weiler; Cohen, Lisanti, Lou; + others]

- Basic scenario: dark matter is part of a confining dark sector, which interacts with the Standard Model via a portal
- This can lead to novel signatures at the LHC, such as emerging or semi-visible jet, which require dedicated search strategies.



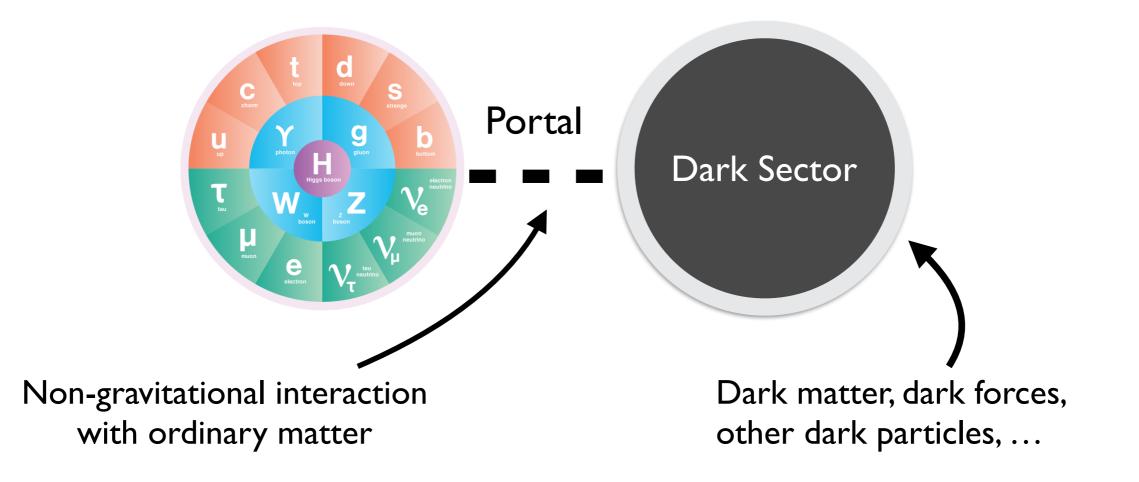
• This is an active field with opportunities for novel searches!

[For an overview, see the Dark Showers Snowmass report, arXiv:2203.09503]

Light dark matter and dark sectors

- Light (sub-GeV) dark sectors are an interesting framework for dark matter
- Relic abundance may be generated via thermal freezeout
- Extension of the WIMP below Lee-Weinberg bound Requires new interactions beyond weak interaction

[Boehm, Fayet] [Pospelov, Ritz, Voloshin] [Feng, Kumar]



Dark Sectors and Portals

• There are three minimal renormalizable portals that connect the visible and dark sectors

$$\begin{array}{ll} \displaystyle \frac{\epsilon}{2\cos\theta_W}F'_{\mu\nu}B^{\mu\nu} & \mbox{Vector Portal} \\ \displaystyle \left(A\,S+\lambda\,S^2\right)H^\dagger H & \mbox{Higgs Portal} \\ \displaystyle yNLH & \mbox{Neutrino portal} \end{array}$$

- There are other interesting options for the mediator:
 - Higher dimension portals, e.g. axion-like particle; anomaly free gauge bosons (e.g., $B L, L_{\mu} L_{\tau}$), etc.
 - In some cases, these mediators come with additional motivations (e.g., neutrino masses, heavy QCD axion, explanations of experimental anomalies, ...)
- The dark sector itself can be minimal or have a rich structure (e.g., dark higgs, inelastic dark matter, dark confinement, ...)

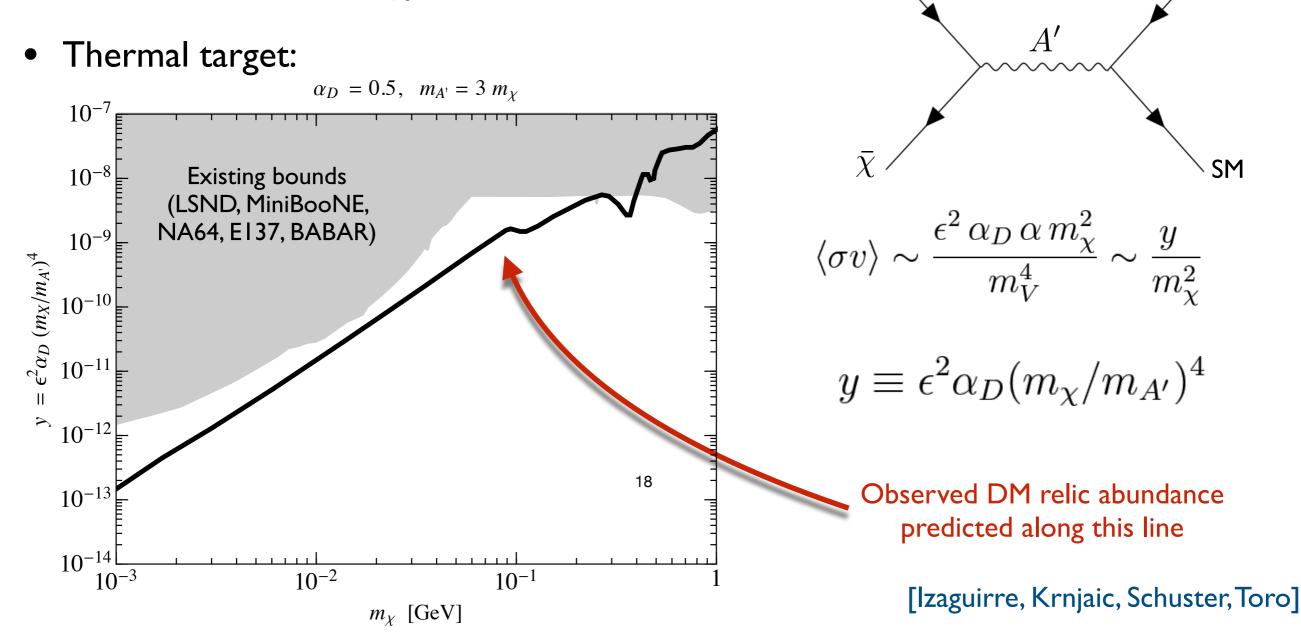
Benchmark model: Vector Portal Dark Matter

$$\mathcal{L} \supset |D_{\mu}\chi|^{2} - m_{\chi}^{2}|\chi|^{2} - \frac{1}{4}(F_{\mu\nu}')^{2} + \frac{1}{2}m_{A'}^{2}(A_{\mu}')^{2} - \frac{\epsilon}{2}F_{\mu\nu}'F^{\mu\nu} + \dots$$

[Holdom] [Pospelov, Ritz, Voloshin]

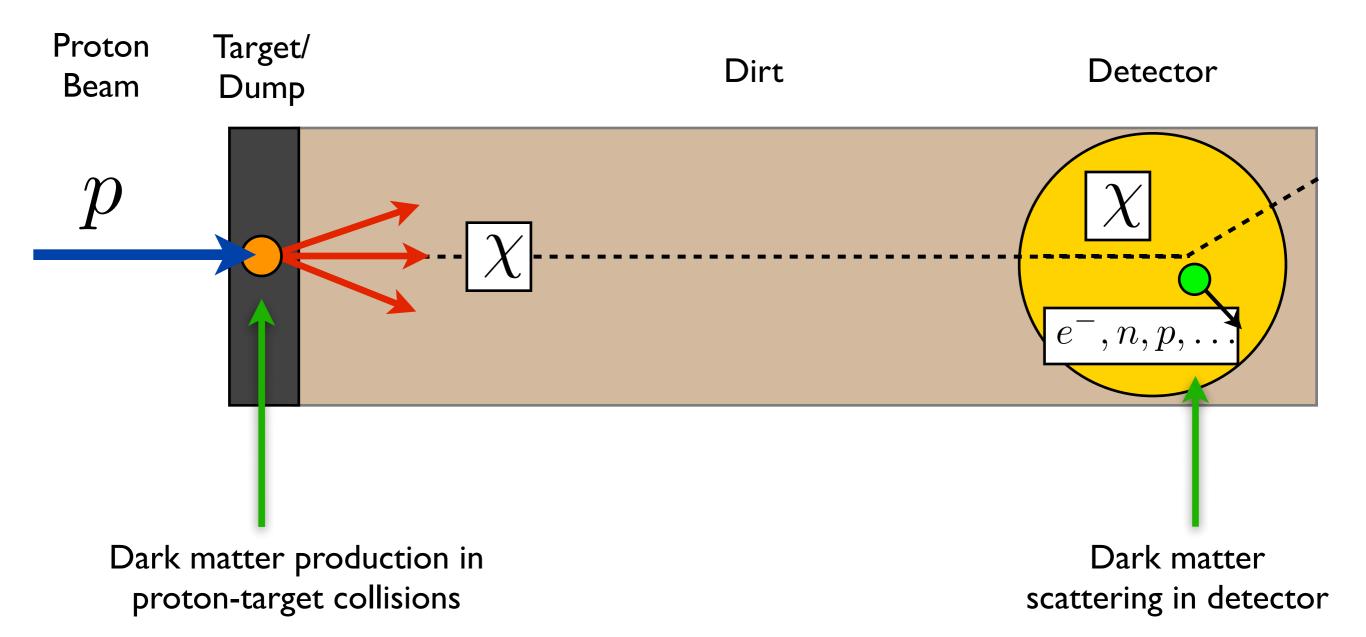
SM

- Dark photon mediates interaction between DM and SM
- 4 new parameters: $m_{\chi}, m_{A'}, lpha_D, \epsilon$

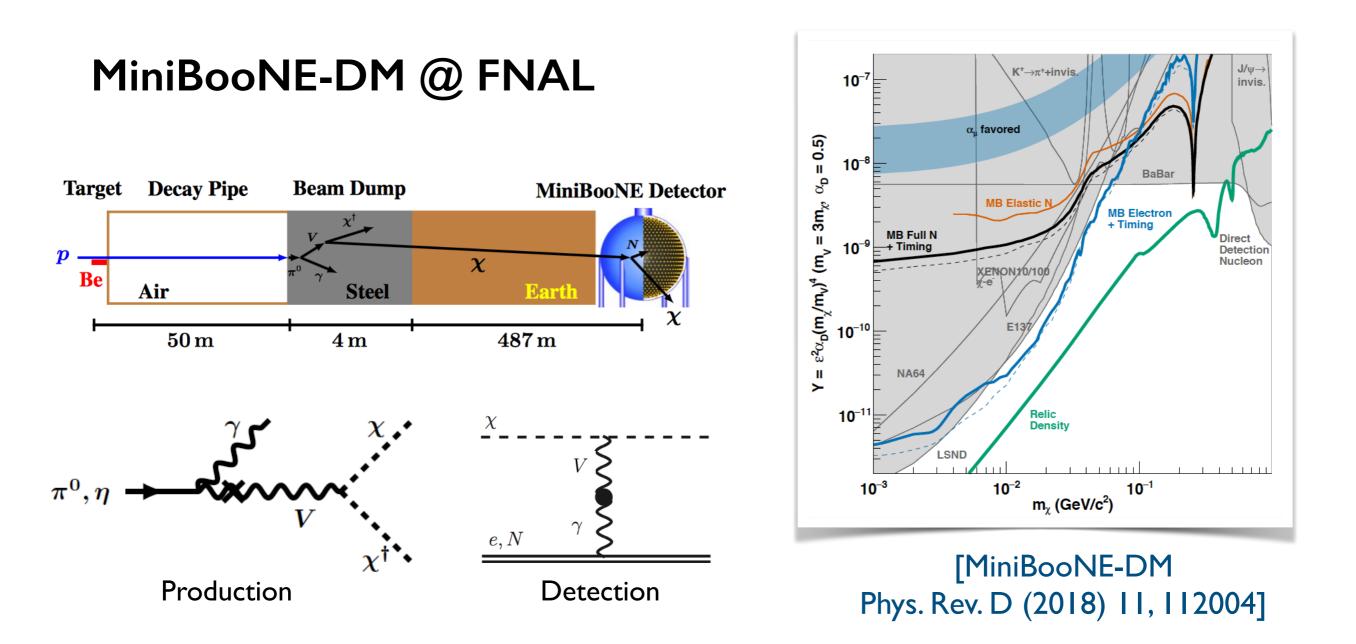


Beam Dump Search for Dark Matter

[BB, Pospelov, Ritz]



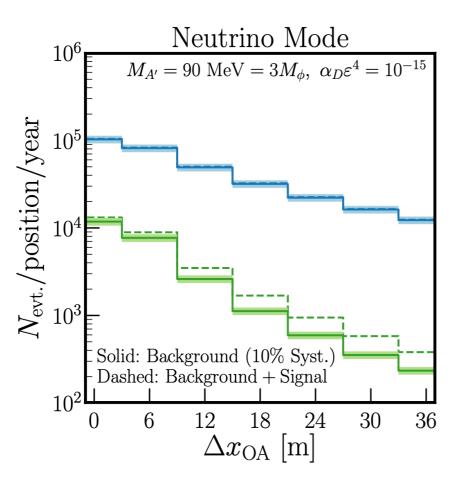
- Unique sensitivity for many models with light dark matter + light mediator
- Can be done with existing and near future accelerator neutrino experiments
 - MiniBooNE, NOvA, T2K, MicroBooNE, SBND, ICARUS, DUNE...



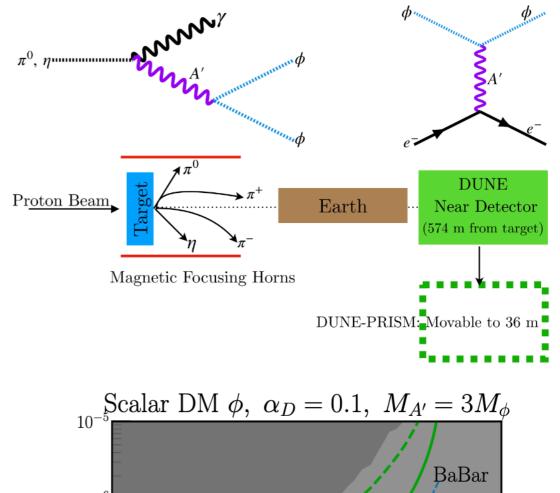
- 8 GeV protons on iron dump; 800 ton mineral oil detector
- Dedicated off target / beam dump run mode, collected 1.9E20 POT
- Leading limits on vector portal dark matter model for ~ 100 MeV mass range
- Demonstrates proton beam dump as an effective search method for light dark matter

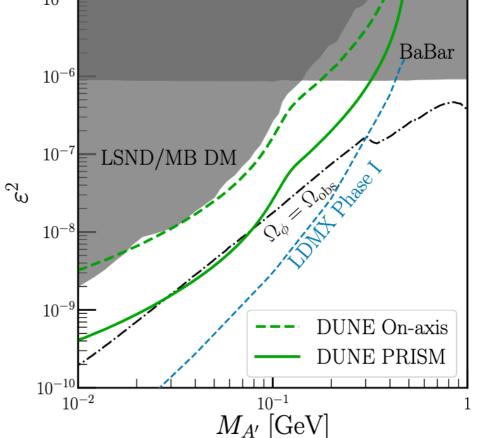
DUNE-PRISM @ FNAL

- 120 GeV protons on graphite target
- DUNE-PRISM movable near detector allows sensitive search to light dark matter
- DM-to-neutrino flux increases as detector is moved off axis
 - Neutrinos produced through decay of charged mesons, which are focused by magnetic horn
 - DM produced through decay of unfocused neutral mesons



[De Romeri, Kelly, Machado] [Breitbach, Buonocore, Frugiuele Kopp, Mittnacht]





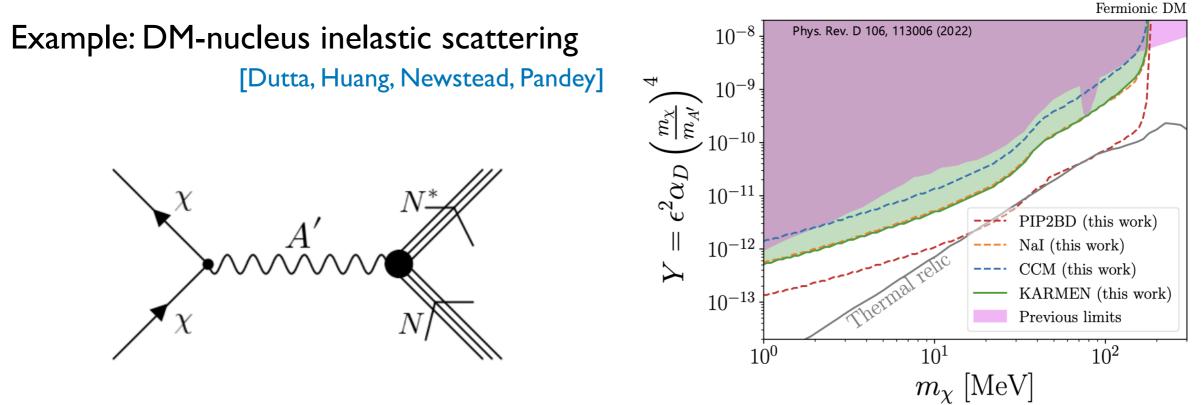
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Opportunities with the FNAL Proton Improvement Project 2 (PIP-2)

- As part of FNAL PIP-2 upgrade, Booster will be replaced, Main Injector will be upgraded
- Excess protons at ~1 GeV, ~10 GeV, 120 GeV will be potentially available for a variety of physics applications, including dark sector studies.

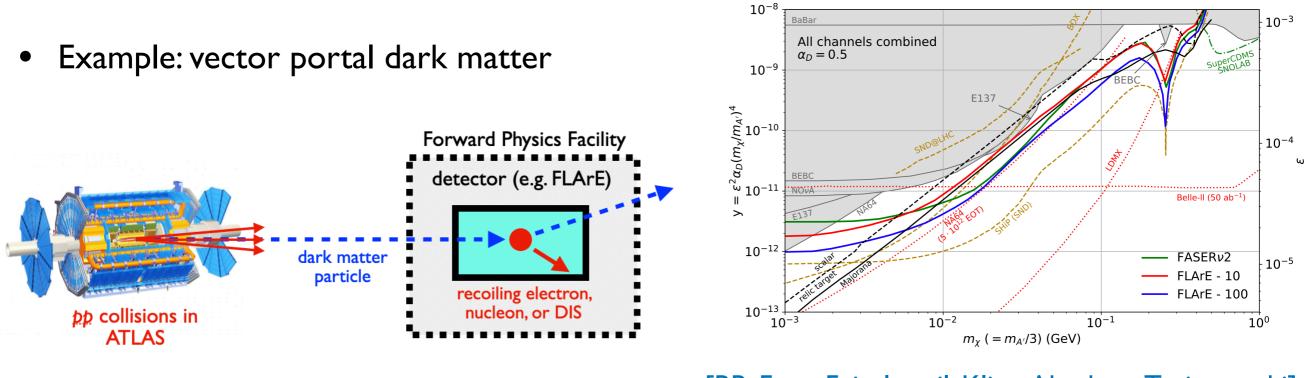
Fermilab Facility for Dark Sector Discovery (F2D2) See talk by M.Toups

- 0.2-2 GeV proton beam dump facility for dark sector searches, 10²²-10²³ POT/year
- Physics opportunities: sub-GeV DM, ALPs, CEvNS studies, light sterile- ν ...



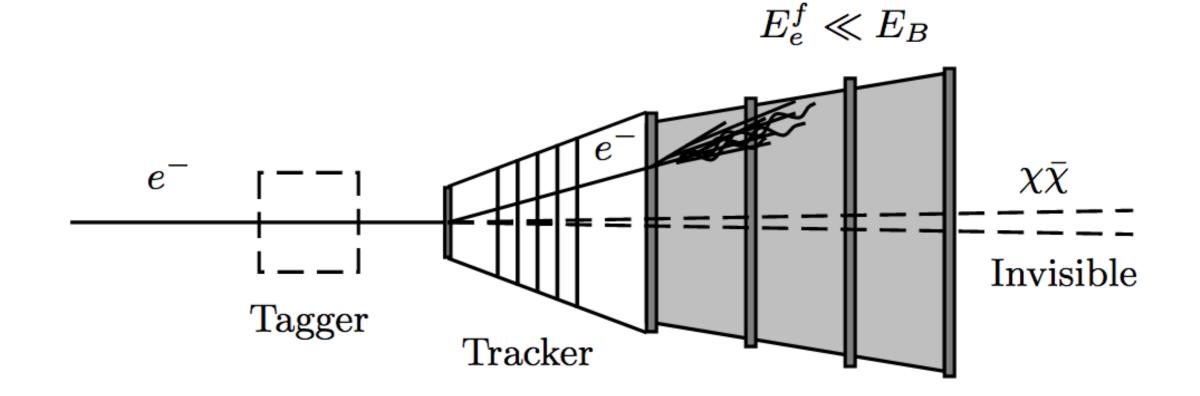
Forward LHC experiments (FASER, FASER ν , FLArE, ...)

- Total LHC pp cross section is ~100 mb, and is directed in the forward region
 - Copious source of TeV energy neutrinos
 - First collider-produced neutrinos detected by FASER ν [arXiv:2105.06197]
 - Exciting prospects at FASER ν , SND@LHC (Run 3) and FASER ν 2, FLArE, FORMOSA (HL-LHC)
- Dark sectors can also be explored with forward LHC experiments
 - For full physics case, see Forward Physics Facility whitepaper [arXiv:2203.05090]



[BB, Feng, Feig, Ismail, Kling, Abraham, Trojanowski] [2101.10338, 2107.00666, 2111.10343]

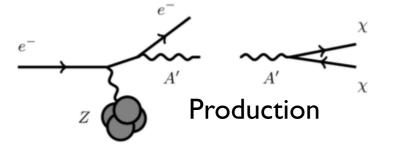
Missing energy/momentum searches for dark matter

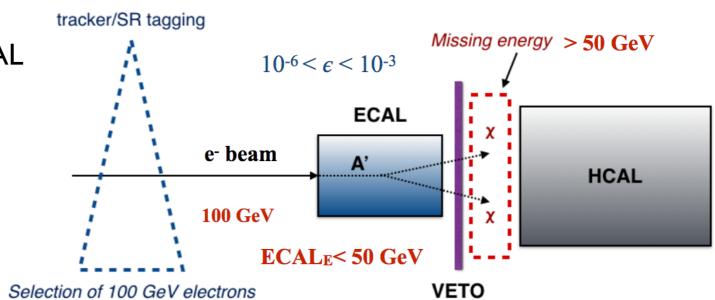


[Andreas. et al] [Izaguirre, Krnjaic, Schuster, Toro]

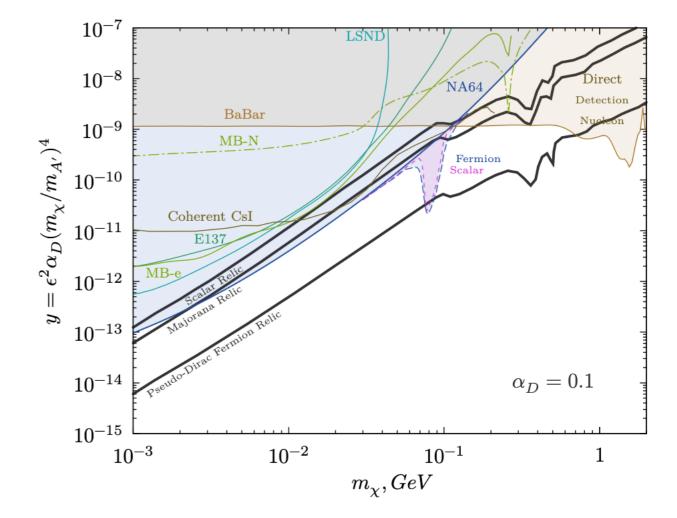
NA64@CERN

- 100 GeV electron beam incident on ECAL
- Dark matter produced in ECAL and carries most of the beam energy



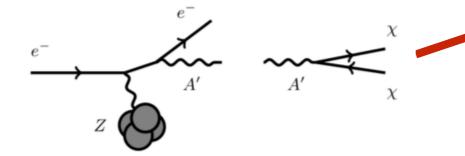


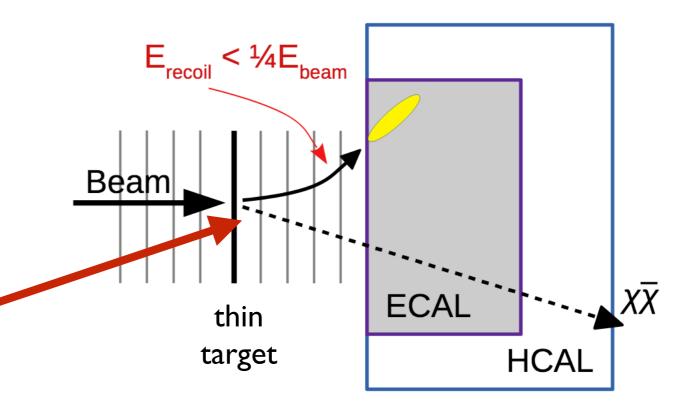
- Large missing energy signature (small energy deposition in ECAL, no energy deposition in HCAL)
- ~10¹² EOT best limits on vector portal dark matter below 300 MeV, starting to probe thermal relic targets



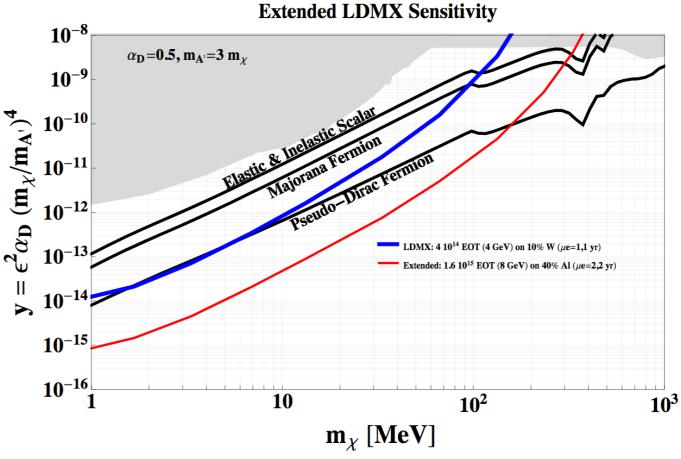
LDMX@SLAC

• Proposed electron beam experiment utilizing missing momentum technique



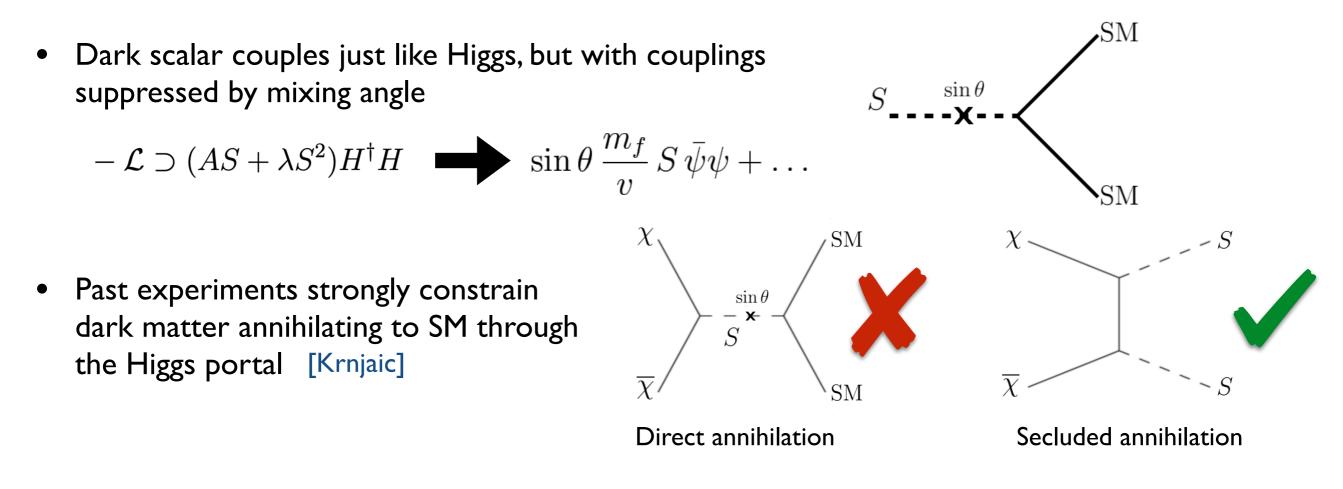


- More kinematic handles to reject backgrounds, discriminate final state electrons from photons 10^{-8} 10^{-9} 10^{-10} 10^{-10} 10^{-10}
- Can cover most thermal targets, irrespective of DM particle nature

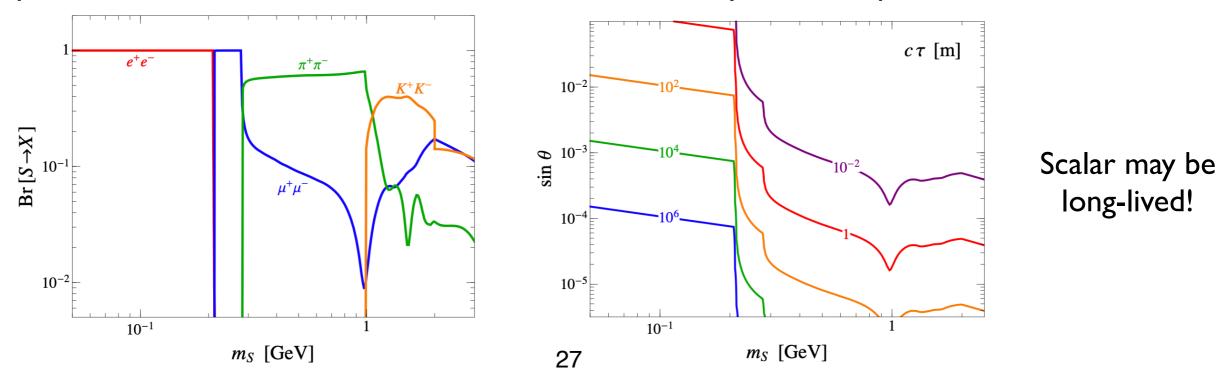


Initial design study, 1808.05219

Visible Signatures of the Higgs Portal

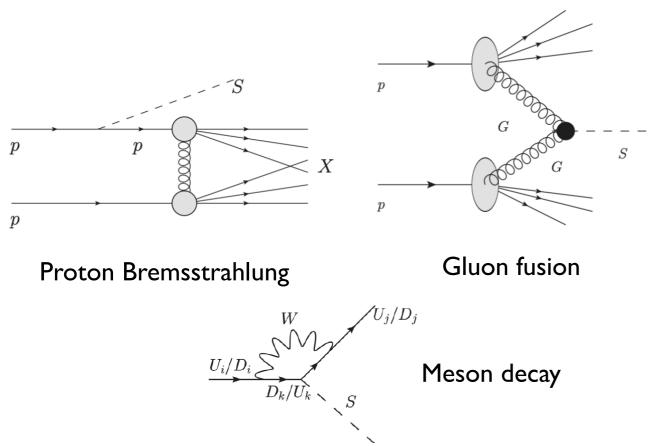


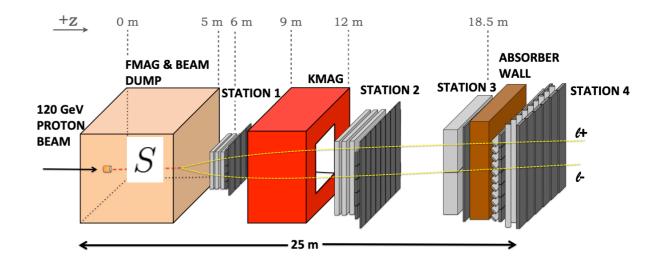
• Expect dark matter is heavier than scalar and scalars decay to visible particles



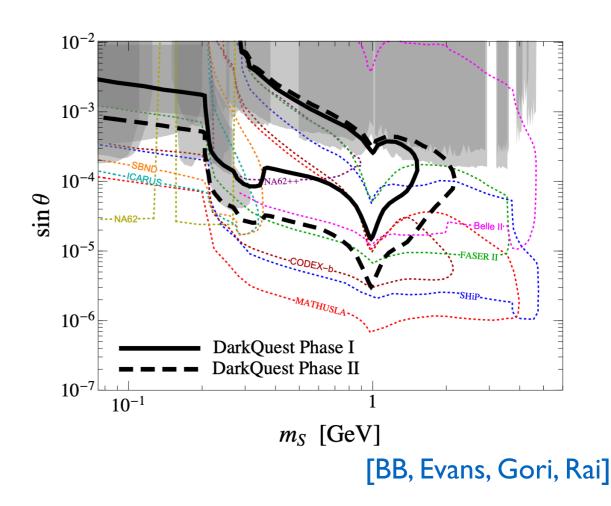
DarkQuest@FNAL

- I20 GeV protons from the Fermilab Main Injector impinge on ~ 5m iron beam dump
- Magnetic field (KMAG), 4 tracking stations, muon ID system, EM calorimeter
- High energies, short baseline allows to probe long-lived particles with moderate lifetimes
- Example: visibly decaying Higgs portal scalar





[Berlin, Gori, Schuster, Toro; Blinov, Kowalczyk, Wynne; Forbes, Herwig, Kahn, Krnjaic, Suarez, Tran, Whitbeck; + others]



Short Baseline ν -Experiments @ FNAL

- MicroBooNE, SBND, ICARUS LArTPC detectors
- Situated along 8 GeV Booster beam line and slightly off axis from 120 GeV NuMi beam line
- Will collect ~ 10²¹ POT over next several years
- These experiments have sensitivity to a variety of dark sector models
- Example: MicroBooNE search for Higgs portal scalar

Hadron

sorber

e⁺e⁻

Decay volume

MicroBooNE

 ℓ,π

 ℓ,π

 π

Top view

BNB

Side view

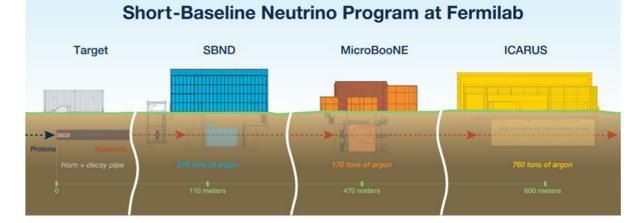
NuMI target

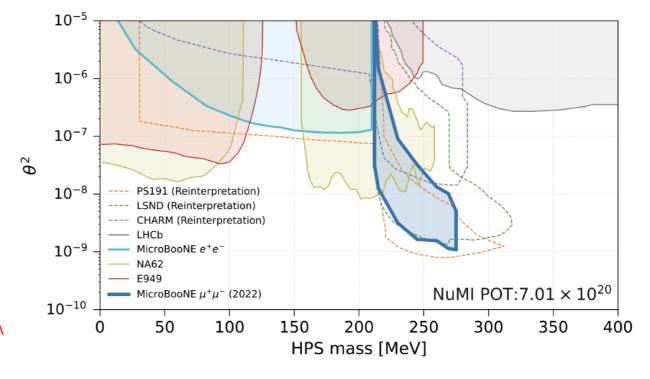
and horns

K

S

c,t



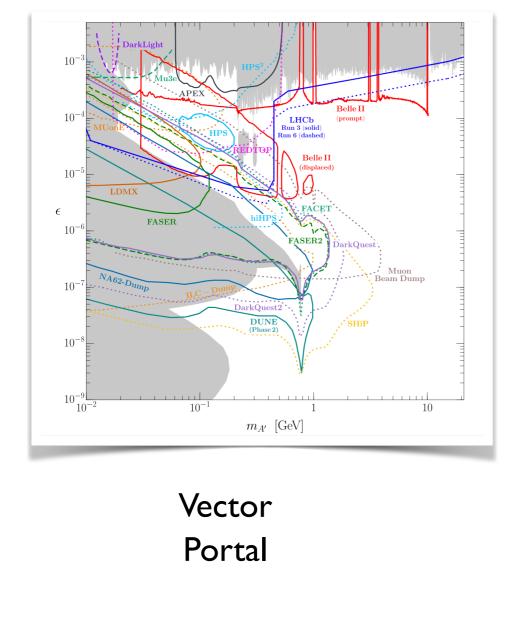


See talk by Lee Hagaman

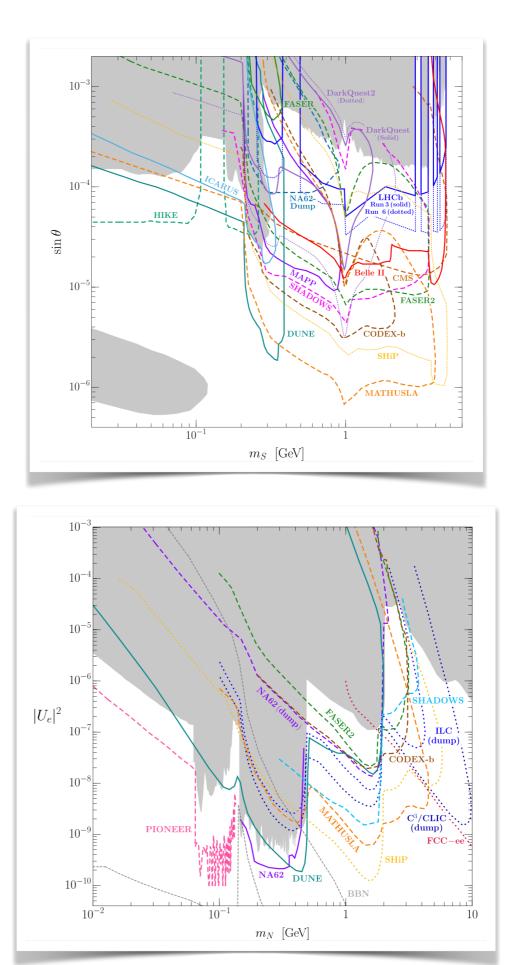
[See also BB, Berger, Ismail, 1909.11670, for prospects at ICARUS and SBND]

NuMI

Visible Portal Benchmark Models



[Snowmass RF6 whitepaper, 2207.06905]



Higgs Portal

Neutrino Portal

Other dark sector models

- I have mainly focused on the simplest vector and Higgs portal scenarios. There are many other interesting dark sector models.
 - Different particle nature/content/dynamics of the dark sector

Spin, coupling structure, excited dark matter states, strongly interacting dark sector sector, spontaneous symmetry breaking, ...

• Different mediator couplings to the SM

Leptophobic, Leptophilic, flavor specific, mediator spin...,

• This leads to interesting alternatives for the cosmological production of dark matter and distinct phenomenological signals

[See the Snowmass RF6 reports for a comprehensive survey of these possibilities, 2207.06905, 2207.08990, 2209.04671]

Summary and Outlook

- Dark matter is a profound mystery!
 - Many candidates, mass scale is largely unconstrained, rich variety of phenomena. Lots of work ahead!
- Accelerators provide an important tool to search for and study dark matter and dark sectors.
 - Complementary to direct detection, indirect detection, astrophysics, cosmology
 - Can test certain thermal dark matter production scenarios
 - Can provide information about dark sector structure
- Still room for exploration in model and signature space, which will provide motivation for new searches at colliders and fixed target experiments
- Many exciting experiments, searches, and results on the horizon!