



LST
COLLABORATION



Detection of OP313 at $z=0.997$ with LST-1: the most distant VHE blazar to date

Mireia Nievas (IAC)
CTAO LST Collaboration
26.08.2024



Chicago 2024



Cofinanciado por
la Unión Europea

CTAO

Cherenkov Telescope Array Observatory.

*~1500 members (~150 institutes, 25 countries).
Aiming to detect $\sim 10^3$ VHE ($E > 100$ GeV)
gamma-ray sources.*

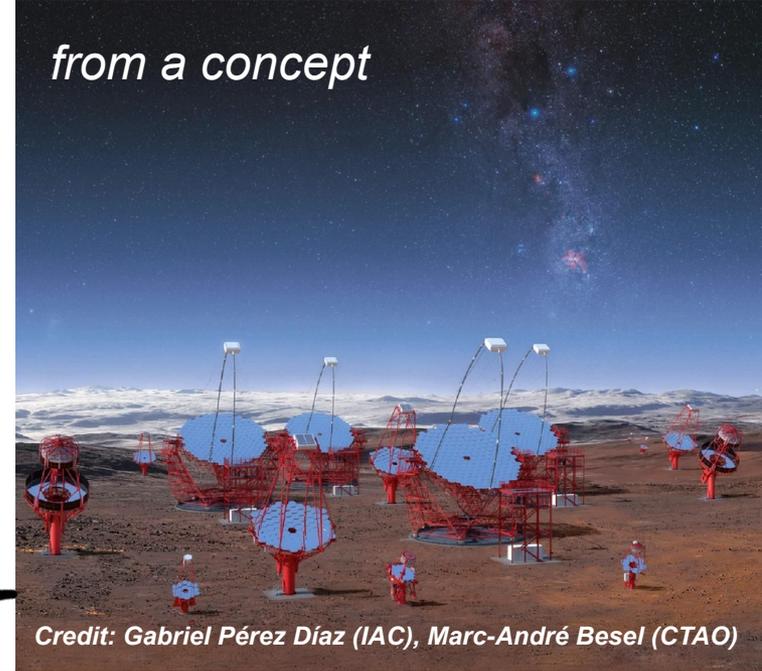
Two sites:

- CTAO-North: Roque de los Muchachos Observatory (La Palma, Spain)
- CTAO-South: Paranal Observatory (Atacama Desert, Chile)

Three telescope sizes

- Large-Sized Telescopes (LSTs)
- Medium-Sized Telescopes (MSTs)
- Small-Sized Telescopes (SSTs)

from a concept



Credit: Gabriel Pérez Díaz (IAC), Marc-André Besel (CTAO)

to reality



LST1

Dec. 2017

Credit: Giovanni Ceribella (MPP)

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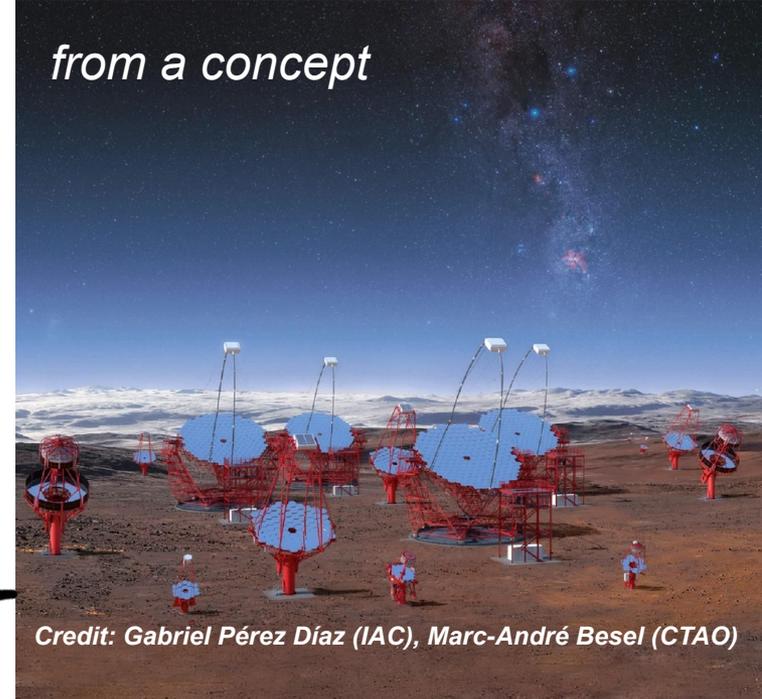
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Credit: Gabriel Pérez Díaz (IAC), Marc-André Besel (CTAO)

to reality



LST1

June. 2018

Credit: Daniel Mazin (ICRR)

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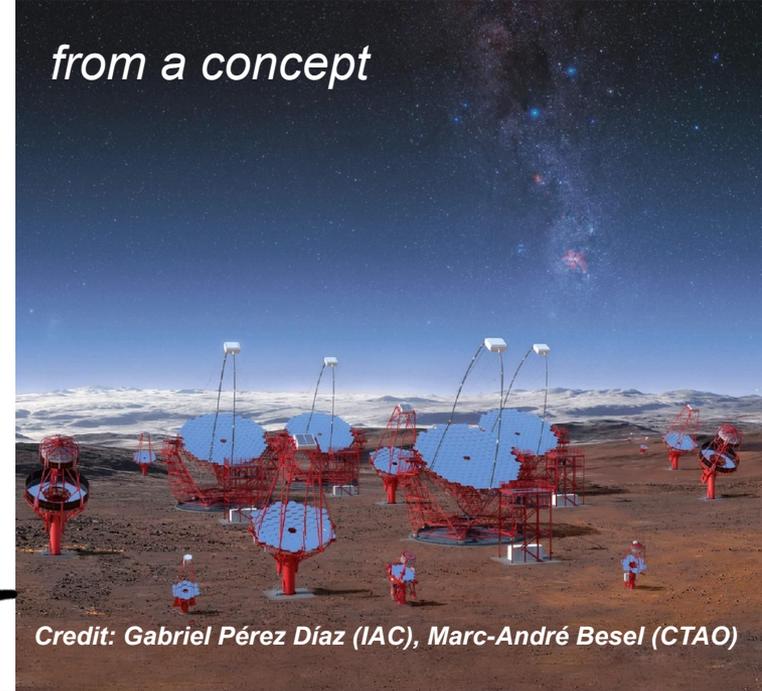
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Credit: Gabriel Pérez Díaz (IAC), Marc-André Besel (CTAO)



to reality



LST1

Sept. 2020

Credit: Mireia Nieves (IAC)

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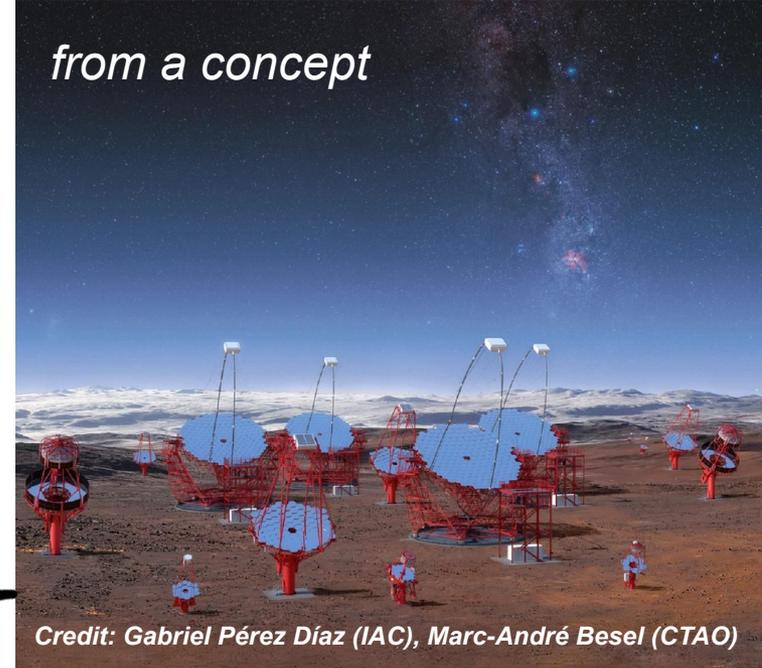
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from a concept



Credit: Gabriel Pérez Díaz (IAC), Marc-André Besel (CTAO)

to reality



LST4

May 2024

Credit: Alice Donini (INAF)

CTAO

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Aiming to detect $\sim 10^3$ VHE ($E > 100$ GeV)
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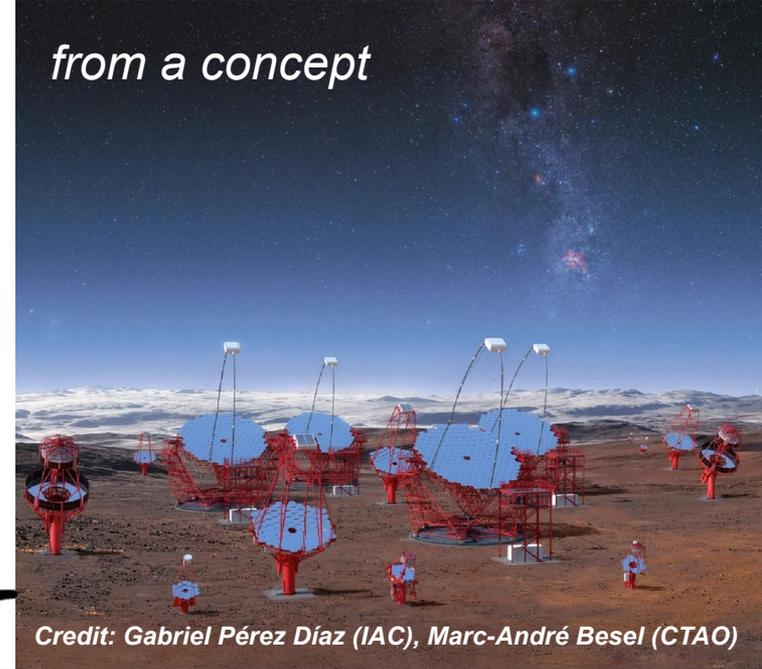
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from a concept



Credit: Gabriel Pérez Díaz (IAC), Marc-André Besel (CTAO)

to reality



LST3

June 2024

Credit: Mireia Nieves (IAC)

CTAO

Cherenkov Telescope Array Observatory.

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- Small-Sized Telescopes (SSTs)

from a concept



Credit: Gabriel Pérez Díaz (IAC), Marc-André Besel (CTAO)

to reality



LST2

Aug. 2024

Credit: Miguel Molero (IAC)



LST1



LST2



LST3



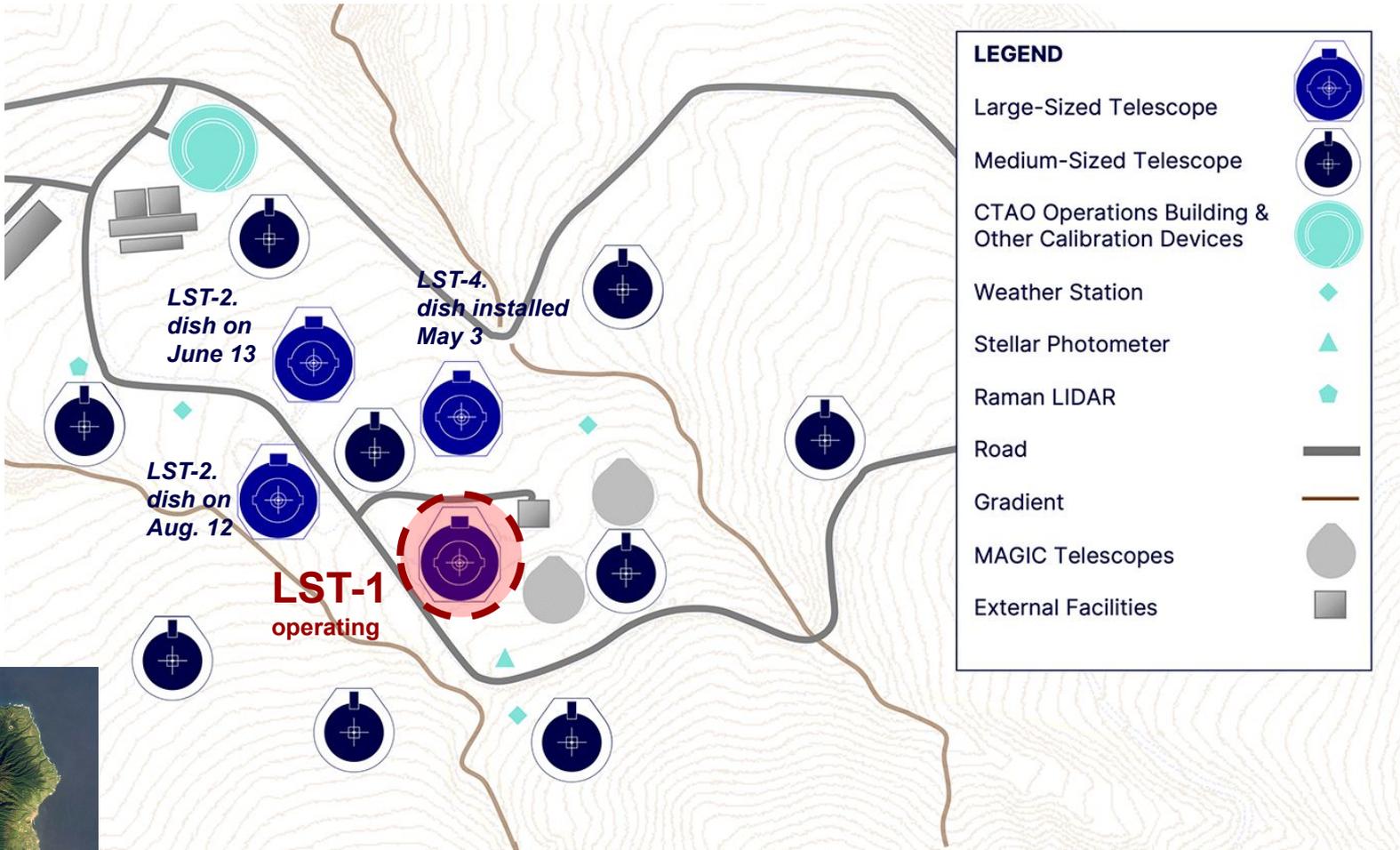
LST4



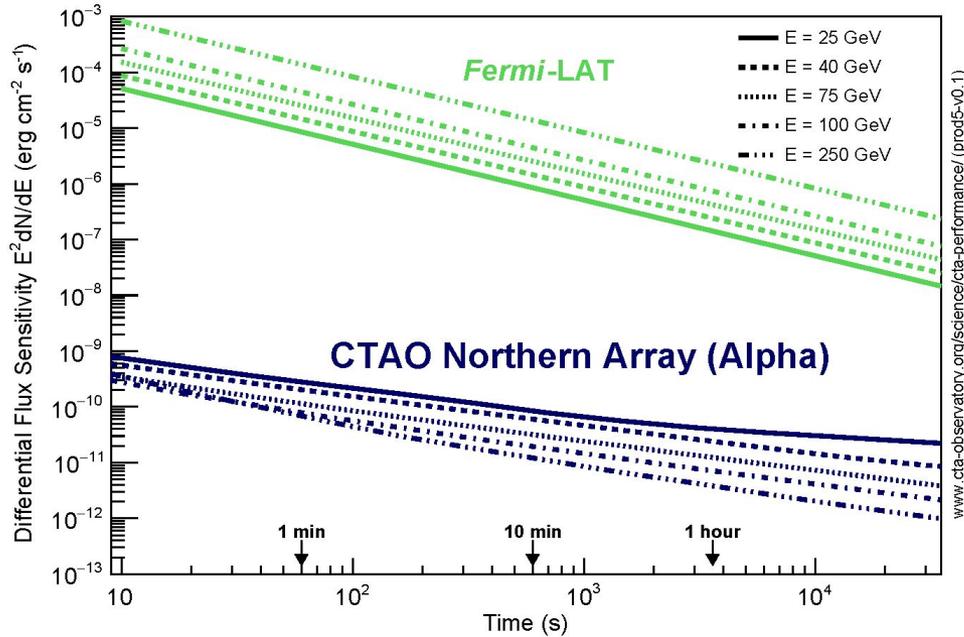
Aug 22, 2024

Credit: Alicia López Oramas (IAC)

CTAO-North

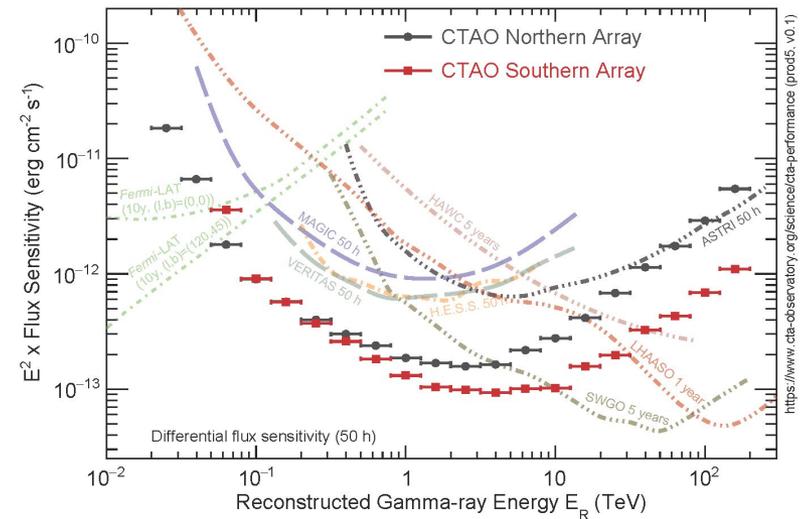


CTAO-North performance



- CTAO-South: More MSTs, +SSTs.
- CTAO-North: More LSTs (En. threshold)

CTAO: Improved sensitivity (x10), angular and energy resolution w.r.t to current generation IACTs (MAGIC, H.E.S.S., VERITAS)



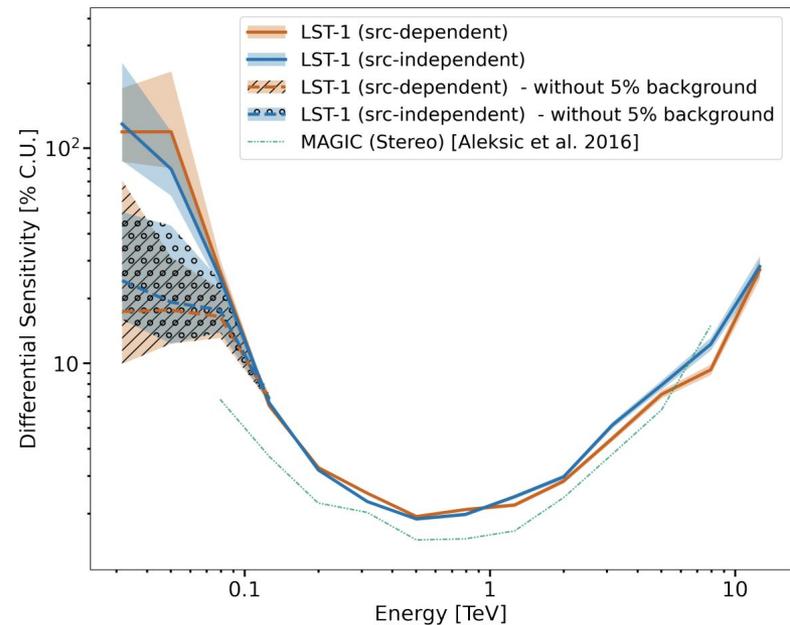
LST-1



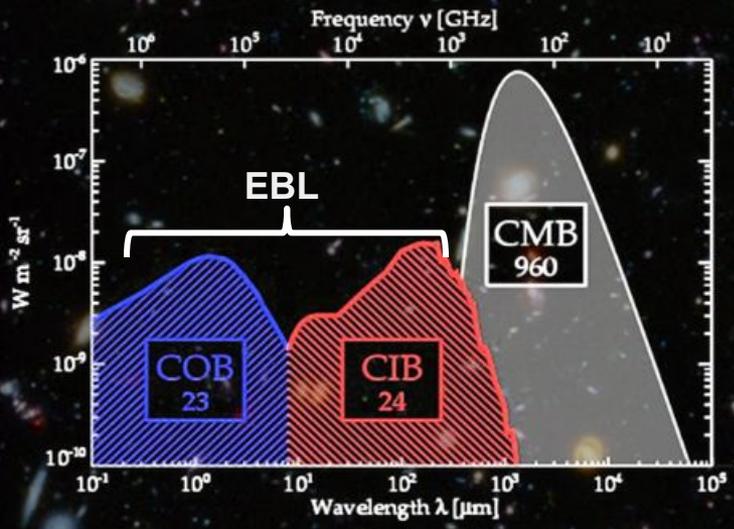
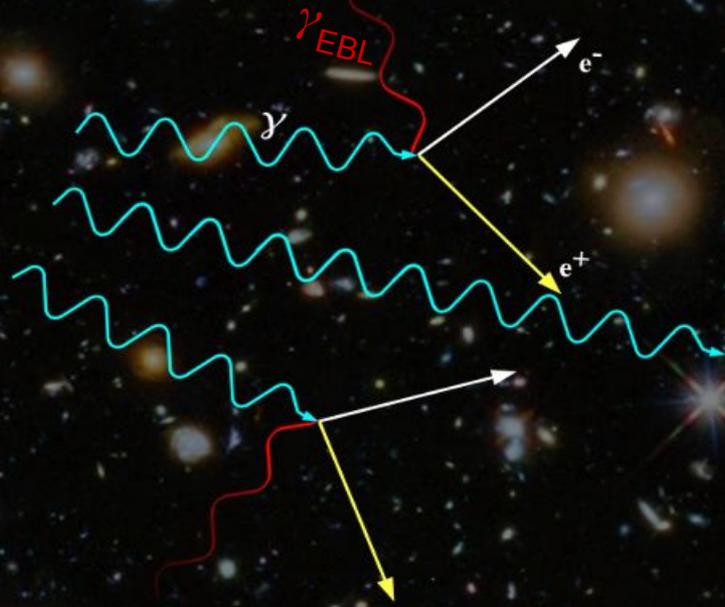
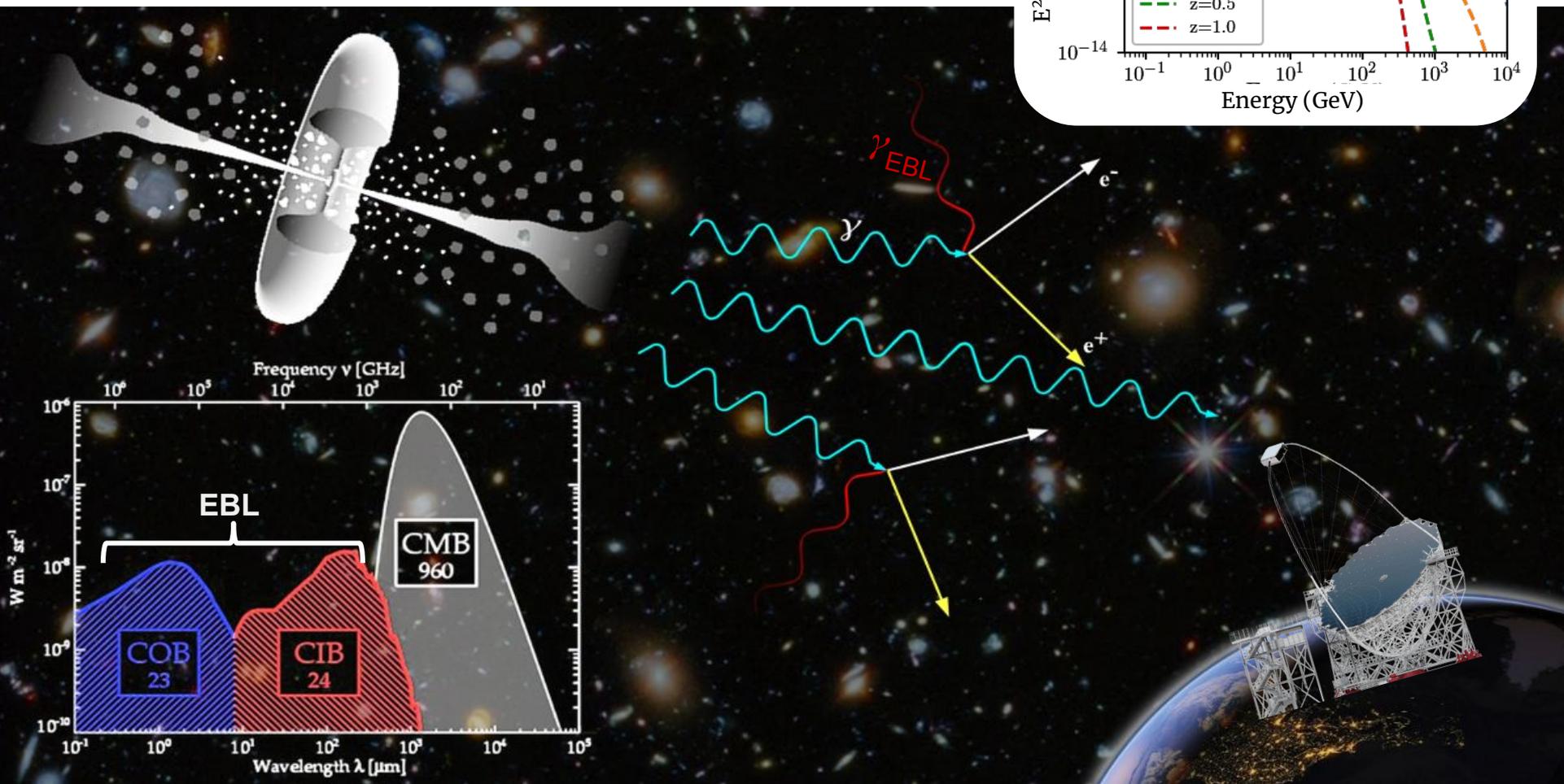
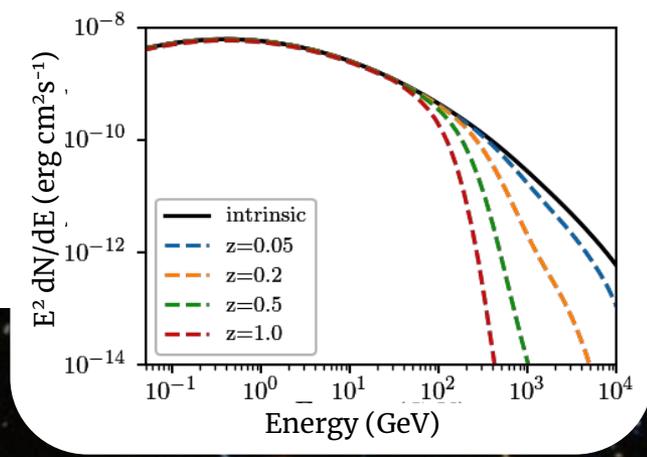
Prototype for CTAO LST and first telescope in the northern array.

- In operation since 2020.
- Lowest energy threshold among Cherenkov telescopes: ~ 30 GeV

Fast rotation: $180^\circ / 20\text{s}$



LST-1: extragalactic science and EBL

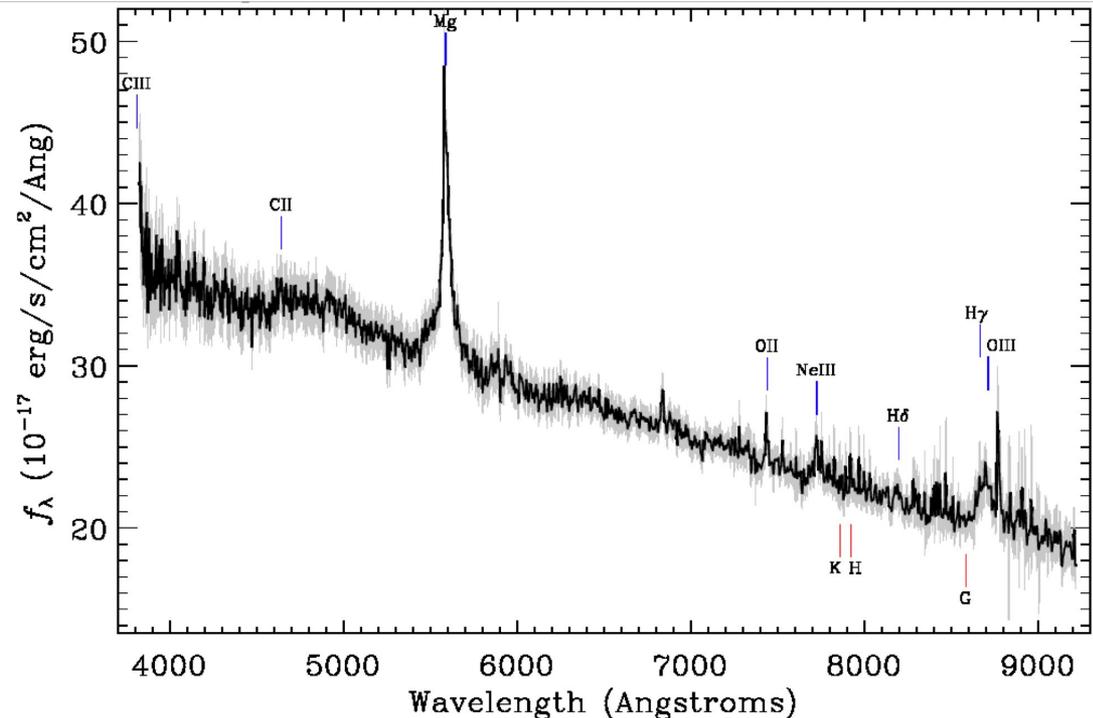


The Flat Spectrum Radio Quasar OP313

Most distant quasar detected in VHE ($z=0.9973$) and 10th in the list. Second most distant VHE source after GRB 201216C.

Not detected in VHE before. Attempted by MAGIC (2014, 2019).

Strong attenuation by EBL in the VHE regime. Possible intrinsic γ - γ absorption.

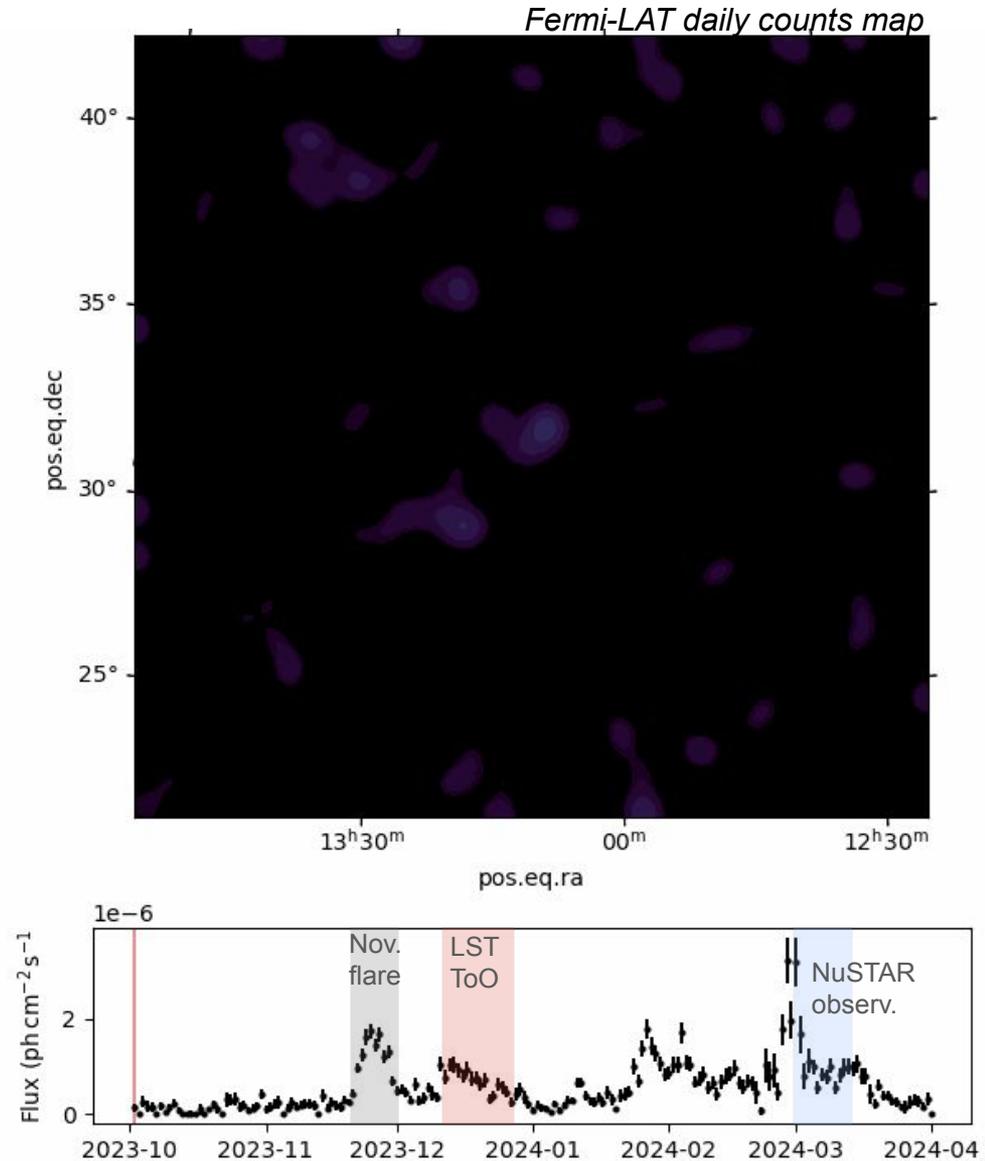


The Flat Spectrum Radio Quasar OP313

Very active in *Fermi*-LAT since November 2023 (during our 'Moon break').

LST-1 observations since December 9th, 2023.

Among the top **most luminous AGN** ever seen by LAT

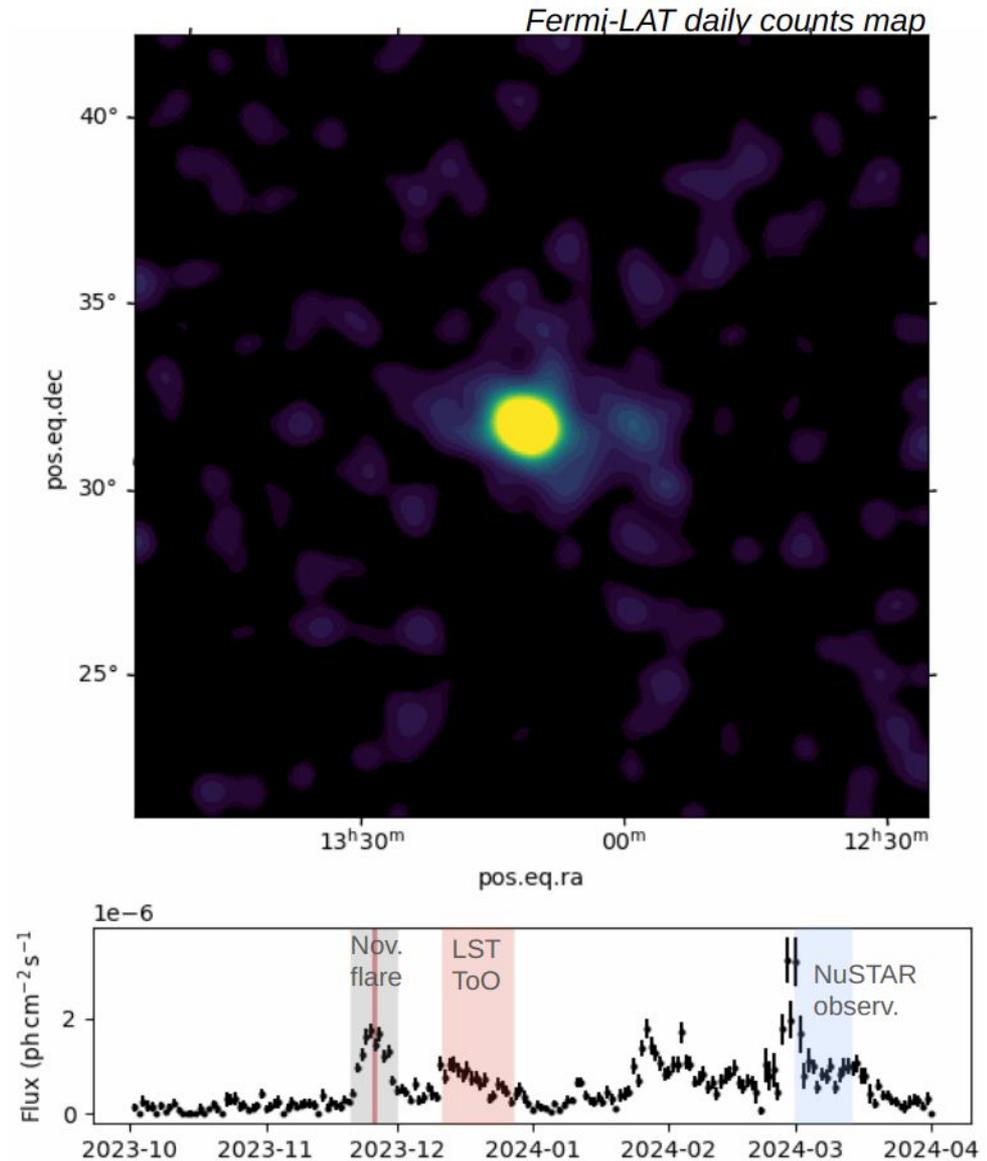


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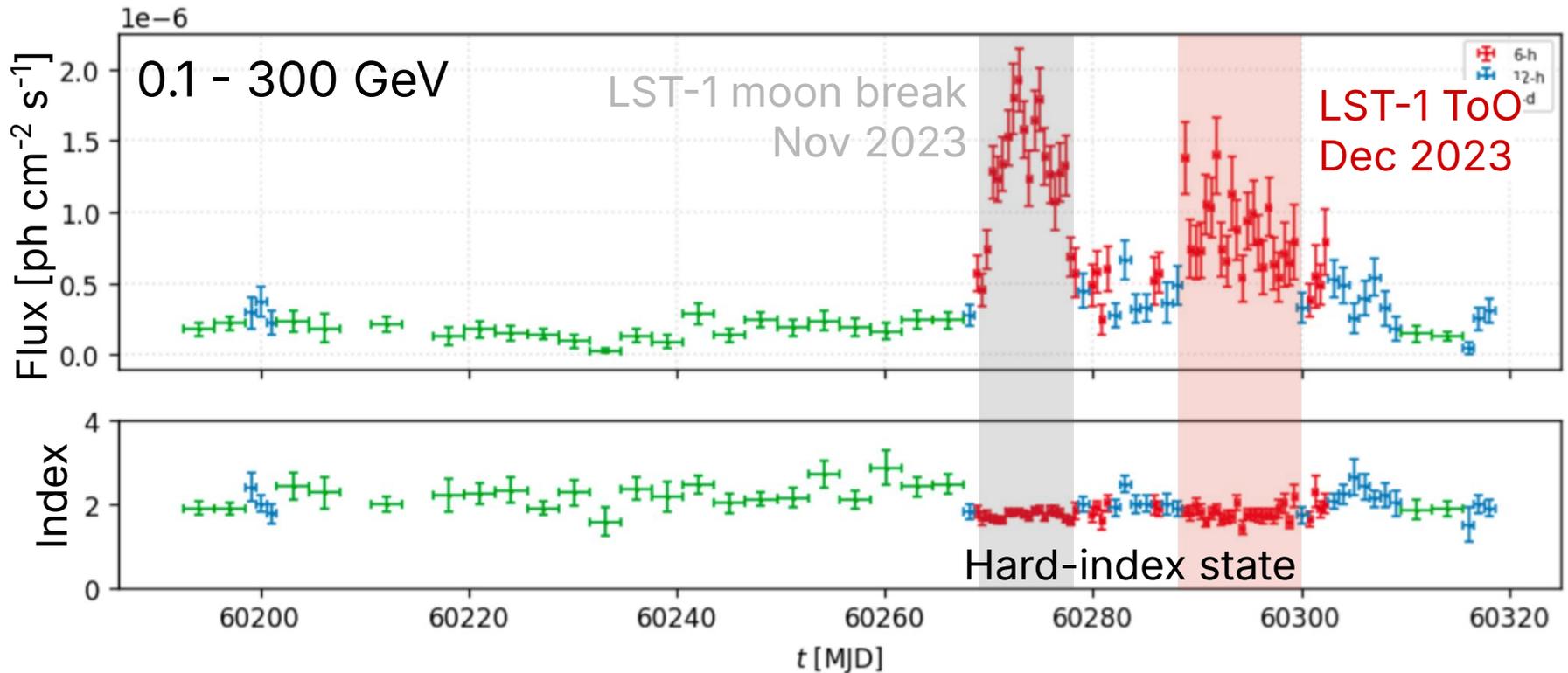
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The Flat Spectrum Radio Quasar OP313

Early observations (this work): daily observations from Dec 9 to 18, 2023 (15 h) [high state] + few days in Jan 2024 (5 h) [low state].

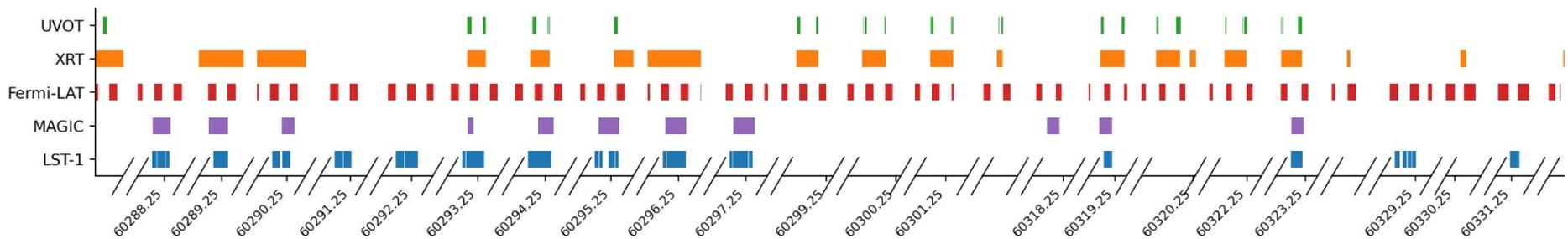
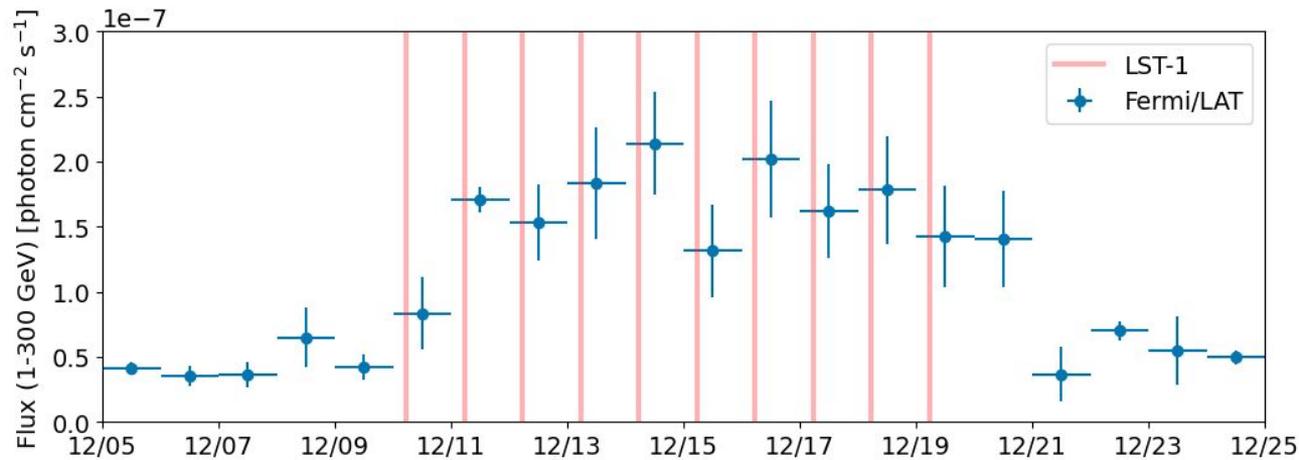
Telescope pointing zenith angle > 30 deg (energy threshold ≈ 40 GeV)



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Telescope pointing zenith angle > 30 deg (energy threshold ≈ 40 GeV)



Detection of **OP313**, ATel and press releases

Detected with $>5\sigma$
(Li&Ma) with data up to
Dec 14th, 2023 (~6 h)

ATel issued by LST-1
(#16381): 10th FSRQ
detected in VHE
gamma rays



The Astronomer's Telegram

*Announced on CTAO,
ATel, Newspaper,
radio and TV*

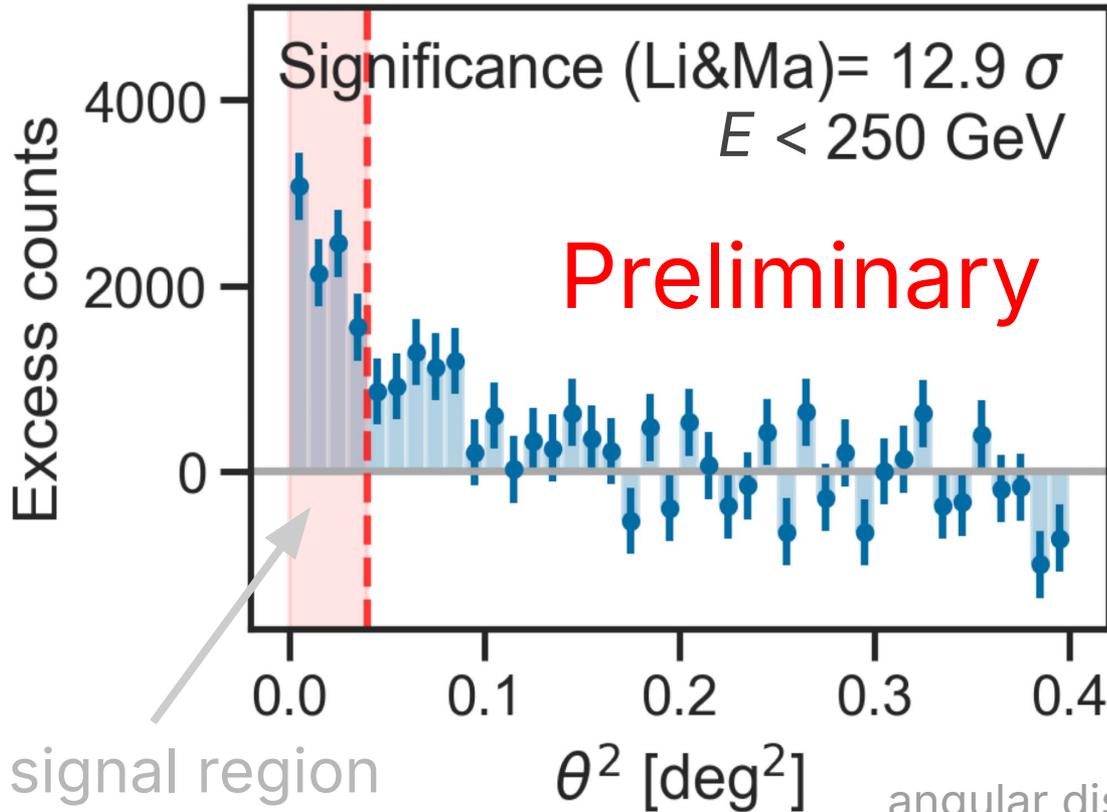
First detection of VHE gamma-ray emission from FSRQ OP 313 with LST-1

ATel #16381; **Juan Cortina (CIEMAT) for the CTAO LST collaboration**
on 15 Dec 2023; 14:31 UT

Credential Certification: Juan Cortina (Juan.Cortina@ciemat.es)

Subjects: Gamma Ray, $>GeV$, TeV, VHE, Request for Observations, AGN, Blazar,
Quasar

Detection of **OP313**: γ -ray excess



December dataset (15h):

Signif. (Li&Ma) $\sim 13\sigma$

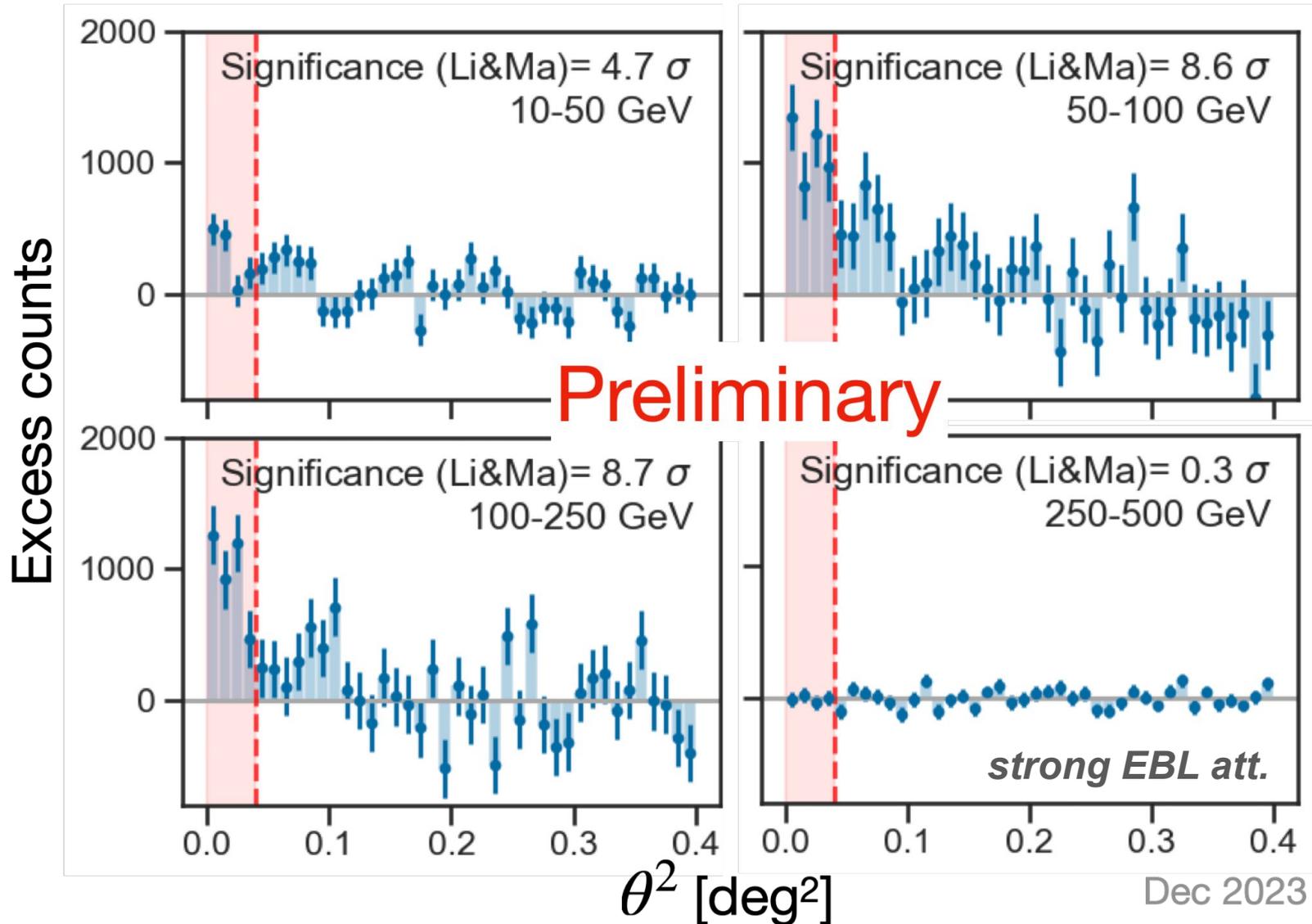
January dataset (5h):

No significant detection

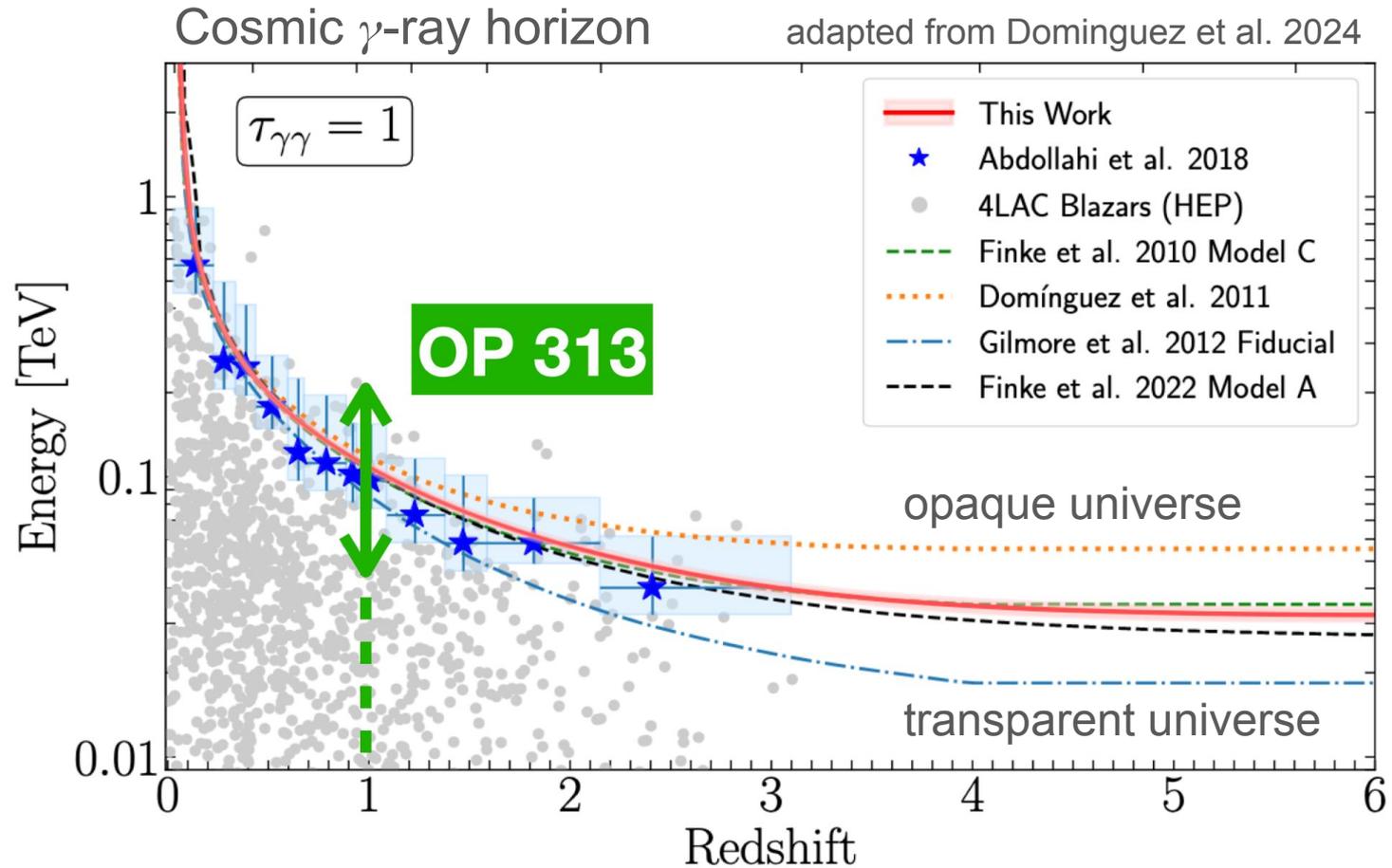
angular distance (squared) between γ -ray like events and source location

Average VHE flux ($>100 \text{ GeV}$): 28% Crab (December 2023)

Detection of **OP313**: in energy bands



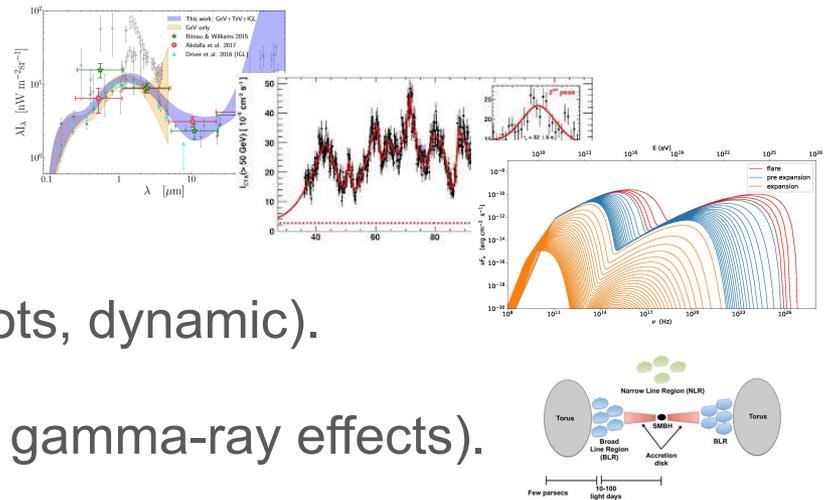
Prospects: EBL constraints



Prospects: Ongoing work and publication plans

Likely among the most intense observation campaigns on a distant FSRQ to date. Large project, with a large team, and many ramifications, including:

- VHE emission & EBL constraints.
- Variability: long-term, short-term.
- Broadband SED modeling (snapshots, dynamic).
- Broad-line-region studies (optical & gamma-ray effects).
- Very deep exposure nights: optical, X-rays & gamma-rays, polarimetry
- Technical papers (new analysis methods, data formats).



Summary

- Major flare from OP313 detected with *Fermi*-LAT in November 2023 (Moon break) and once again December 2023 (LST-1 begins observations)
- First detection of VHE emission from quasar OP 313 ($z=0.997$) by LST-1 + more detections in the next months (very structured flare)
 - Prompt reaction of LST Collaboration (observations and analysis)
 - First VHE source discovered by LST-1 (important milestone)
 - ... and the most distant VHE AGN.
 - ... and among the most luminous ones.
- Very fruitful observation campaign. Many projects started.
- First publication from LST-1 coming soon, stay tuned.

Bonus

A new multi-wavelength analysis and data management workflow based on ...



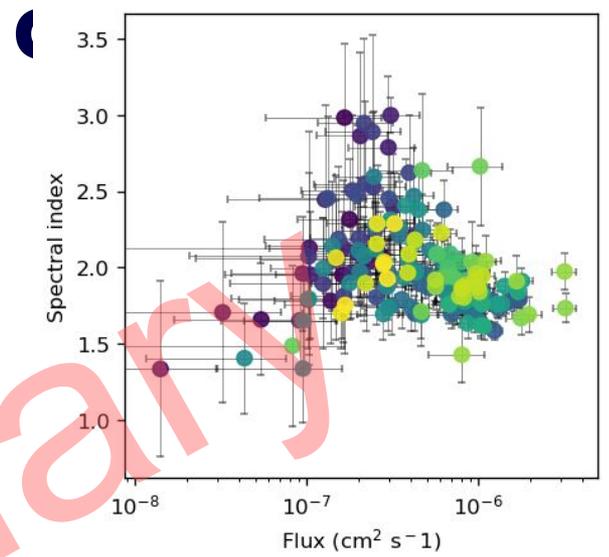
A method **to store** (*binned*) event data from optical to gamma-rays (>11 orders of magnitude)

A method **to analyse** (full forward folding with events + IRFs): no more flux points !

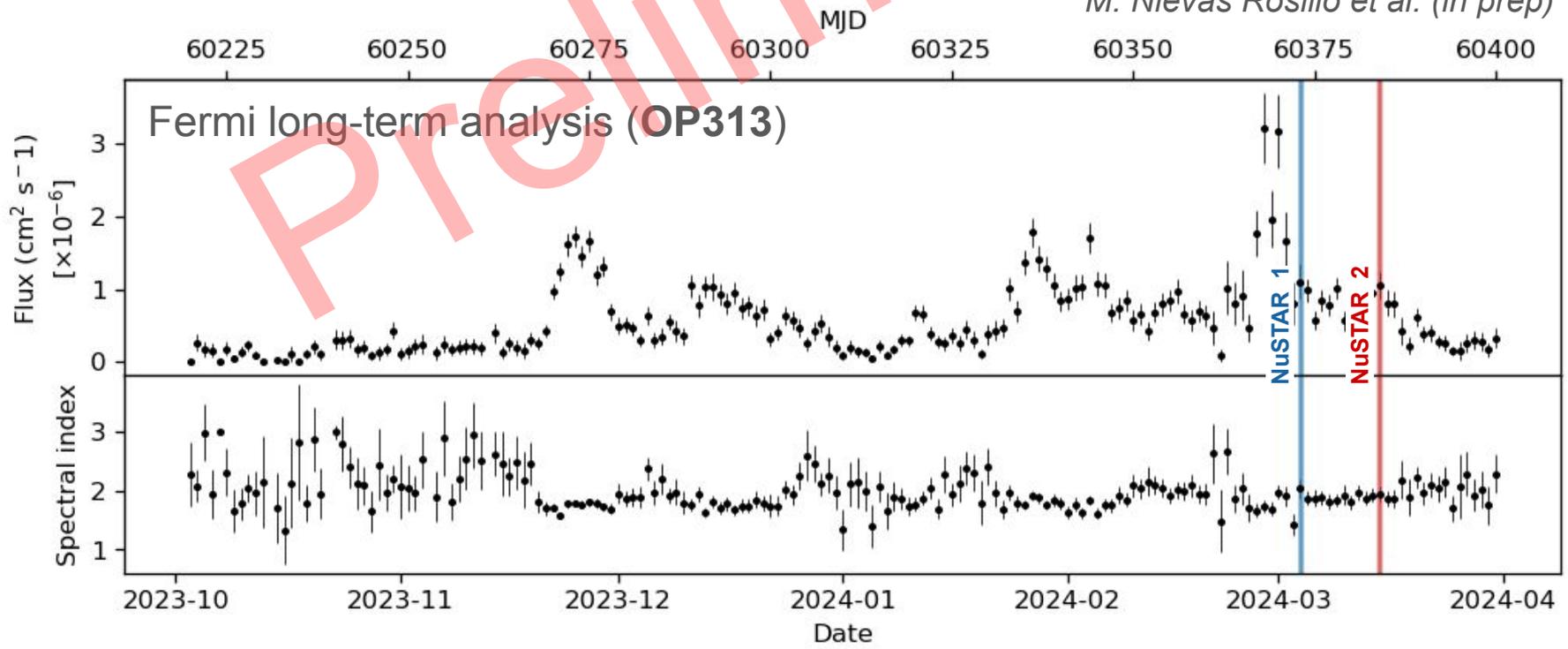
MWL *gammapy* workflow

1st Goal: publish *binned event data* from optical to gamma-rays (11 orders of magnitude)

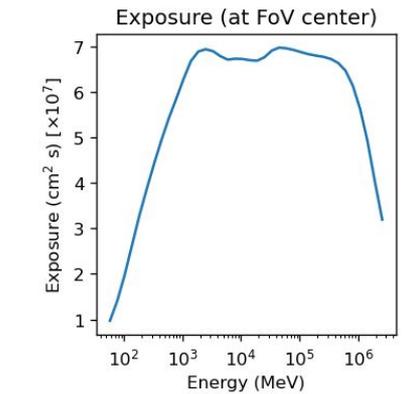
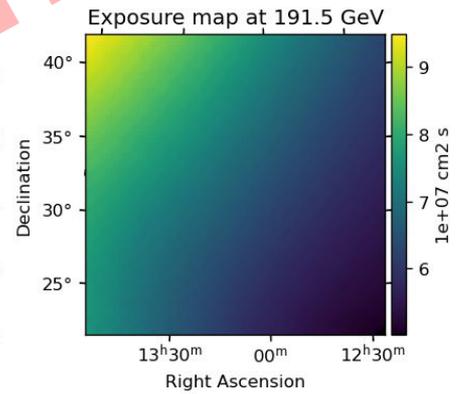
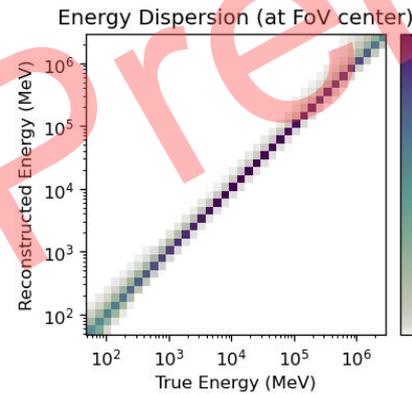
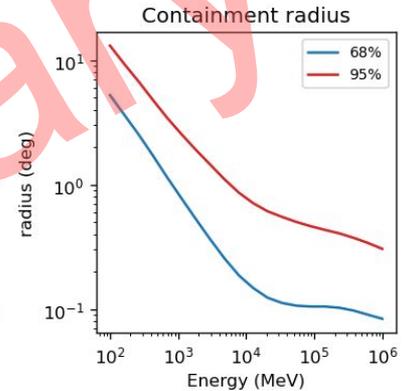
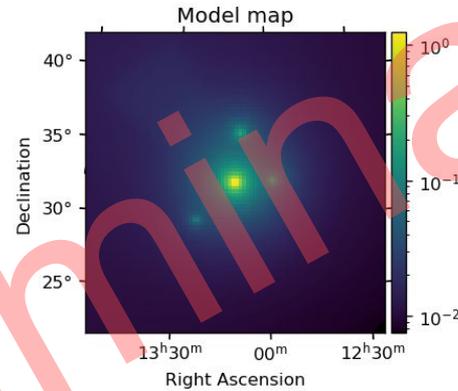
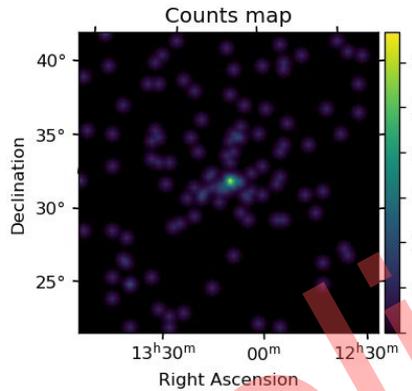
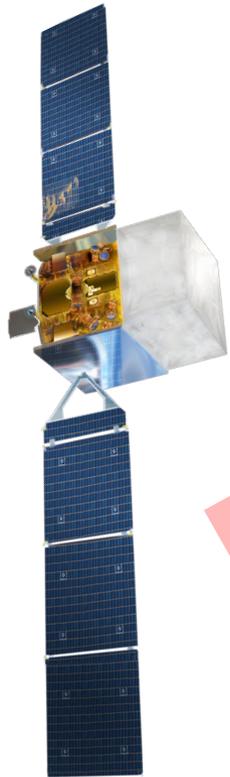
2nd Goal: forward folding fit of models (no more flux point fitting)



M. Nieves Rosillo et al. (in prep)



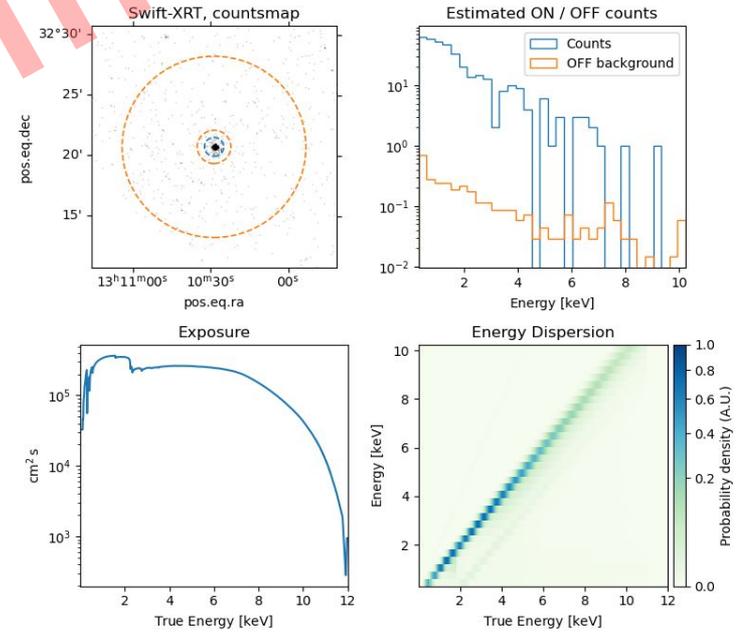
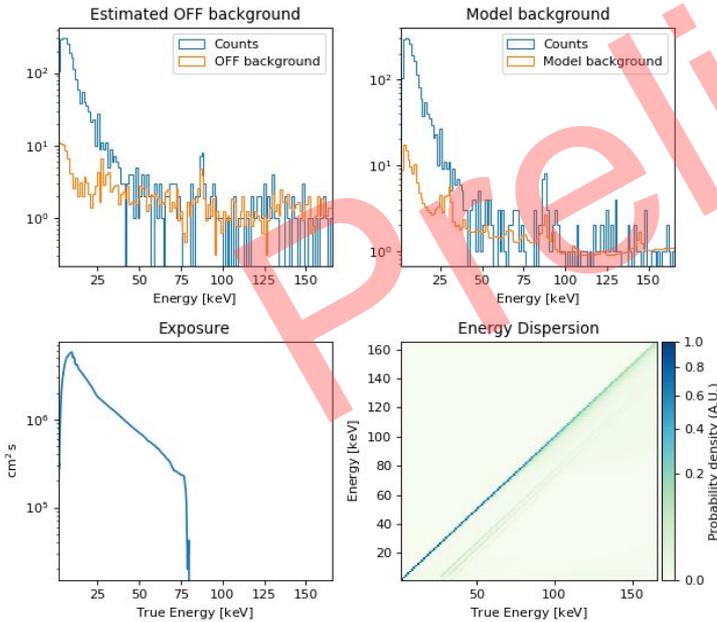
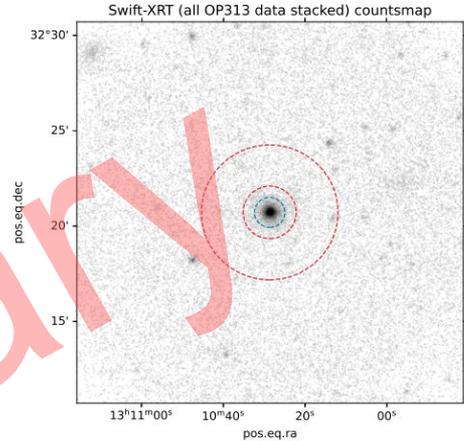
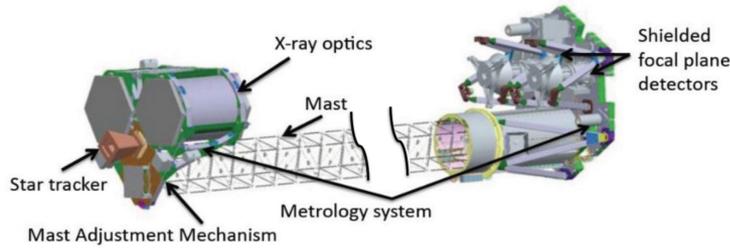
MWL *gamma*mapy workflow



Fermi-LAT

(space-based, HE gamma rays, 3D)

MWL *gammapy* workflow



NuSTAR
(space-based, hard X-rays, 1D)

Swift (XRT & UVOT)
(space-based, X-ray and optical-UV, 1D)

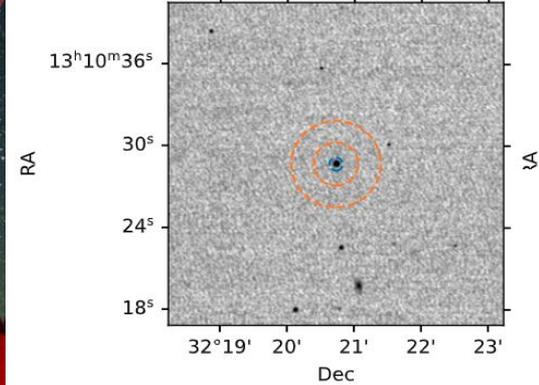
MWL *gammapy* workflow

Liverpool Telescope (ground based, optical-IR photometry, 1D)

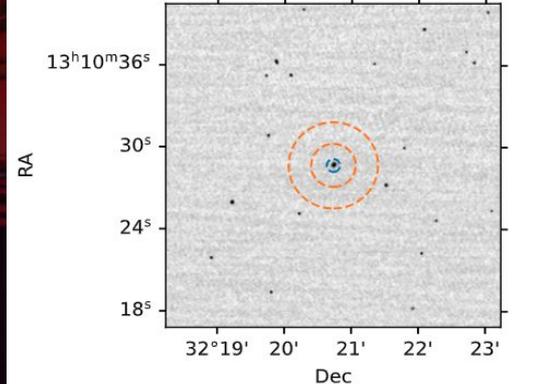


Daniel López / IAC

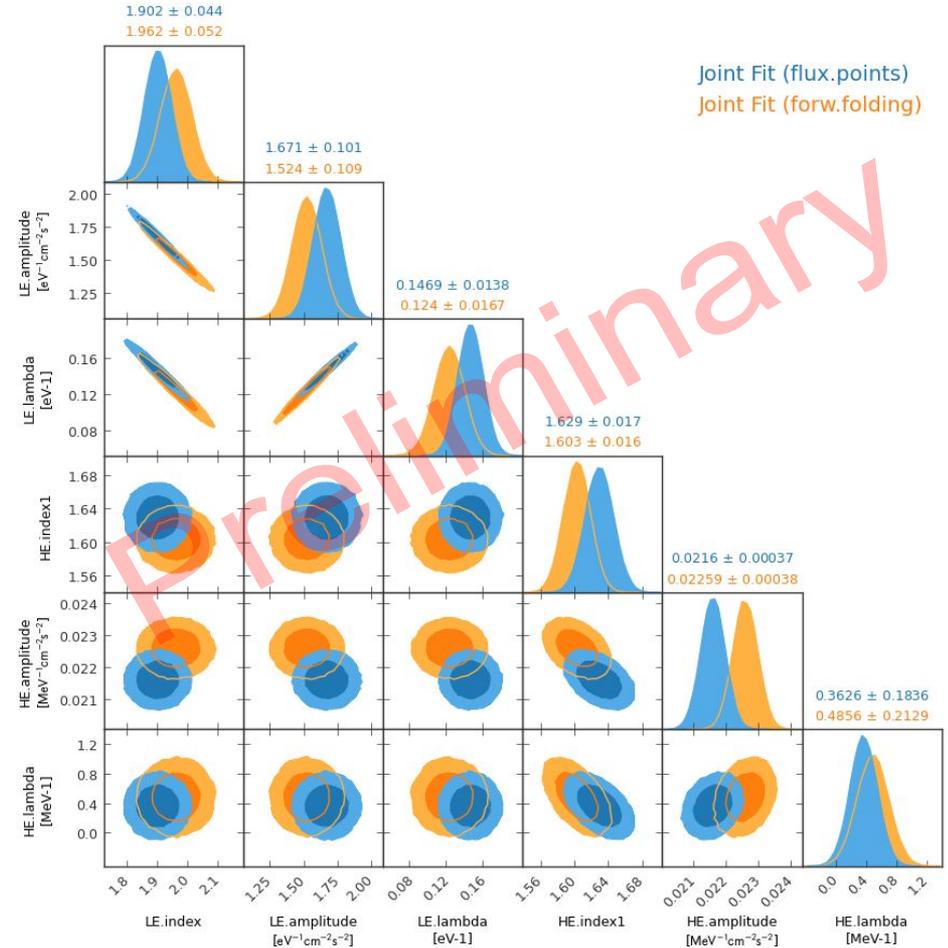
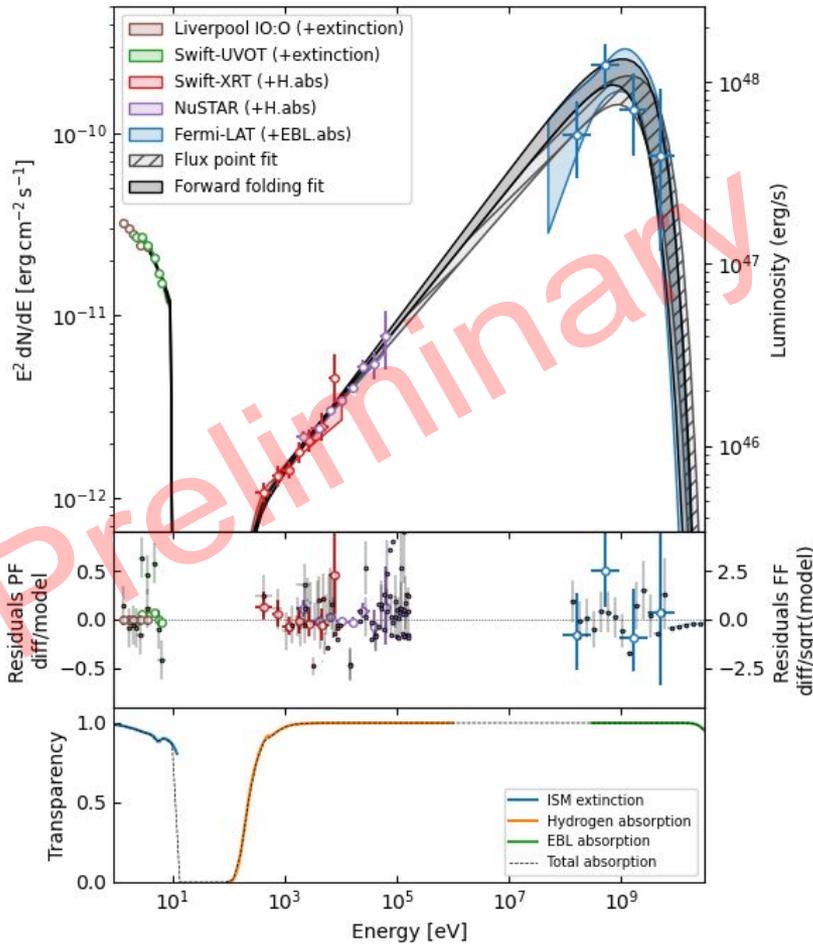
Liverpool Telescope-IO:O.SDSS-Z

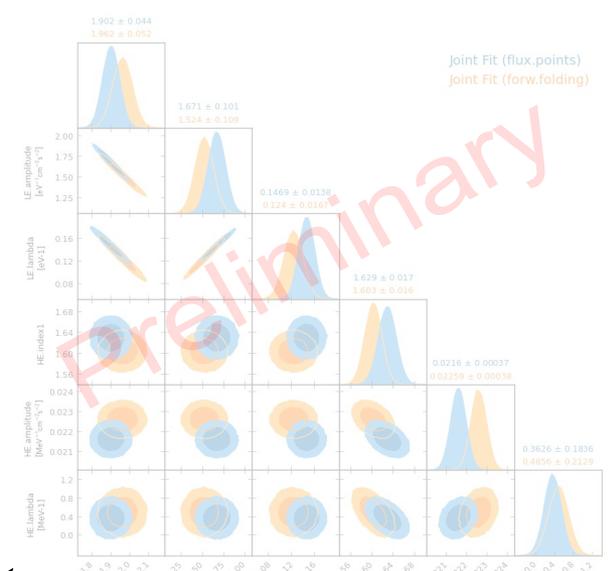
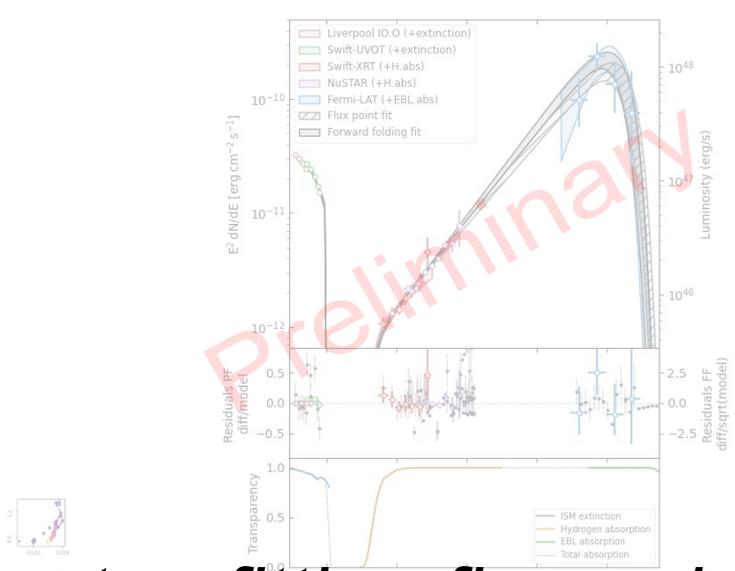
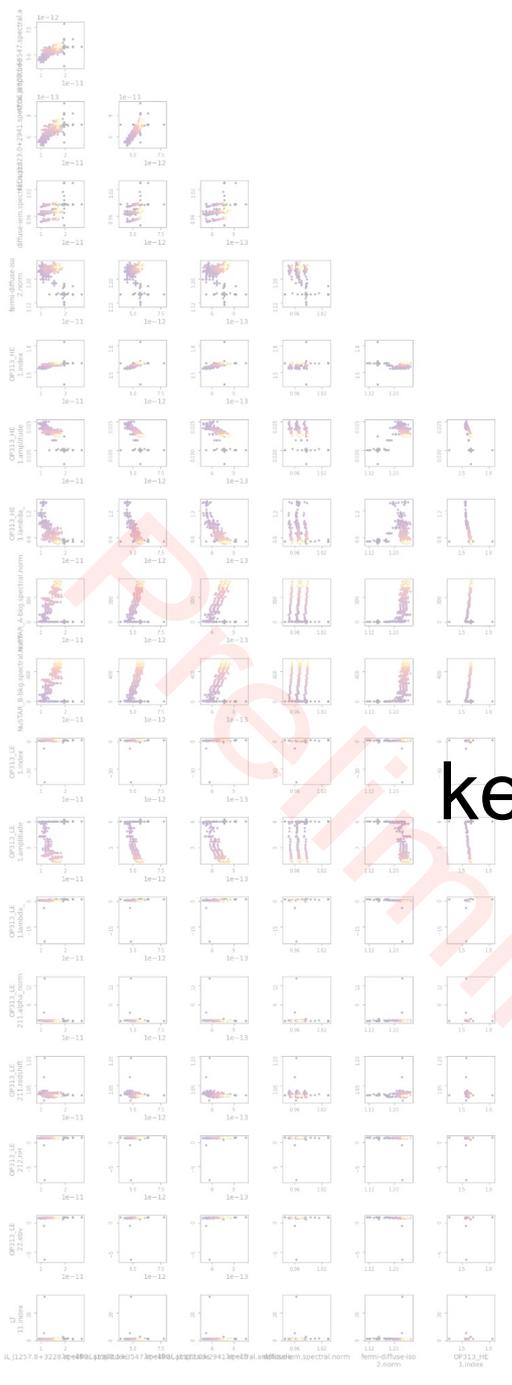


Liverpool Telescope-IO:O.SDSS-U

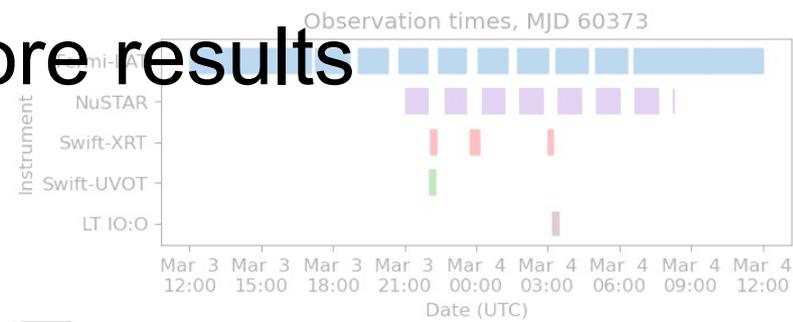


MWL *gammapy* workflow





**stop fitting flux points
&
keep tuned for more results**

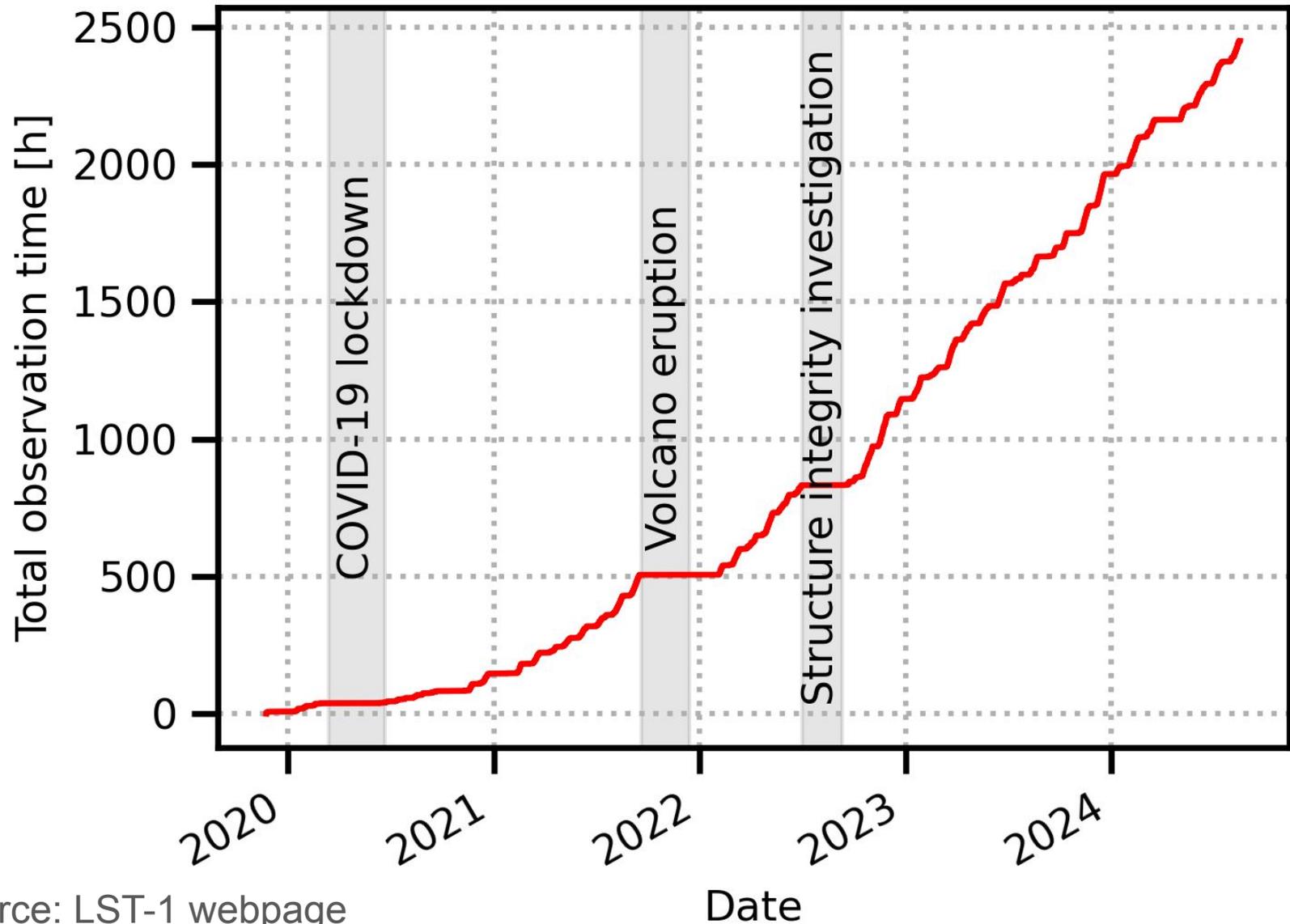


M. Nievas Rosillo et al. (in prep)

Backup

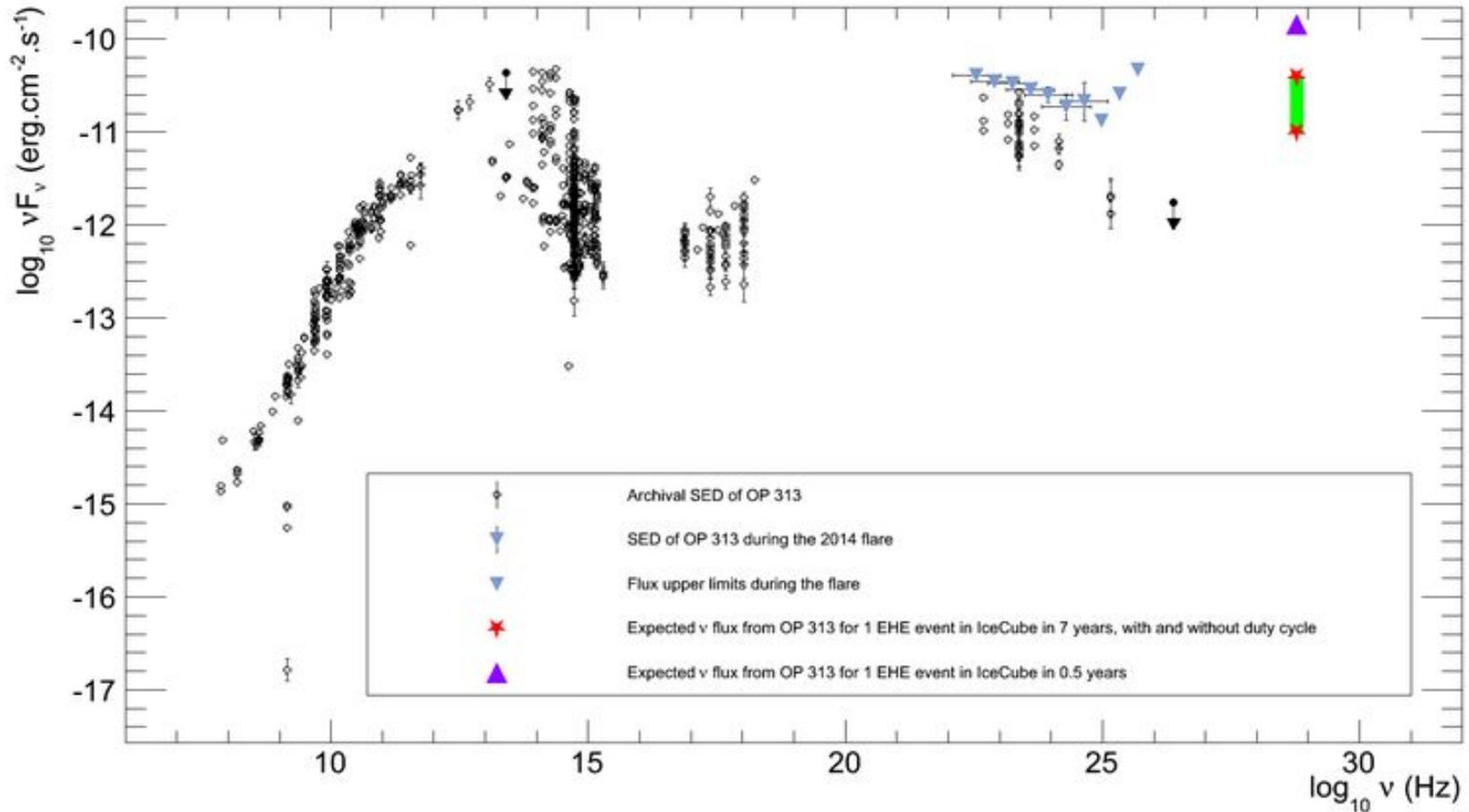
Total obs. time: 2449 h

Last updated: 2024-08-23



2014 flare

SED for OP 313



Marinelli+ 21 MNRAS 506, 3,
pp.3760-3772

2014 & 2019 flares

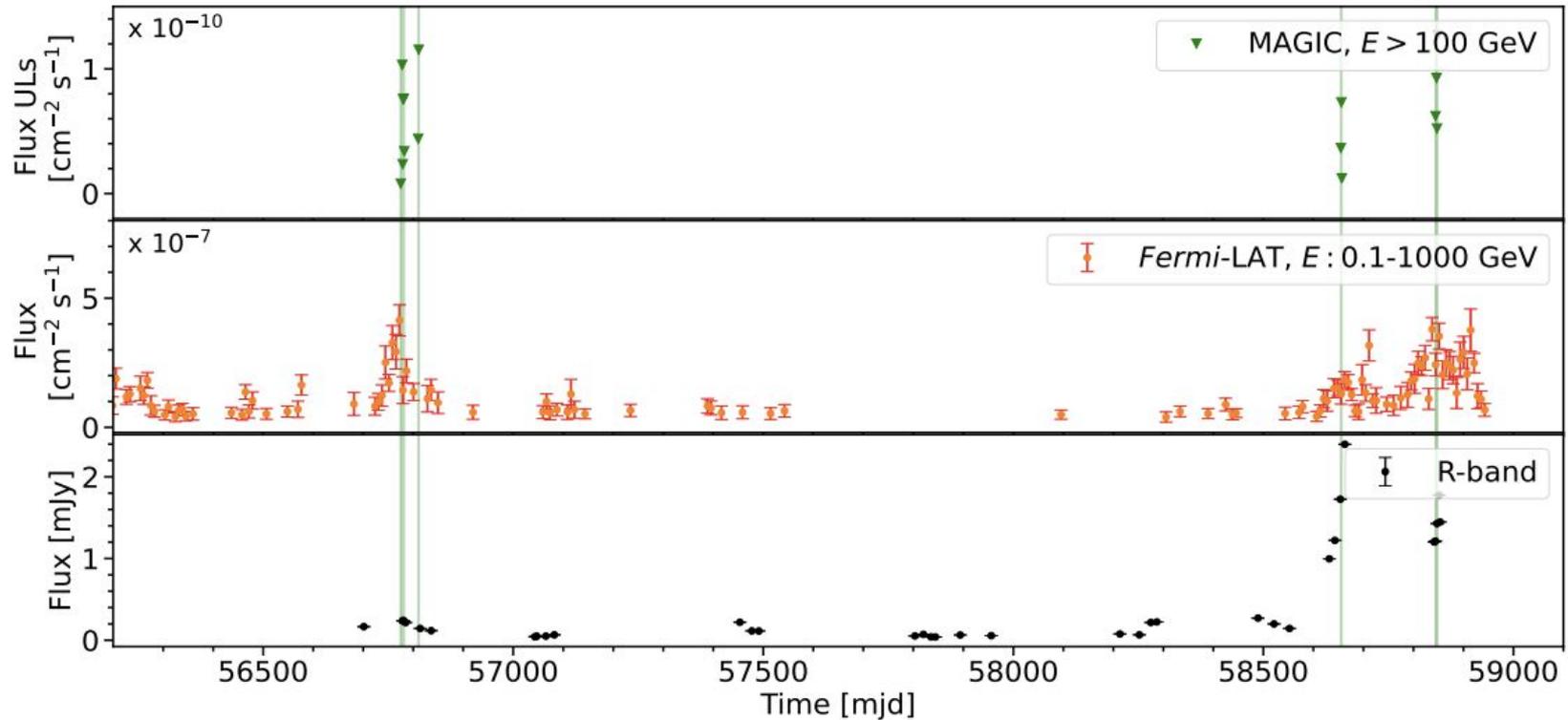


Figure A3. MWL light curve of OP 313. The labels are the same as Fig. A2.