

The Radio Neutrino Observatory in Greenland (RNO-G): Overview and Status

Kaeli Hughes, for the RNO-G Collaboration
June 11, 2024



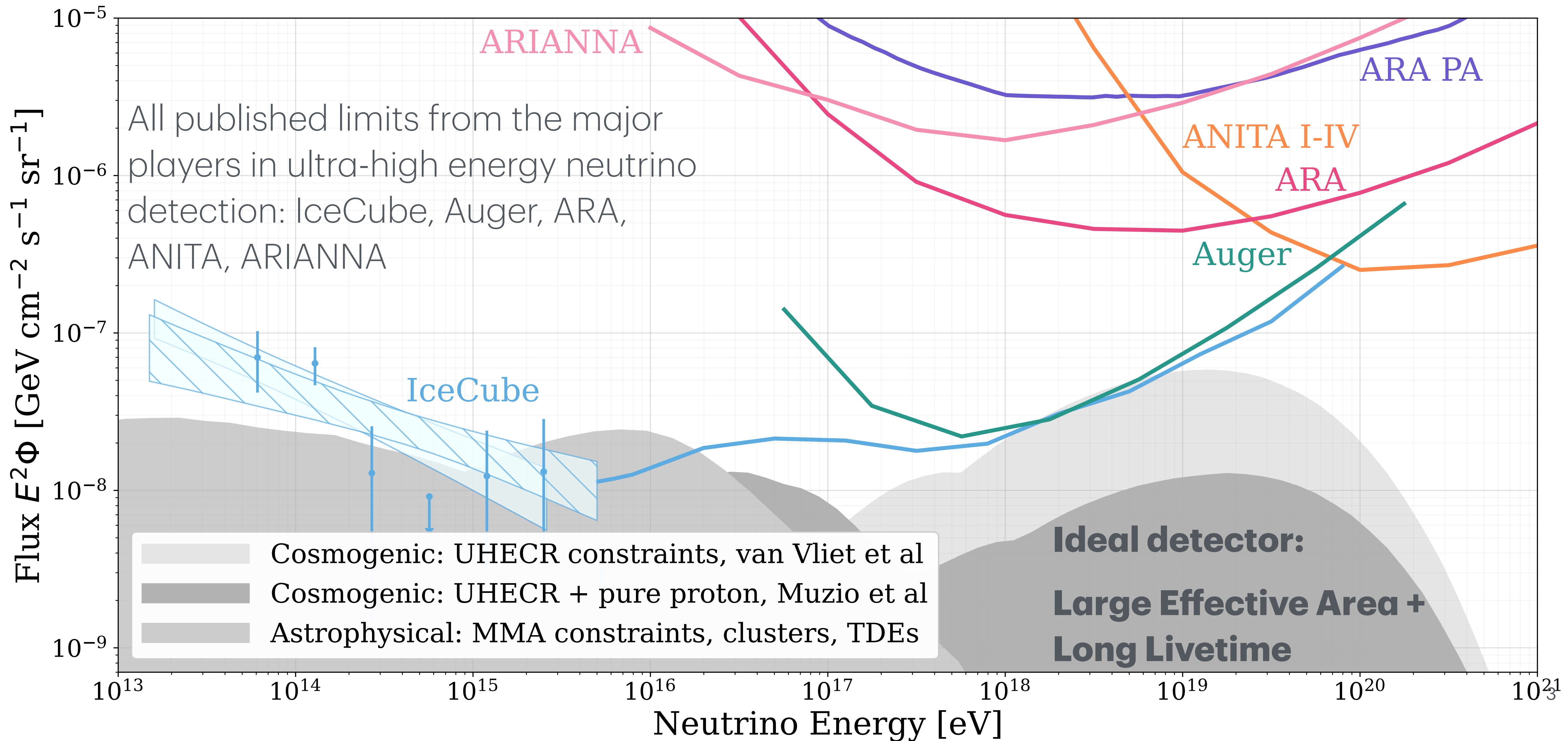
The RNO-G Collaboration



RNO-G
Collaboration
February 2023

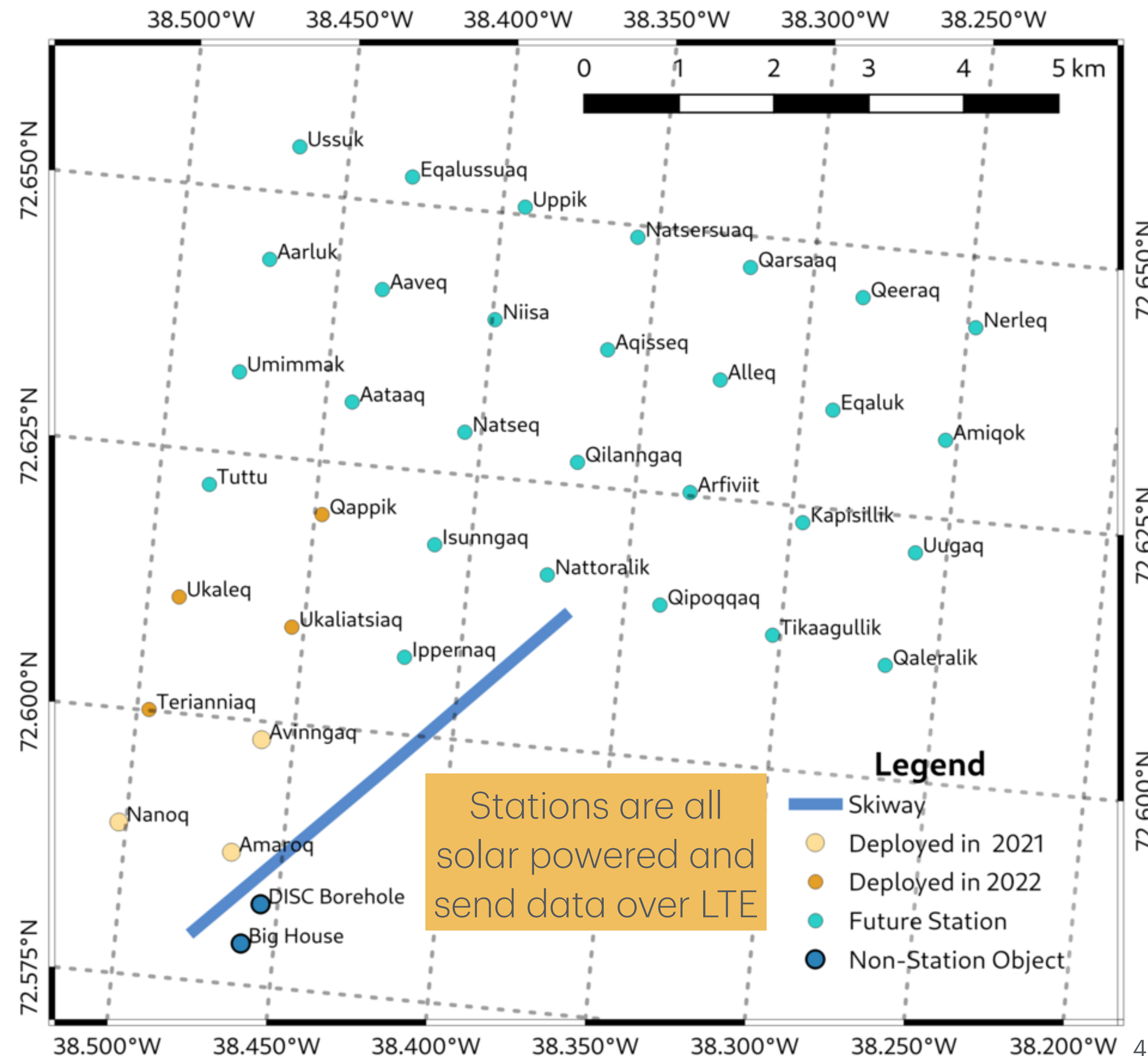
			
			
			
			
			

The state of the field: June 2024

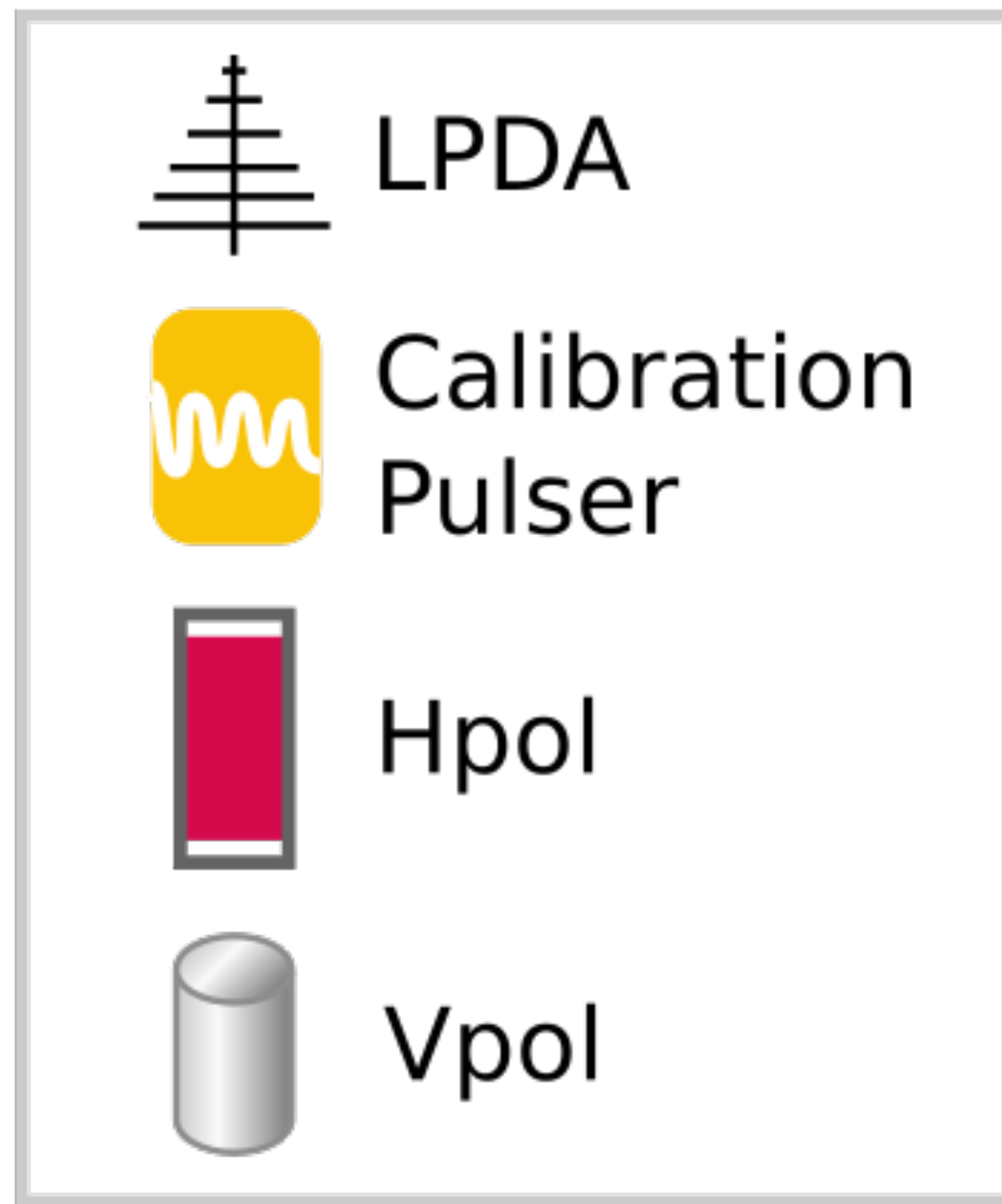


RNO-G: a new experimental effort

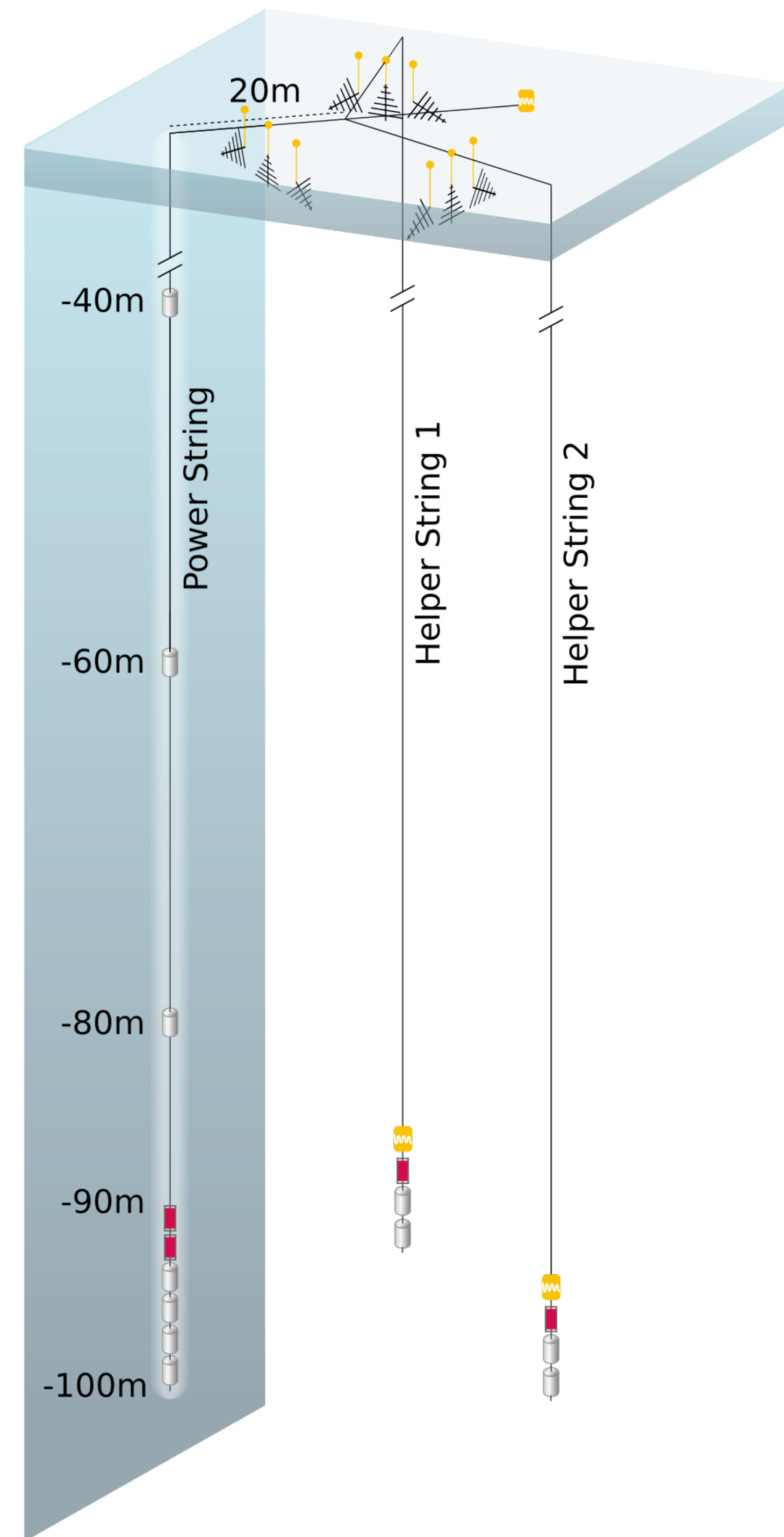
- Deployed near Summit Station, Greenland
- Hardware is fully funded to reach 35(+) stations: already the largest in-ice radio neutrino detector by area!
- Currently in building phase: holes for 7 more stations are being drilled this year, with DAQs installed next year
- Science team is also working on calibration, simulation, instrument performance- see following talks from RNO-G team members!



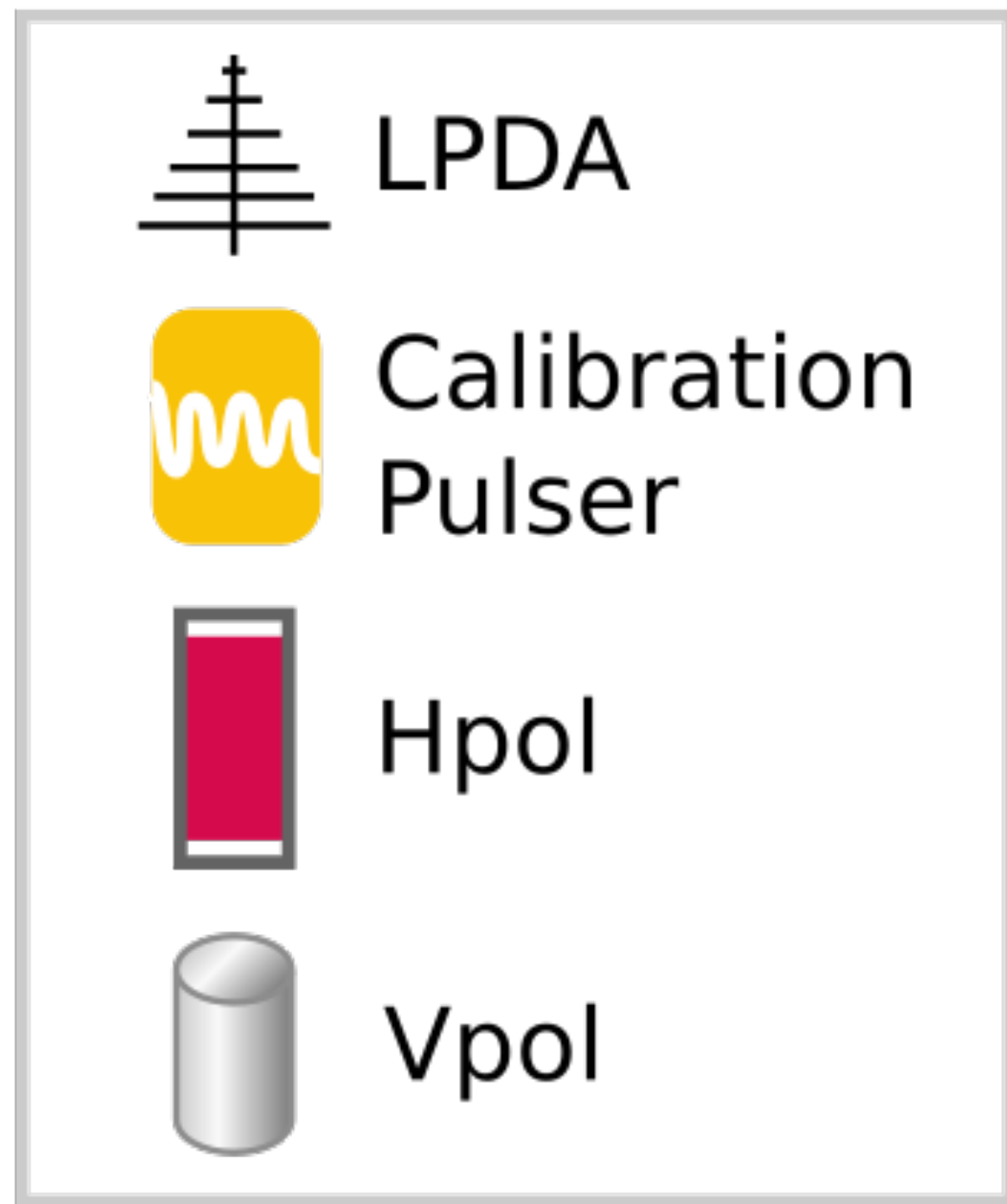
A single RNO-G Station



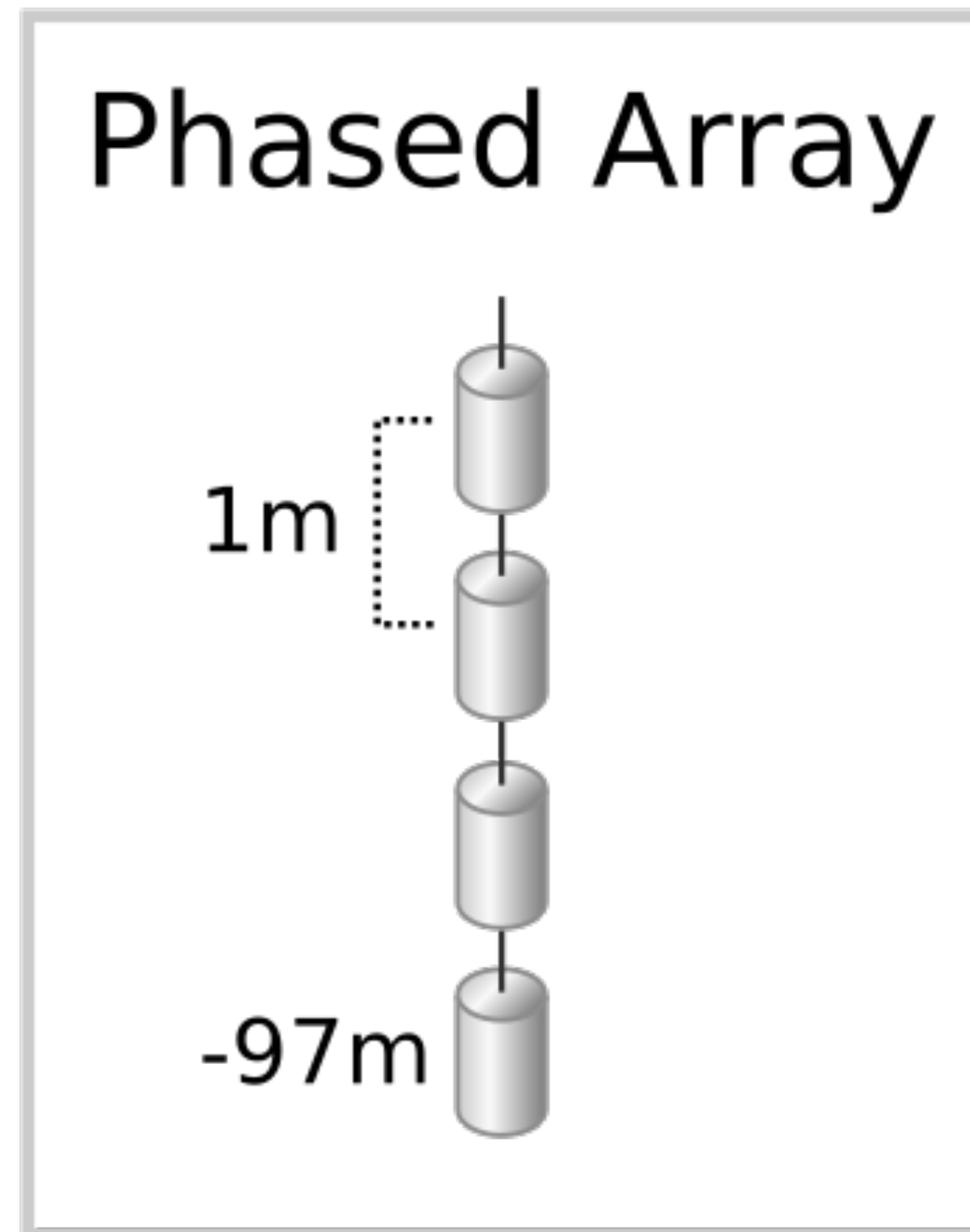
Antennas types: Log-periodic dipole antenna (LPDA), Horizontally-polarized (Hpol), and Vertically-polarized (Vpol)



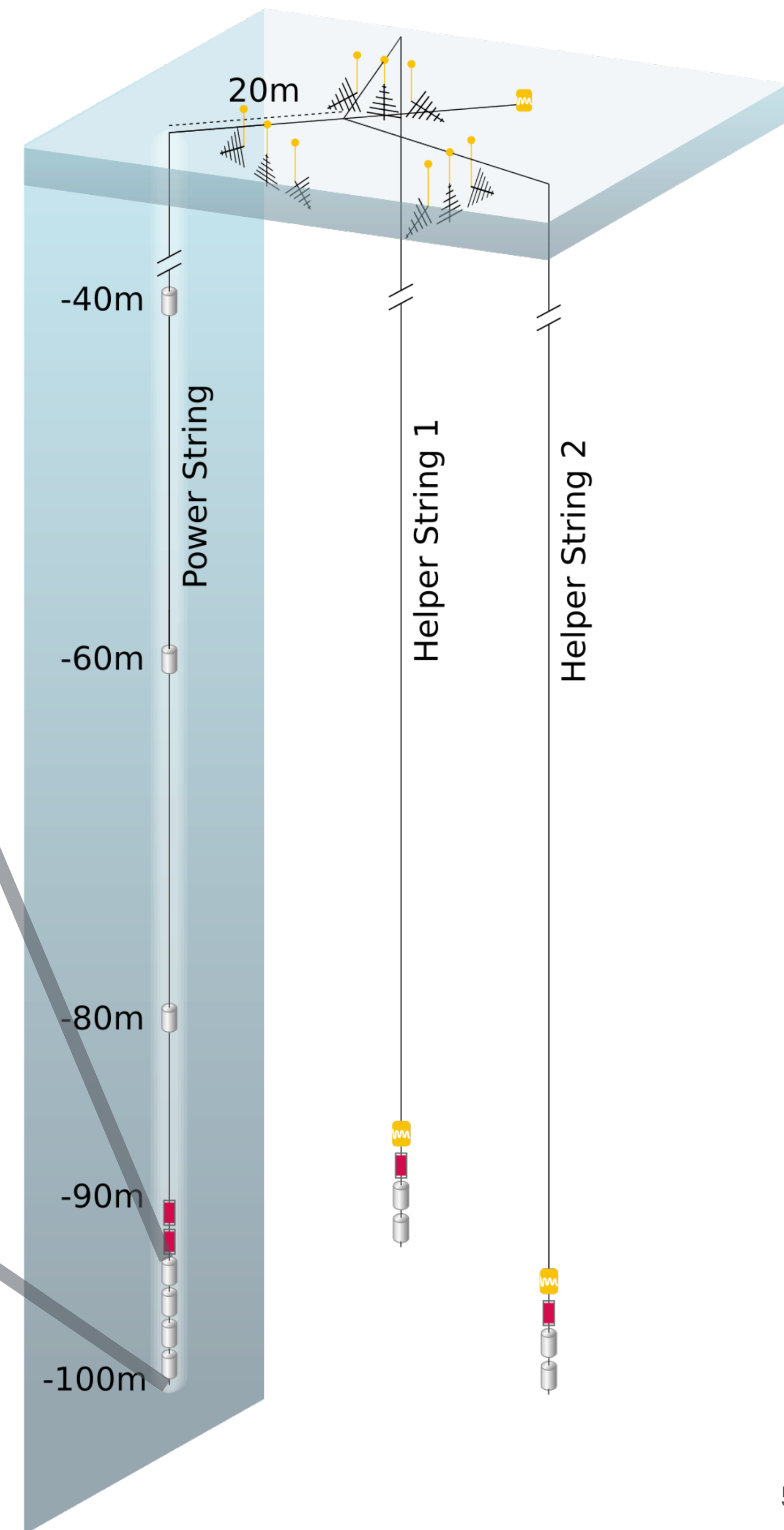
A single RNO-G Station



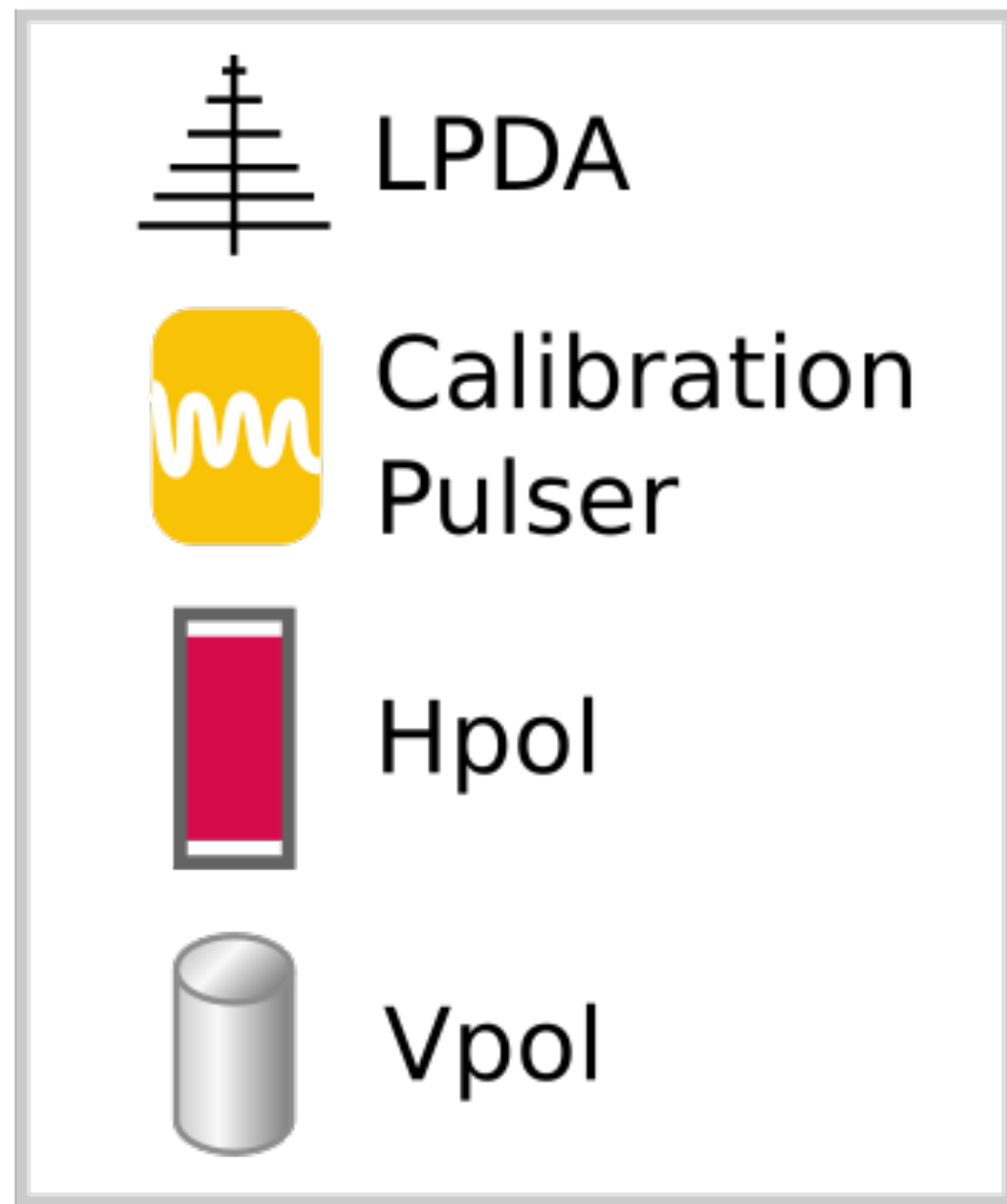
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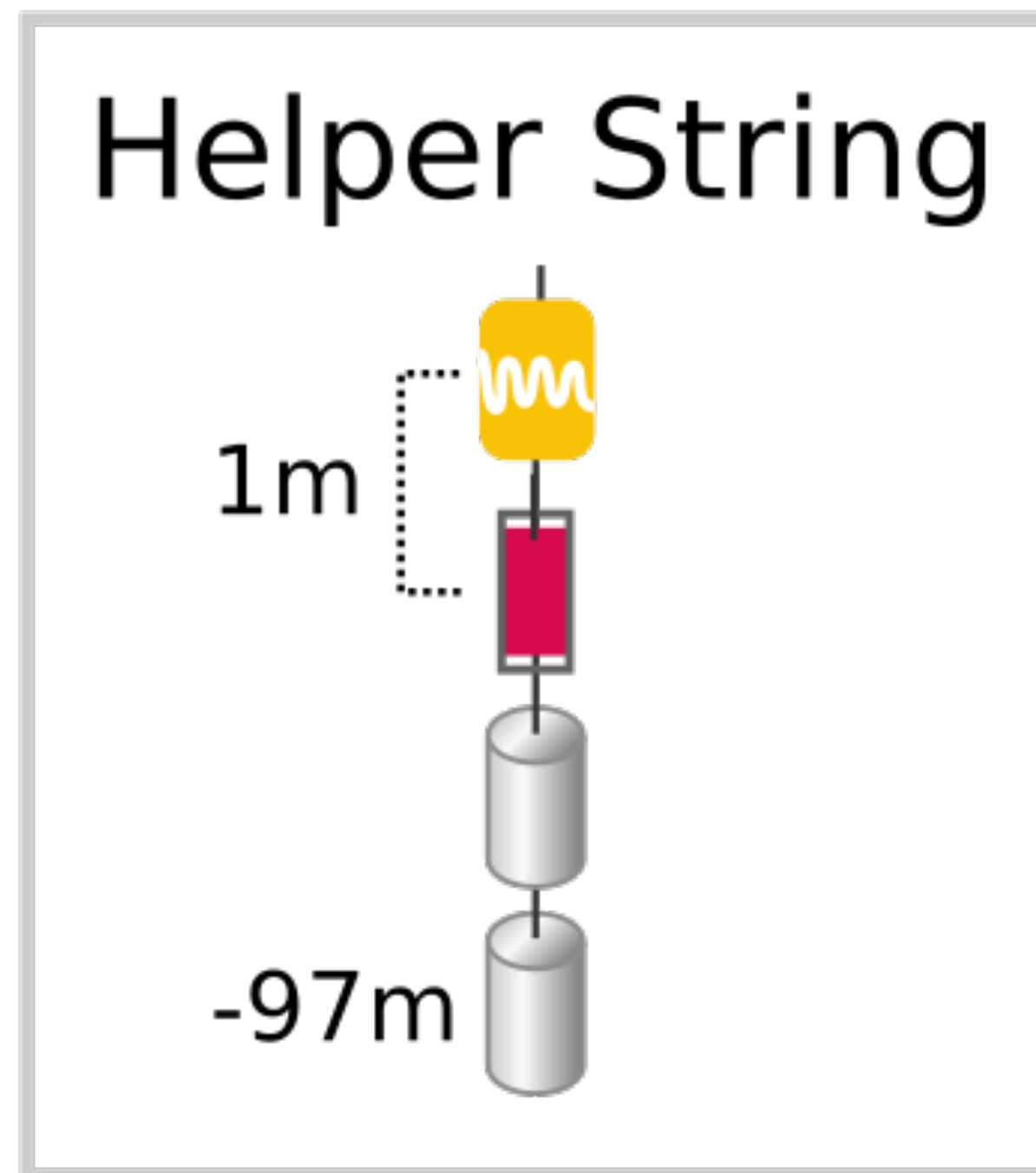
Densely packed phased array of vertically polarized (Vpol) antennas will make up our most sensitive trigger.



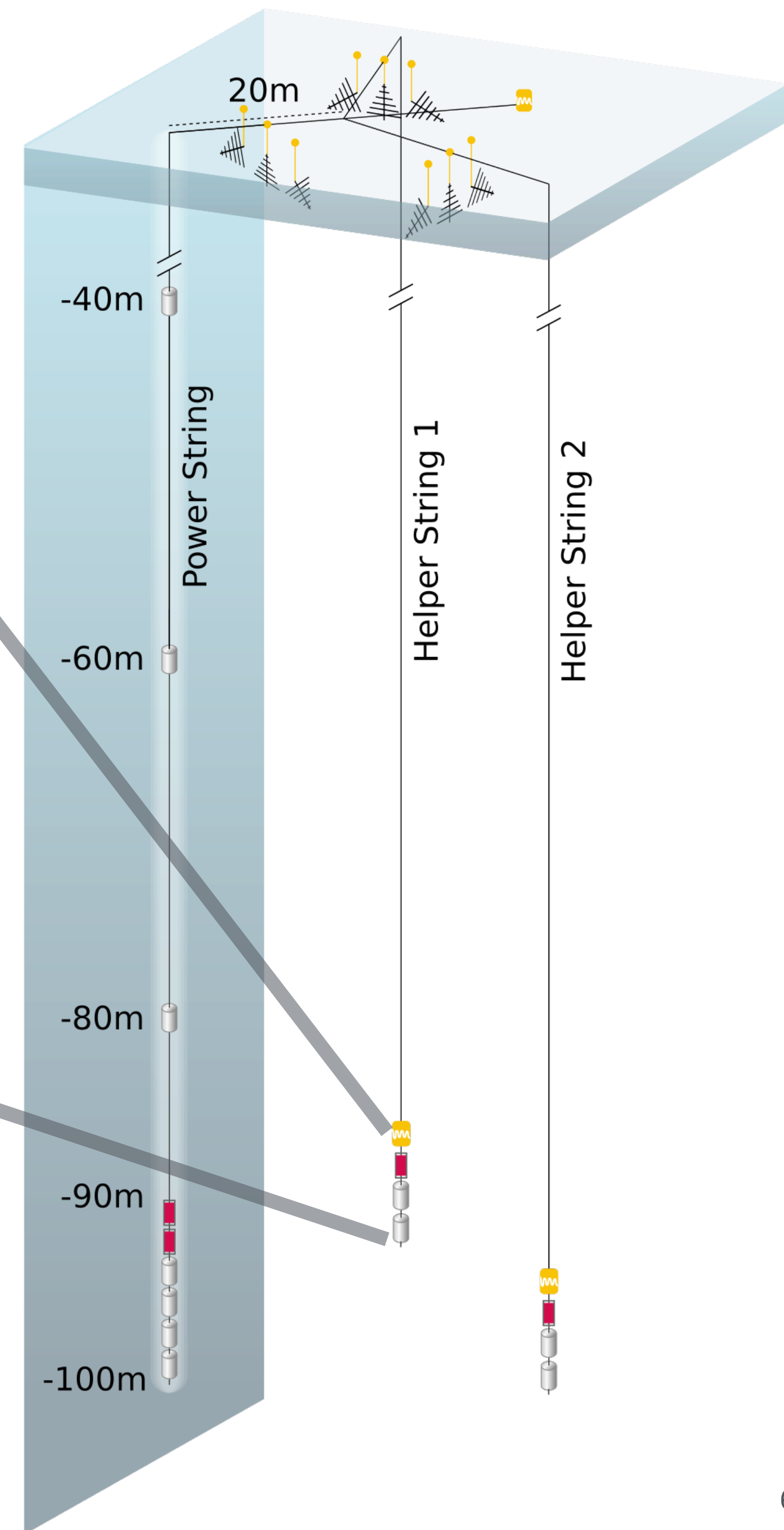
A single RNO-G Station



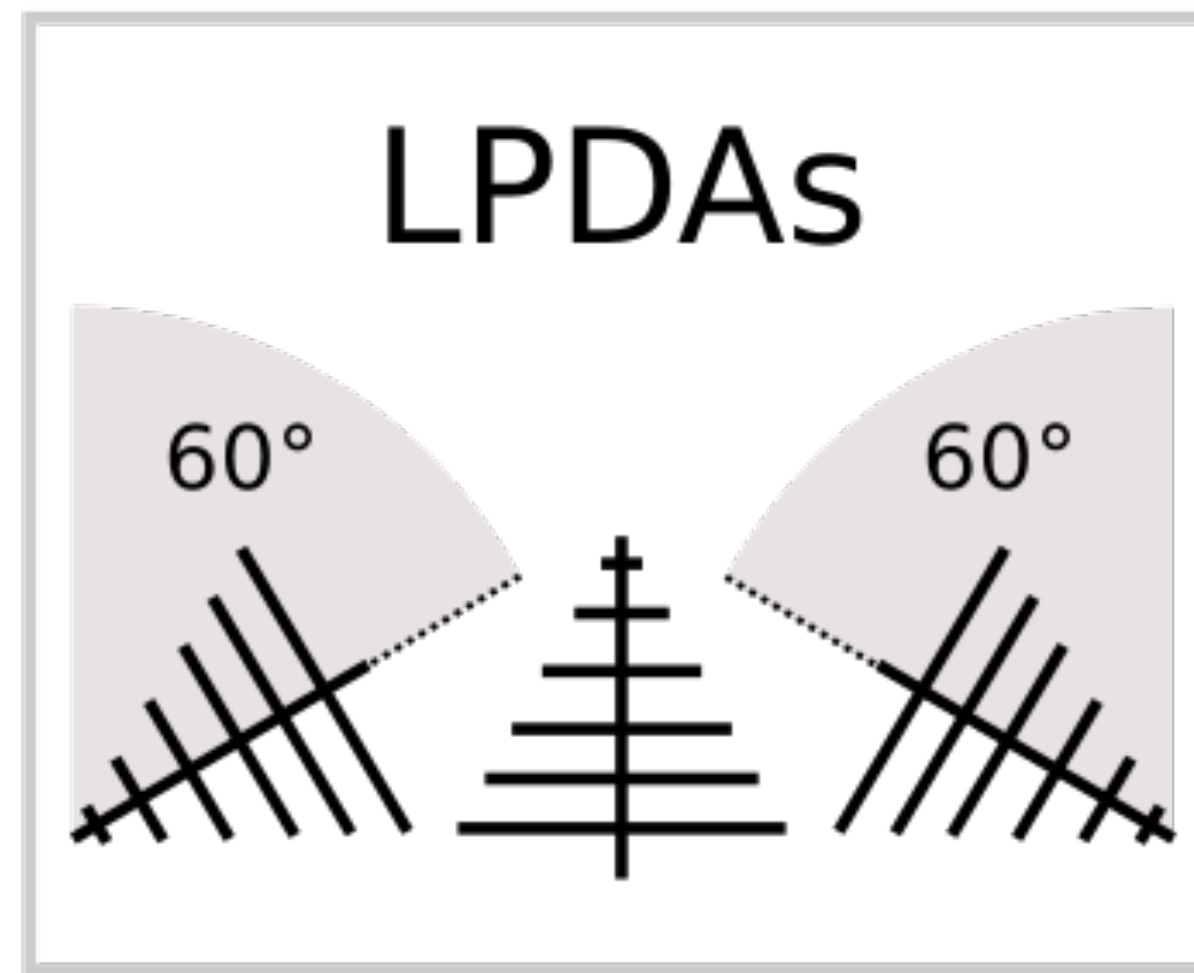
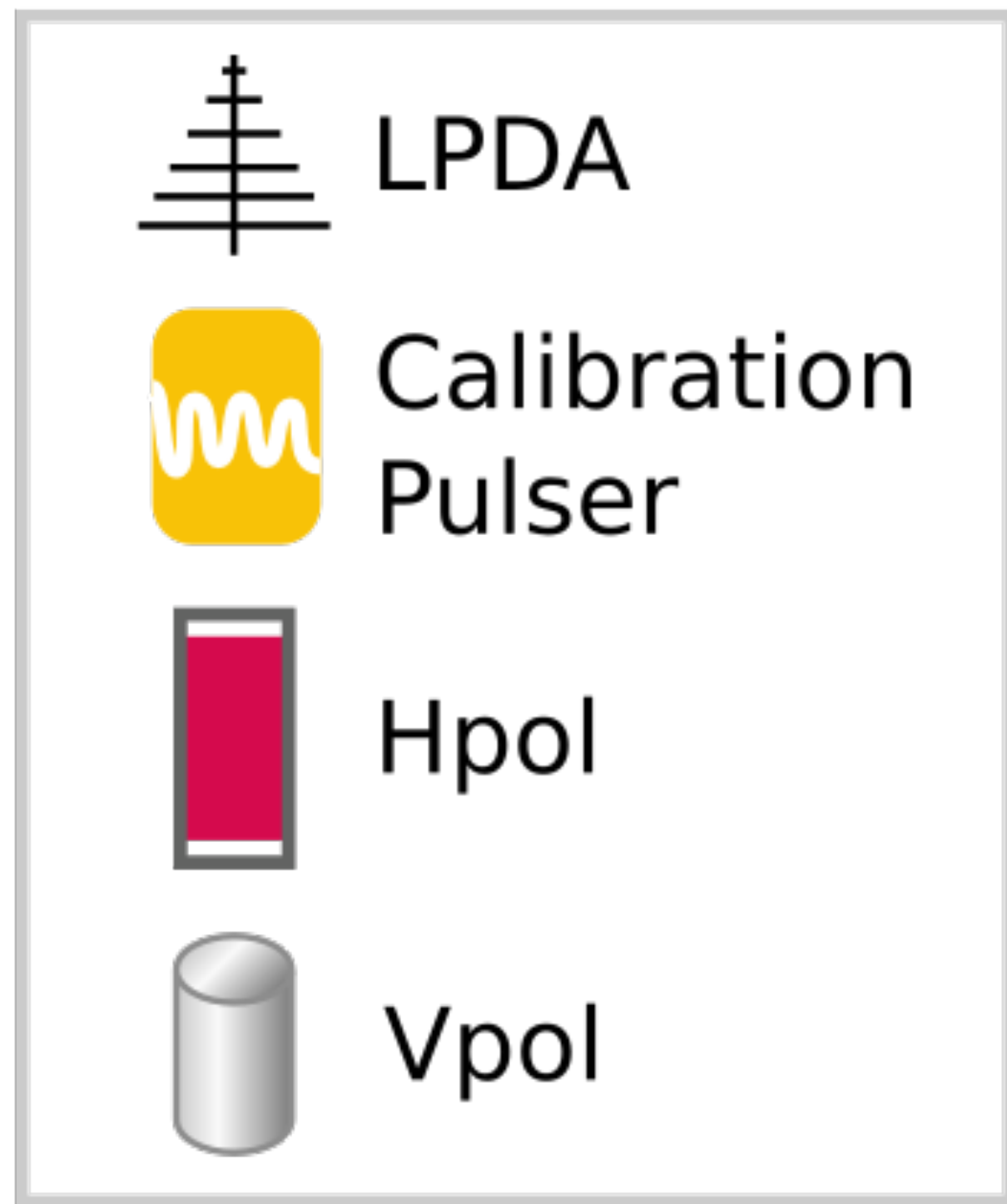
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Helper strings include Vpol and Hpol antennas for signal reconstruction, as well as calibration pulsers that transmit signals as needed.

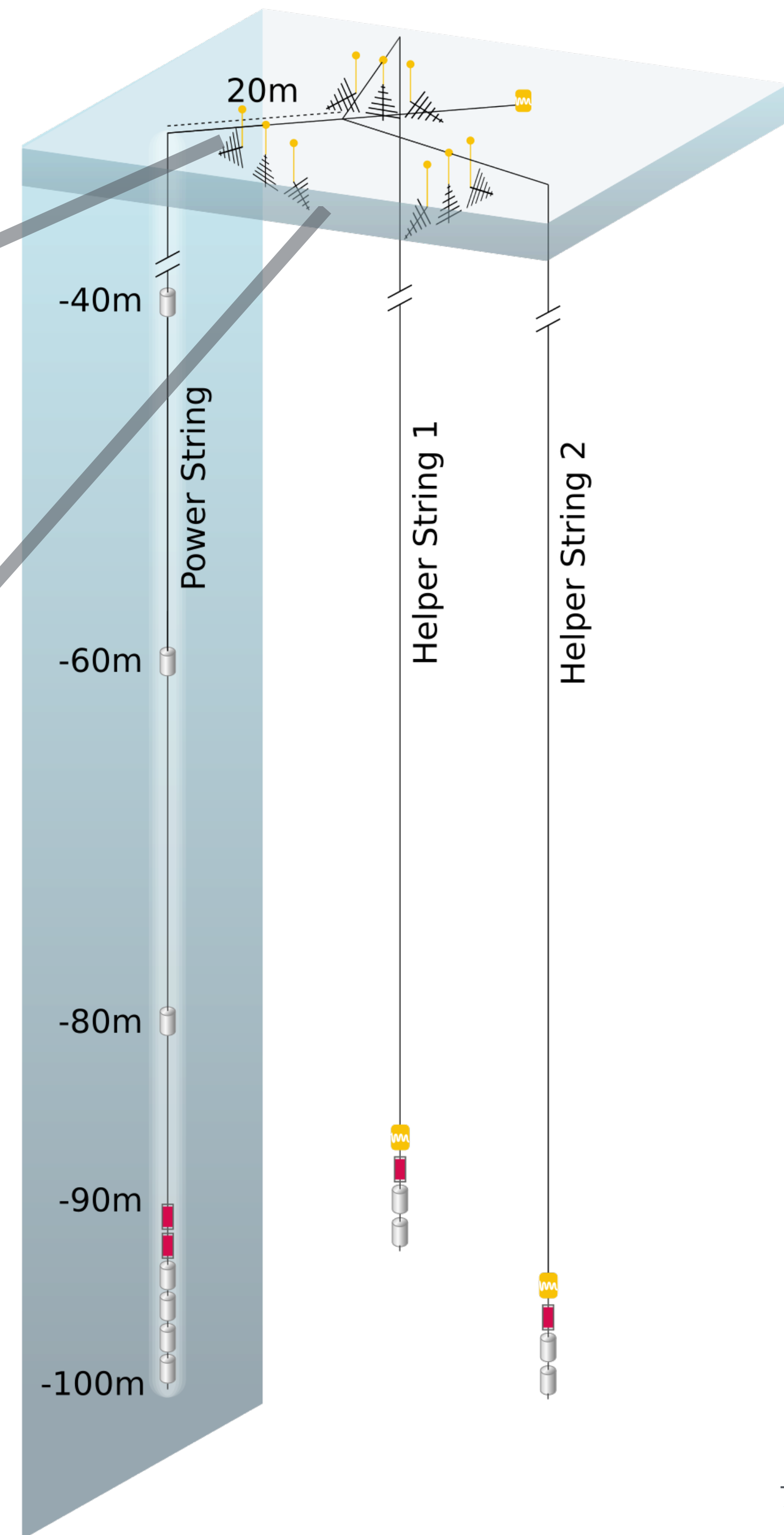


A single RNO-G Station



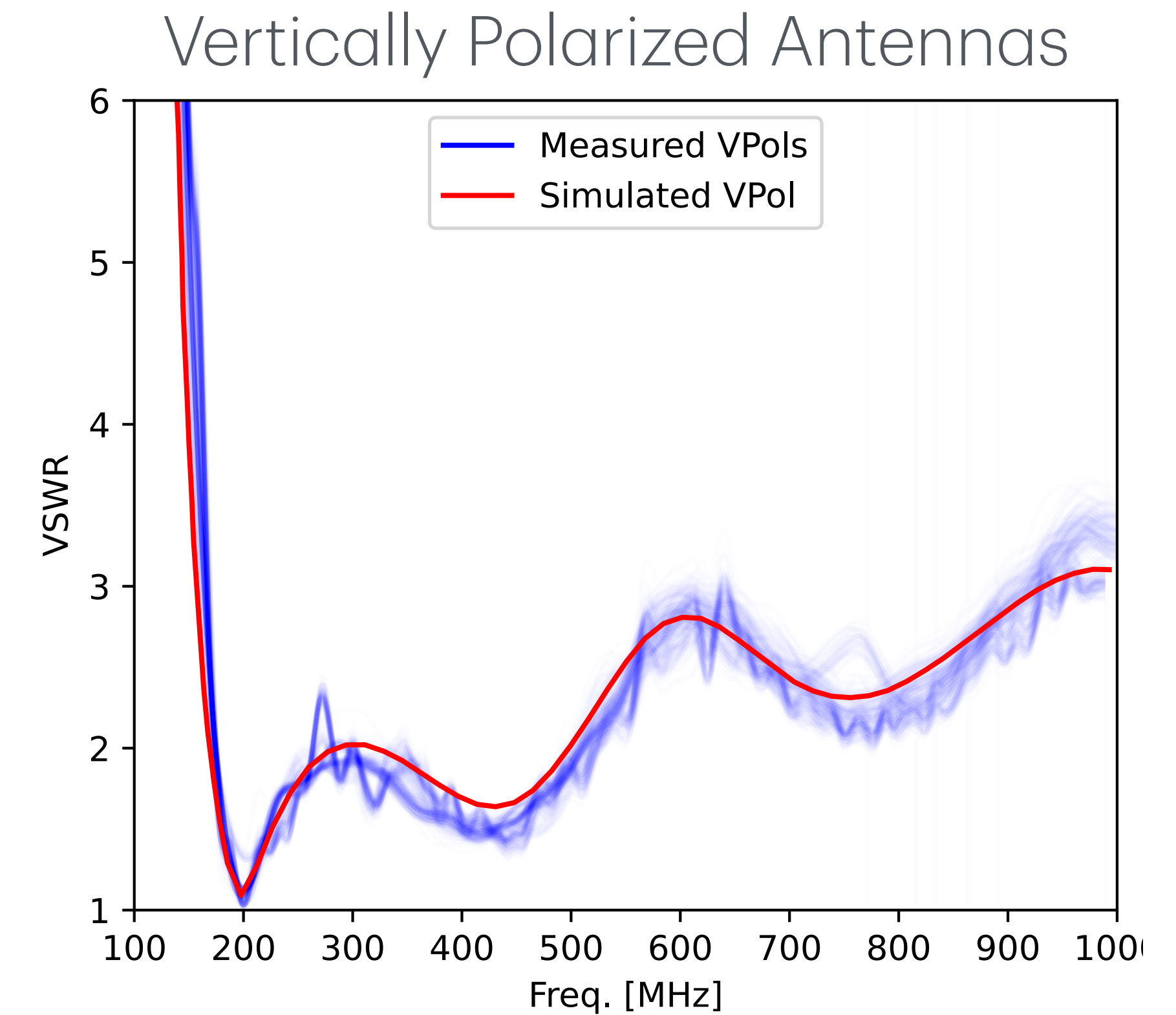
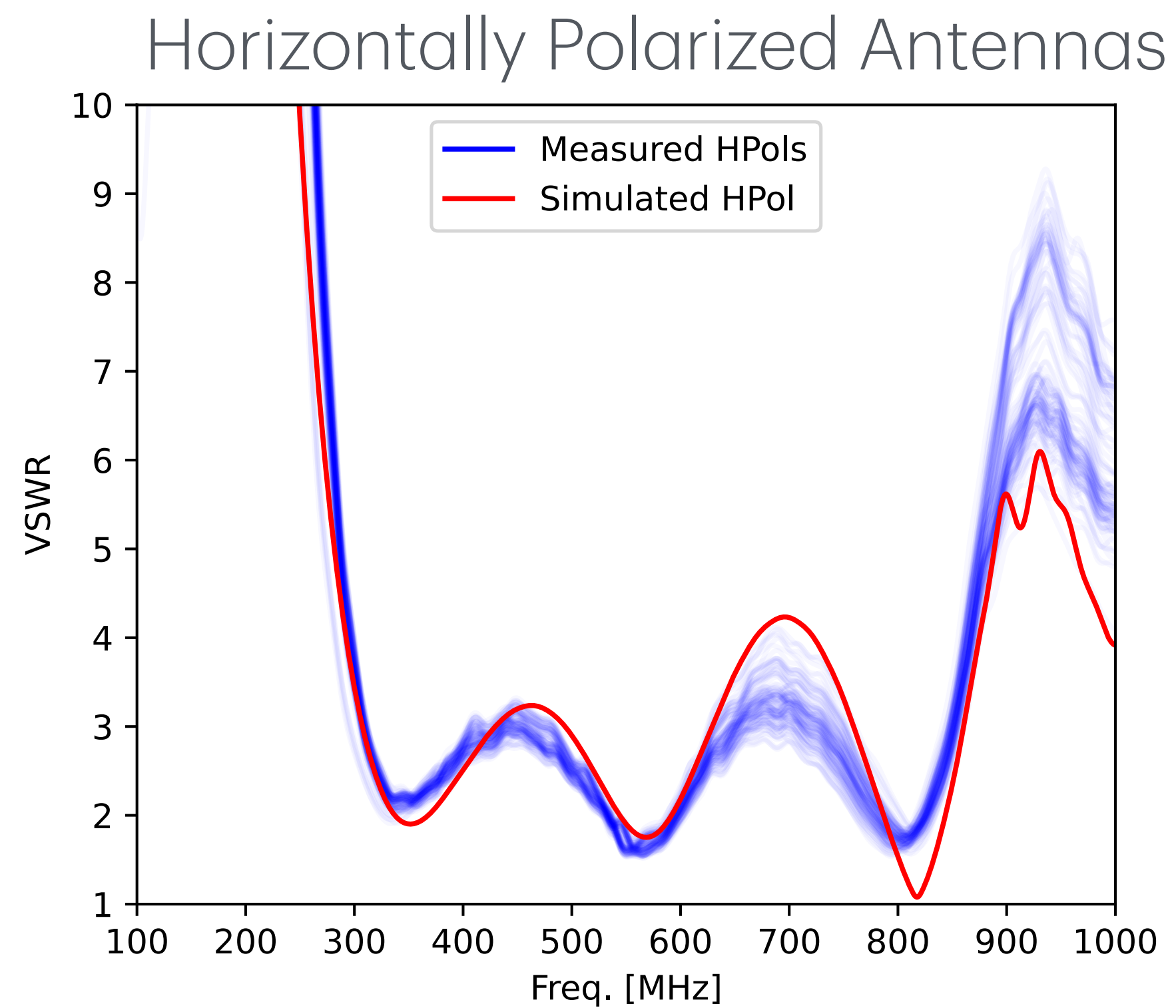
LPDAs at the surface help with signal reconstruction and act as a surface veto. They also contribute a small amount to neutrino effective volume.

Antennas types: Log-periodic dipole antenna (LPDA), Horizontally-polarized (Hpol), and Vertically-polarized (Vpol)



Modeling our antennas

Good agreement between simulation and measurement (in air) will improve our detector simulation accuracy. Currently, we extrapolate this to an in-ice response.

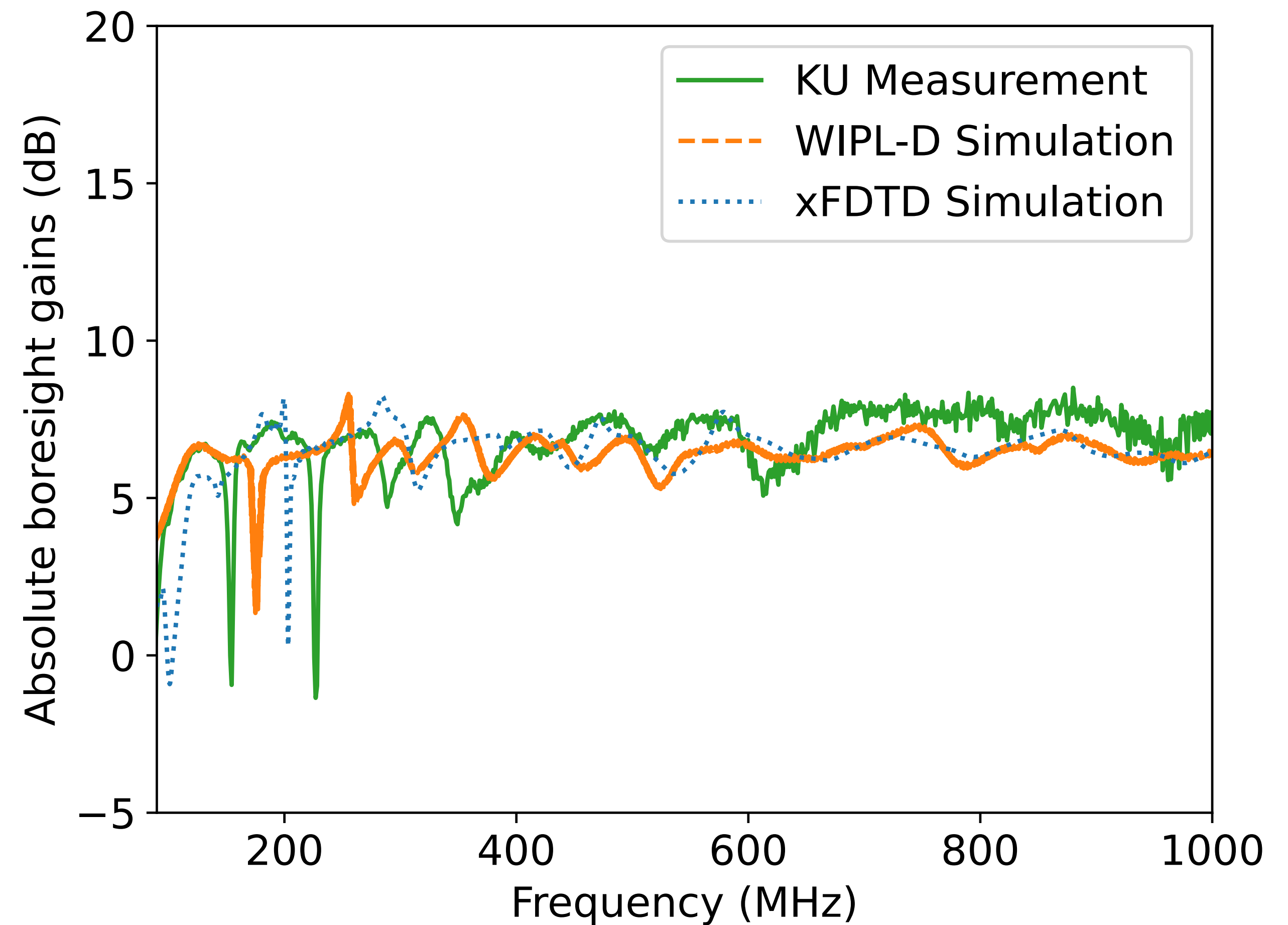


Credit: Bryan Hendricks (Penn State)

Modeling our antennas

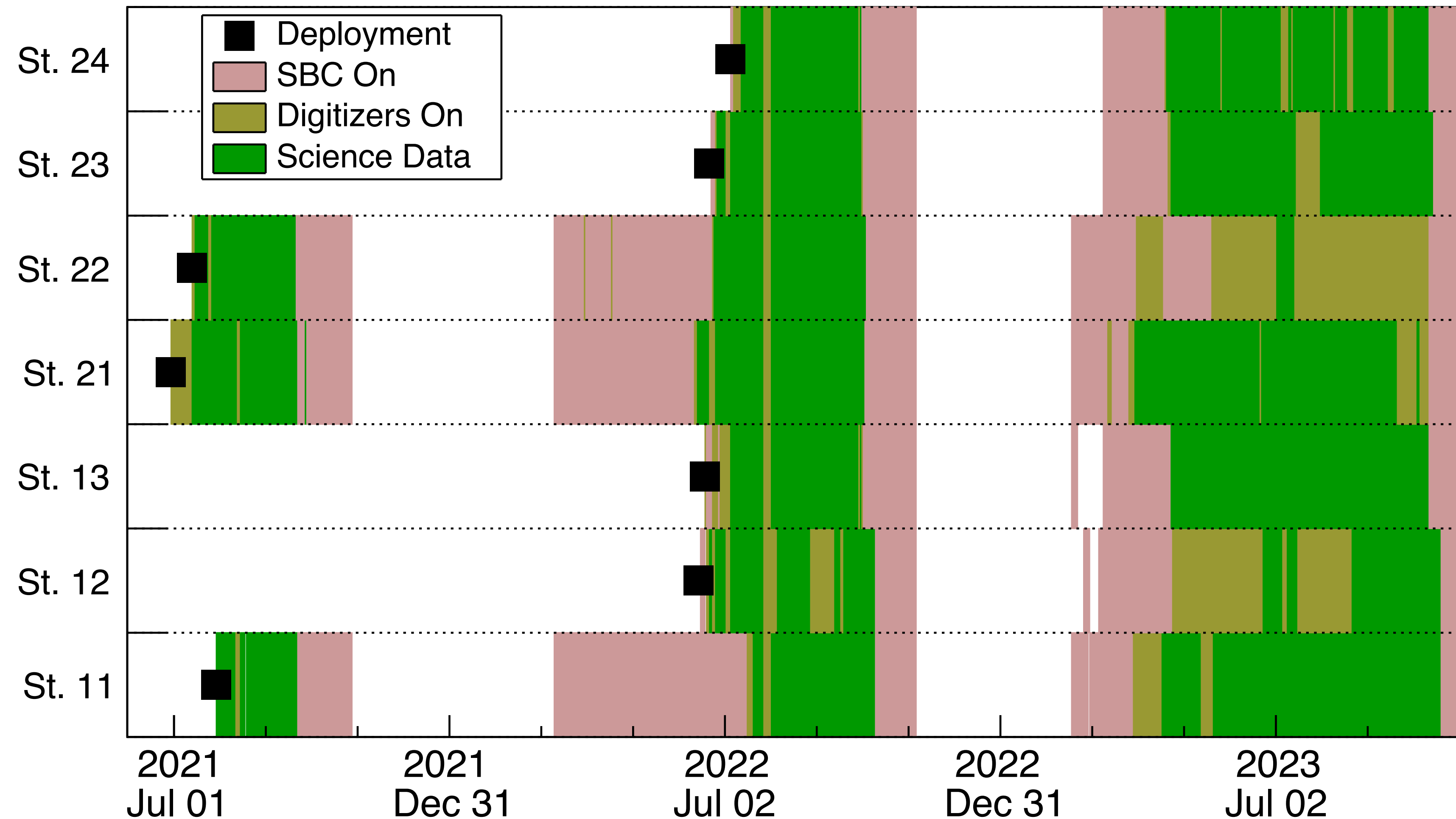
LPDAs

Credit: Mohammad F. H. Seikh



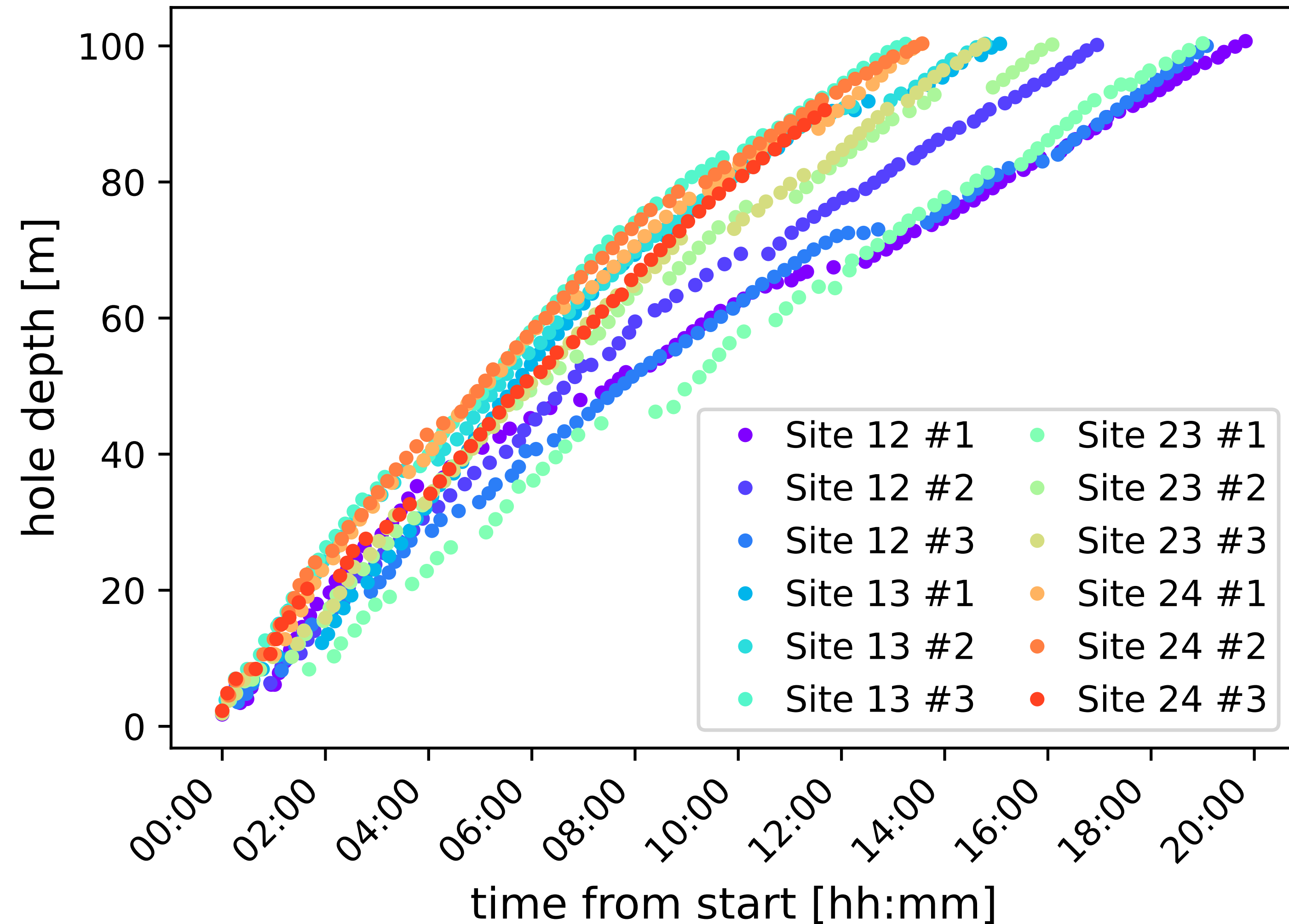
Deployment Progress

- RNO-G is solar powered; limited to taking data during summer months when sun and batteries can support power needs (~25 W per station)
- Investigating wind power as an alternative to extend lifetime into winter months (see Ryan Krebs' talk)

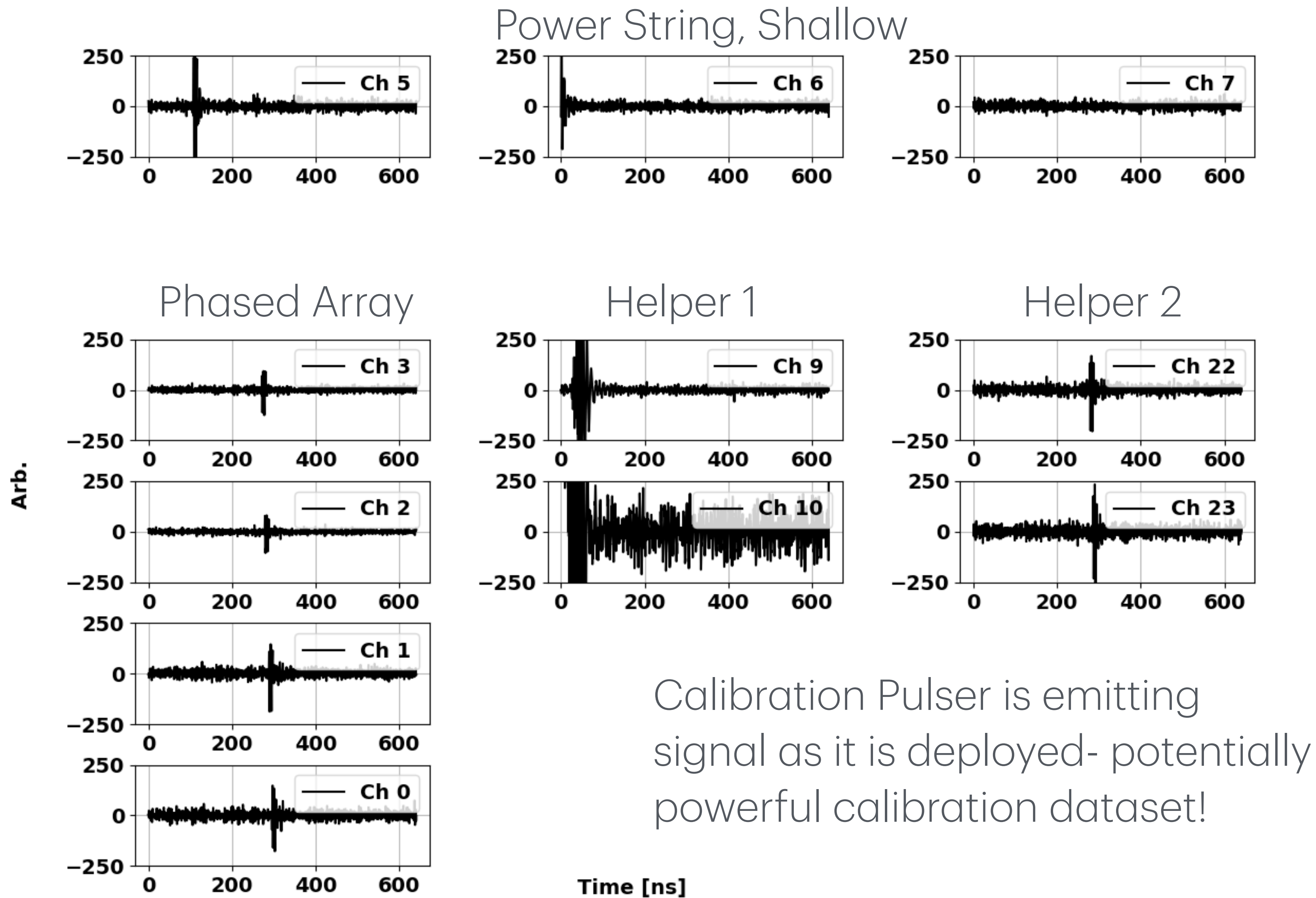


Deployment Plan: Summer 2024

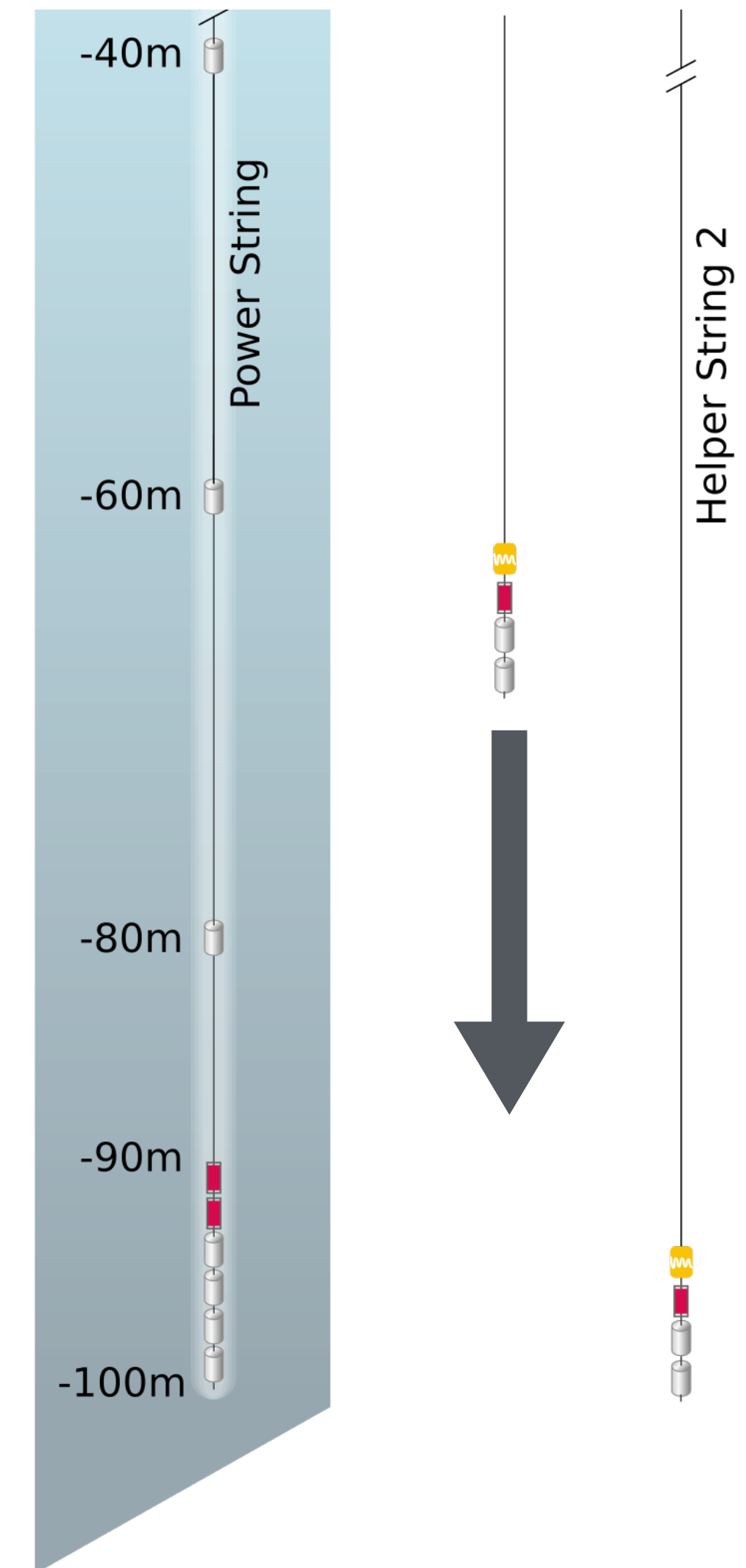
- **Drilling:** enough for 7 new stations
 - This would double our rate of hole production compared to previous years; seems feasible with drill upgrades and longer drill season
- **Retrofitting** existing stations with updated RADIANT trigger and digitizer board:
 - Improved surface trigger efficiency + fix for non-working filter banks
- **Calibration:** pulsing using nearby Saltzman hole to light up RNO-G Station 21 and test holes



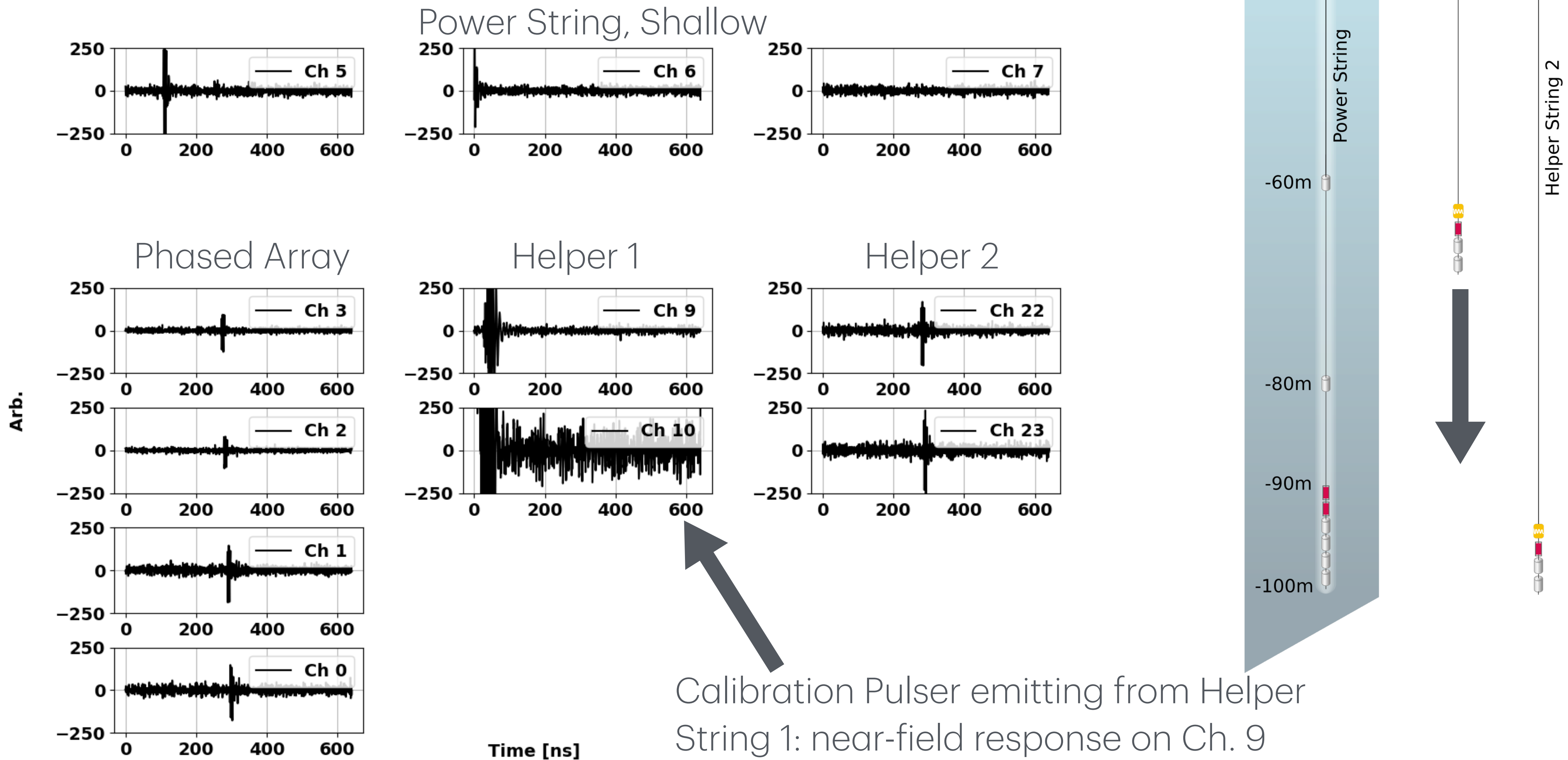
Example Calibration Event: Vpol channels



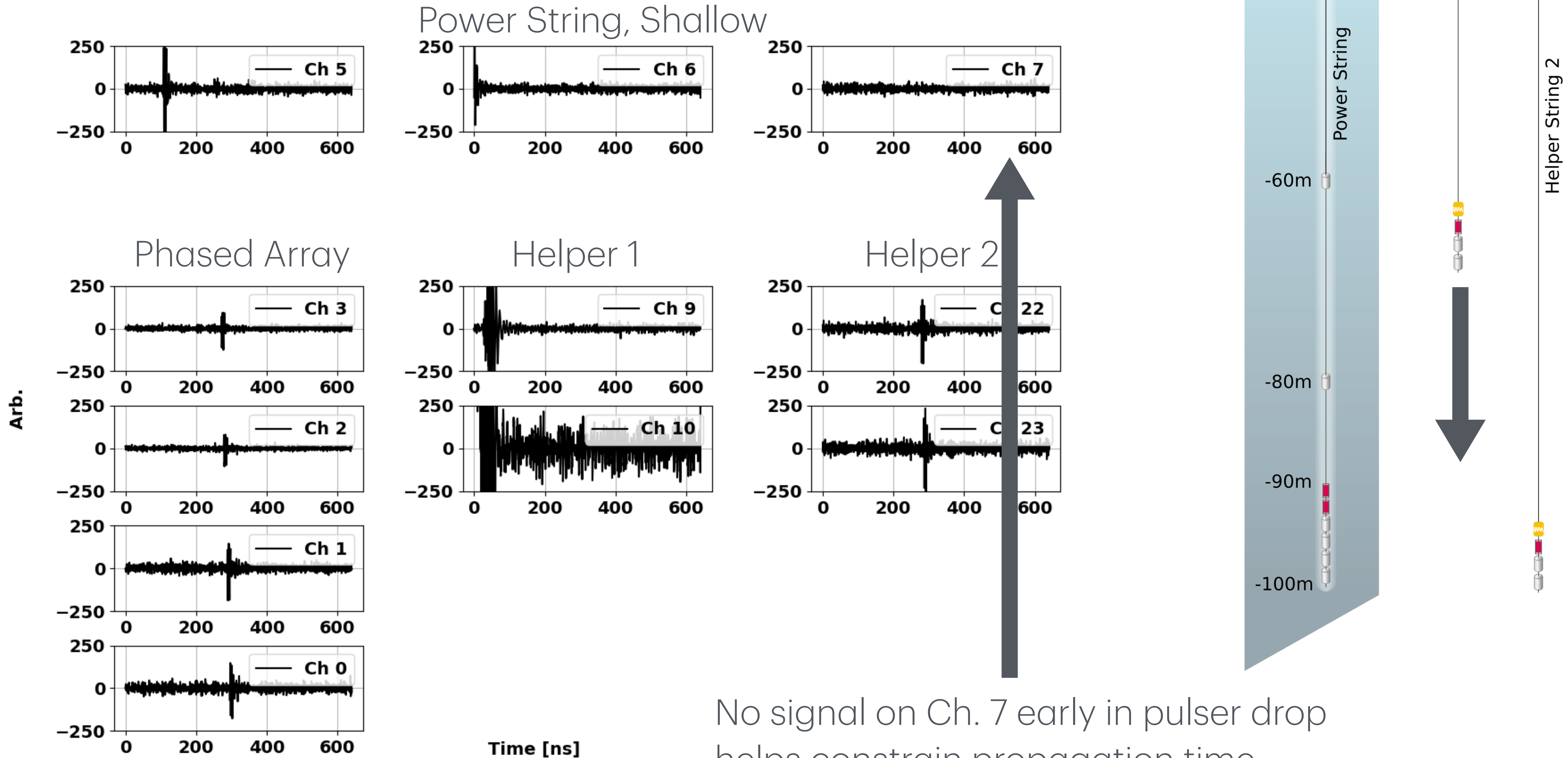
Calibration Pulsar is emitting signal as it is deployed- potentially powerful calibration dataset!



Example Calibration Event: Vpol channels

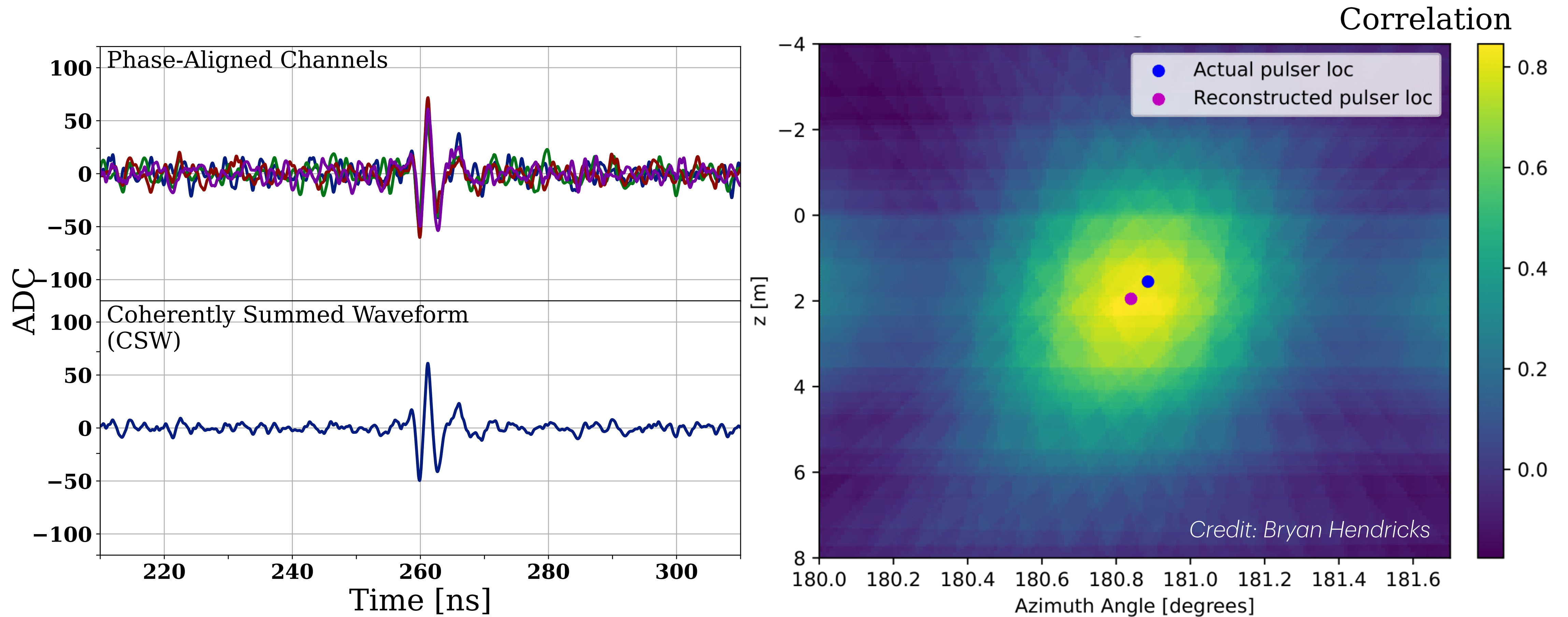


Example Calibration Event: Vpol channels



Calibration Pulsar Reconstruction

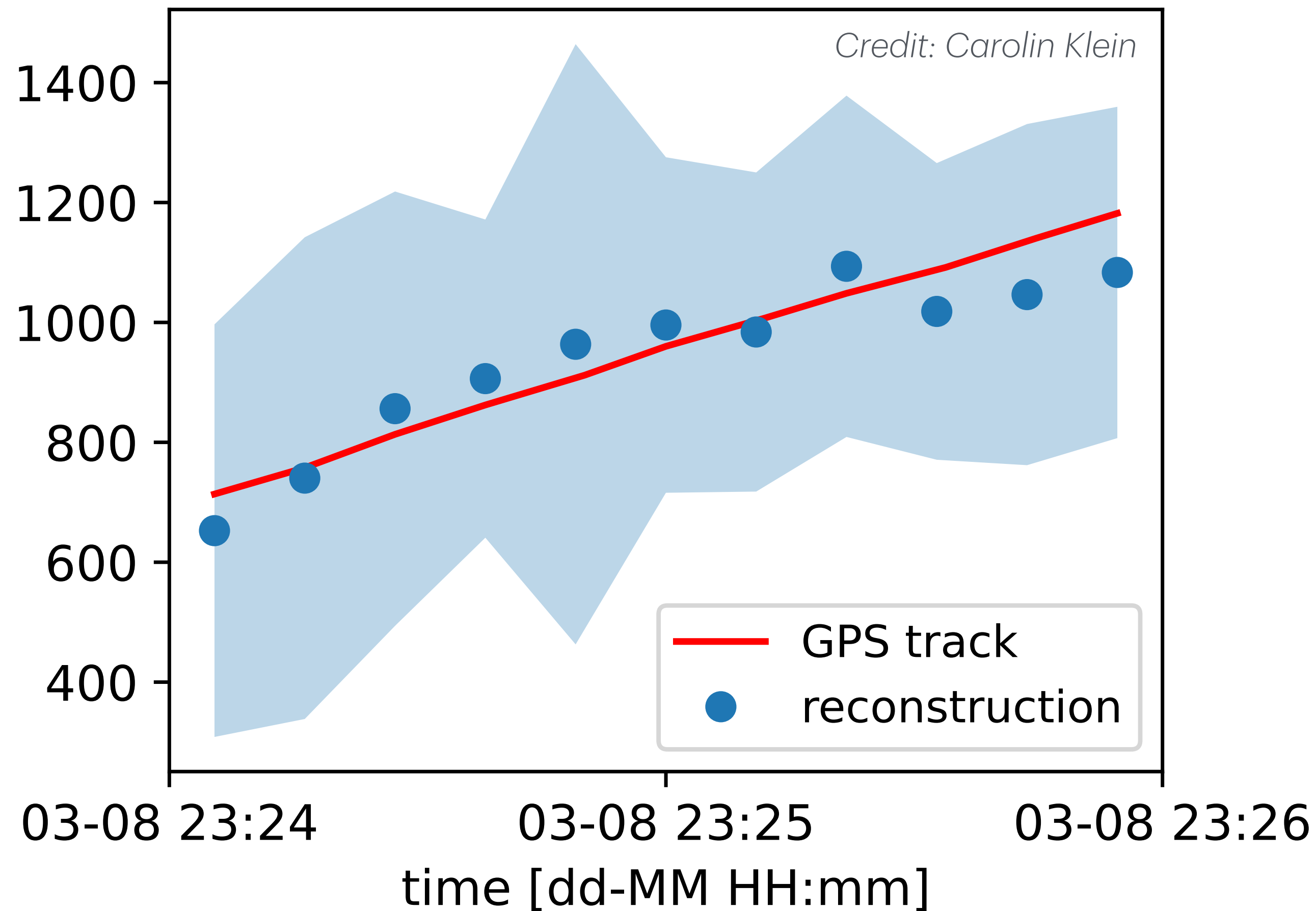
Local calibration pulser plays a crucial role in station position calibration (see Philipp Windischhofer's talk)



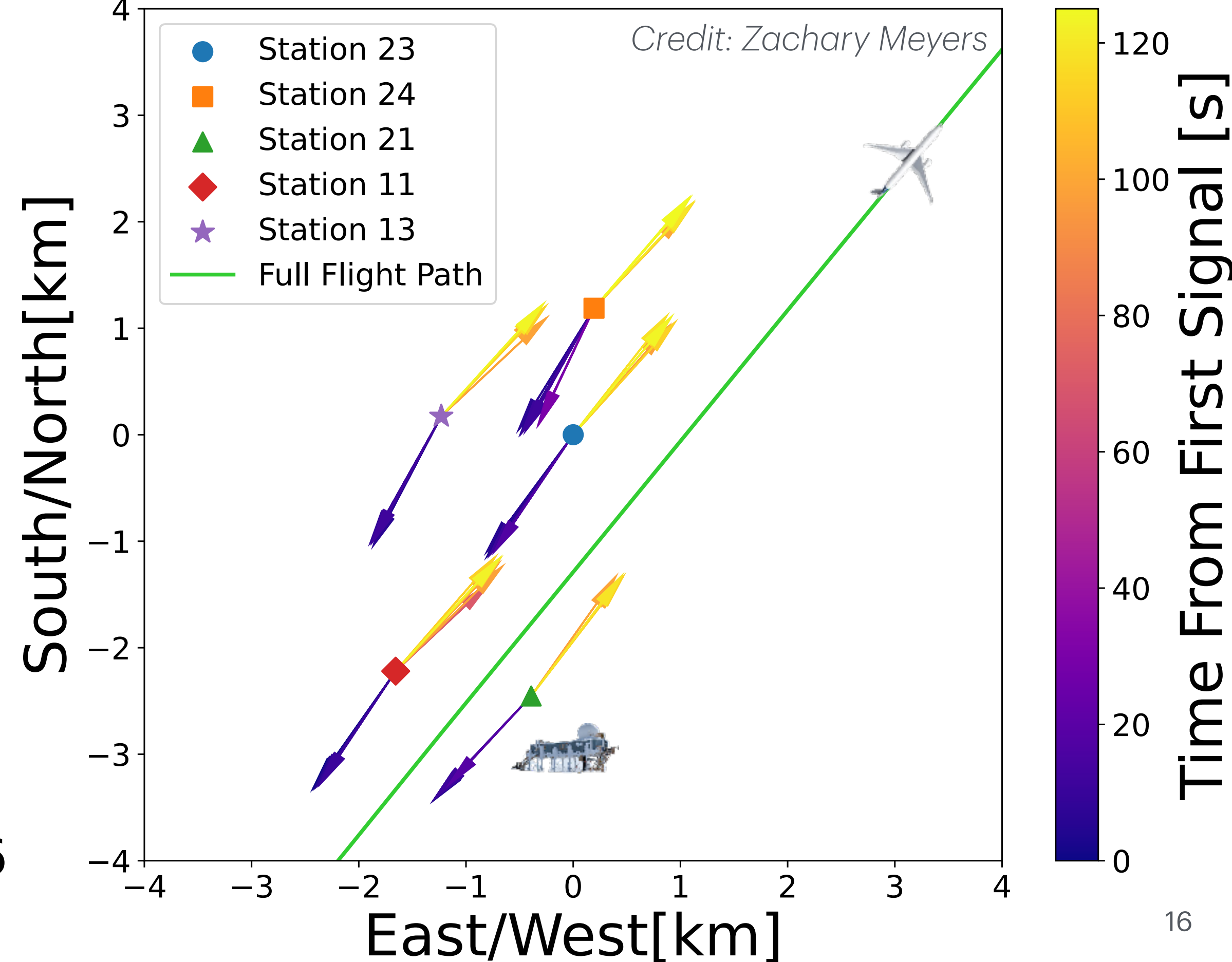
External Radio Sources

Anthropogenic signals like radiosonde weather balloons and airplanes are detected as well; this is a potential future calibration source

Radiosonde Balloon Reconstruction



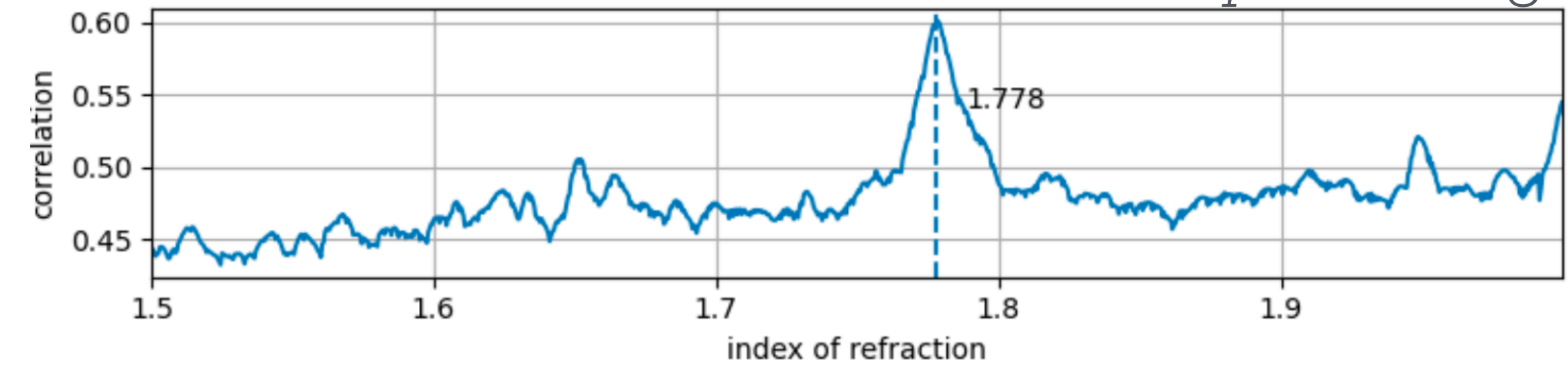
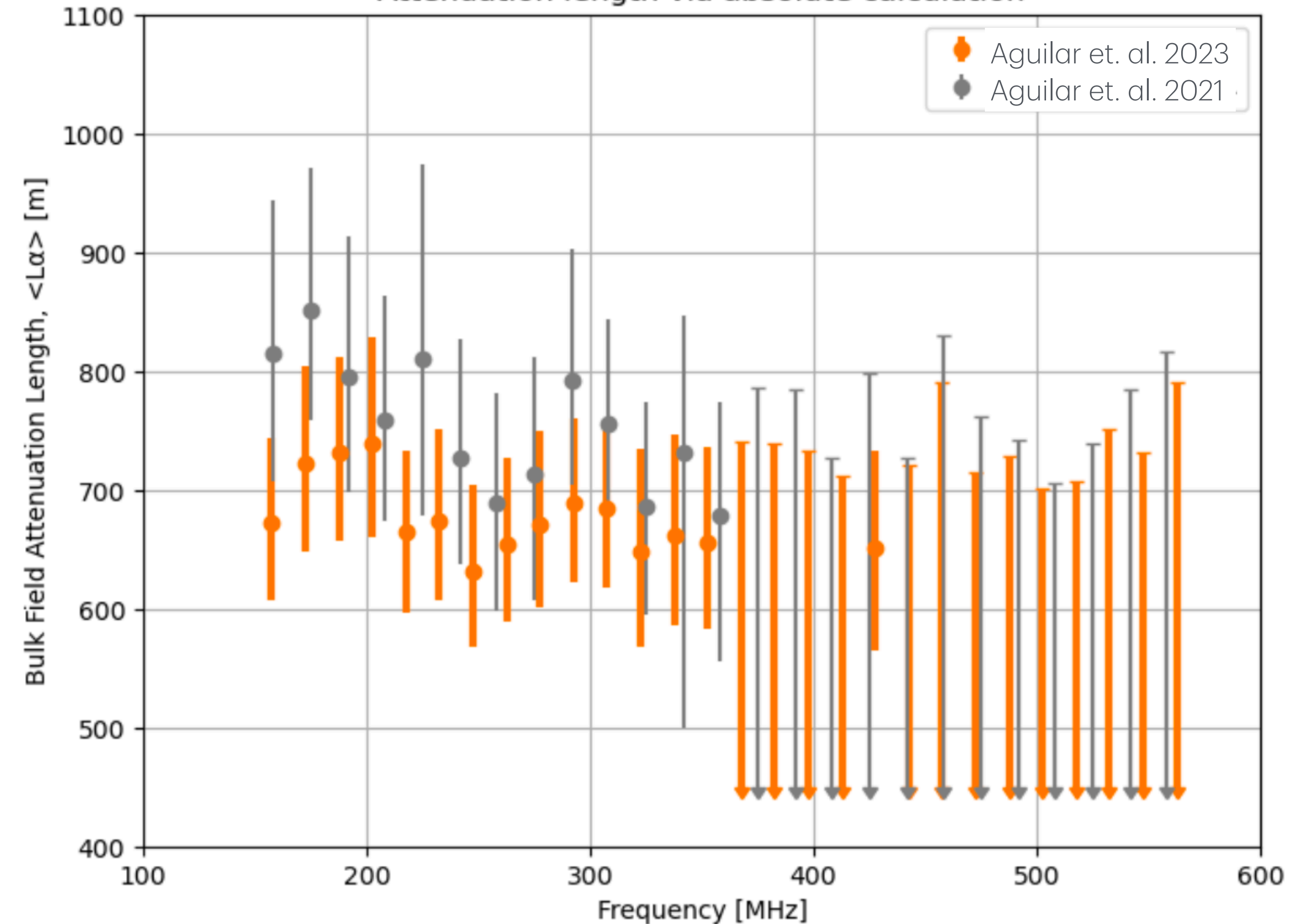
Airplane Reconstruction



Ice Measurements

Christoph Welling

Attenuation length via absolute calculation



- Lots of effort has gone into studying the ice around RNO-G; critical for successful calibration and reconstruction
- The attenuation length and bulk index of refraction for radio frequencies have been measured; index of refraction in firn still a work in progress (see Philipp's talk)
- No significant birefringence effects observed on vertically propagating signals

Building towards the future

- RNO-G is currently being constructed and is carefully building tools needed to conduct a neutrino search
- Currently using cosmic rays to determine instrument performance (see Anna Nelles' talk)
- Lots of advancements have been needed to make this happen, on every front: drilling, antenna design, hardware/firmware, and calibration
- 35 stations + 5 years of data will make RNO-G sensitive to most optimistic cosmogenic flux models

