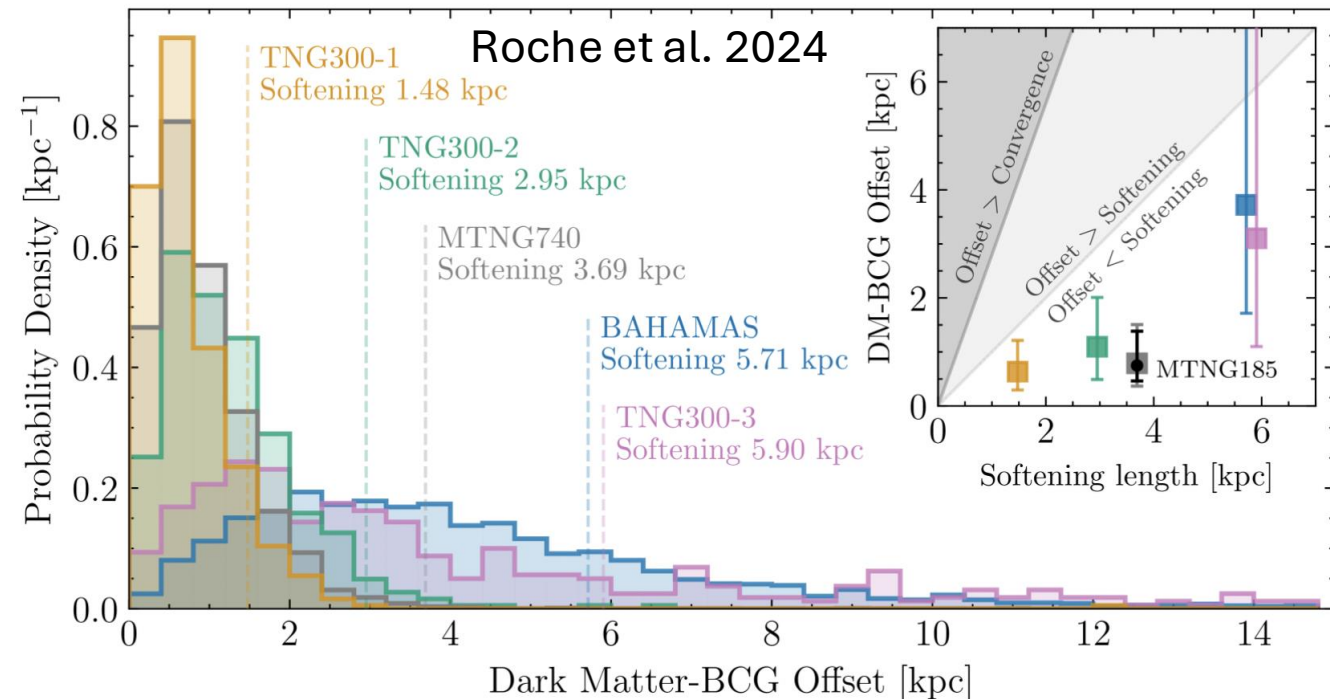
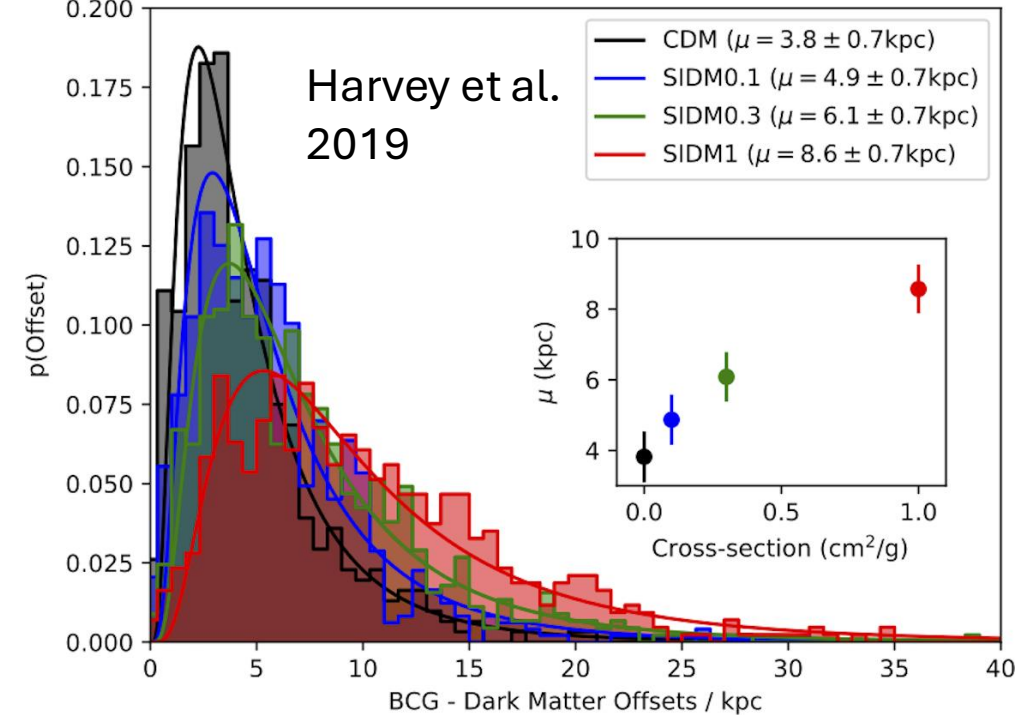


# **The Peculiar Motions of the BCGs in SPT Clusters**

Raven Gassis + SPT Collaboration – mm Universe  
– June 23<sup>rd</sup>, 2025

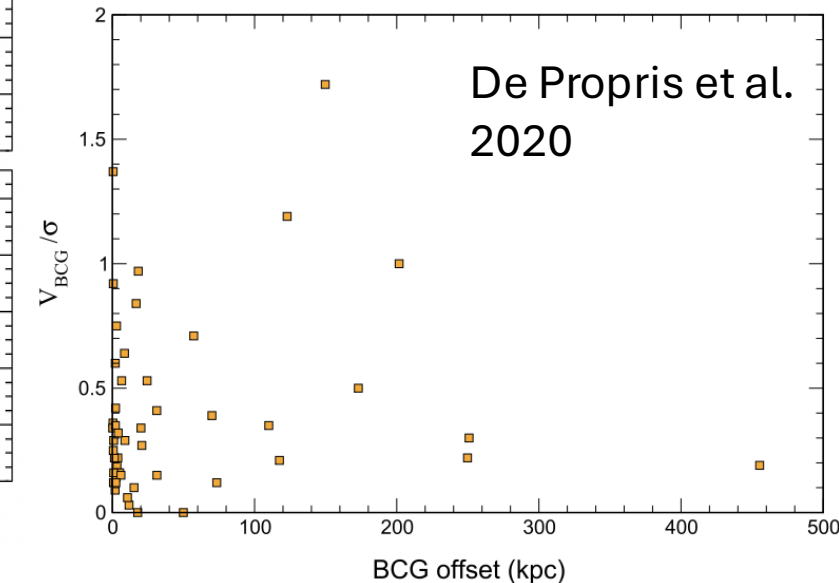
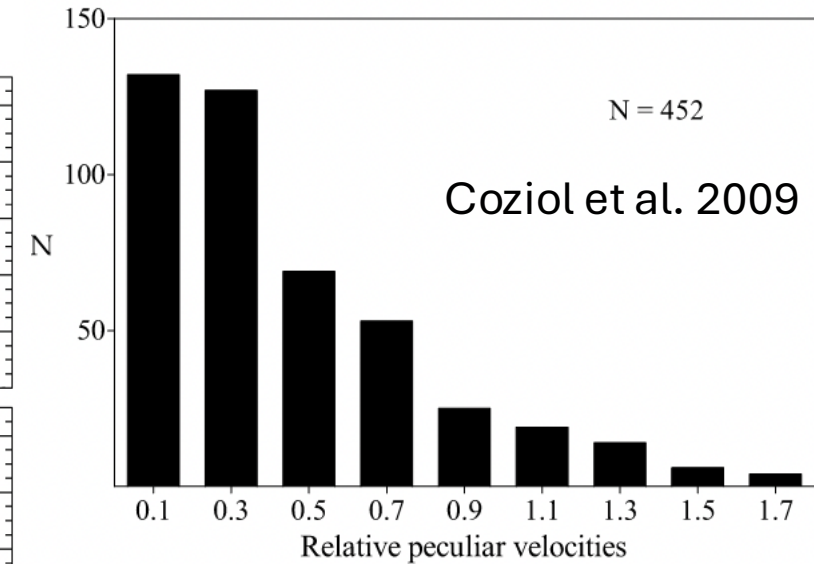
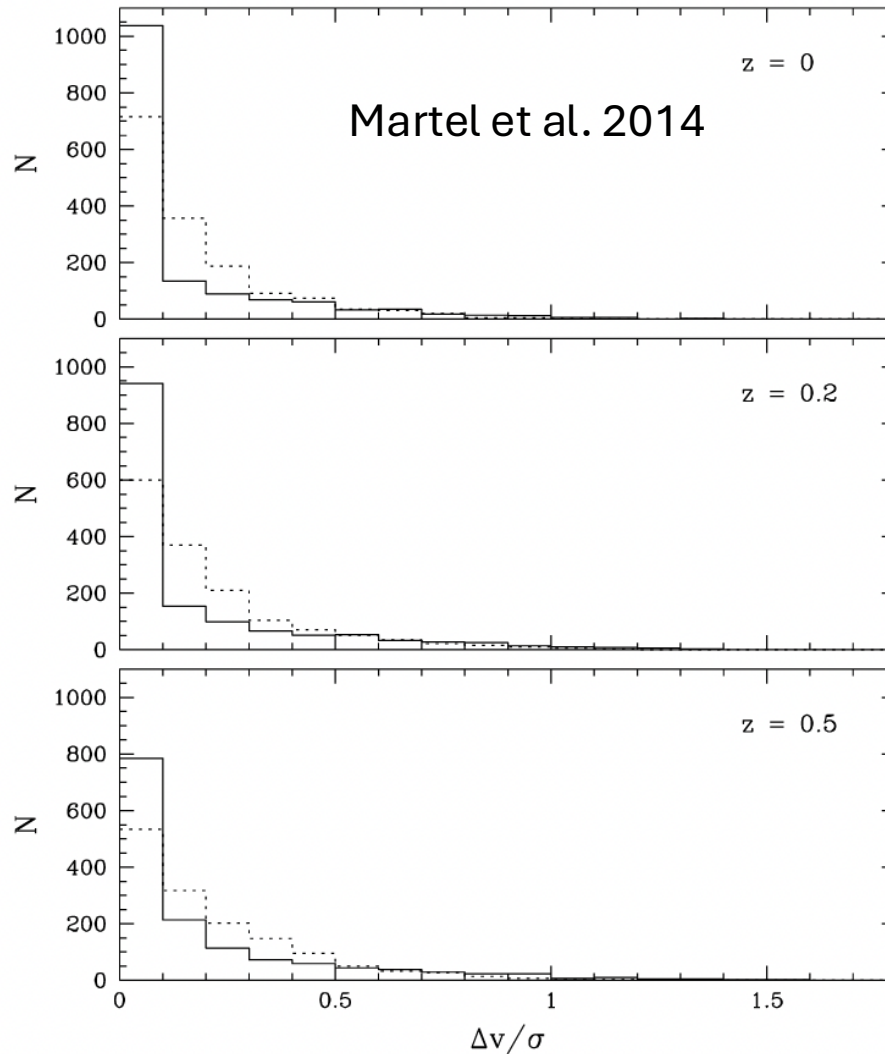
# The Central Galaxy Paradigm

- Galaxy clusters form via hierarchical merger processes
- Dark Matter Halo defines cluster structure and gravitational potential
- Central Galaxy Paradigm (CGP): the Brightest Cluster Galaxy (BCG) should be at rest in the center of its host galaxy cluster
- BCG formation framework: a high gas accretion phase at high  $z$ , a peak star formation phase and moderate  $z$ , and a merger and accretion phase at low
- Departure from the CGP  $\rightarrow$  post-galaxy formation mergers or alternative DM models



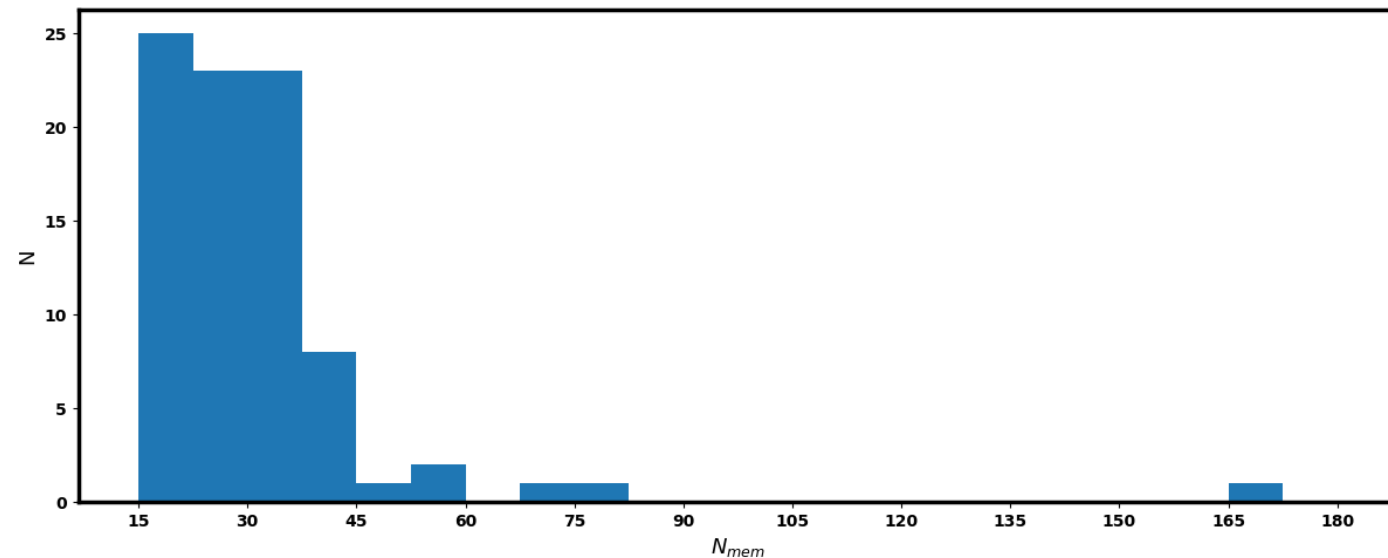
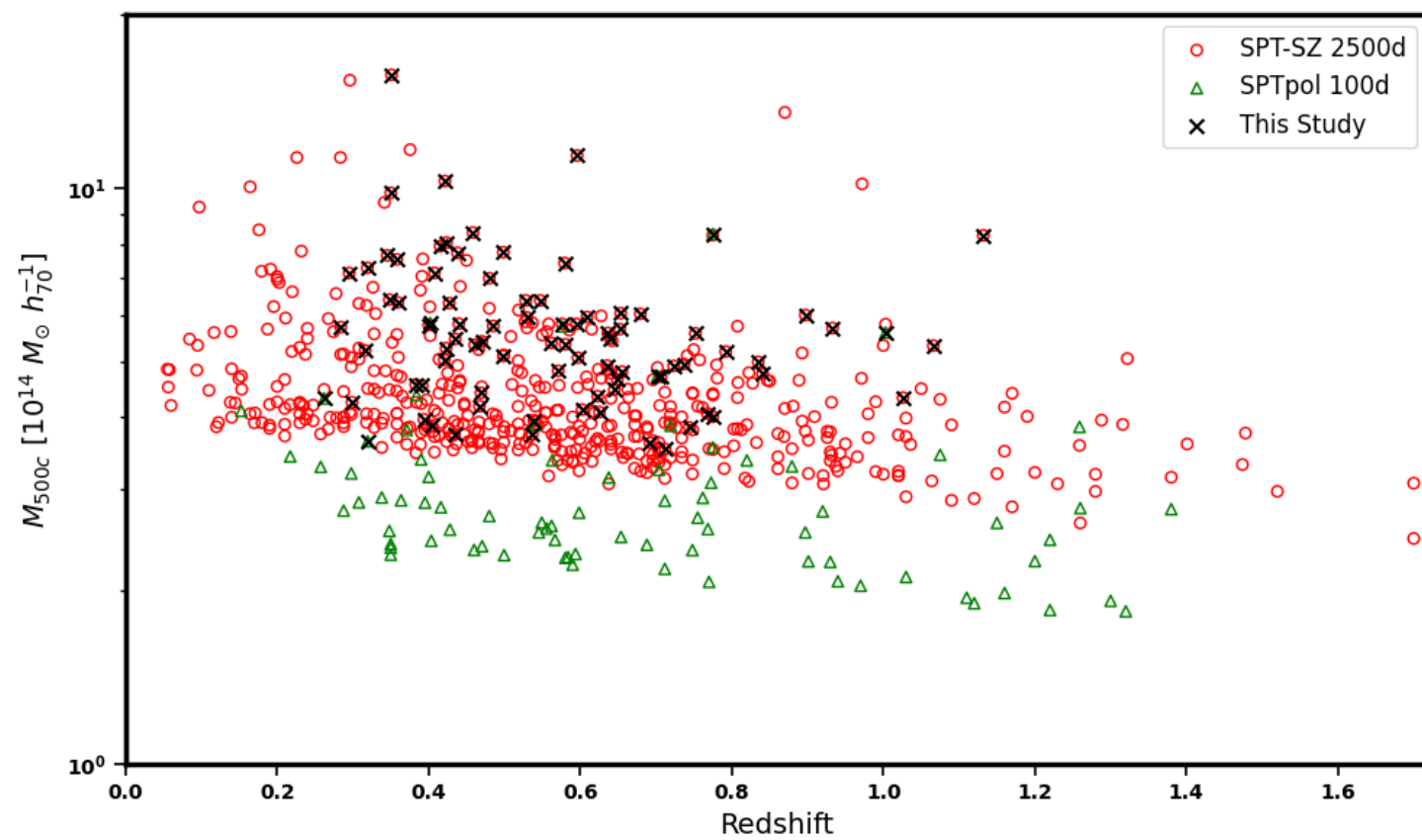
# Peculiar Velocity as a test of $\Lambda$ CDM

- 2 tests of CGP:
  - BCG peculiar velocity offset
  - BCG centroid offset
- De Propriis sample:
  - 164 clusters from Planck Early SZ sample ( $z \leq 0.35$ ) + a flux-limited X-ray sample of 100 clusters ( $z \leq 0.3$ ).
- Martel Simulations:
  - Dashed line => values derived from all matter
  - Solid lines => values from galaxies only.
- Coziol sample:
  - 452 Abell clusters w/  $N_{\text{mem}} > 10$  ( $z \leq 0.2$ )



# Project Subsample

- SPT spectroscopic catalogues:
  - $v_p$ : BCG peculiar velocity
  - $\sigma$ : cluster velocity dispersion
- BCG selection:
  - using automated algorithm (Noble et al. in prep, Somboonpanyakul et al. 2022)
  - lensing arcs
  - X-ray peaks
  - By-eye inspection
- 85 clusters with  $N_{\text{mem}} > 15$ .



# X-ray Subsample

$$c_{[R500]} = \frac{Flux(r < 0.2 * R500)}{Flux(r < R500)}$$

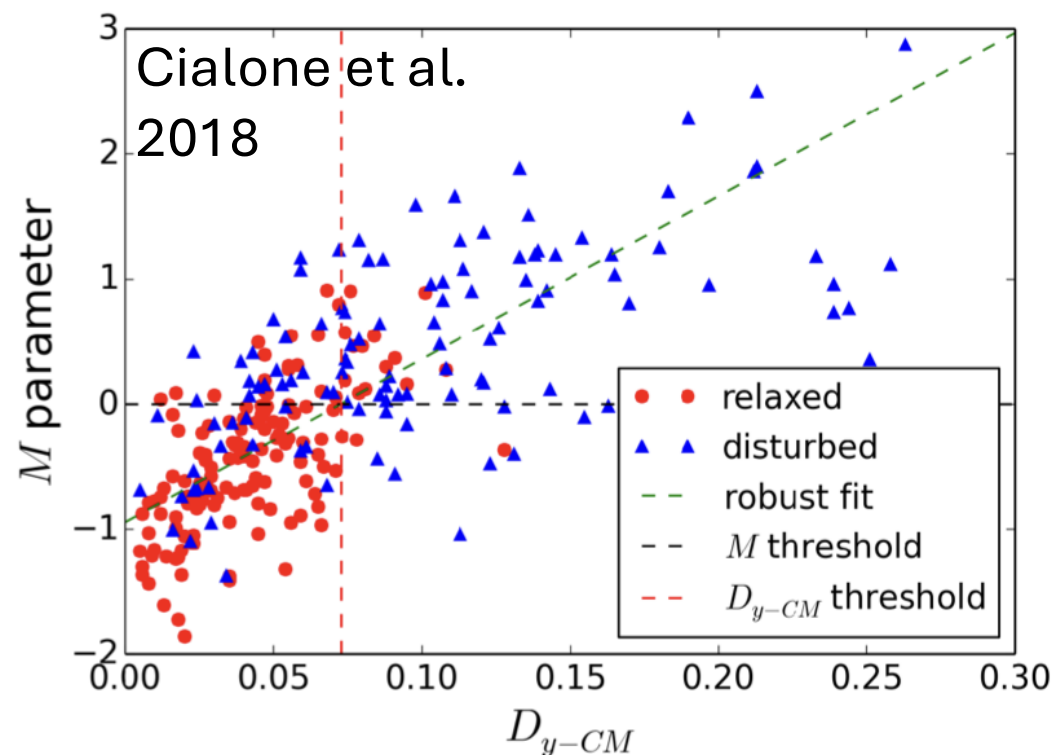
$$D[kpc] = d_A \sqrt{\frac{((\alpha_{ICM} - \alpha_{BCG}) * \cos(\delta_{BCG}))^2}{+(\delta_{ICM} - \delta_{BCG})^2}}$$

- 73 Clusters have either XMM and/or Chandra follow-up data

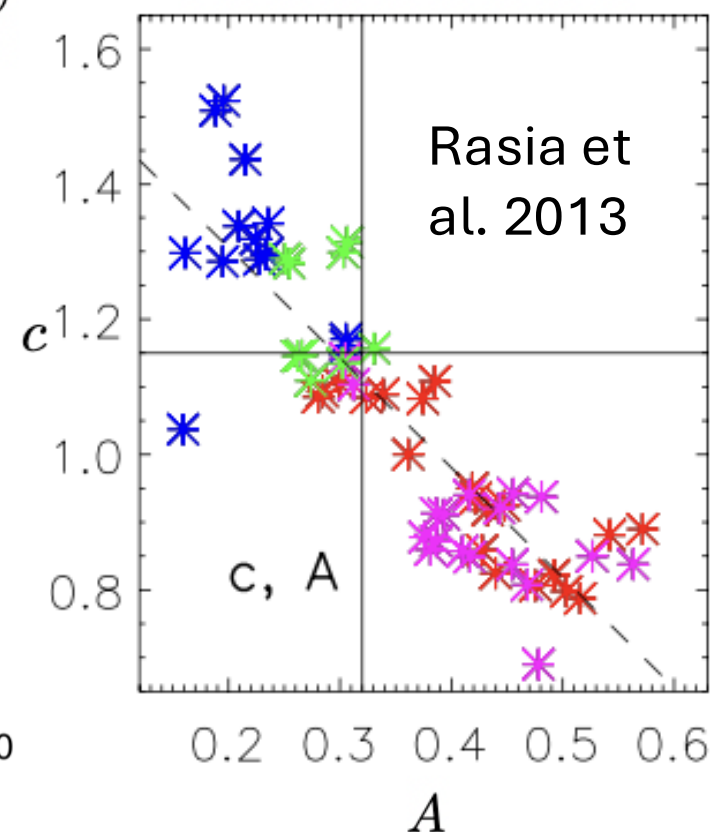
## Measure:

- A: asymmetry parameter
- c: concentration parameter
- D: centroid separation between BCG and X-ray distribution
- M: combined relaxation parameter
- $L_{par}$ : cut value
- $q_{par}$ : 1<sup>st</sup>/3<sup>rd</sup> quartile ( $par < m_{par} / par > m_{par}$ )

$$M = \frac{1}{N_{par}} \sum_{par} (A_{par} \times \frac{par - L_{par}}{|q_{par} - m_{par}|})$$

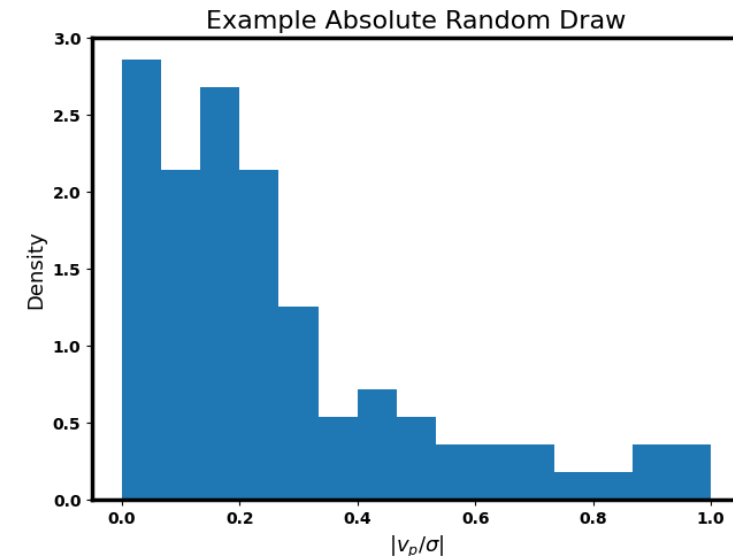
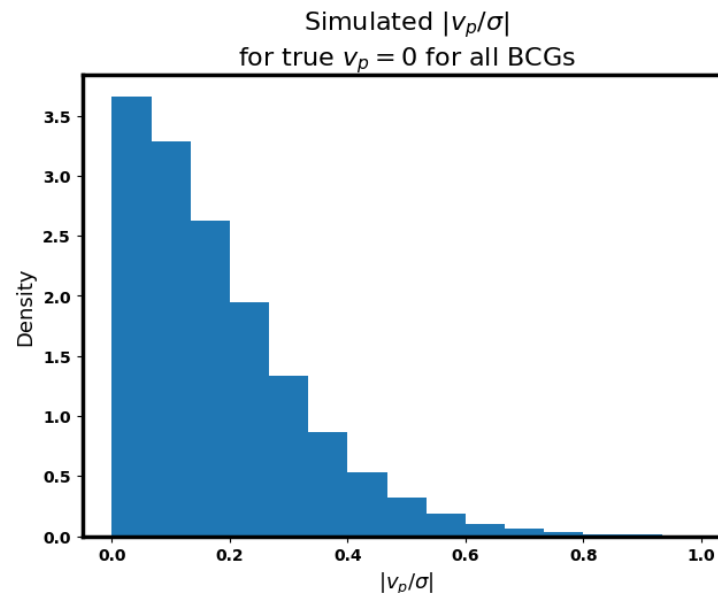
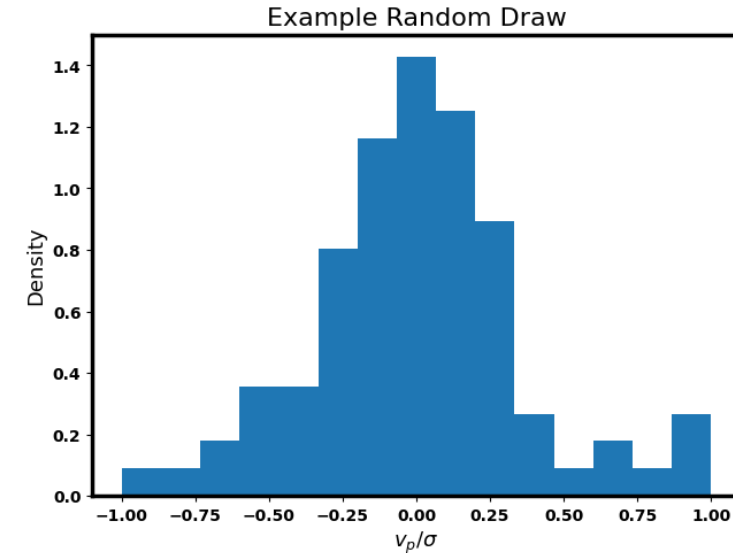
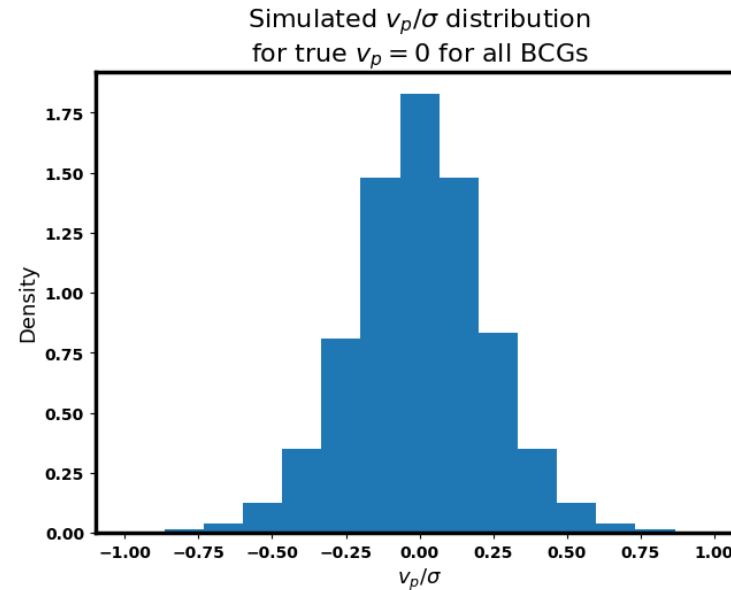


$$A = \frac{\sum (|I - R|)}{\sum I}$$



# Measuring BCG Peculiar Motion

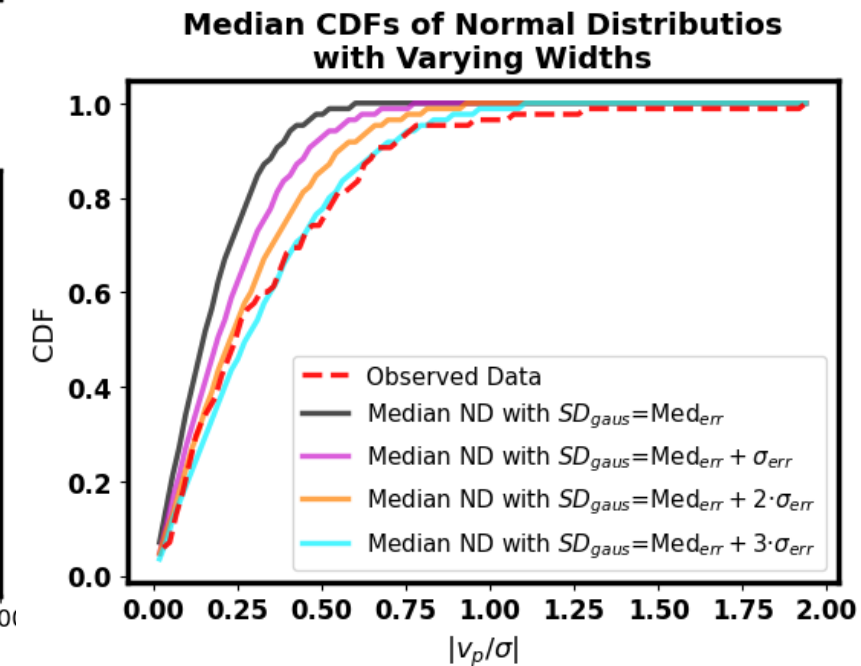
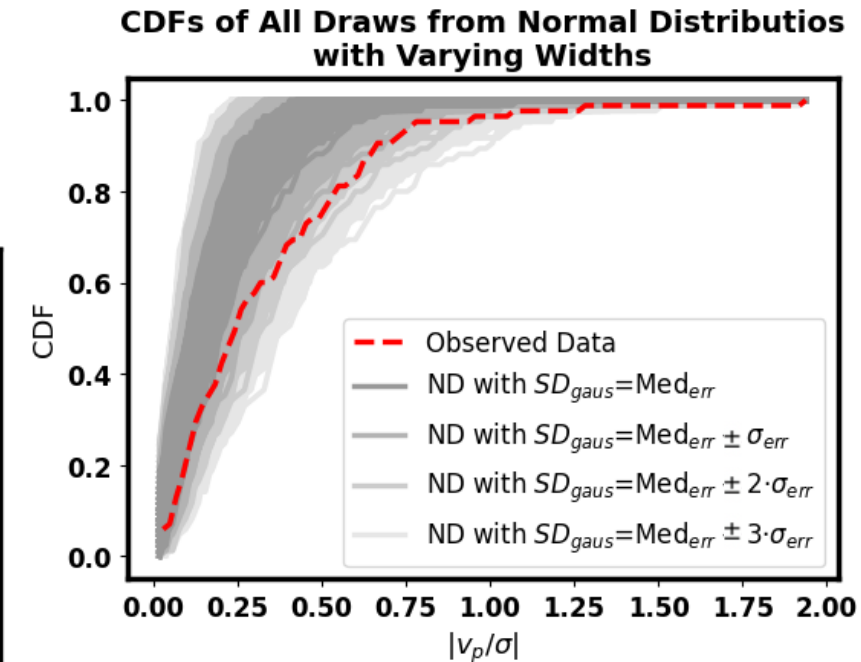
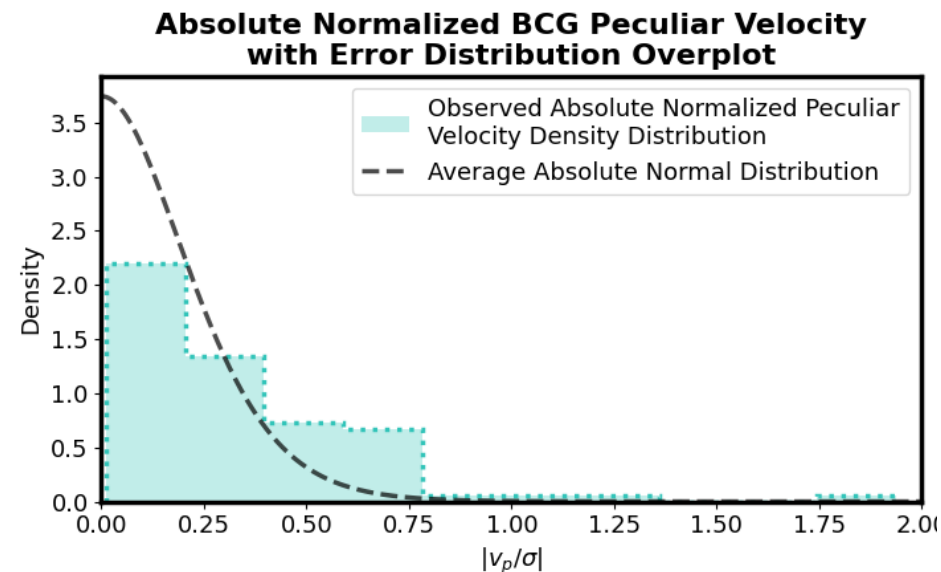
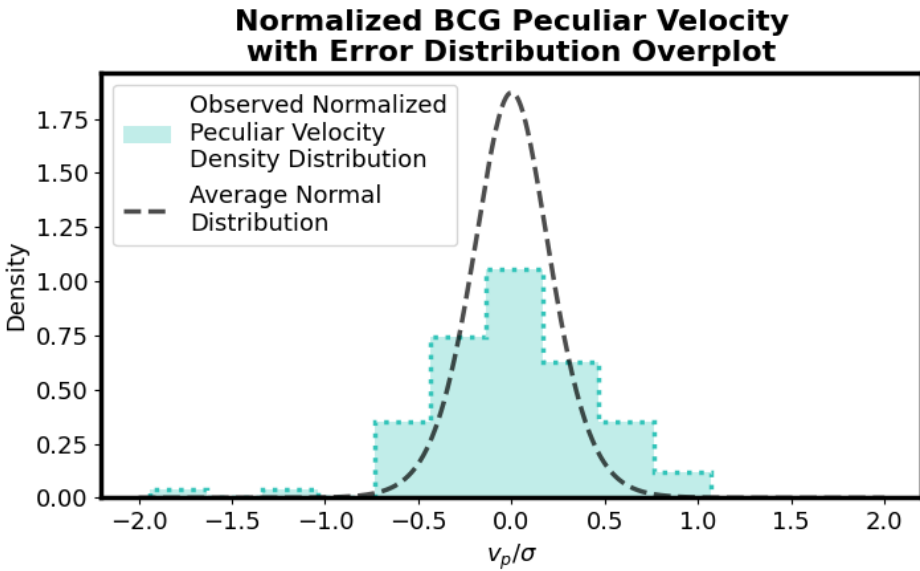
- $v_p / \sigma$ : “BCG peculiar motion”
- Normal Distribution ND with  $\mu=0$ : the Idealized case.
- K-S test of 1000 random draws and 100 different time seeds.
- $\text{Med}_{\text{err}}$ : median of the measurement error distribution
- $\sigma_{\text{err}}$ : the confidence interval from the non-Gaussian measurement error distribution





# Observed vs. Theorized Offsets

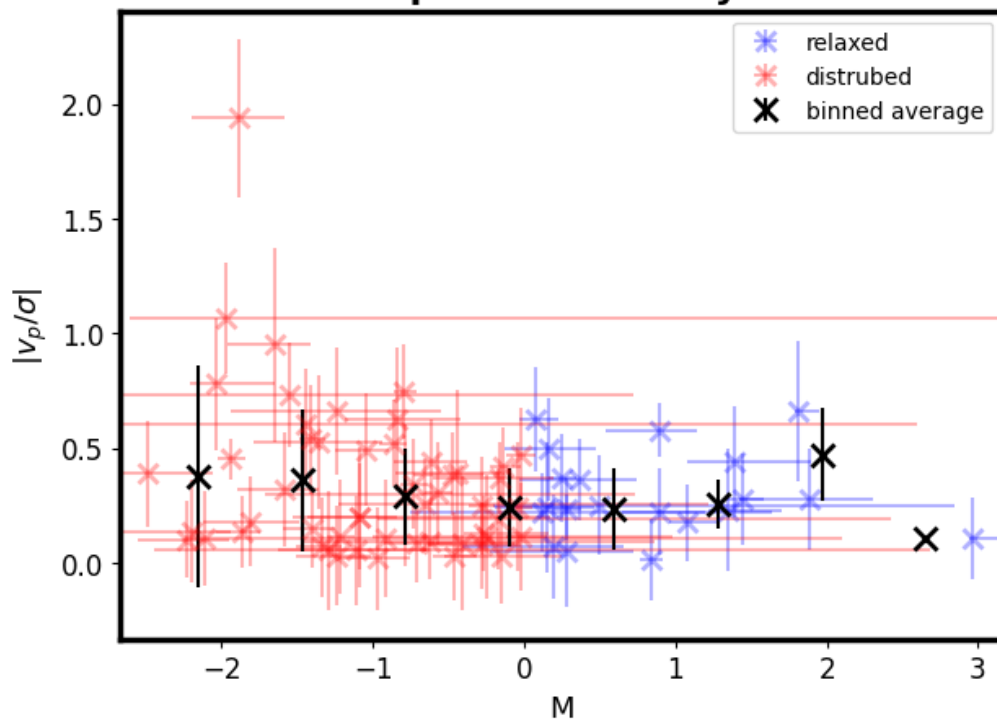
- $SD_{\text{gaus}}$ : standard deviation of the Normal Distribution (ND)
- K-S Null Hypothesis Rejection Rate:
  - ND and observed values are from same parent distribution:  $(25.1^{+1.2}_{-1.2})$
  - Absolute ND values are greater than the observed values:  $(100^{+0.0}_{-0.1})$



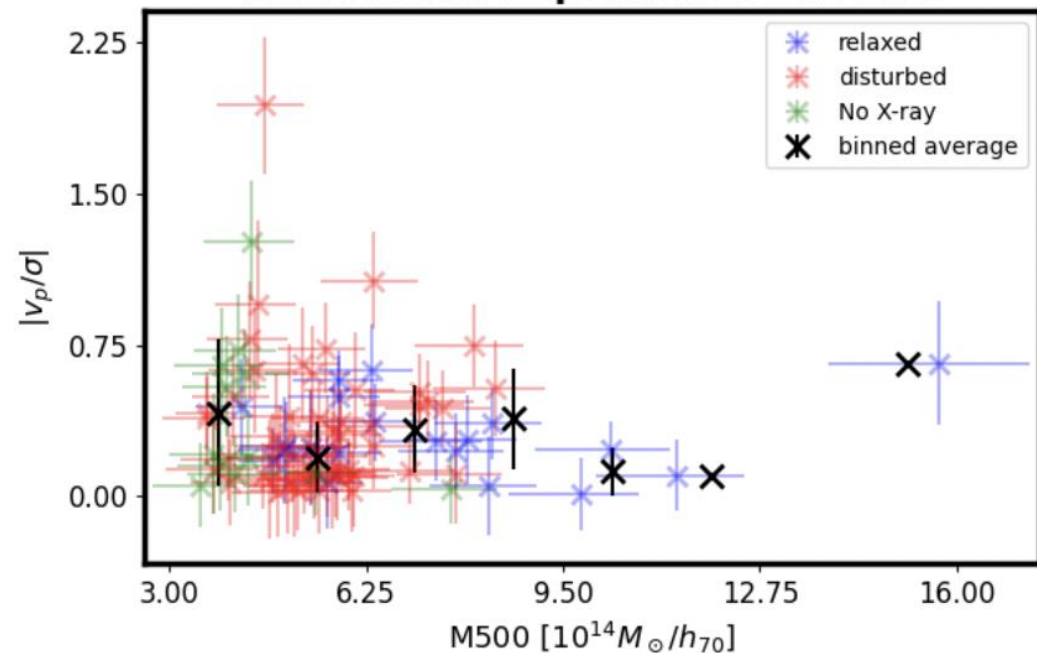
# Dynamical State, Redshift, and Mass

parameters	r	p
$ v_p/\sigma $ vs. M	0.054	0.651
$ v_p/\sigma $ vs. z	-0.114	0.297
$ v_p/\sigma $ vs. M500 [ $10^{14} M_\odot/h_{70}$ ]	-0.009	0.935

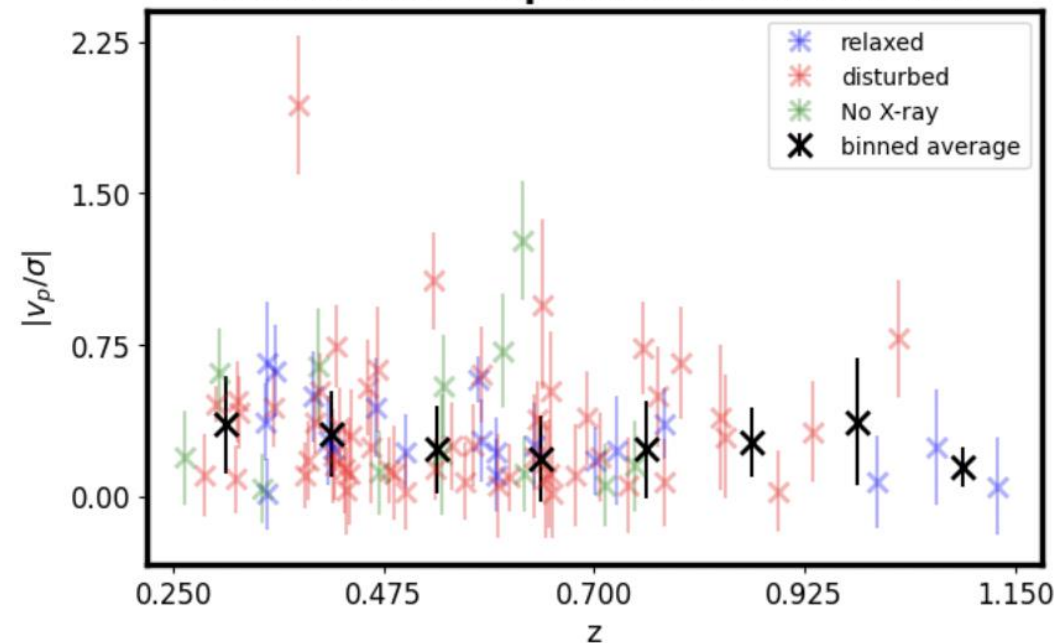
**BCG Motion Dependence on Dynamical State**



**BCG Motion Dependence on Mass**



**BCG Motion Dependence on Redshift**

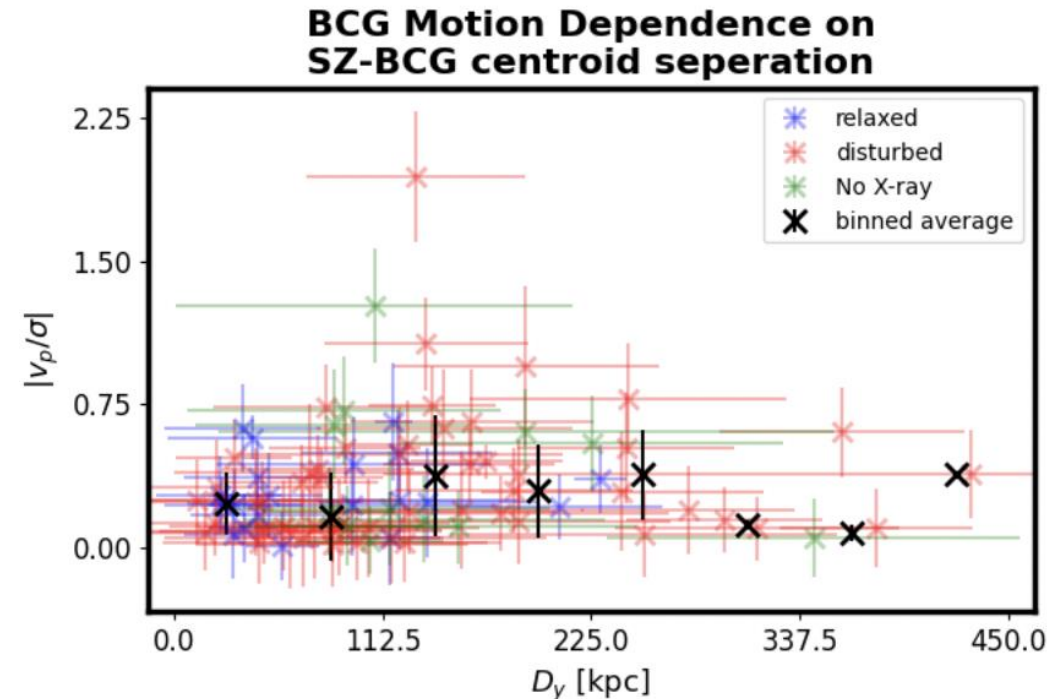
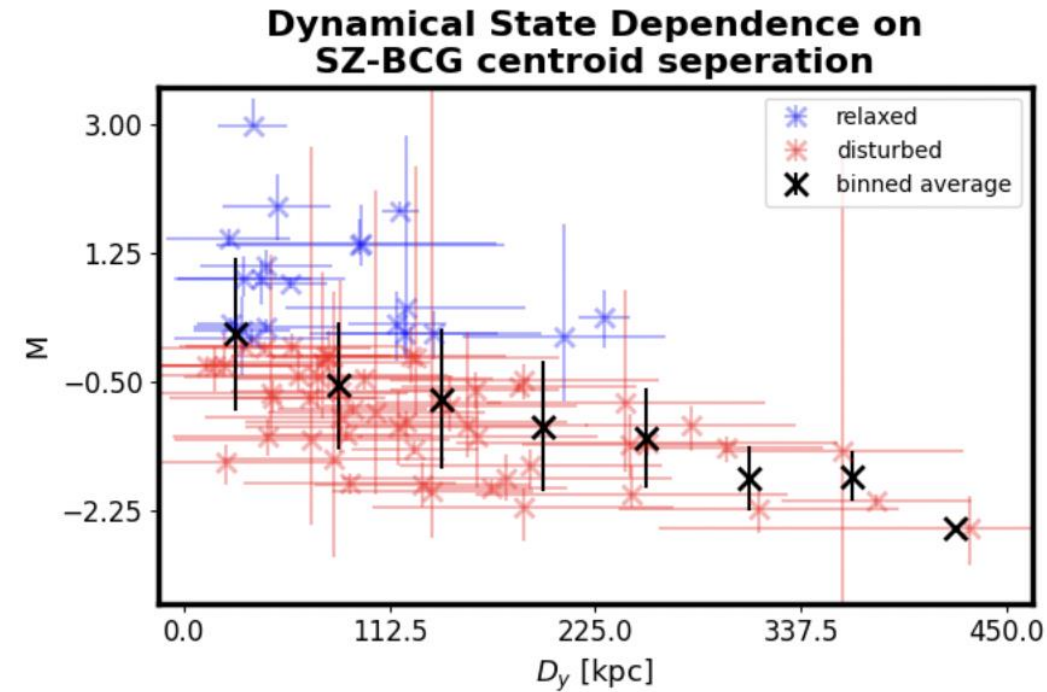
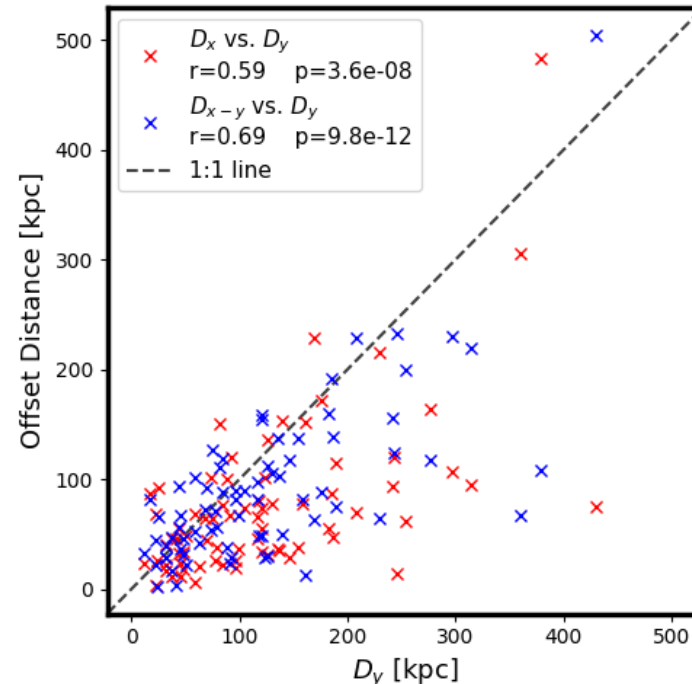




# BCG-SZ Offset

- $D_x$ =BCG-Xray centroid offset
- $D_y$ =BCG-SZ centroid offset
- $D_{x-y}$ =Xray-SZ centroid offset

parameters	r	p
$ v_p/\sigma $ vs. $D_y$ [kpc]	0.022	0.839
M vs. $D_y$ [kpc]	-0.534	1e-06
$ v_p/\sigma $ vs. A	0.016	0.893
$ v_p/\sigma $ vs. c	0.076	0.523
$ v_p/\sigma $ vs. $D_x$ [kpc]	-0.081	0.496
$ v_p/\sigma $ vs. M	0.054	0.651



# Conclusions

- We find that the observed BCG peculiar motion is larger than what it would be if the observed  $v_p / \sigma$  was due to measurement uncertainty alone
- We do not find a very significant correlation between BCG peculiar motion and dynamical state, mass, or redshift.
- $D_y$  acts as a viable dynamical state proxy in cases where we do not have X-ray follow up data.

# Secret Bonus Slides

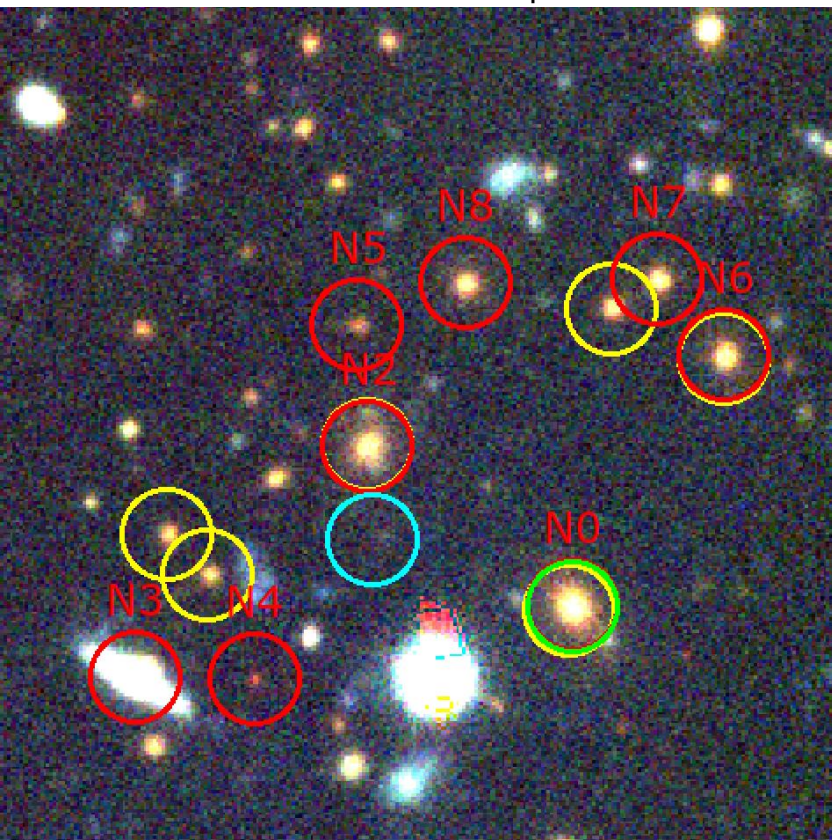


# High $|v_p / \sigma|$

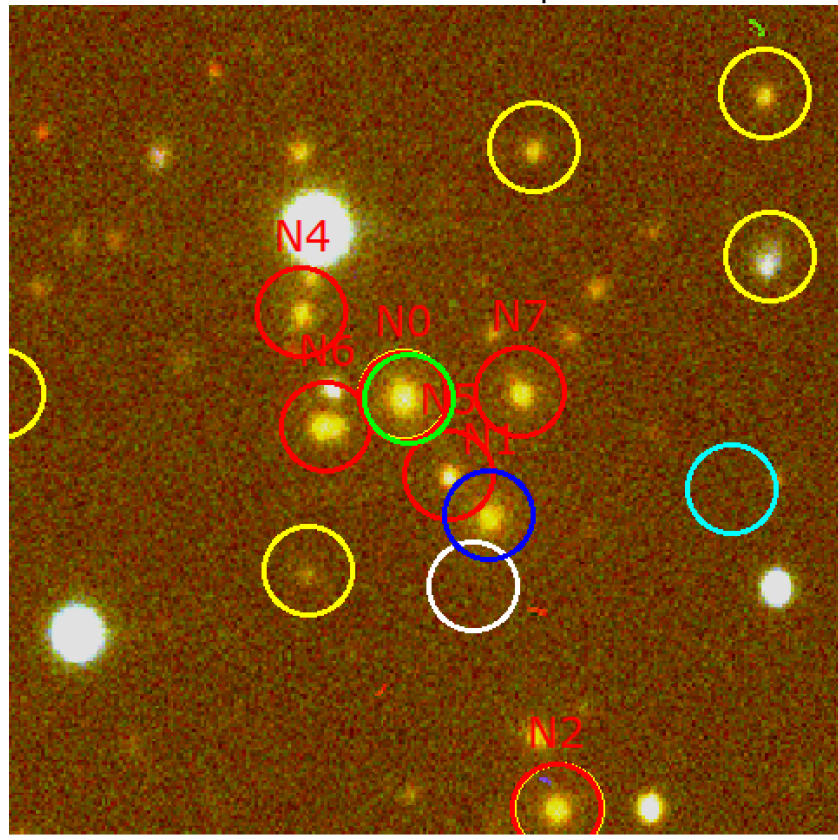
Name	Selected BCG	Probability	2 <sup>nd</sup> Choice BCG	Probability
SPT-CLJ2146-4846	N0	57.14%	N2	13.8%
SPT-CLJ2022-6323	N0	69.93%	N1	8.01%
SPT-CLJ2035-5251	N1	27.59%	N0	39.16%

red	Automated algorithm selection
green	Blind selection
blue	Zenteno et al. 2020 selection
magenta	Calzadilla et al. 2023 selection
yellow	Redshift measured
cyan	SZ center
white	X-ray center

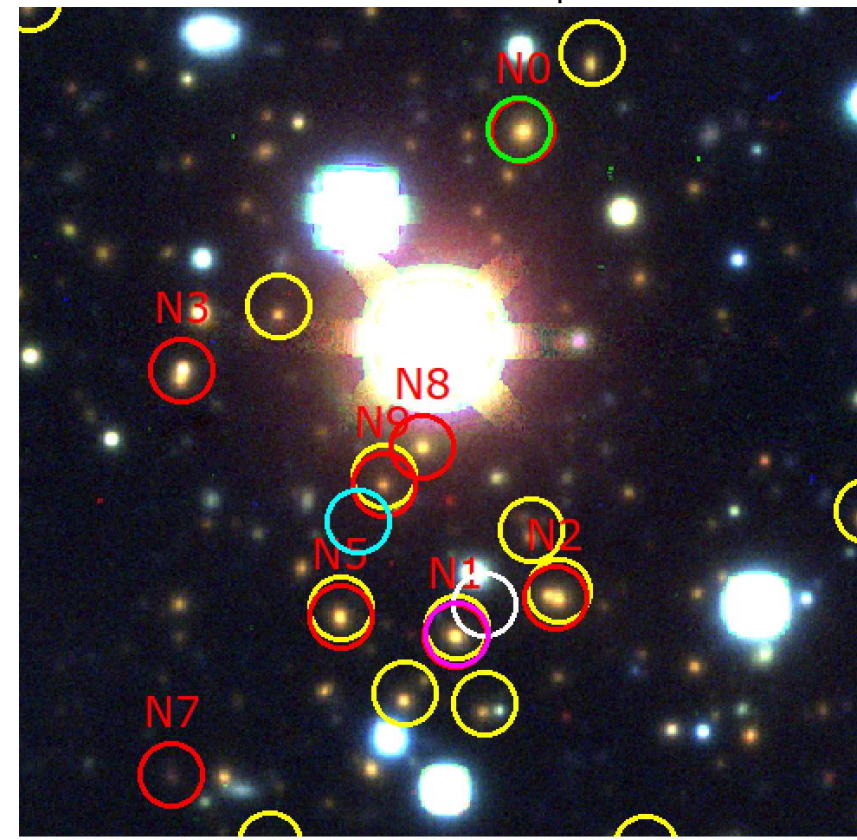
SPT-CLJ2146-4846,  $v_p/\sigma = -1.27$



SPT-CLJ2022-6323,  $v_p/\sigma = -1.94$

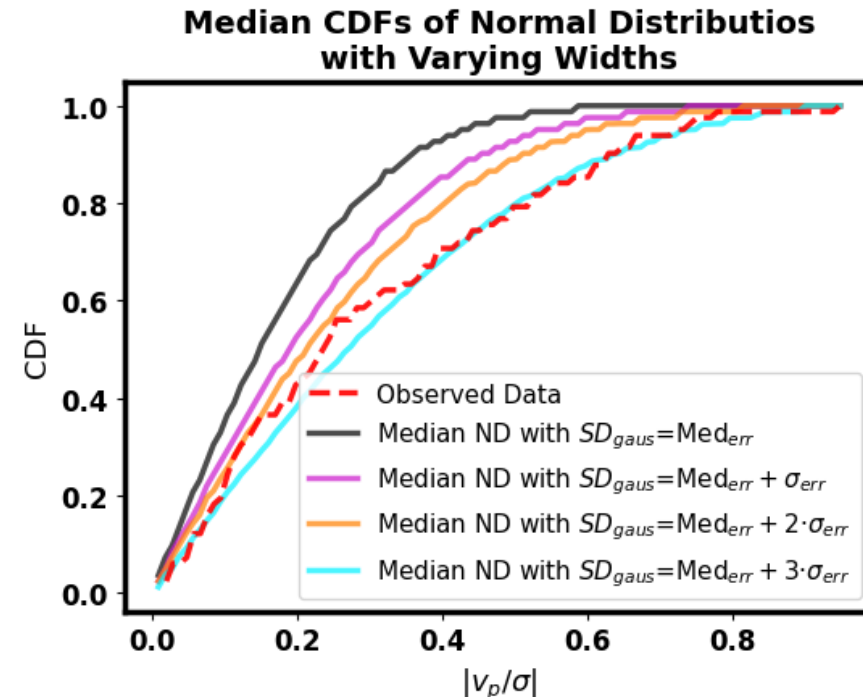
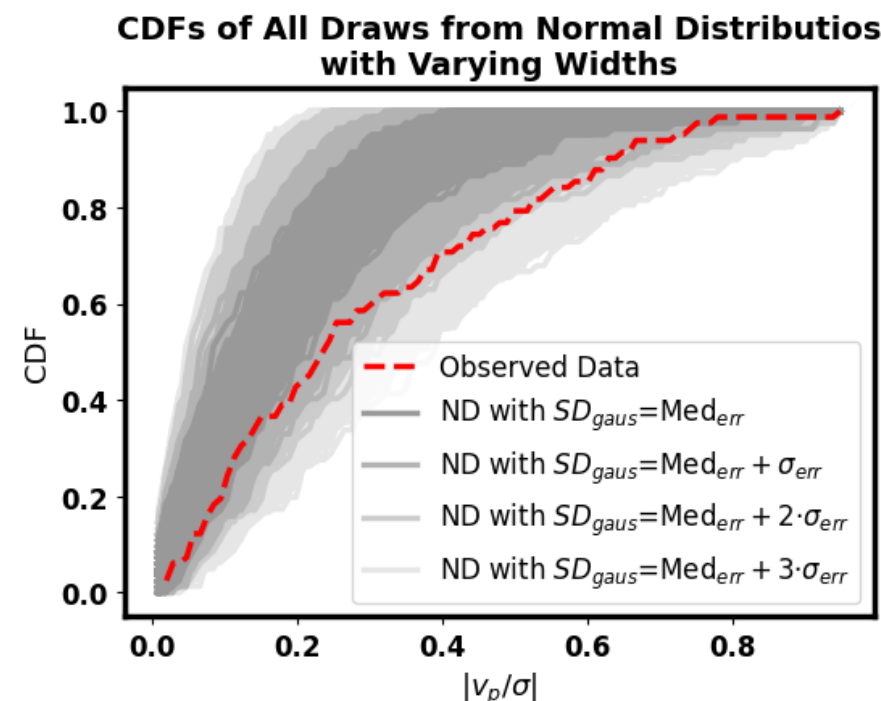
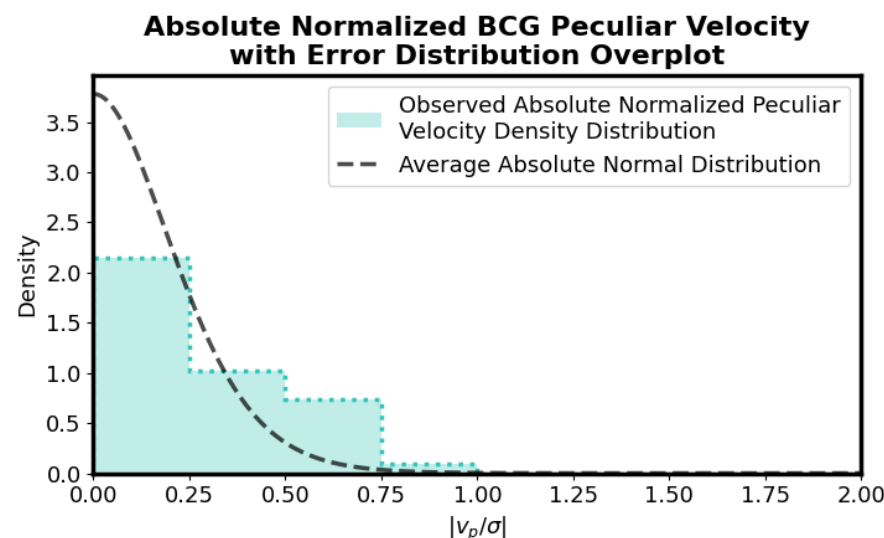
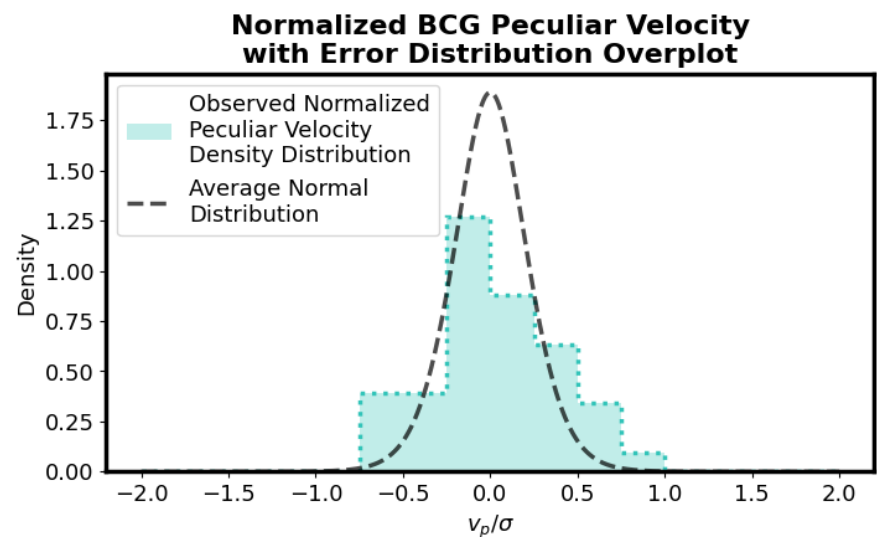


SPT-CLJ2035-5251,  $v_p/\sigma = 1.07$



# Observed vs. Theorized Offsets (High $|v_p/\sigma|$ Removed)

- The removal of the 3 high  $|v_p/\sigma|$  does not qualitatively change our conclusions.
- K-S Null Hypothesis Rejection Rate:
  - ND and observed values are from same parent distribution:  $(13.65^{+1.08}_{-1.15})$
  - Absolute ND values are greater than the observed values:  $(99.4^{+0.23}_{-0.2})$

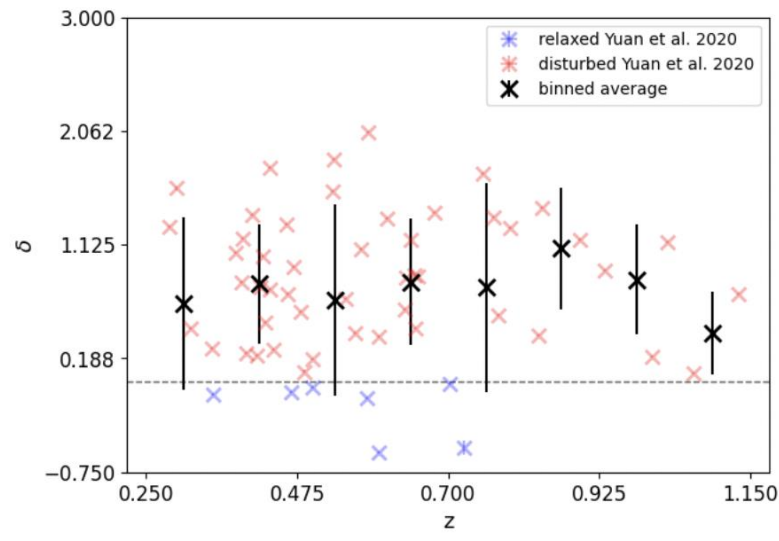
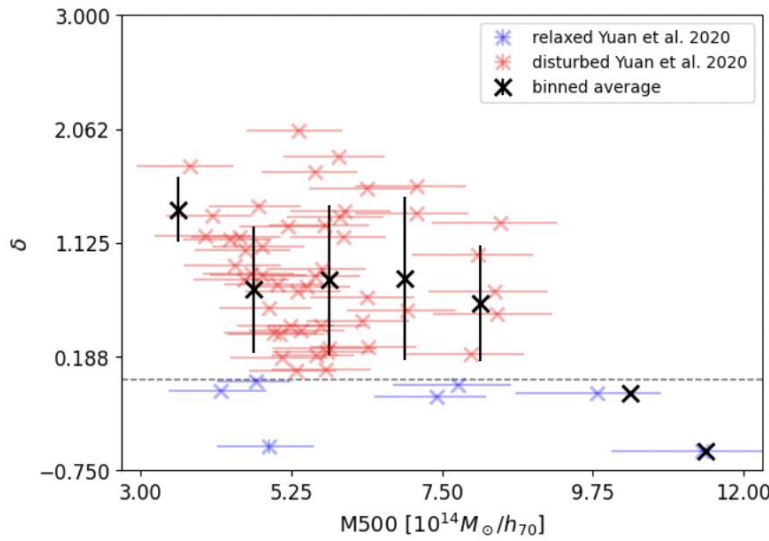
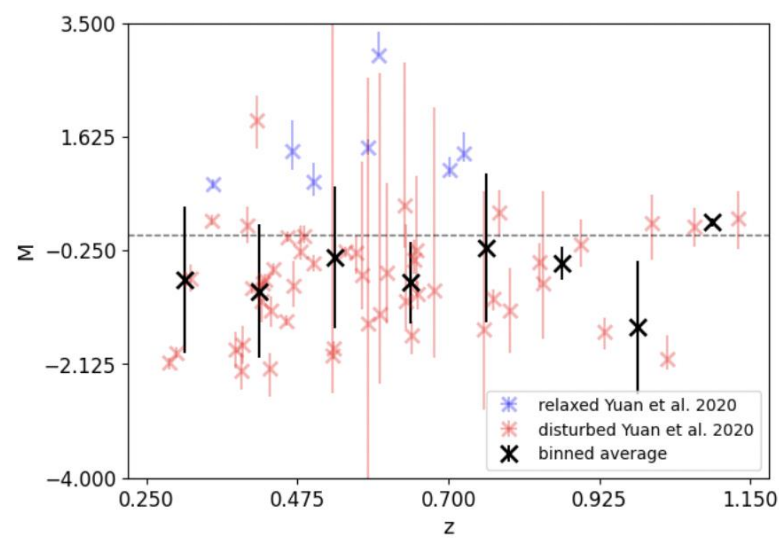
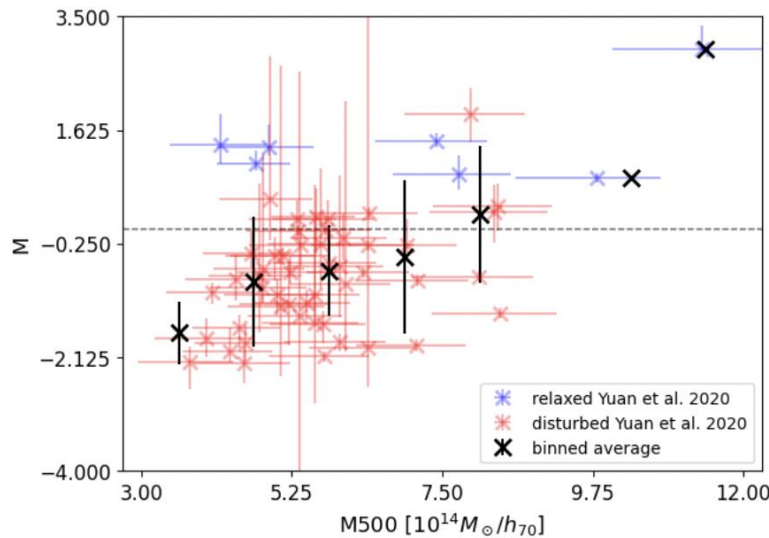
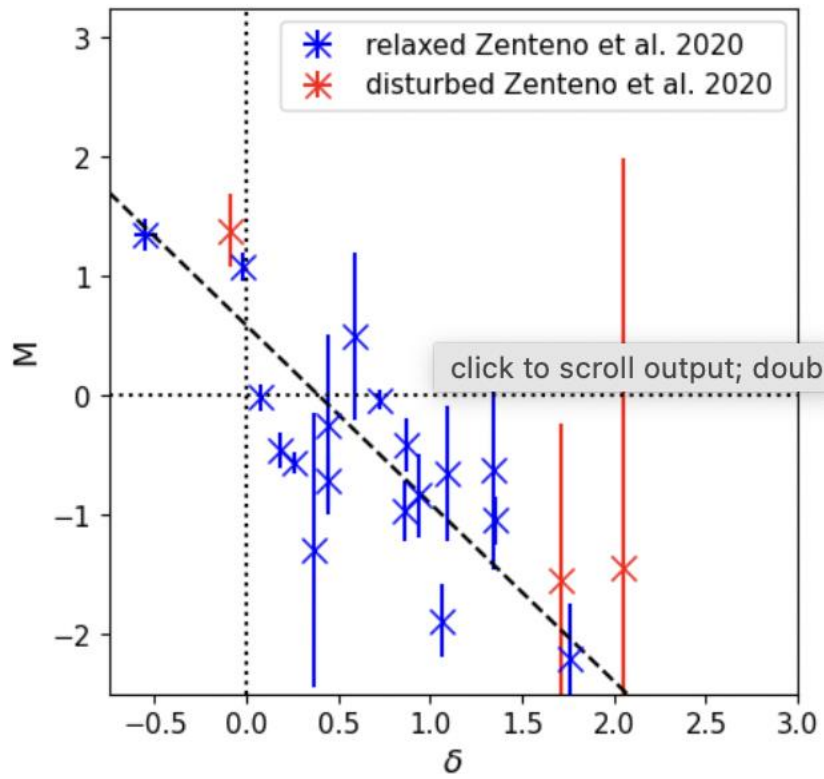




# X-ray Comparisons and Evolution

Catalogue A – Catalogue B	# Common Clusters	% Classification Match	% Classification Mismatch
This Work - Yuan et al. 2020	56	85.71%	14.29%
This Work - Zenteno et al. 2020	29	31.03%	68.97%
Yuan et al. 2020 - Zenteno et al. 2020	21	19.05%	80.95%

parameters	r	p
M vs. M500 [ $10^{14} M_{\odot}/h_{70}$ ]	0.486	0.00014
$\delta$ vs. M500 [ $10^{14} M_{\odot}/h_{70}$ ]	-0.289	0.031
M vs. z	0.11	0.42
$\delta$ vs. z	-0.007	0.96



# Important parameters and bins

Parameter	Value
$\text{Min}_{err}$	0.09
$\text{Med}_{err} - 3\sigma_{err}$	0.119
$\text{Med}_{err} - 2\sigma_{err}$	0.134
$\text{Med}_{err} - \sigma_{err}$	0.177
$\text{Med}_{err}$	0.222
$\text{Med}_{err} + \sigma_{err}$	0.285
$\text{Med}_{err} + 2\sigma_{err}$	0.344
$\text{Med}_{err} + 3\sigma_{err}$	0.413
$\text{Max}_{err}$	0.42

bin #	M bins	z bins	M500 bins [ $10^{14} M_{\odot}/h_{70}$ ]
1	$-2.5 < M < -1.8125$	$0.25 < z < 0.3625$	$3.0 < M500 < 4.625$
2	$-1.8125 < M < -1.125$	$0.3625 < z < 0.475$	$4.625 < M500 < 6.25$
3	$-1.125 < M < -0.4375$	$0.475 < z < 0.5875$	$6.25 < M500 < 7.875$
4	$-0.4375 < M < 0.25$	$0.5875 < z < 0.7$	$7.875 < M500 < 9.5$
5	$0.25 < M < 0.9375$	$0.7 < z < 0.8125$	$9.5 < M500 < 11.125$
6	$0.9375 < M < 1.625$	$0.8125 < z < 0.925$	$11.125 < M500 < 12.75$
7	$1.625 < M < 2.3125$	$0.925 < z < 1.0375$	$12.75 < M500 < 14.375$
8	$2.3125 < M < 3.0$	$1.0375 < z < 1.15$	$14.375 < M500 < 16.0$