



Calibration challenges in the MAIA calorimeters

Rose Powers (Princeton University)

2nd Annual USMCC Meeting

7 August 2025



Outline

1. Introduction
2. The MAIA calorimetry system and solenoid
3. Particle behavior in the solenoid
4. Analytical calibration approach
5. 2D calibration matrix – benefits and challenges
6. Conclusion and next steps

Introduction

*What is **MAIA**?*

- **M**uon **A**ccelerator **I**mplemented **A**pparatus
- One of two detector concepts proposed for a Muon Collider at $\sqrt{s}=10$ TeV
- Composed of shielding nozzles, trackers, a 5T solenoid, calorimeters, and a muon system

Why are calorimeters important?

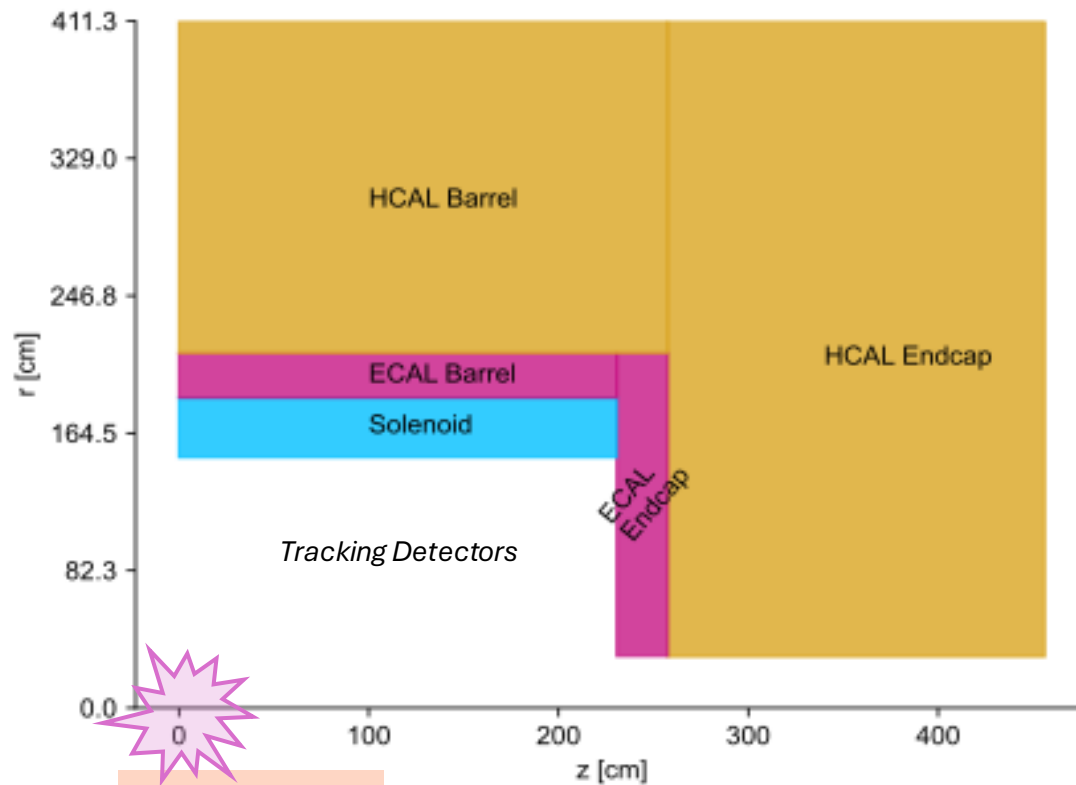
- The electromagnetic and hadronic calorimeters are designed to **stop** EM and hadronic particles, respectively, and **measure their energy**
- **5D calorimetry**—information about energy, timing, and position

Why and how do we calibrate our calorimeters?

- Accurate energy measurements are crucial to reconstruction
 - **Particle ID**
 - Identifying **missing energy**
 - Finding **resonances**
- We can assess the calorimeters' performance with the simplest neutral objects (**photons and neutrons**) as **standard candles**

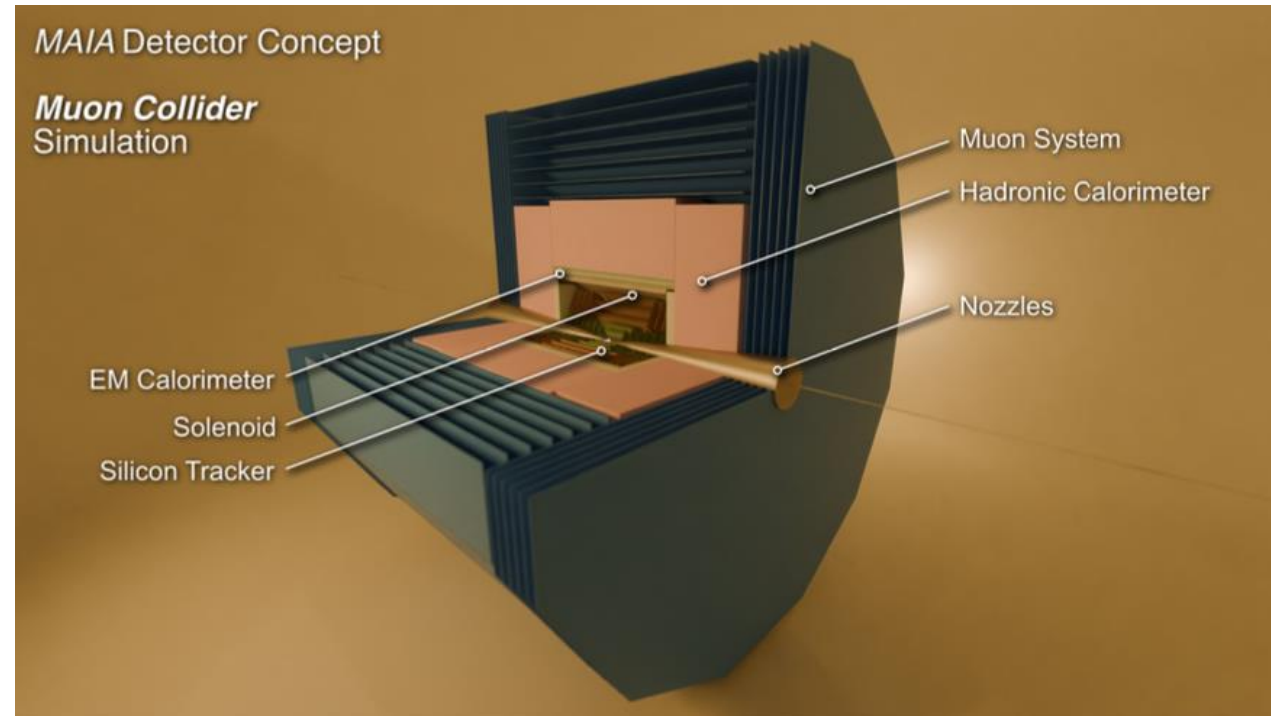
MAIA Calorimeters

Cross-section of the first quadrant of the MAIA detector in the RZ plane, focusing on the solenoid and calorimeters.



Interaction Point

8/7/25



ECAL:

- Silicon and Tungsten
- 5x5 mm² cells
- 50 layers

HCAL:

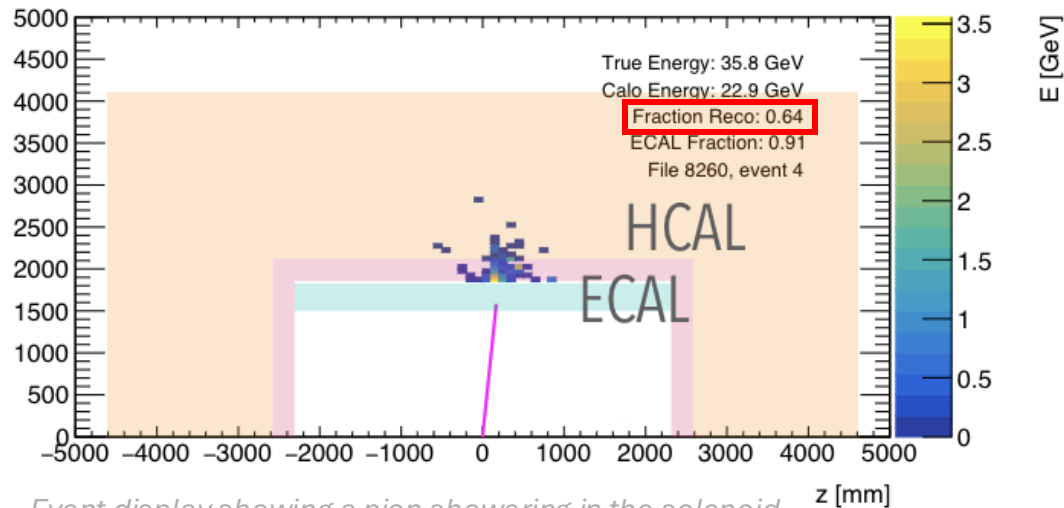
- Iron and scintillator
- 30x30 mm² cells
- 75 layers

The solenoid: BIB bodyguard or black box?

The **solenoid** adds approximately **265 mm of aluminum** between the **trackers** and the **calorimeters**. Two interpretations...

265 mm of **additional shielding**, drastically reducing BIB occupancy in the calorimeters!

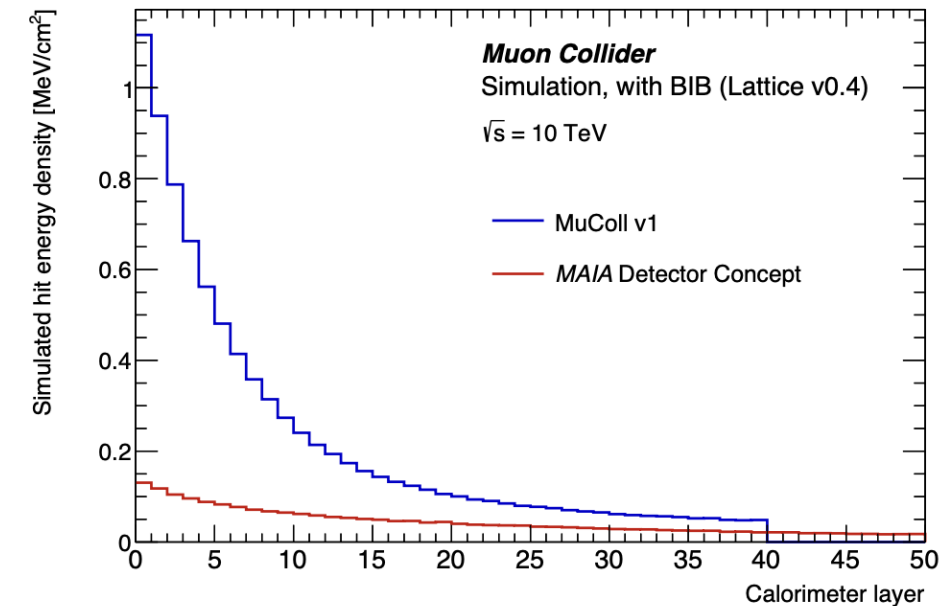
265 mm of **detector-free material** where particles **interact** and **lose energy**.



Event display showing a pion showering in the solenoid.

Plot courtesy of Tova Holmes.

8/7/25



Comparison of BIB energy density deposition in ECAL for different solenoid placements.

[1] [arXiv:2502.00181](https://arxiv.org/abs/2502.00181)

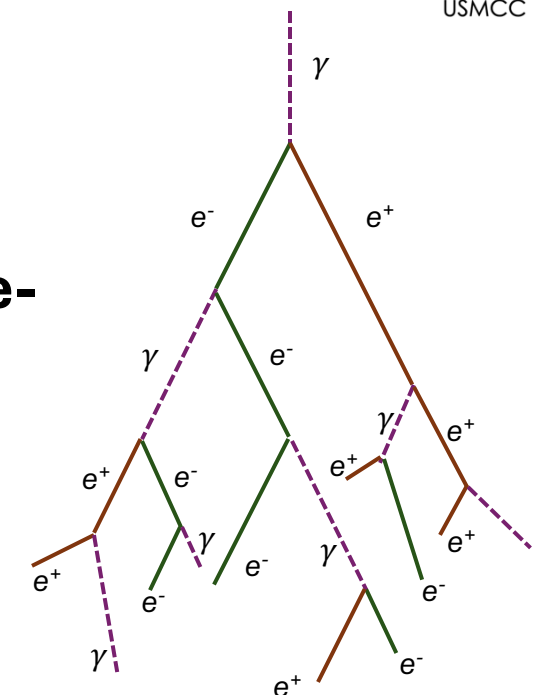
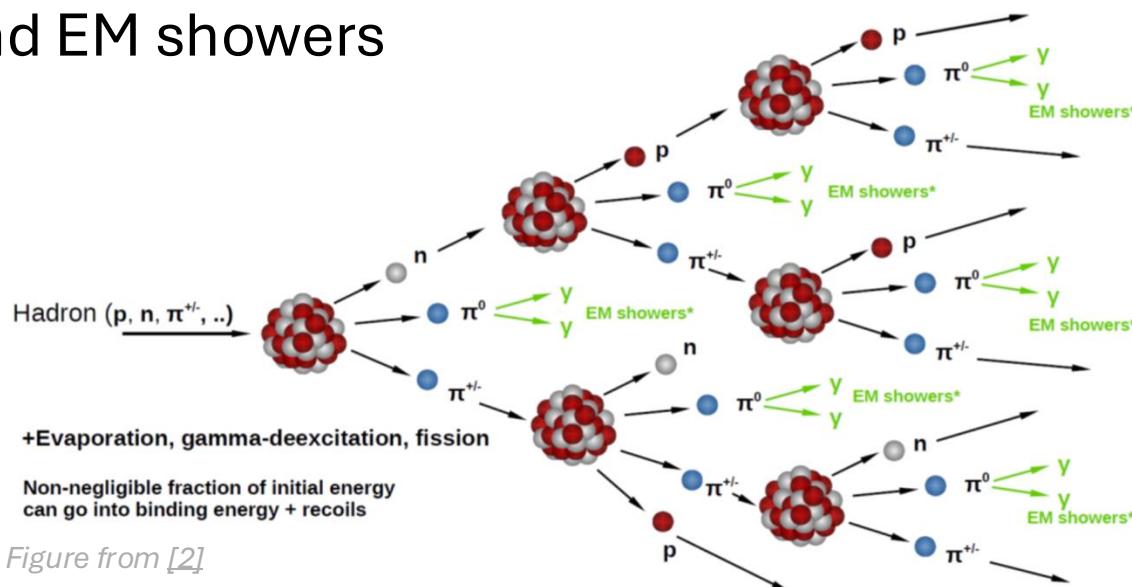
Material interactions in the solenoid

Photons lose energy via **electromagnetic showers**

- Cascade of **pair production**, **Bremsstrahlung**, and **e^+/e^- annihilation**

Neutrons lose energy via **hadronic showers**

- Cascade of nuclear processes, collisions, and EM showers



Both EM particles and hadrons can shower in aluminum, resulting in energy loss between the trackers and the calorimeters.

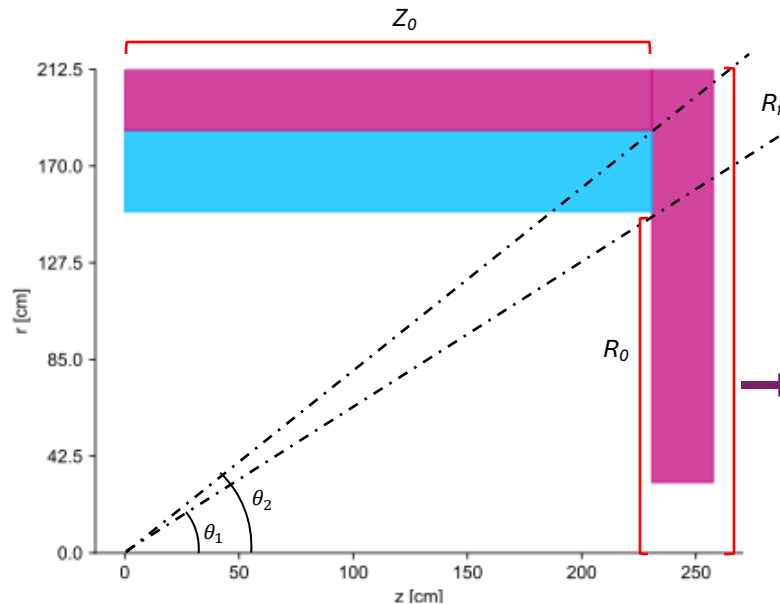
Analytical energy loss function

Let's take an analytical approach to EM showering

- Much simpler to model than hadronic showering

We derive energy response shape as a function of:

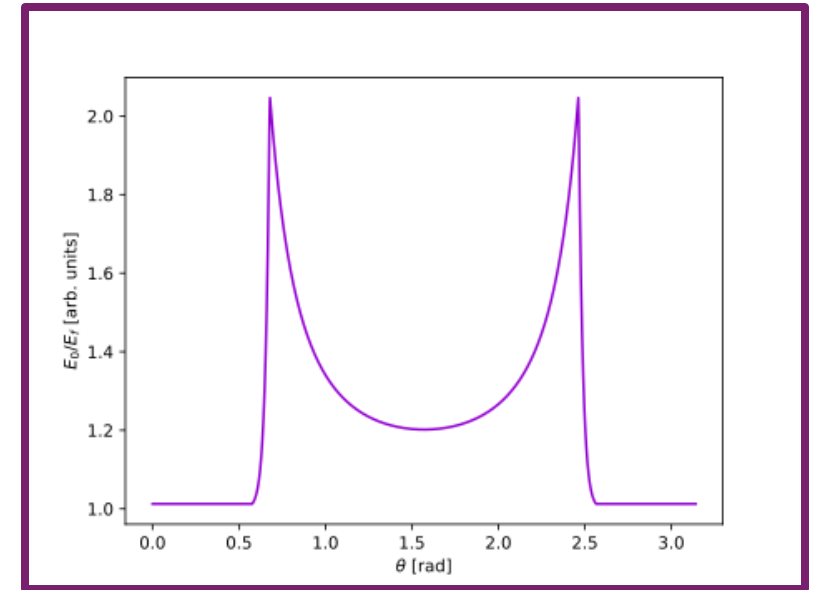
- Geometric parameters of detector
- Photon polar angle



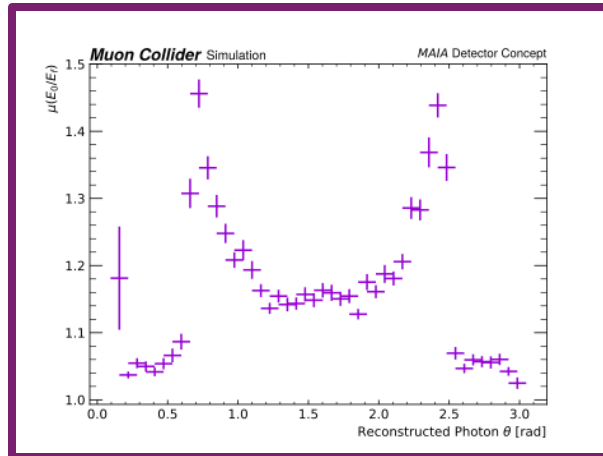
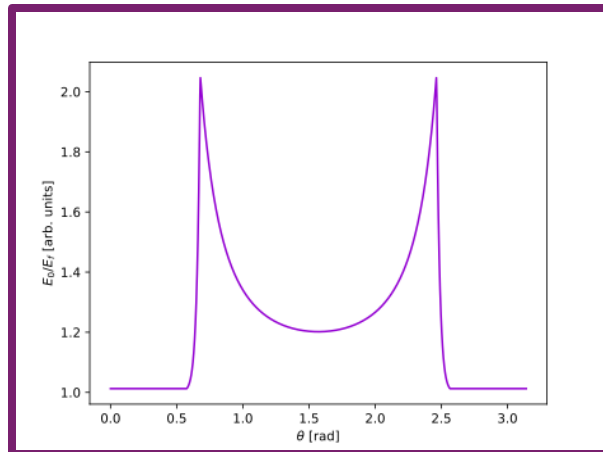
Radiation length $X_0 = 8.89$ cm

$$N(\theta) = \begin{cases} 0 & \theta < \theta_1 \\ (Z_0 |\sec(\theta)| - R_0 \csc(\theta)) / X_0 & \theta_1 < \theta < \theta_2 \\ (R_f - R_0) \csc(\theta) / X_0 & \theta > \theta_2 \end{cases}$$

$$E_0/E_f \propto 2^{N(\theta)}$$

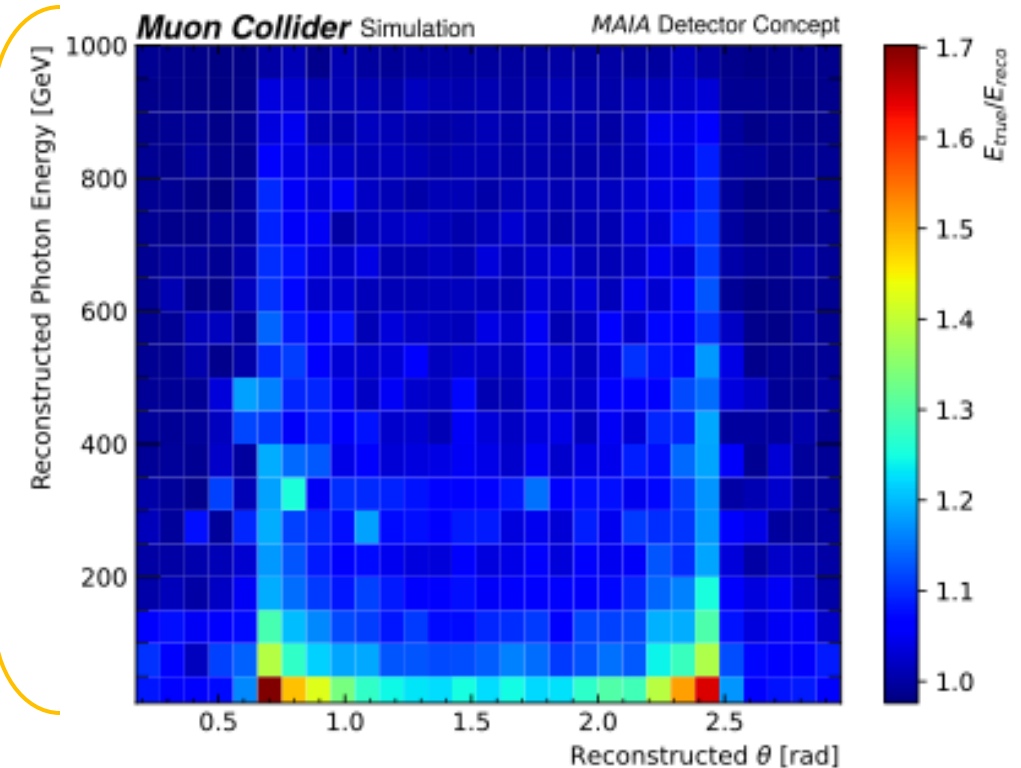


Indeed, the *theta*-dependence of photonic energy response follows this form...



Response is also **energy dependent**—
higher-energy particles
lose **fractionally less**
energy in the material.

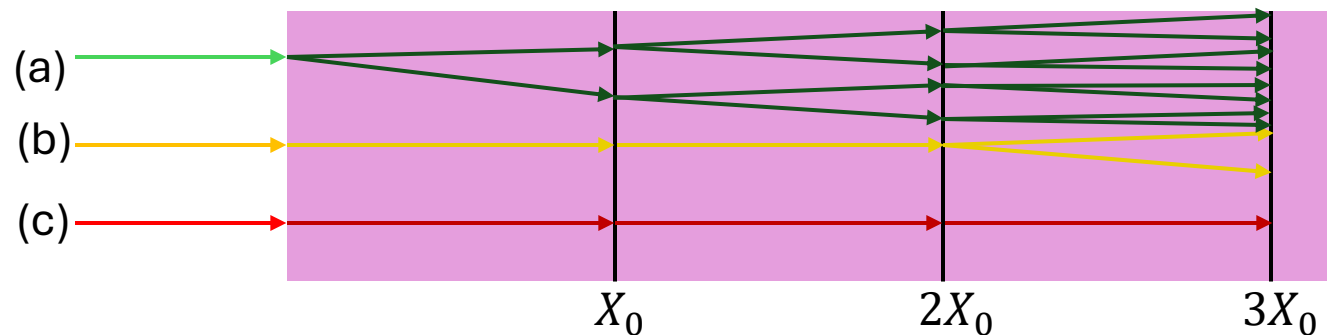
... but *theta*-dependence is not the whole story.



Stochasticity of showering

*Unlike geometric theta dependence, energy dependence of response is highly **nontrivial** to model.*

- Our $2^{N(\theta)}$ functional form assumes **immediate** showering
- At high energies, may see MIP-like behavior or **stochastically delayed** showering

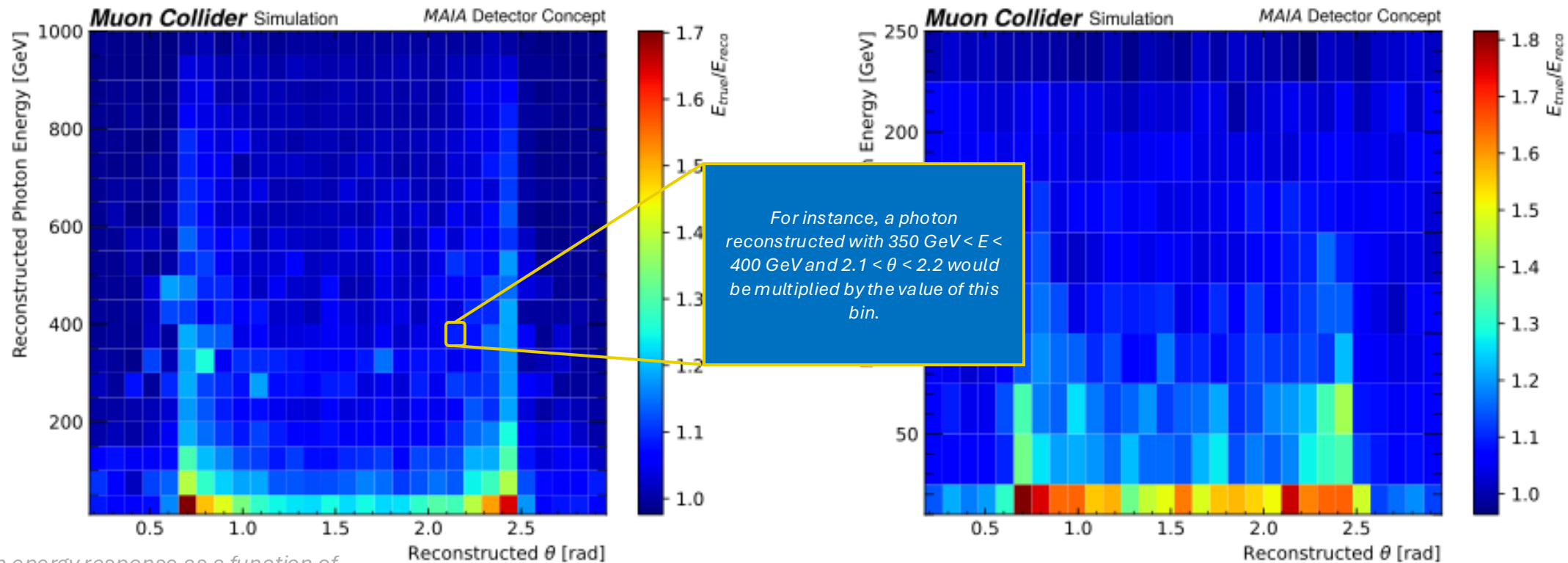


Cartoon of three particles beginning to shower at (a) the entrance to the solenoid, (b) after two interaction lengths, and (c) not at all.

Bottom line: we cannot rely on a realistic analytical calibration.

2D response as calibration function

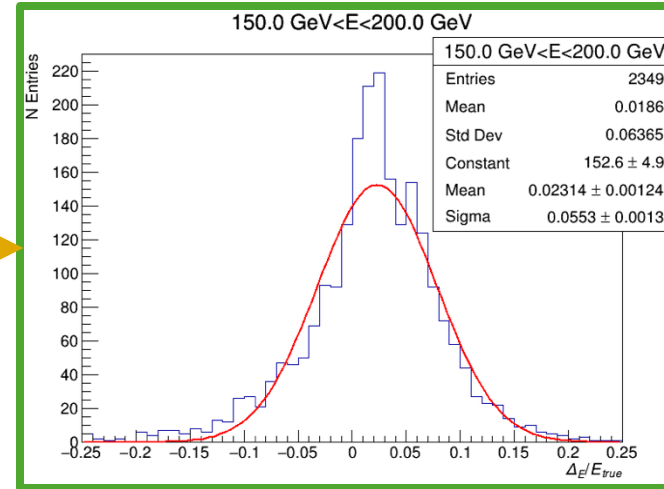
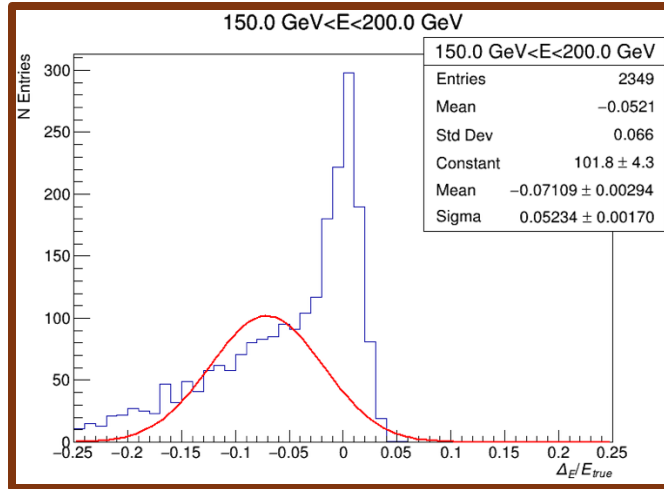
In lieu of an **analytical** calibration, we can use the response plots themselves as **calibration matrices**.



Photon energy response as a function of reconstructed energy and azimuthal angle.

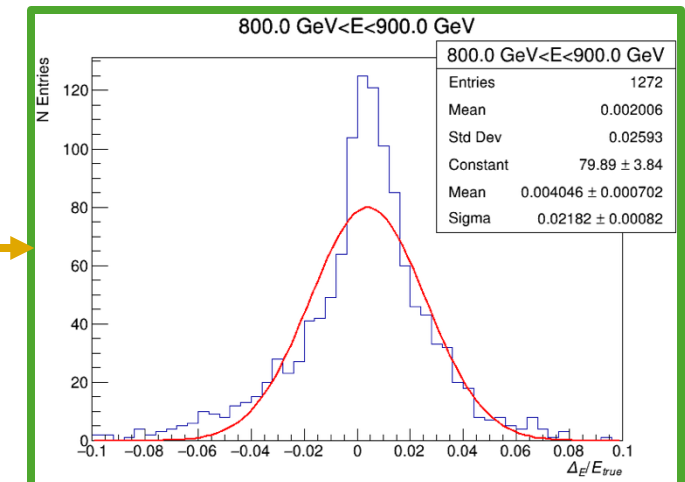
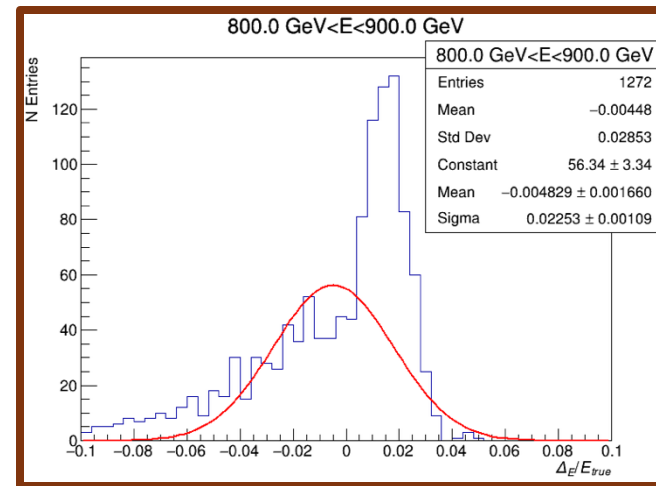
Analogous plot for neutrons.

Calibrated fractional response



Two examples of the binned $\frac{\Delta E}{E_{true}}$ distribution for photons in an inclusive θ range, both before and after 2D response calibration is applied.

Calibration effectively **eliminates** bimodality due to geometric effects.

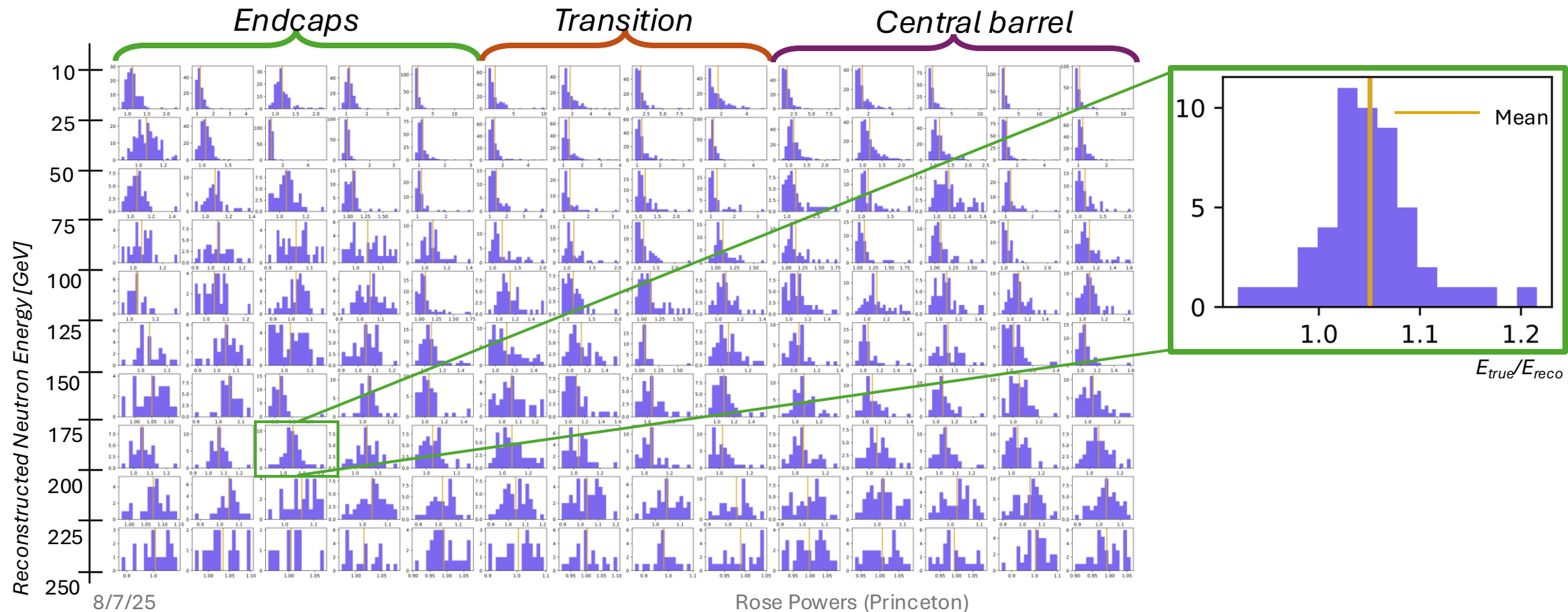


Rose Powers (Princeton)

But is this really a valid approach?

These response plots take the **arithmetic mean** in each bin

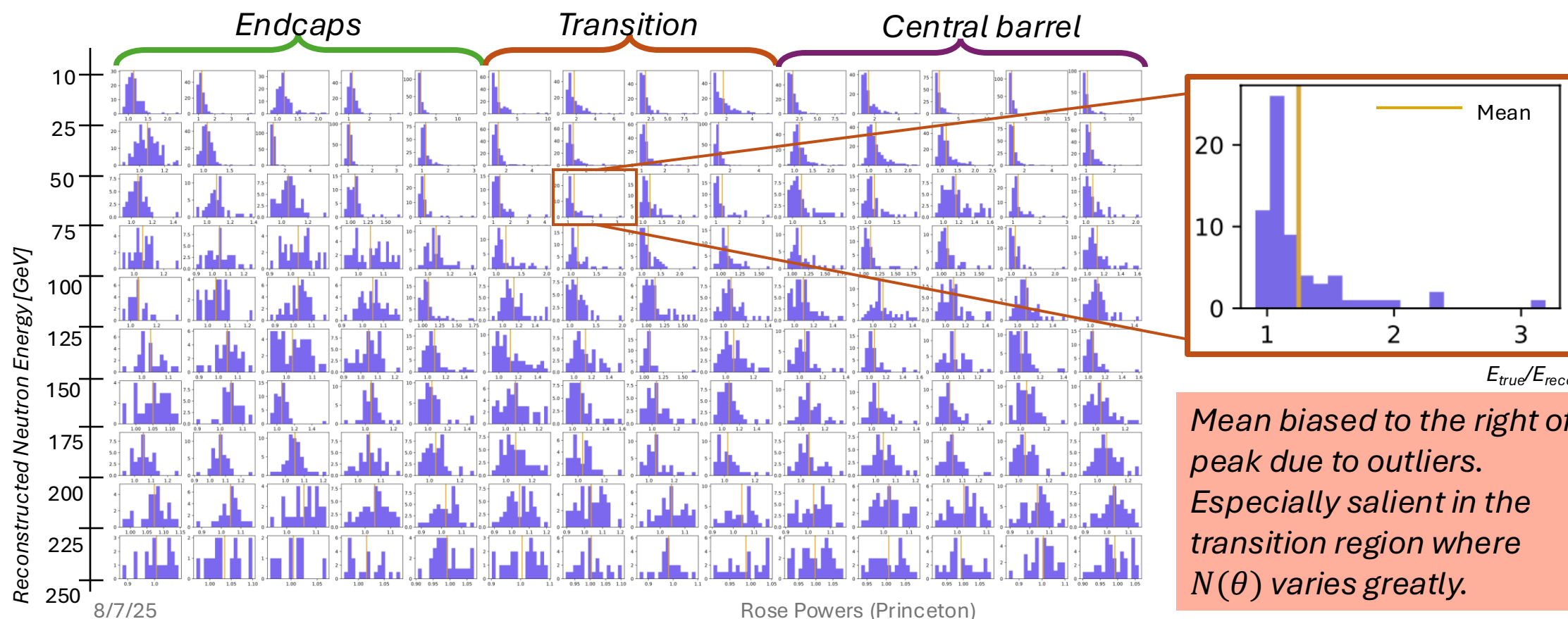
- Works well in bins with **Gaussian** distributions...



But is this really a valid approach?

These response plots take the **arithmetic mean** in each bin

- However, is susceptible to outliers
- Especially troublesome in **low-N** or **wide-range** regions



Alternative methods

A robust calibration method should not be susceptible to the influence of outliers.

- **Fitting** binned response distributions
 - Many bins exhibit near-Gaussian peaks: fitting those peaks would give a more stable mean
- A dedicated **large-statistics** calibration dataset
 - Could also stabilize and allow us to bin calibration matrix more finely
- A detailed study of particles' interactions in the solenoid
 - Might make an interesting tangential investigation with applications in detector performance studies → opportunities here to get involved!

Conclusions and outlook

Calibration cannot completely remove stochastic broadening of energy response due to the solenoid.

- However, a 2D response-based calibration can **mitigate** the most salient geometrical and energetic dependencies
- A more **nuanced assessment** of binned response may further improve accuracy of calibration
- While the solenoid adds challenges to the calibration process, these challenges can largely be mitigated and are **far outweighed** by the solenoid's BIB reduction efficacy